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SPSLC P282  
October 10, 1994

CM-P00059063

**Proposal**

**Neutron halo and antiproton - nucleus potential  
from antiprotonic X-rays**

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**ABSTRACT**

It is proposed to investigate the characteristics of antiprotonic X-rays in a number of heavy nuclei. The objective of this study is a combined analysis of observables depending on the nuclear periphery and the antiproton-nucleus optical potential. These observables would be gathered during the realisation of the proposed program (X-rays level shifts and widths) and will also come from the previous neutron halo investigation performed within the PS203 experiment. The present Proposal demonstrates that such an analysis will substantially improve our knowledge of the nuclear periphery in heavy nuclei and, at the same time, will substantially delimitate the parameters defining the antiproton-nucleus optical potential. In 1995 we request 3 weeks of the beam time for this experiment (about  $10^5 \bar{p}/s$ , 200 MeV/c).

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\* Presently at CERN, PPE Division, Isolde Group

# Nuclear surface study

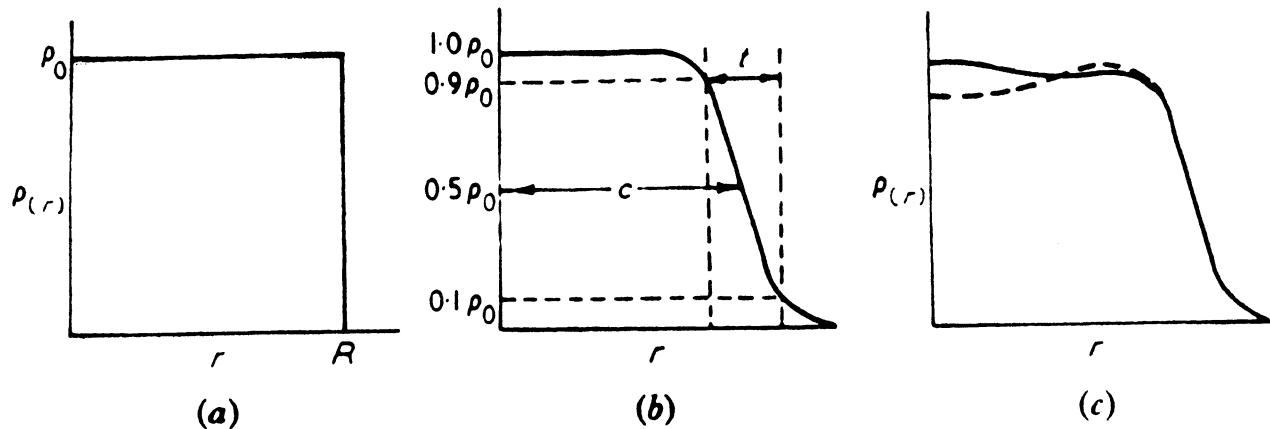


Figure 1. Examples of nuclear matter distributions. (a) Billiard ball nucleus with a well-defined radius  $R$ . (b) The fermi distribution. (c) Shell model distribution.

$$A g_m(r) = Z g_p(r) + N g_n(r)$$

Fermi:  $g = g_0 \left[ 1 + \exp(r - c)/a \right]^{-1}$

$$t = 4.4 a$$

$$\langle r^2 \rangle = \int_0^\infty r^2 g(r) \cdot 4\pi r^2 dr$$

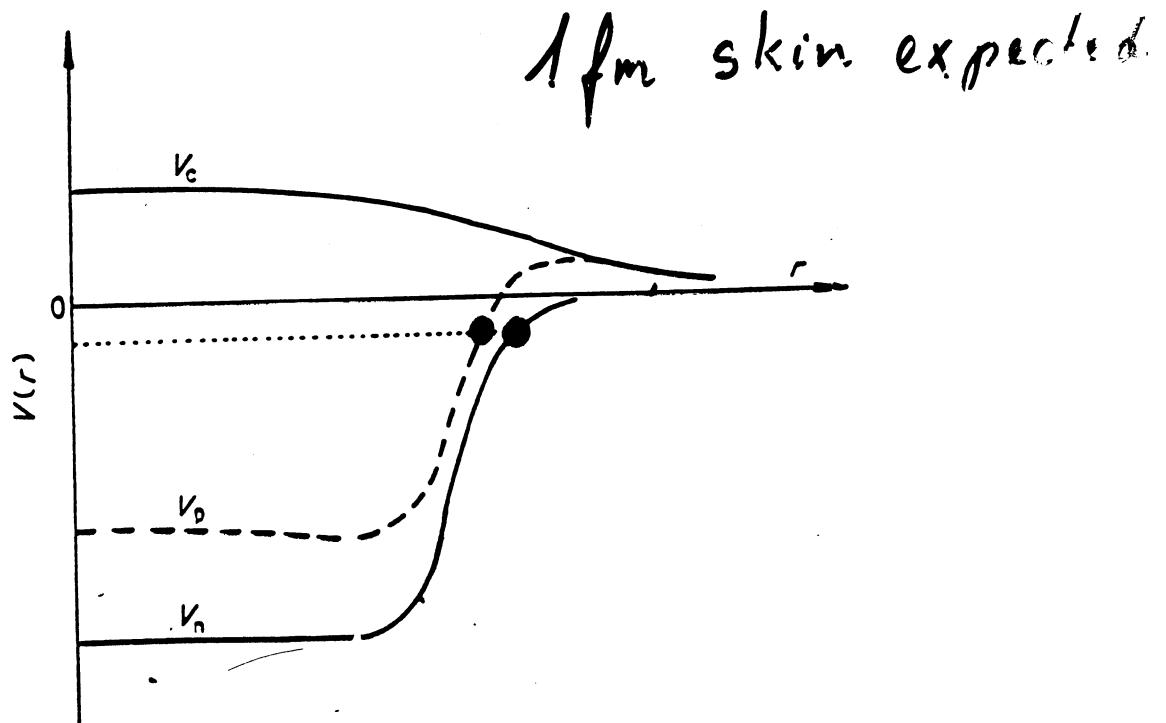


Figure 2. Behaviour of neutron and proton potentials on the assumption that  $V_p = \underline{V_n + V_c}$ .

M. H. Johnson and E. Teller

Phys. Rev. 93, 357 (1954)

however  $|V_p| > |V_n|$

Is the Surface of the  
Nucleus Neutron-Rich?

D. H. Wilkinson  
(1967)

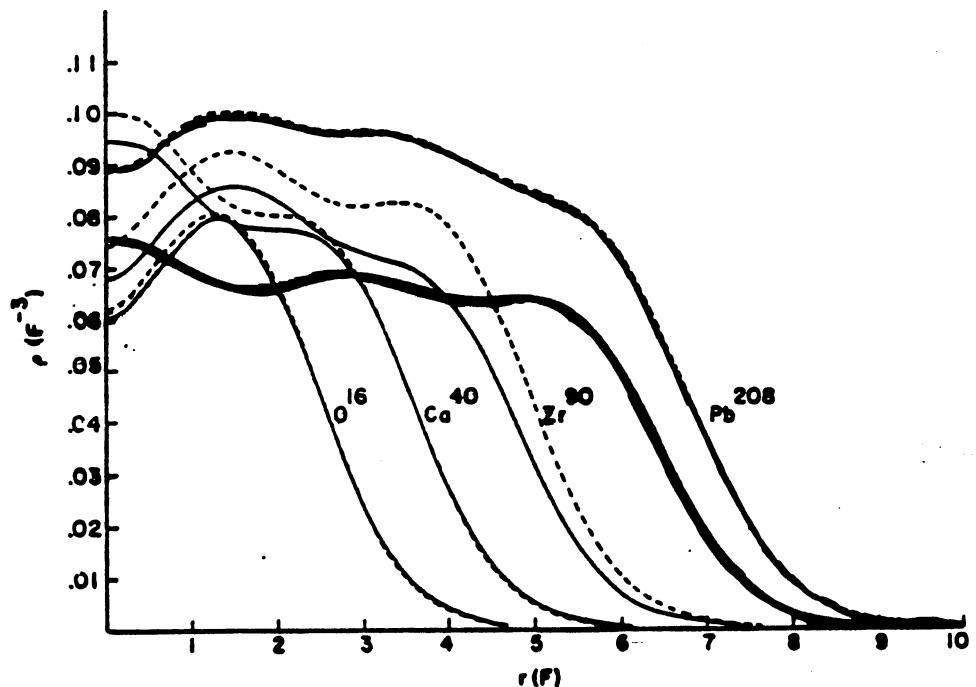


Fig. 1. Neutron (dashed lines) and proton (solid lines) density distributions, with no nucleon size or c.m. motion corrections.

J. D. Negele Phys. Rev. C1, 126 (1970)

$$\text{For } \text{Th} \quad \Gamma_{n}^{\text{rms}} > \Gamma_{p}^{\text{rms}} \\ 4\%$$

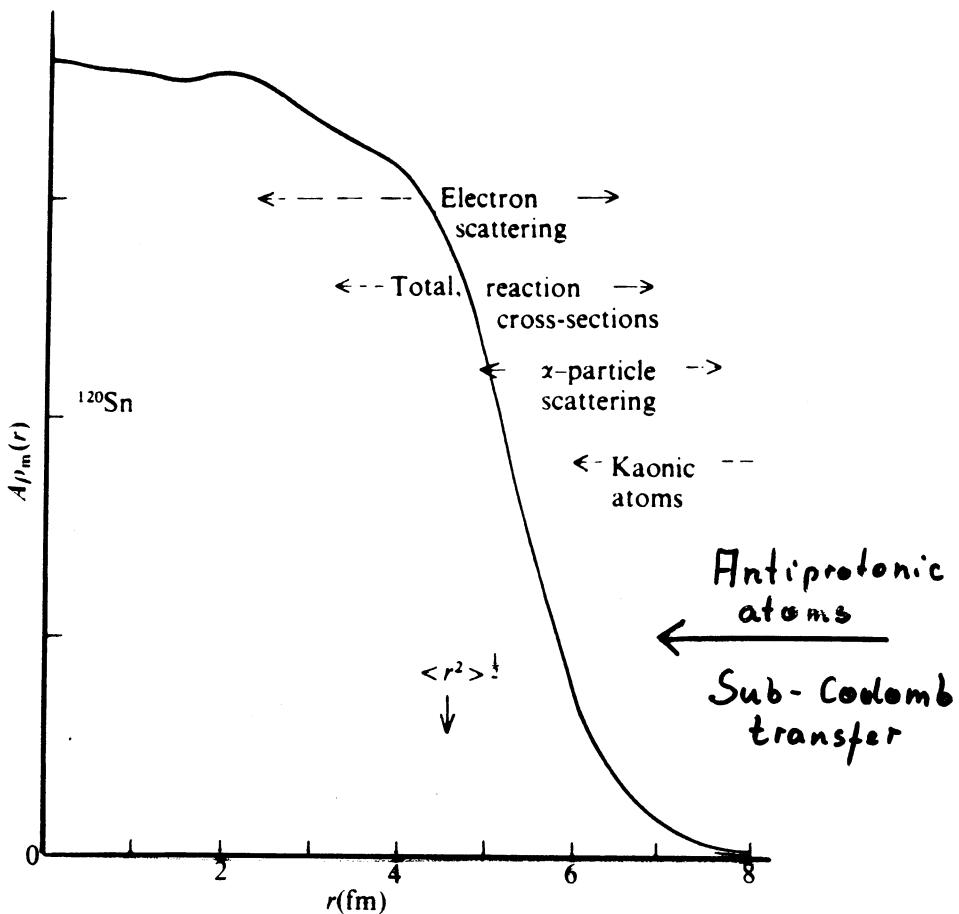


FIG. 11.1. The nuclear matter distribution  $A_{p_m}(r)$  for  $^{120}\text{Sn}$  given by the BG (proton) and ZD (neutron) distributions shown in Fig. 2.2. The horizontal lines show the region of the nucleus probed by various processes; where the lines are broken the extent of the region is still uncertain.  
 (From Jackson 1975.)

R.C. Barrett, D.F. Jackson  
 Nuclear Sizes and Structure  
 (1977)

Sub-Coulomb Pick-up on  $^{208}\text{Pb}$

(d,t)

(p,d)

H.J. Körner and J.P. Schiffer

Phys. Rev. Lett. 27  
(1971) 1457

determined for a number of bombarding energies and angles

$\mathcal{E} \rightarrow |\chi|^2$  values of neutron hole states

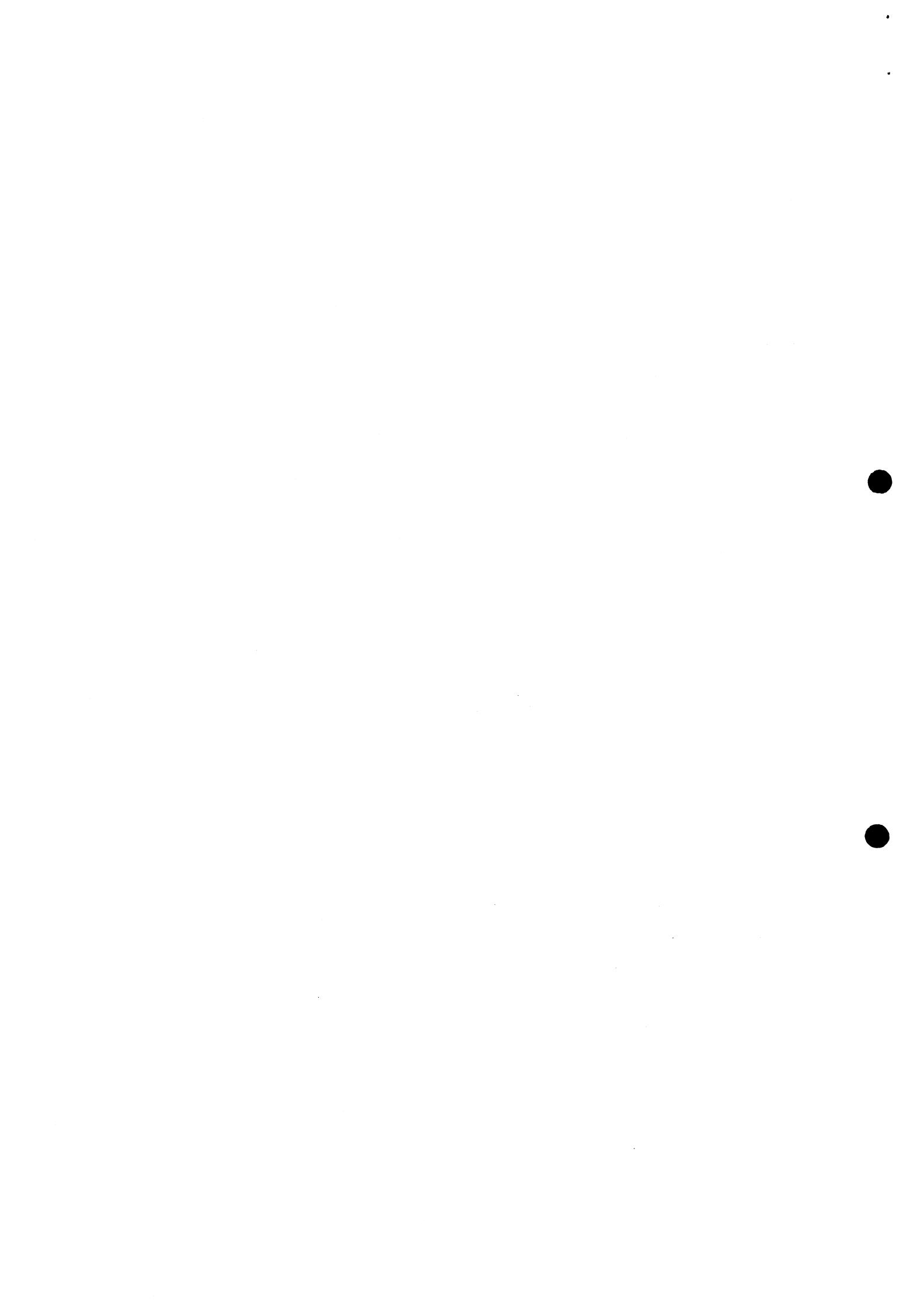
$^3P_{1/2}, ^3P_{3/2}, ^2f_{5/2} \dots$

at 15 fm

Saxon-Woods potential shape assumed

$$S_{\text{tot}}(r) = \sum_i |\chi_i|^2$$

$S_p(r) \rightarrow$  Saxon-Woods parameters  
adjusted to fit  $\langle r^2 \rangle$   
measured for the charge  
distribution



Sub-Coulomb Pickup on  $^{208}\text{Ti}$   
(d, t) (p, d)

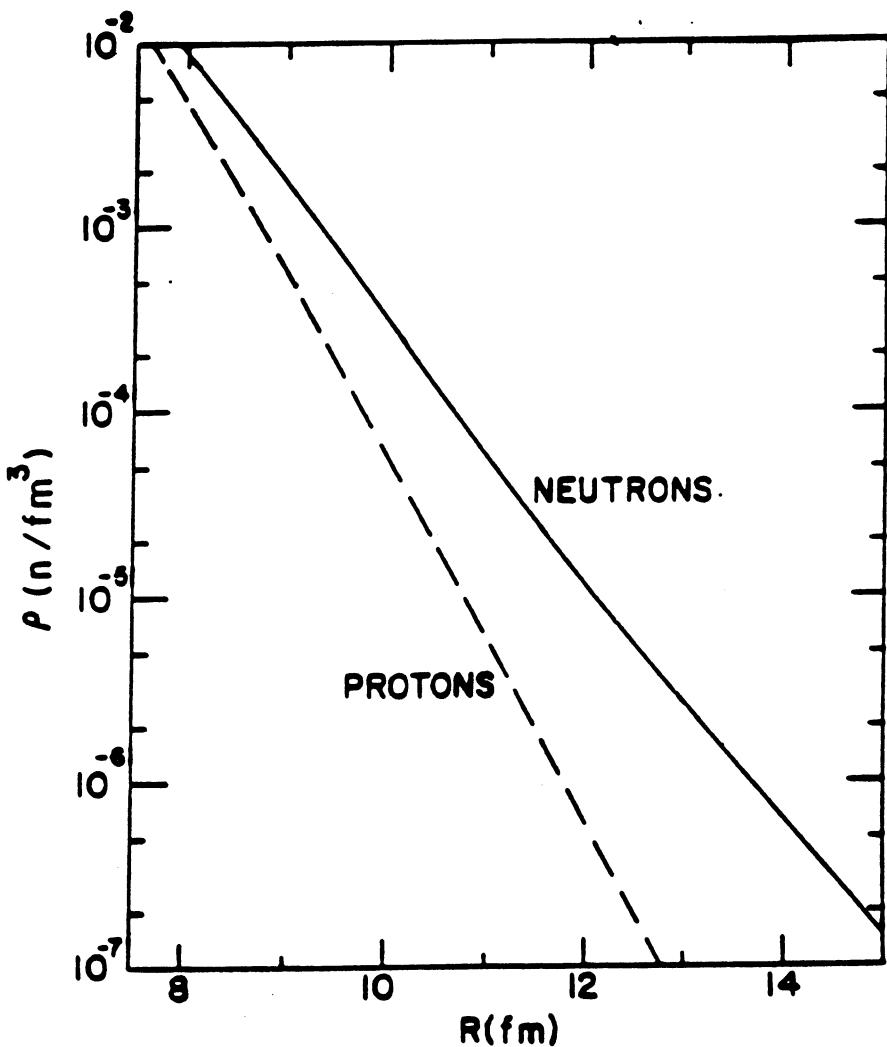


FIG. 1. Neutron and proton densities, derived as described in the text.

H. J. Körner and J. P. Schiffer

Phys. Rev. Lett. 27 (1971) 1457

L $\bar{P}$

W. M. Bugg et al.

Phys. Rev. Lett.  
31 (1973) 475

BNL 30" hydrogen bubble chamber

stopped  $\bar{p}$  on C, Ti, Ta, Pb

total charge of mesonic prongs investigated.

$\bar{p}p$  net charge : 0

$\bar{p}n$  -1

$$f_{\text{halo}} = \frac{N(\bar{p}n)}{N(\bar{p}p)} \frac{Z_e}{N_e} \frac{\text{Im } a_p}{\text{Im } a_n}$$

$^{12}C$

$f_{\text{ex def}} = 1.0$

$^{208}Pb$

$f = 2.3 \pm 0.5$

multiple corrections necessary

K<sup>-</sup>

D. H. Wilkinson Phil Mag. 4 (1959) 215

D. H. Davis et al. N. Phys. B1 (1967) 431

E. H. Burhop N. Phys. B1 (1967) 43

## Nuclear emulsions

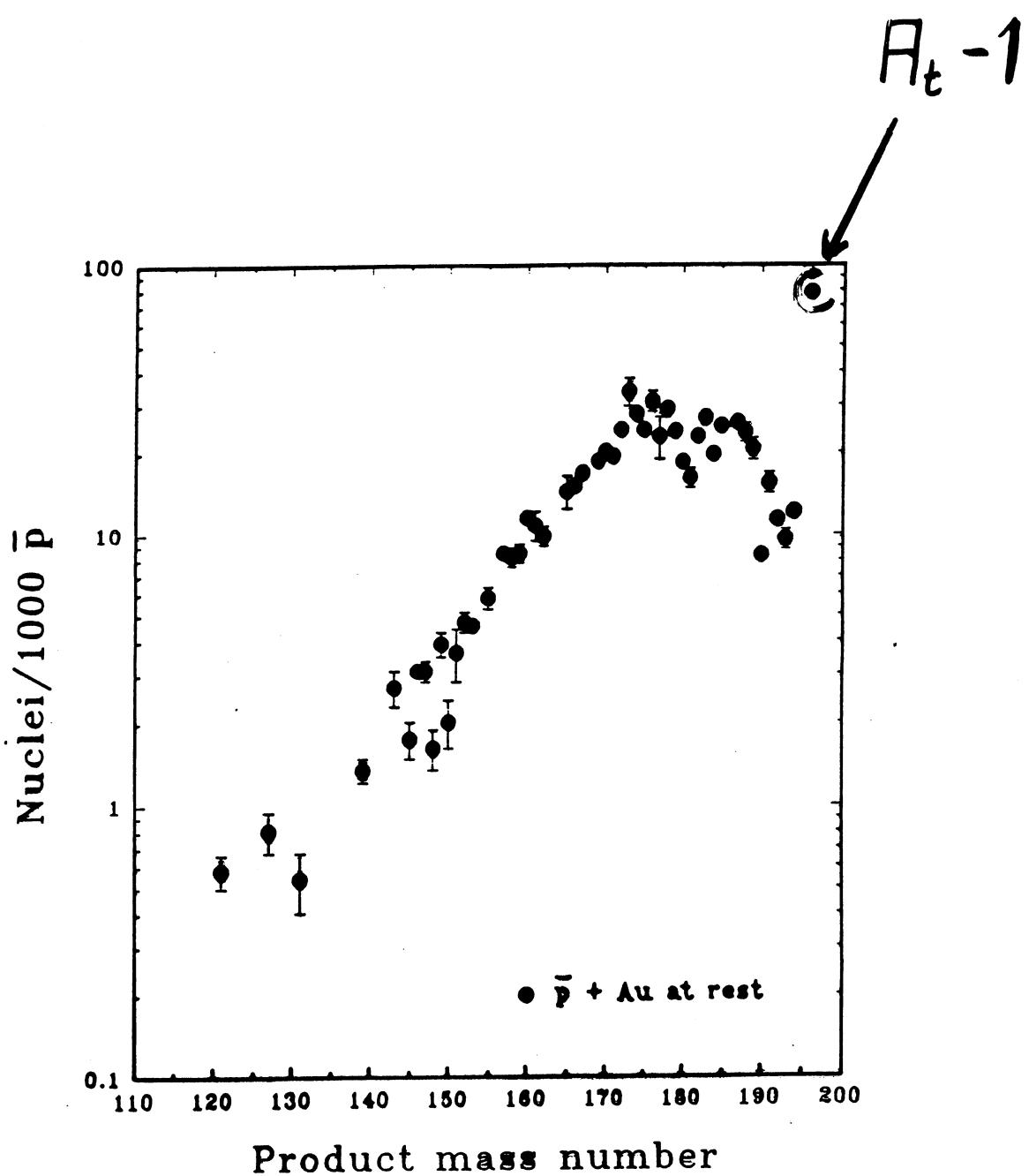
Similar principle as aniprotonic  
experiment of Bugg



A. Bohr , B. R. Mottelson

## Nuclear Structure

"In heavy nuclei, because of the Coulomb forces and associated neutron excess, there may be some difference between the density distributions of neutrons and protons. Tentative evidence bearing on this interesting problem is discussed ... "

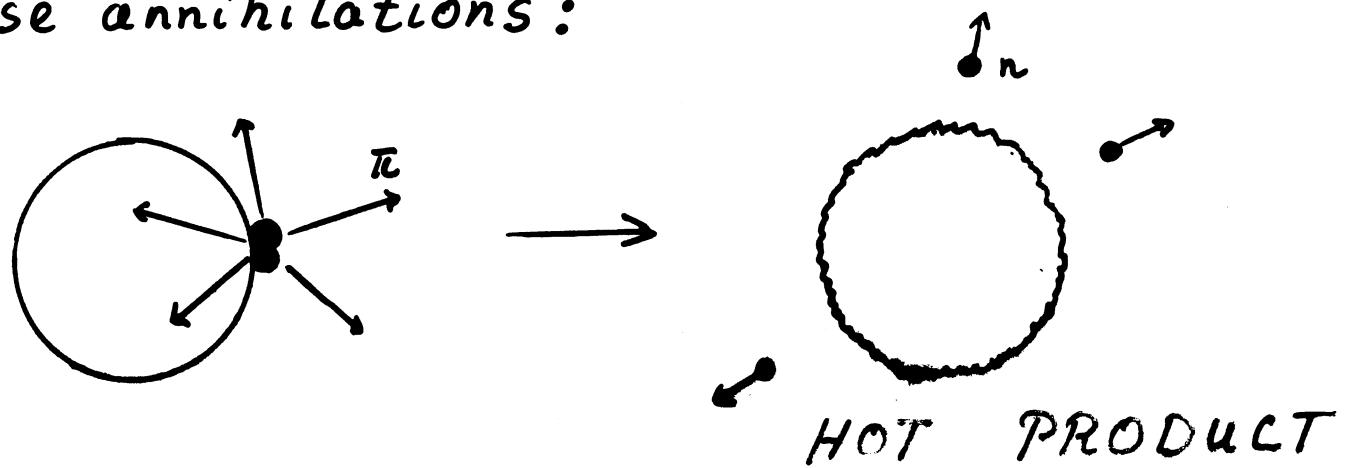


P.S 203

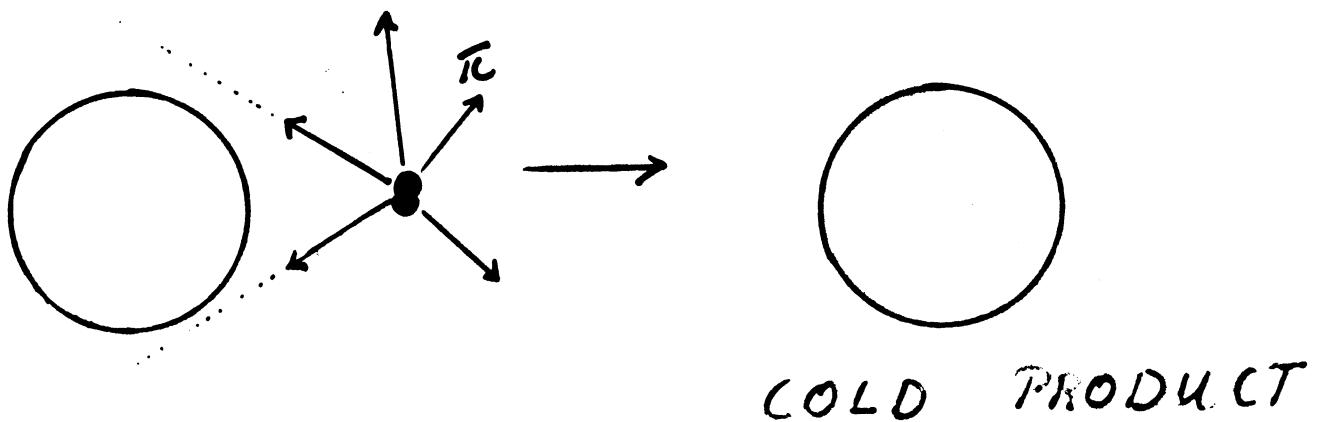
Physics of the phenomenon of  
 $A_{\text{target}} - 1$

Production : void cascade events

Close annihilations:



Distant annihilations:



If, in a distant annihilation  $\bar{\rho}$  encounters a neutron :

produced nucleus is  $N_t - 1$

encounters a proton :

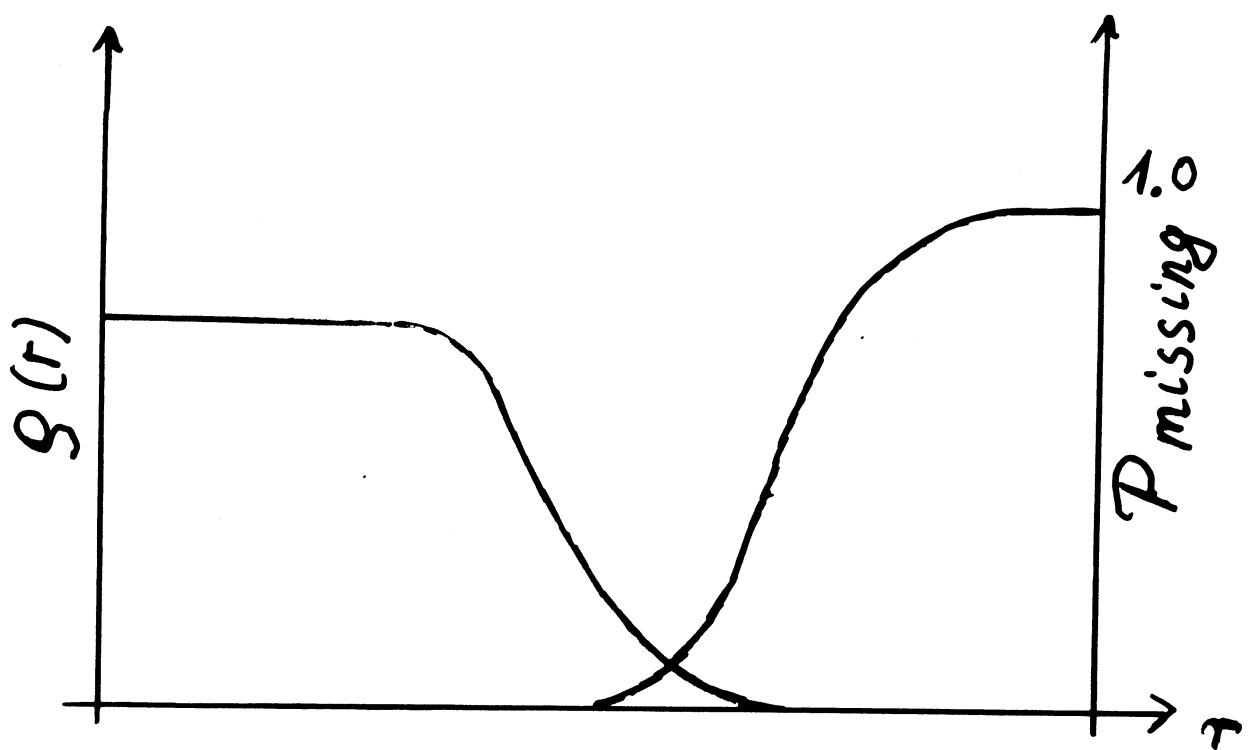
produced nucleus is  $Z_t - 1$

if  $(N_t - 1)$  and  $(Z_t - 1)$  products are both radioactive their yield can be determined by the radiochemical methods.

The ratio of their yields :

$$\frac{Y(N_t - 1)}{Y(Z_t - 1)} = f\left(\frac{S_n}{S_p}\right)$$

for  $r$  where  $P_{nuc} \neq 0$



<b>U 232</b> 70.0 a .. 5.320; 5.263... SF Y(58) 129... e n 73.1; or 75.2	<b>U 233</b> $1.592 \cdot 10^4$ a .. 4.824; 4.783... Y(42; 97...) e e 47.7; or 531.1	<b>U 234</b> 0.005 2.445 $\cdot 10^3$ a .. 4.773; 4.723... Y(43; 121) e e 48.2; m < 0.05	<b>U 235</b> 0.720 .. 7.026; 6.979 .. 4.400 .. 106 .. 100/11; 98.38 .. 583.34	<b>U 236</b> $2.342 \cdot 10^4$ a .. 4.490; 4.449 .. 15 .. 109; 1131 .. 57	<b>U 237</b> 6.75 d .. 0.2... Y 60; 208... e n 411; m < 0.35	<b>U 238</b> 99.275 195 ms; 4.448; 10 <sup>-3</sup> .. 4.19; .. 2514; 2515 1870... e d; .. 4270
<b>Pa 231</b> $3.276 \cdot 10^4$ a .. 3.014; 4.952 5.020 .. 27.303; 300 e n 210; m 0.010	<b>Pa 232</b> 1.31 d .. 0.3; 1.3 .. 900; 884; 190 e e 700; m 700	<b>Pa 233</b> 27.0 d .. 0.3; 0.6 .. 312; 300; 341 e e 21; 20; m - 0.1	<b>Pa 234</b> 1.17 m .. 23 .. 1001 .. 12 .. 174; .. 1703 .. 500; .. 4900	<b>Pa 235</b> 24.2 m .. 1.4... .. 120; 659 m	<b>Pa 236</b> 9.1 m .. 2.9; 3.1... Y 642; 637; 1703... Np; q	<b>Pa 237</b> 8.7 m .. 1.4; 2.3... Y 644; 636 529; 541...
<b>Th 230</b> $7.54 \cdot 10^4$ a .. 4.600; 4.621 .. 168; 144 e .. 23.2 m .. 0.0012	<b>Th 231</b> <u>25.5 h</u> .. 0.3; 0.4... Y 26; 84 e	<b>Th 232</b> 100 .. 405 $\cdot 10^{14}$ a .. 4.013; 3.964 .. 1001 .. 7.40; .. 6.000000000000000	<b>Th 233</b> 22.3 m .. 1.2 .. 87; 29; 459 .. 1500; m 15	<b>Th 234</b> 24.10 d .. 0.2... .. 92.63 e m .. 1.8; m - 0.01	<b>Th 235</b> 6.9 m .. 1.0... .. 418; 932	<b>Th 236</b> 37.1 m .. 10... .. 111; 113
<b>Ac 229</b> 62.7 m .. 1.1 .. 165; 549 262; 146; 133	<b>Ac 230</b> 122 s .. 27... .. 455; 508 1244 e	<b>Ac 231</b> <u>7.5 m</u> .. 282; 307 281; 106; 309	<b>Ac 232</b> 35 s ..			

140

142

144

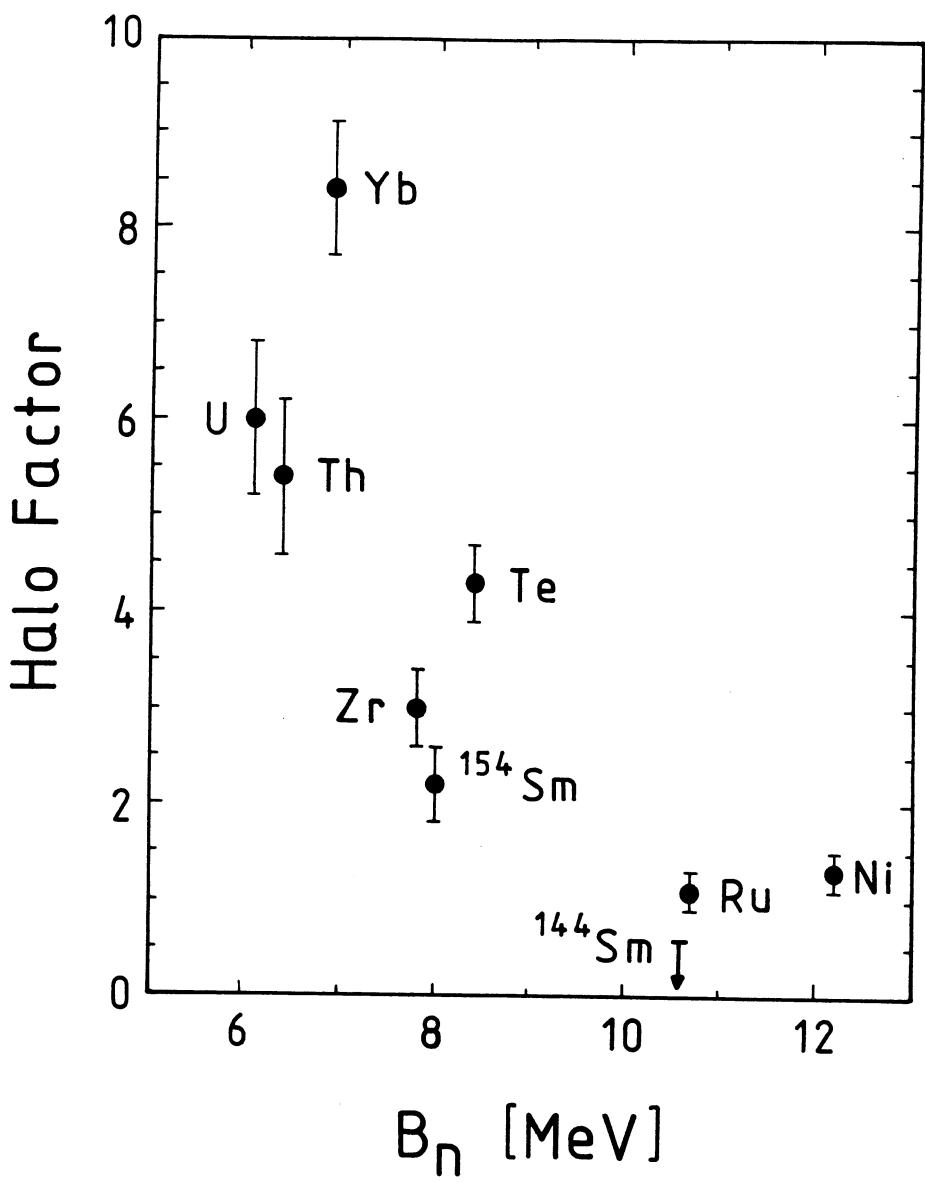
146

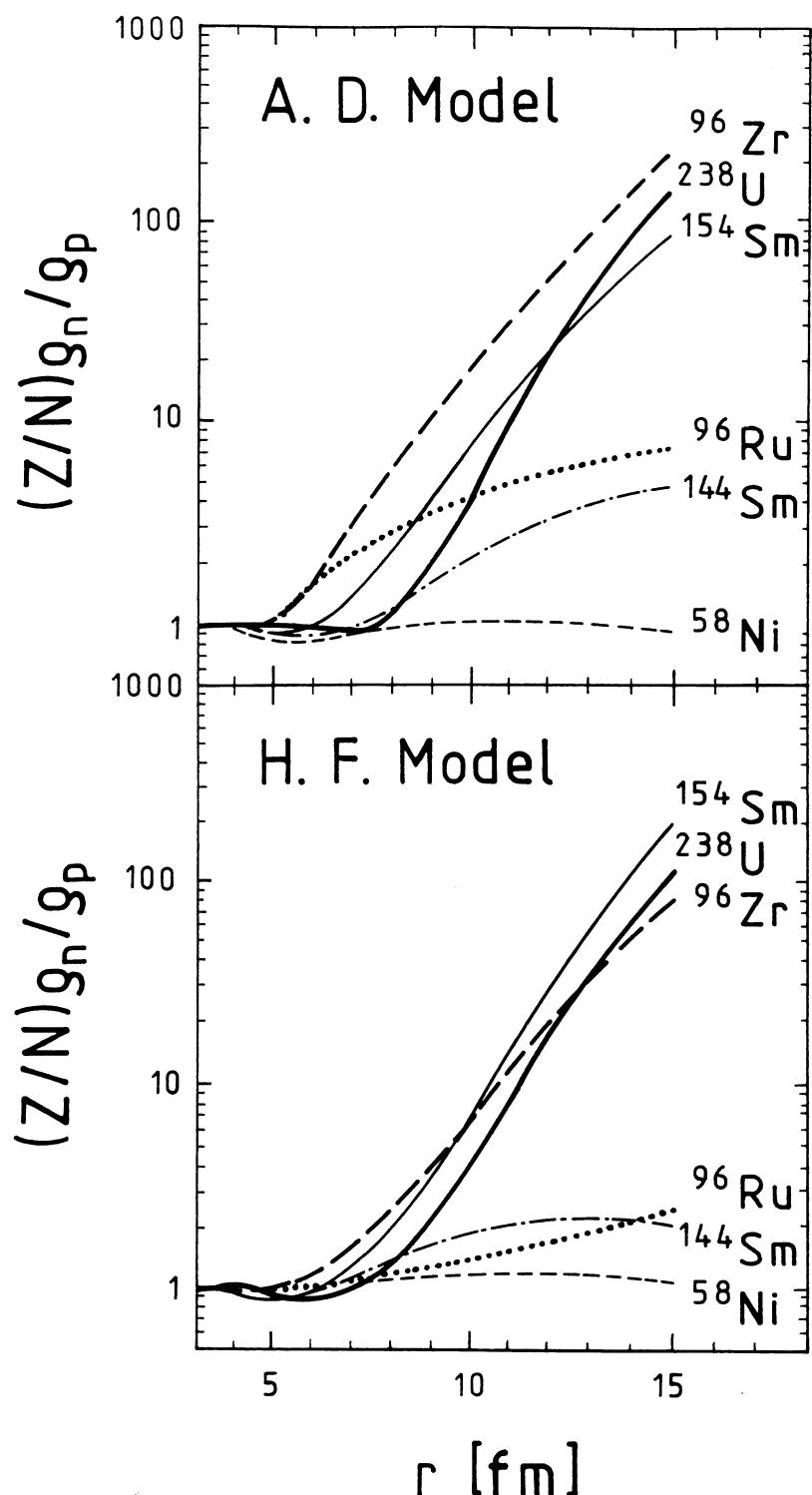
$$f_{\text{halo}}^{\text{peripheral}} = \frac{N(\bar{p}n)}{N(\bar{p}p)} \frac{Im a_p}{Im a_n} \frac{Z_t}{N_t}$$

$$\frac{1}{(0.63)}$$

## Halo factor

expresses the (normalized) ratio  
 of antiproton encounters with neutrons  
 to encounters with protons in the orbitals  
 from which absorption takes place





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DEPT: Heavy Ion Laboratory PHONE # 22-222-123

FROM: Physical Review Letters MS. CODE #: LW5096

TITLE: Neutron halo in heavy ...

AUTHORS: P Lubinski et al.

Dear Dr. Jastrzebski:

The above manuscript has been accepted for publication in Physical Review Letters; there are no reviewer comments to address. The paper is now being prepared for the printer. Note that proofs are NOT sent. If any questions arise in the production process we will attempt to contact you.

Thank you for your cooperation.

Sincerely,

Editors  
Physical Review Letters

AMERICAN PHYSICAL SOCIETY EDITORIAL OFFICE FAX # 516-924-5294

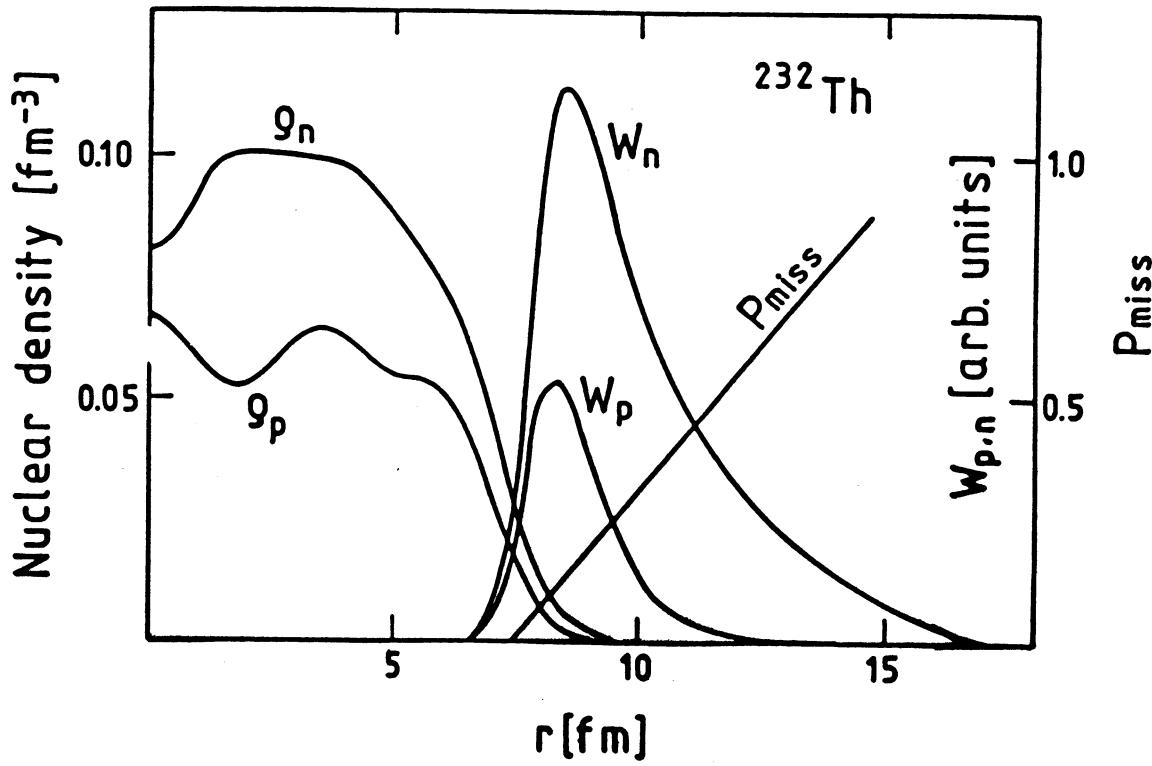


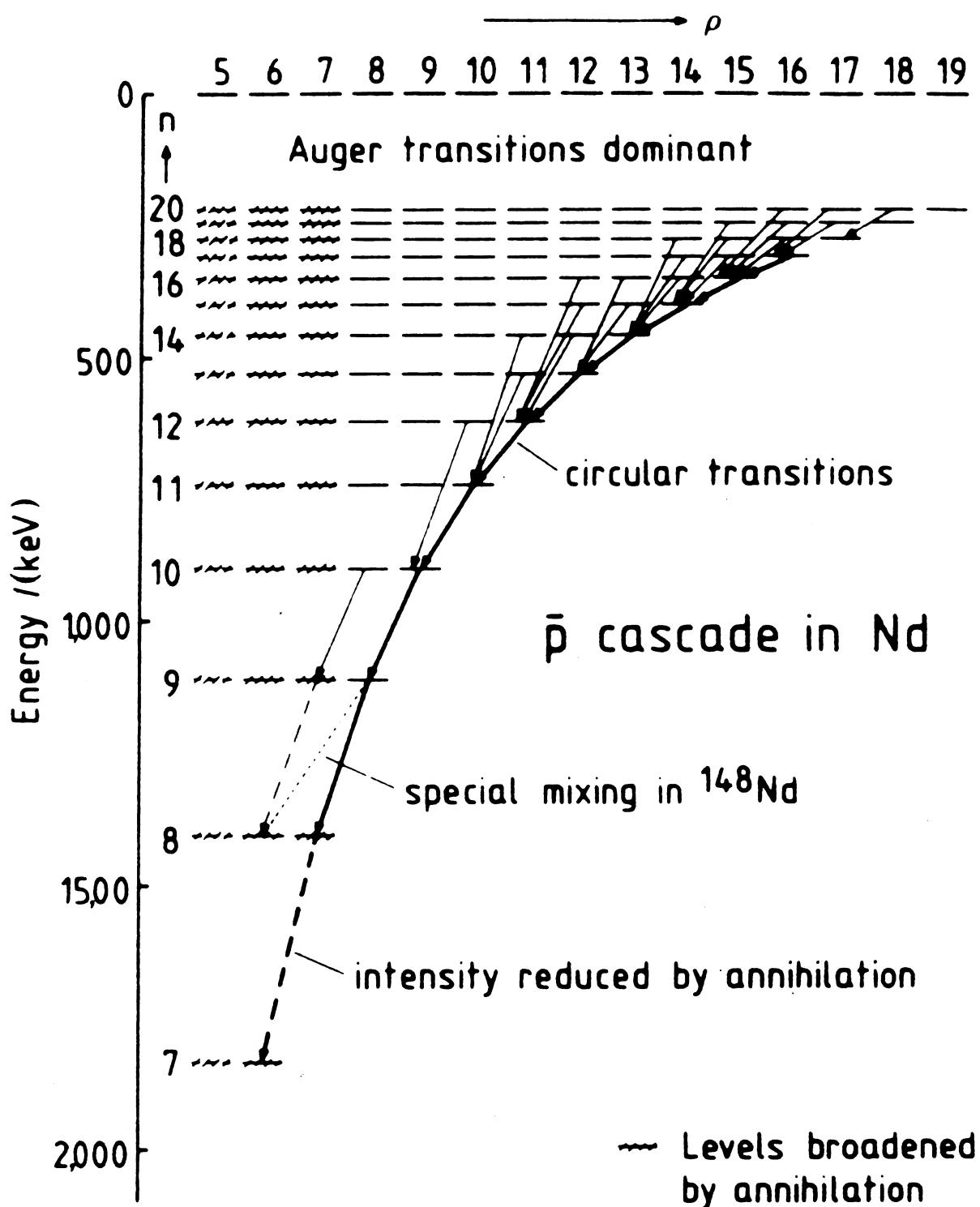
Figure 5. Neutron and proton densities,  $\rho_n$  and  $\rho_p$  (see also Fig. 3), the anti-absorption probability on neutrons,  $W_n$ , and on protons,  $W_p$ , and the "missing probability"  $P_{\text{miss}}$  (from Ref. 27, see also text).

$$\frac{P}{2} = \int W(r) dr = \int \text{Im } V(r) / |\chi_{n,e}(r)|^2 r^2 dr$$

$$V(r) = \frac{2\pi}{\mu} a \, g(r)$$

$$a = (1.53 \pm 0.27) + i (2.50 \pm 0.25) \text{ fm}$$

(fit to light nuclei X-ray data)



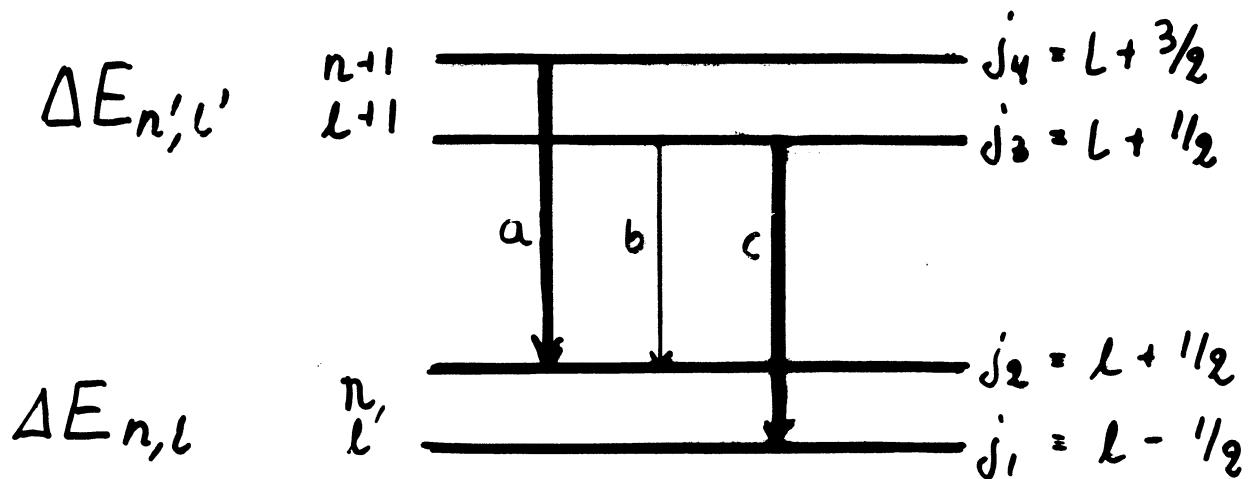
Antiprotonic levels in the Nd atom.

Fine structure splitting:

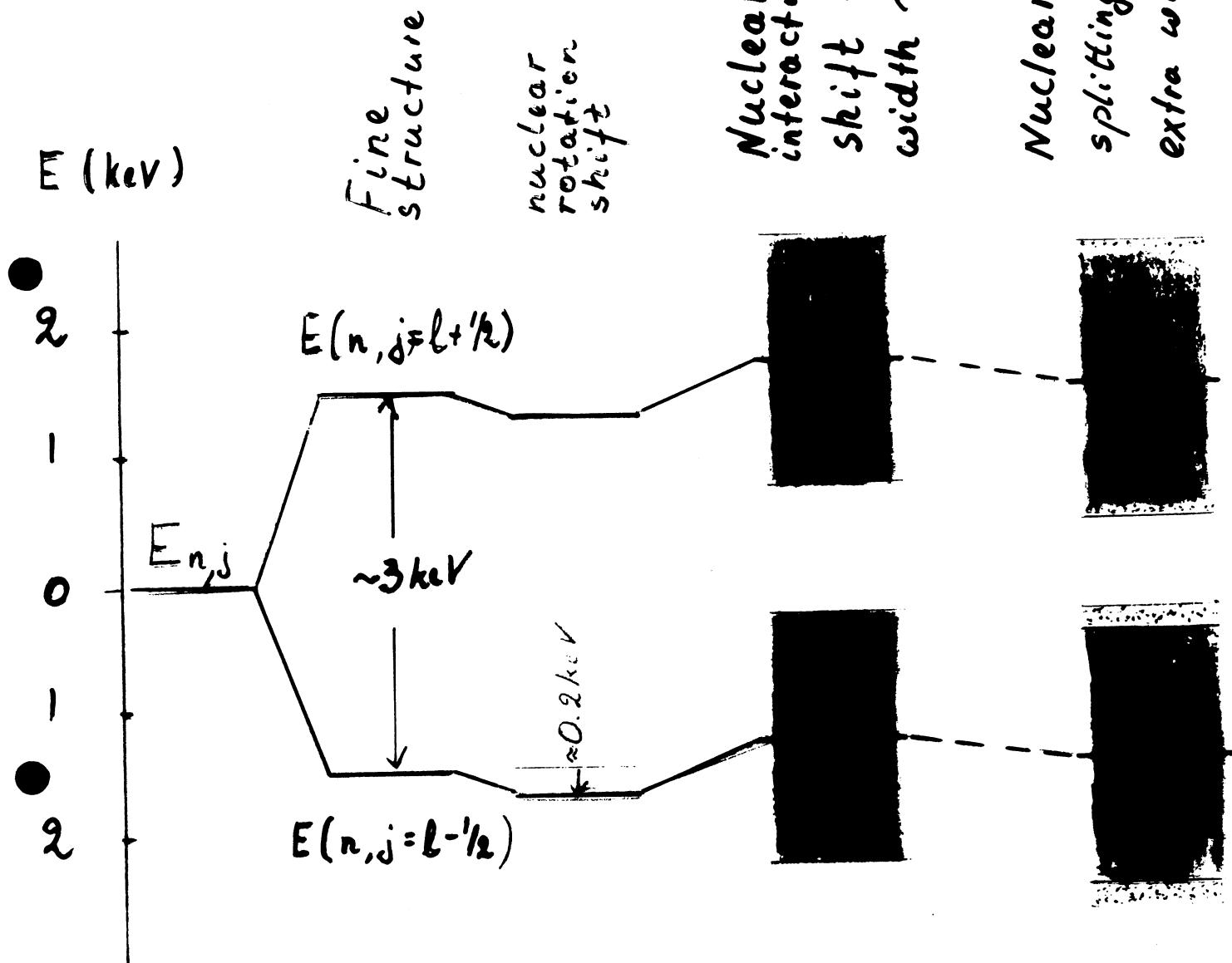
$$\Delta E_{n,l} = (g_0 + 2g_1) \frac{(\alpha Z)^4}{2n^3} \frac{\mu}{l(l+1)}$$

$$g_0 - \text{Dirac} = -1$$

$g_1$  - anomalous g factor (Pauli)



# Lower level structure



Electromagnetic -  
calculable

Nuclear -  
to measure

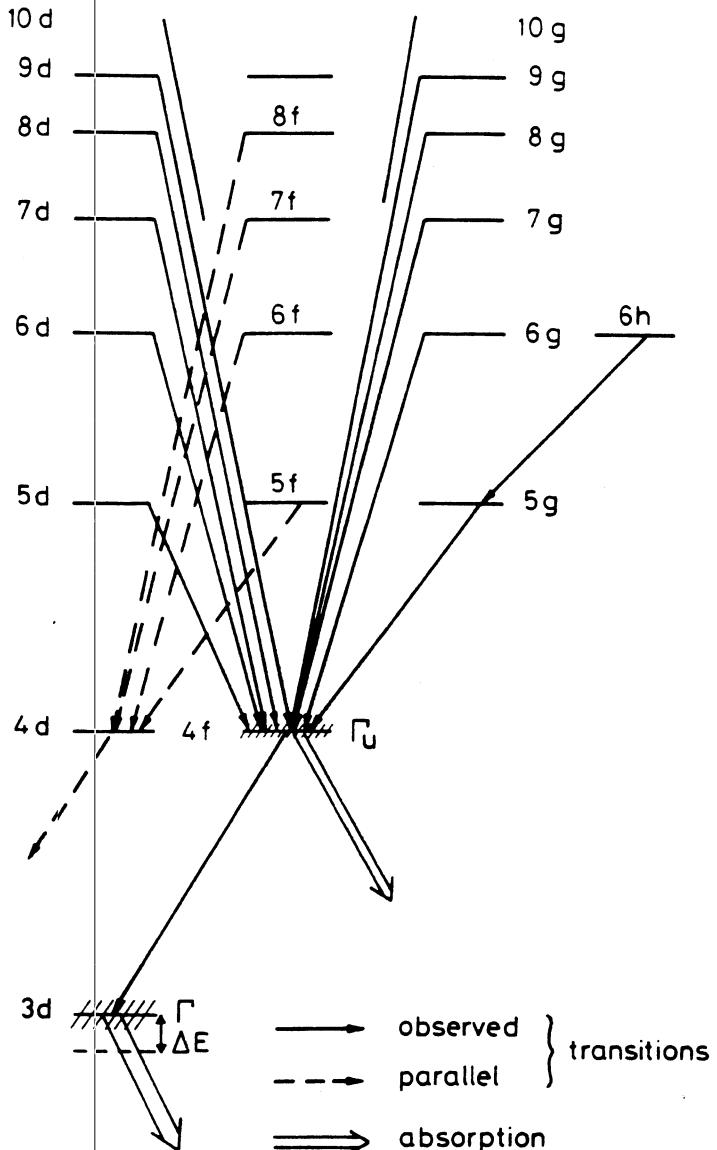


Fig. 1. Antiprotonic atom level scheme showing how the width ( $\Gamma_u$ ) of the upper level can be determined.

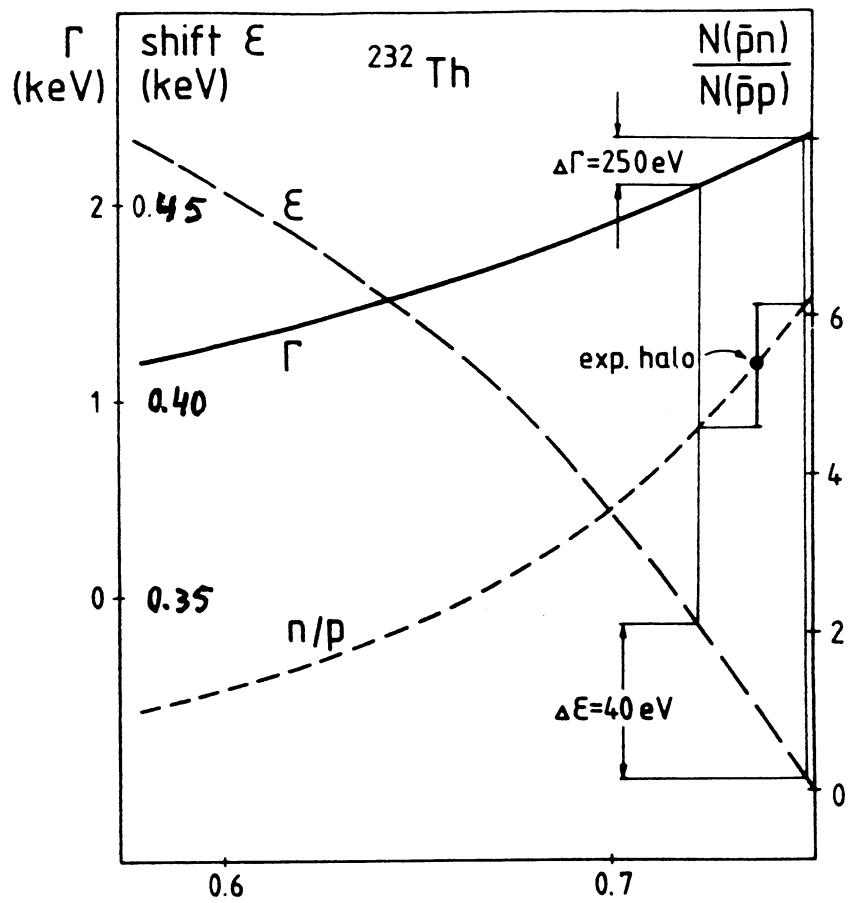
Observables of the proposed experiment :

Lower level shift  $\epsilon$

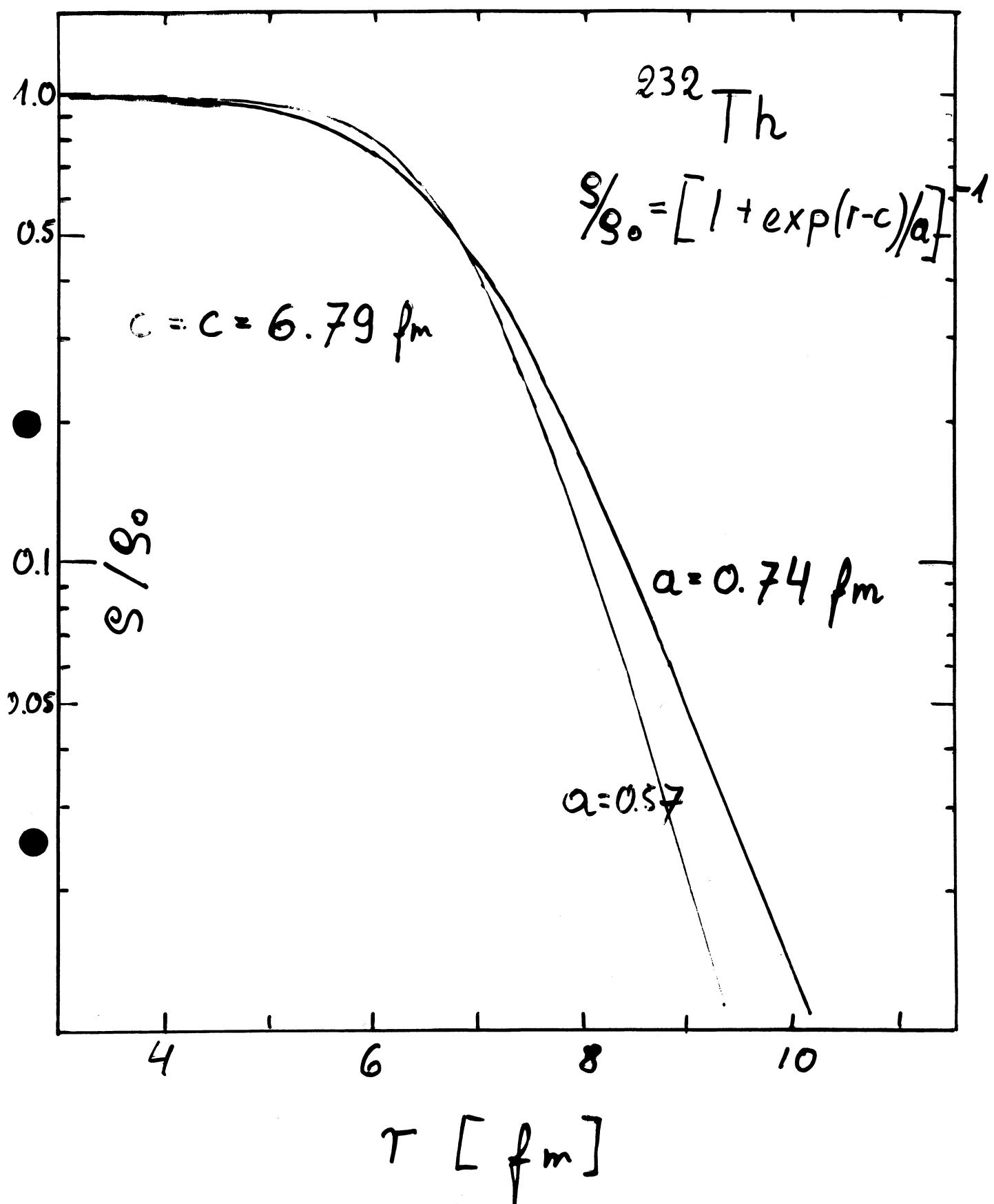
Lower level width  $\Gamma_{low}$

Upper level width  $\Gamma_{up}$

Absolute transition intensities  $I_x$



Neutron surface parameter  $a_n$   
 [fm]



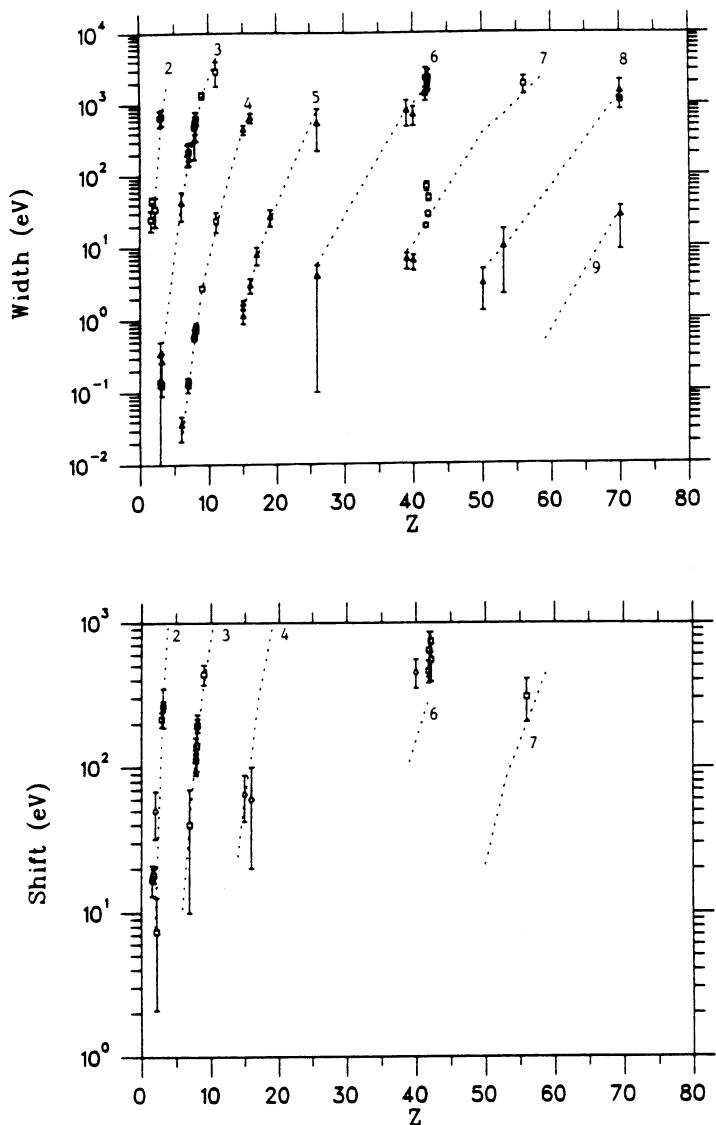
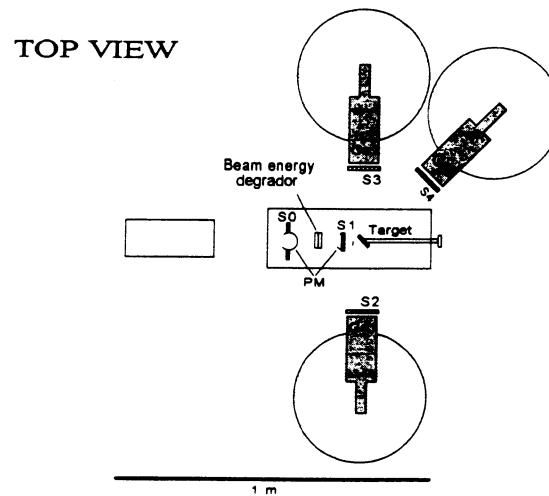
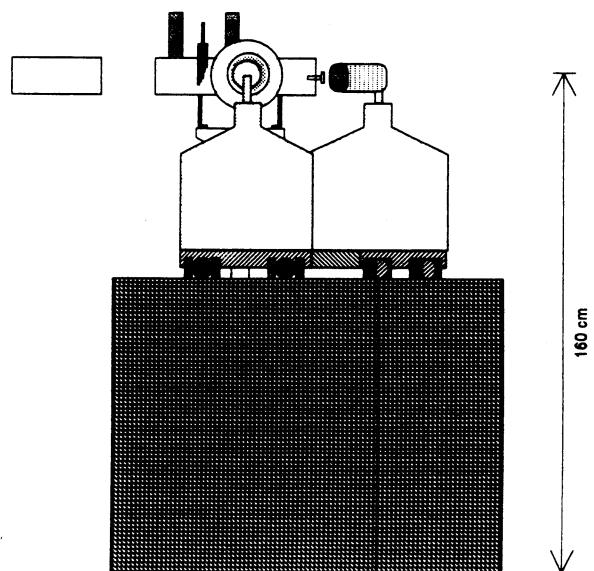


Fig. 4. Measured shifts (with reversed sign) and widths for  $\bar{p}$ -atoms.  
The dashed curve shows values calculated using the simple  
optical model (eq. (5)) with  $\bar{a} = 1.53 + i2.50$  fm.

Batty, 1990



**SIDE VIEW**



**Fig. 6. The general view of the main parts of the experimental setup.**

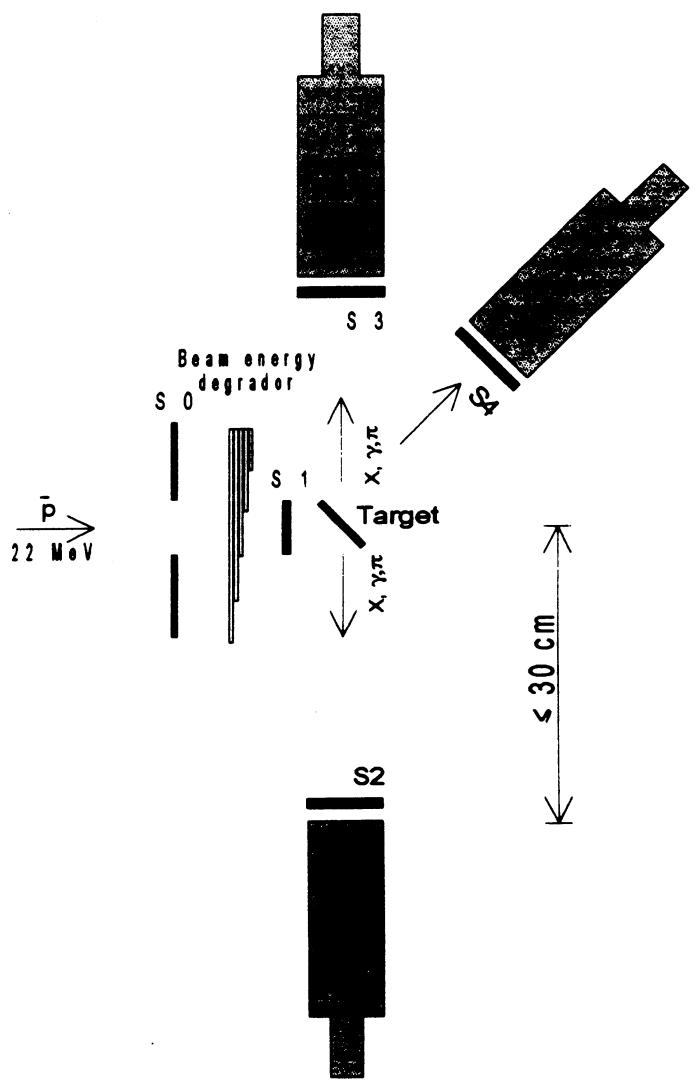
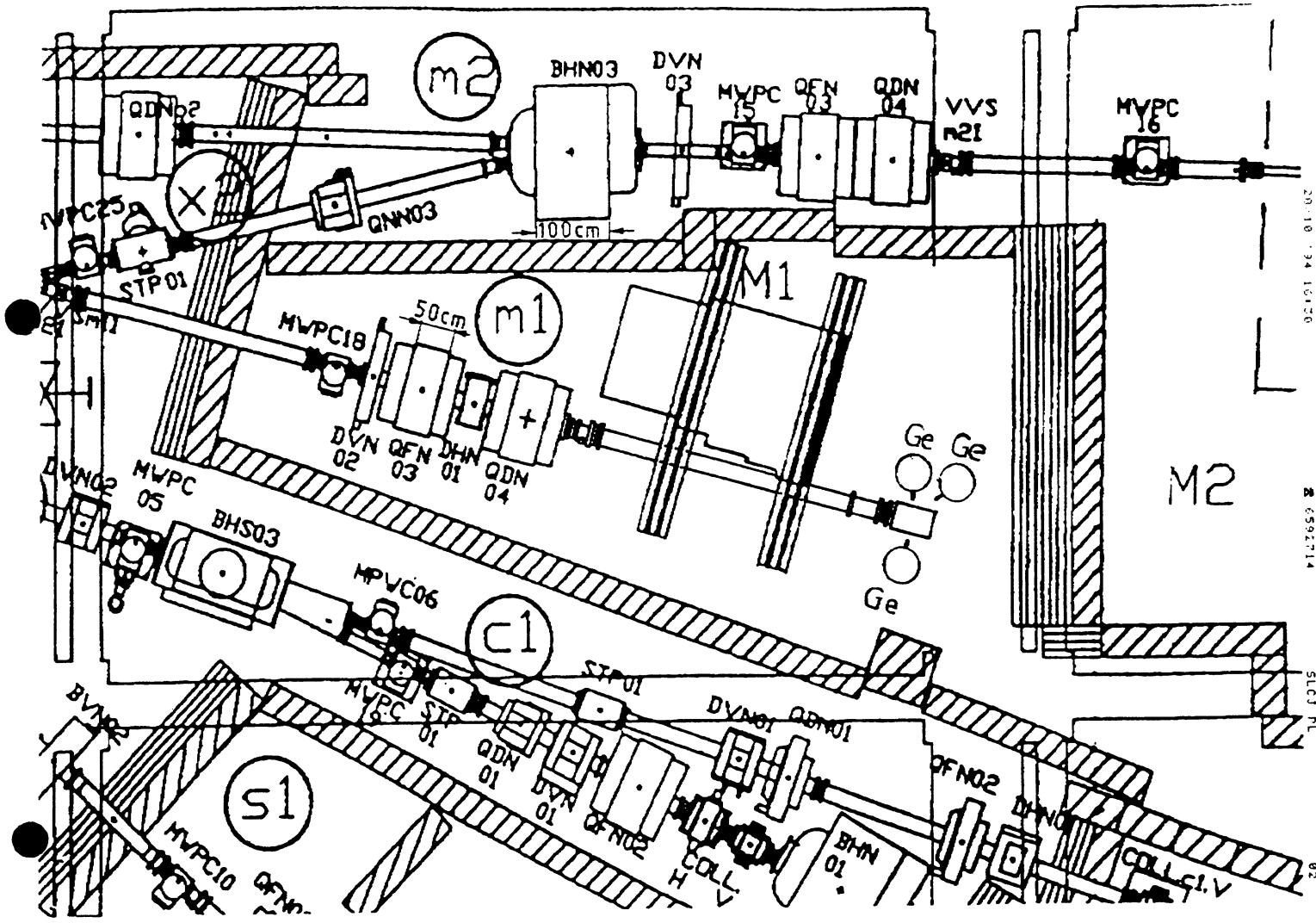


Fig. 7. The schematic view of the target - counter ensemble.



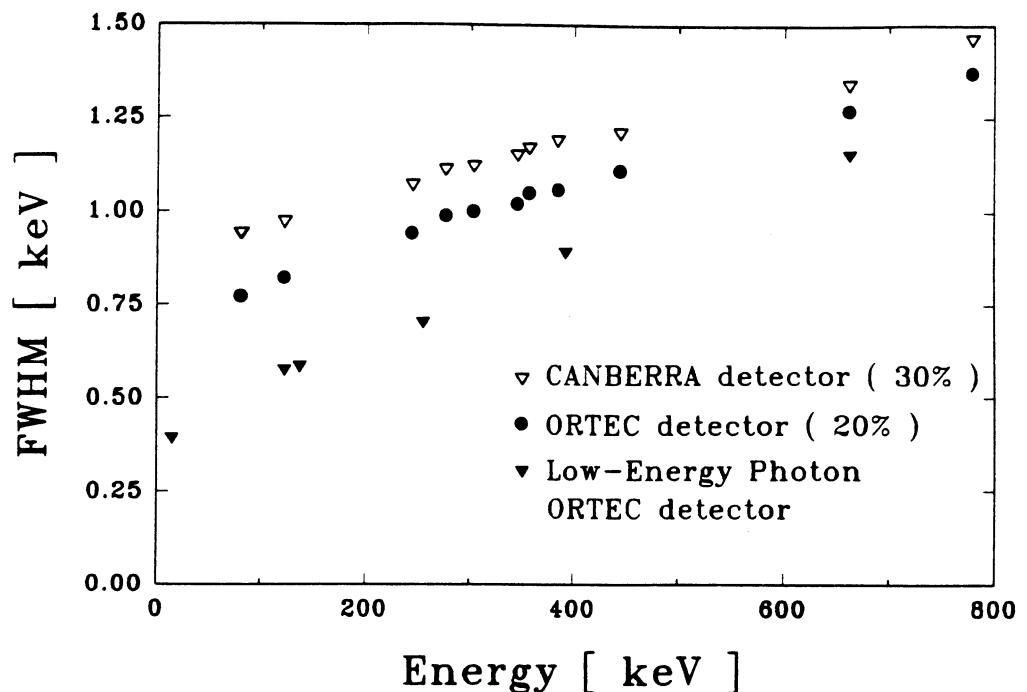


Fig. 11. Photo-peak resolution (FWHM) as a function of the gamma-ray energy for three available detectors.

The following table compares the estimate for gamma-X ORTEC and Low-Energy Photon ORTEC detectors previously described:

	Gamma-X ORTEC detector (Distance 30 cm)	Low-Energy Photon ORTEC detector (Distance 10 cm)
Detector surface	1892 mm <sup>2</sup>	491 mm <sup>2</sup>
Detector thickness	49 mm	13 mm
Solid angle (% 4Π)	0.16	0.34
γ-ray efficiency (200 keV)	1.0	0.8
π efficiency	1	1
p efficiency	1	1
π energy loss (E <sub>π</sub> = 200 MeV)	37 MeV	10 MeV
p energy loss (E <sub>p</sub> = 50 MeV)	50 MeV	50 MeV
counting rates (10 <sup>5</sup> $\bar{p}$ stopped in the target)		
X-rays	1860/sec	2860/sec
π	400/sec	850/sec
p	160/sec	340/sec
Energy deposition rate	23 000 MeV/sec	25 000 MeV/sec
Preamplifier rate limit (from producer)	140 000 MeV/sec	≈ 30 000 MeV/sec
γ Peak /Total ratio (200 keV)	0.55	0.32
Single photopeak counting rate	102/sec	92/sec

About 10<sup>6</sup> counts in a photopeak will be collected during 5 h of measuring time with 10<sup>5</sup>  $\bar{p}$ ps. Pions and charged particles will overload the X and γ-X detectors. For both detectors the recovery time will be comparable so the γ-X detector will perform better due to lower rate. However energy resolution is better for the X detector by about 25 %. Plastic scintillators in front of detectors will prevent triggering data acquisition during overload (anticoincidence).

# Expected results

- Fix the atomic cascade and determine states of the nuclear capture in nuclei where the neutron halo was detected - precision of the annihilation site and absorption probability distribution
- From shifts and widths in a number of targets determine the phenomenological  $\bar{p}$  Nucleus optical potential in heavy nuclei  $a_R, a_i, (r_0)_R, (r_0)_i$
- Use the neutron halo information as a supplementary constraint on this potential.

## OBSERVATION OF ANTIPIROTOMIC ATOMS

A. BAMBERGER, U. LYNEN, H. PIEKARZ\*, J. PIEKARZ\*\*, B. POVH and H. G. RITTER  
*Max-Planck-Institut für Kernphysik, Heidelberg, Germany*

*and CERN, Geneva, Switzerland*

and

G. BACKENSTOSS, T. BUNACIU, J. EGGER\*\*\*, W. D. HAMILTON ‡ and H. KOCH  
*Institut für Experimentelle Kernphysik der Universität und des Kernforschungszentrums,  
Karlsruhe, Germany*  
*and CERN, Geneva, Switzerland*

Received 28 August 1970

X-rays originating from the antiprotonic cascade have been observed for several elements. From the energies of the measured antiprotonic transitions of  $^{81}\text{TL}$ , the mass of the antiproton could be determined to be  $938.3 \pm 0.5$  MeV. With this result the masses of the antiproton and the proton are equal within 0.5 MeV with a confidence of 68 %. The magnetic moment of the antiproton shows up in a broadening of the  $n = 10 \rightarrow n = 9$  transition.

In this letter we report the first measurement of X-rays originating from the atomic transitions of antiprotons bound in a nuclear Coulomb field.

Antiprotons stopped in a target will finally be captured into Bohr orbits of the target nuclei. This capture is expected to take place into orbits with high main quantum numbers. Subsequently, the antiprotons will cascade down to lower orbits releasing their energy by emission of either Auger electrons or X-rays. For sufficiently low quantum numbers, the overlap between the atomic antiproton wave function and the nucleus becomes so large that nuclear absorption dominates the radiative processes and no further X-rays are observed. According to perturbation calculations for nuclei in the oxygen region, the cascade is expected to terminate with the transition  $n = 4 \rightarrow n = 3$  at about 70 keV, whereas in heavy nuclei such as lead the transition  $n = 9 \rightarrow n = 8$  with an energy of 550 keV should be the last with a strongly reduced yield. The energies of these X-rays can be well calculated because the nuclear interaction is important only for the lowest populated levels.

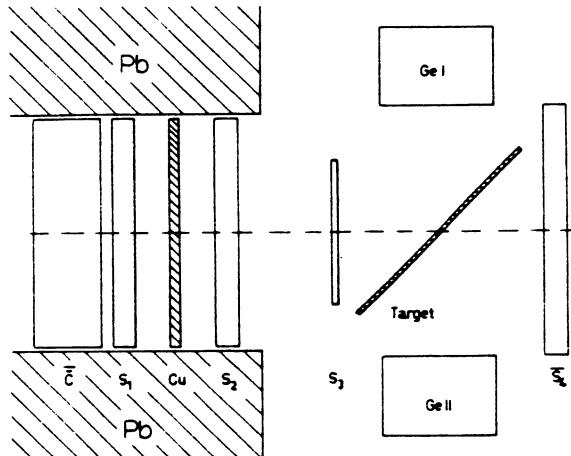


Fig. 1. Experimental set-up of the beam telescope, the target and two Ge(Li)-detectors.

A low momentum separated beam has been constructed at the CERN Proton Synchrotron for experiments with stopped particles [1]. In this beam about 300 antiprotons can be stopped in a  $4 \text{ g/cm}^2$  thick target with  $7 \times 10^{11}$  protons hitting the 6 cm long tungsten production target of the slow ejection. The antiproton beam is strongly contaminated with  $10^3$  passing pions per stopped antiproton. A careful identification of the antiprotons is made with the beam tele-

\* On leave of absence from: Institute of Nuclear Research, Warsaw, Poland.

\*\* On leave of absence from: Institute of Experimental Physics, University of Warsaw, Warsaw, Poland.

\*\*\* Visitor from ETH Zürich, Switzerland.

† Present address: University of Sussex, Brighton, England.