The submitted manuscript has been authored by a contractor of the U. S. Government under contract No. W-31-109-ENG-38. Accordingly, the U. S. Government retains a nonexclusive, royalty-free license to publish or reproduce the published form of this contribution, or allow others to do so, for U. S. Government purposes.

ANL-HEP-CP-95-10

For the Proceedings of the 11th Int. Symposium on High Energy Spin Physics and 8th Int. Symposium on Polarization Phenomena in Nuclear Physics Indiana University, Bloomington, IN 15-22 Sep. 1994

Drell-Yan Pairs, W^{\pm} and Z^{0} Event Rates and Background at RHIC

A. A. Derevschikov¹, V. L. Rykov^{1,2}, K. E. Shestermanov¹ and A. Yokosawa³

¹ Experimental Physics Division, Institute of High Energy Physics, Protvino, Moscow District 142284, Russia

² Department of Physics and Astronomy, Wayne State University, Detroit, Michigan 48202, USA

³ High Energy Physics Division, Argonne National Laboratory, Argonne, Illinois 60439, USA

Abstract. The estimates for the Drell-Yan pairs, W^{\pm} and Z^{0} acceptances and event rates in the STAR and PHENIX detectors at RHIC are presented. The background to W^{\pm} in STAR evaluated. The results were obtained by Monte-Carlo simulations with the PYTHIA/JETSET and GEANT programs.

Introduction

The measurements of Drell-Yan lepton pairs, W^{\pm} and Z^{0} production in polarized proton collisions appear to be the only way to determine polarization properties of the sea quarks. Moreover, the intermediate bosons are produced due to the parity violating mechanism providing an unique opportunity to study a parity violation effects in the hard hadron interactions. All these phenomena have been discussed in details elsewhere (see, for example, (1-4)), and a quintessence of these discussions is reflected in the Proposal on Spin Physics at RHIC (5).

In this report we provide the event rate estimations (6) for the processes above for the two major detectors, STAR and PHENIX. We also present the first results of the background study for the W^{\pm} detection in STAR (7). All data were obtained by Monte-Carlo simulations with the PYTHIA/JETSET and GEANT~V3.15 programs. The EHLQ1 (8) set of proton structure functions has been used. The parameters for PYTHIA were tuned up in order to produce results in a reasonable agreement with the available experimental data from E772, ISR, UA1, UA2 and CDF (see (6) and referencies therein).

Work supported in part by the U.S. Department of Energy, Division of High Energy Physics, Contract W-31-109-ENG-38.



SW 9591

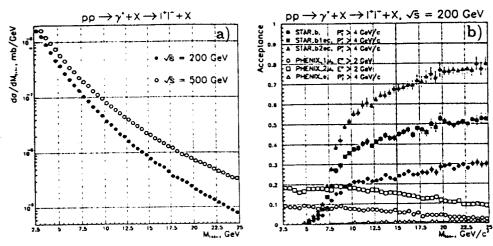


Figure 1: Differential Drell-Yan pairs production cross section in pp collisions (a) and STAR and PHENIX acceptances (b).

Acceptances and Event Rates

The differential cross sections for Drell-Yan pairs production at $\sqrt{s}=200$ and 500~GeV are shown in Fig. 1a. For W^\pm and Z^0 , it has been assumed that only lepton decay modes are detected, i.e. one high P_T lepton, for $W^\pm \to l^\pm \nu$, and two high P_T leptons, for $Z^0 \to l^+ l^-$, where l^\pm are either electron or muon. PYTHIA~V5.6 provides the following estimates for W^\pm and Z^0 production cross sections in pp interactions at $\sqrt{s}=500~GeV$:

$$\sigma \cdot B(pp \to W^+ + X \to l^+\nu_l + X) = 120 \text{ pb};$$

$$\sigma \cdot B(pp \to W^- + X \to l^-\overline{\nu}_l + X) = 43 \text{ pb};$$

$$\sigma \cdot B(pp \to Z^0 + X \to l^+l^- + X) = 10 \text{ pb}.$$

The integrated luminosity, during the exposition time of $4 \cdot 10^6$ seconds (100 days, 50% efficiency), has been taken as:

```
- At \sqrt{s} = 500 \, GeV, \int L \, dt = 800 \, pb^{-1} (L = 2 \cdot 10^{32} \, cm^{-2} \cdot sec^{-1}); - At \sqrt{s} = 200 \, GeV, \int L \, dt = 320 \, pb^{-1} (L = 8 \cdot 10^{31} \, cm^{-2} \cdot sec^{-1}).
```

The STAR and PHENIX setups had been described elsewere (9,10). Here we considered six geometries for the detection single leptons and lepton pairs: STAR_b: STAR detector with the Barrel Electromagnetic Calorimeter (EMC) of the acceptance (fiducial): $-0.95 < \eta < 0.95$ and 2π coverage in φ ; STAR_blec: STAR detector with the Barrel + one End Cup EMC of the acceptance (fiducial): $-0.95 < \eta < 1.9$ and 2π coverage in φ ;

STAR_b2ec: STAR detector with the Barrel + two End Cup EMC of the acceptance (fiducial): $-1.9 < \eta < 1.9$ and 2π coverage in φ ;

PHENIX_1 μ : PHENIX one-arm muon detector of the summary acceptance (geometry): 1.1 < η < 2.5 and 2π coverage in φ ;

PHENIX_2μ: PHENIX two-arm muon detector of the summary acceptance

Table 1: Drell-Yan pairs event rates, integrated over mass intervals $M_1 < M_{l+l-} < M_2$, at $\sqrt{s} = 200 \, GeV$ for $\int L \, dt = 320 \, pb^{-1}$; Cuts: $P_T^e > 4 \, GeV/c$ or $E_\mu > 2 \, GeV$.

| $M_1 - M_2$, GeV/c^2 | 2-5 | 5-9 | 9-12 | 12-15 | 15-20 | 20-25 | 2-25 |
|-------------------------|---------|--------|--------|-------|-------|-------|---------|
| STAR_b | 550 | 11,500 | 7,200 | 2,900 | 1,800 | 600 | 24,500 |
| STAR_blec | 800 | 19,000 | 13,000 | 5,400 | 3,400 | 1,100 | 43,000 |
| STAR_b2ec | 1,000 | 27,000 | 19,000 | 8,100 | 5,100 | 1,700 | 62,000 |
| PHENIX_1μ | 95,000 | 19,400 | 2,600 | 850 | 360 | 60 | 118,000 |
| PHENIX ₋₂ μ | 190,000 | 41,000 | 6,300 | 2,100 | 1,000 | 200 | 240,000 |
| PHENIX_e | - | 450 | 260 | 110 | 70 | 25 | 915 |

Table 2: The same as Table 1, but at $\sqrt{s} = 500 \, GeV$ for $\int L \, dt = 800 \, pb^{-1}$; Cuts: $P_T^e > 5 \, GeV/c$ or $E_\mu > 2 \, GeV$.

| $M_1 - M_2$, GeV/c^2 | 2-5 | 5-9 | 9-12 | 12-15 | 15-20 | 20-25 | 2-25 |
|-------------------------|---------|---------|--------|--------|--------|--------|---------|
| STAR_b | 500 | 12,000 | 21,000 | 12,300 | 9,000 | 3,800 | 59,000 |
| STAR_blec | 600 | 21,500 | 37,000 | 23,500 | 17,500 | 7,300 | 107,000 |
| STAR_b2ec | 660 | 31,000 | 53,000 | 35,000 | 27,000 | 11,300 | 107,000 |
| PHENIX_1µ | 235,000 | 67,500 | 14,000 | 6,200 | 4,100 | 1,400 | 328,000 |
| PHENIX_2µ | 470,000 | 140,000 | 30,000 | 14,000 | 9,500 | 3,500 | 670,000 |
| PHENIX_e | - | 600 | 900 | 430 | 340 | 140 | 2,400 |

(geometry): $1.1 < \eta < 2.5$ plus $-2.5 < \eta < -1.1$ with 2π coverage in φ ; PHENIX EMC, consisting of two separated segments of the acceptance (geometry): $-0.35 < \eta < 0.35$ with the coverage in φ : $-112.5^{\circ} < \varphi < -22.5^{\circ}$ and $22.5^{\circ} < \varphi < 112.5^{\circ}$.

The STAR and PHENIX acceptances to the Drell-Yan pairs are shown in Fig. 1b. Besides geometries, the thresholds to the lepton energies E_l and/or transverse momenta P_T , which are supposed to be applied at the lowest trigger levels, have also been taken into account. The expected event rates are presented in Tables 1 and 2.

One can observe, PHENIX μ covers well the low pair mass region, while STAR is more sensitive to the higher mass di-electrons. The sensitivities, in terms of the "hardness" of the primary hard-colliding proton constituents producing detected lepton pair, are shown in Fig. 2. PHENIX μ acquires better muon pairs, originated from the hard interaction of quark with high x_q but of rather soft antiquark, while STAR has a better sensitivity to the antiquark sea with the higher $x_{\overline{q}}$.

The expected event rates for W^{\pm} and Z^0 are shown in Table 3. Apparently, Z^0 is almost unreachable in PHENIX due to its low acceptance to the high mass dilepton pairs. High background level in the forward and backward directions might also make difficult to detect $W^{\pm} \to \mu \nu$ signal with the PHENIX- μ setups.

ground. A number of the STAR detector features can be used to separate charged hadrons from electrons in order to get a necessary rejection power of $\sim 10^3$.

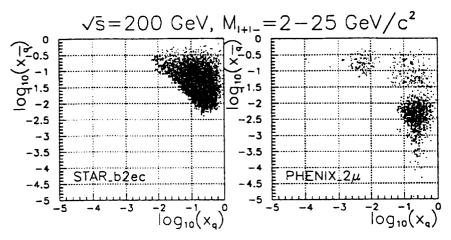


Figure 2: The x_q vs $x_{\overline{q}}$ plot of the primary constituents, producing lepton pairs within STAR_b2ec and PHENIX_ 2μ acceptances.

Background for W^{\pm} in STAR

Any event with a high, local energy deposition in the EMC, matching a high- P_T charged particle, and observed in the tracking system, may be considered a potential candidate for a W^{\pm} decay. Other sources can also provide such a signature, however, for example: Z^0 decays with one missing electron; electrons from π^0 Dalitz decays; misidentified high- P_T charged hadrons as electrons; overlapping in the EMC γ -quanta from π^0 decays with charged high P_T hadrons.

For the STAR barrel EMC the background from $Z^0 \to e^+e^-$, with one electron missing, is expected to be quite low: $\sim 2.5\%$ for W^+ , and about 10% for W^- . With the end cup(s), this background will be even lower due to a better detection efficiency to electron pairs.

The P_T spectrum of e^{\pm} originated from $\pi^0 \to \gamma e^+e^-$ Dalitz decays drops rapidly when P_T increases (2,7). As a result, with the only requirement being $P_T^e > 20 - 25 \; GeV/c$, the background from Dalitz pairs will be reduced to about the same level or even lower than from Z^0 decays.

Table 3: W^{\pm} and Z^{0} event rates at $\sqrt{s} = 500$ GeV for $\int L \ dt = 800$ pb^{-1} ; Cuts: $P_{T}^{e,\mu} > 20$ GeV/c.

| I | | STAR_b | STAR_blec | STAR_b2ec | PHENIX_1µ | PHENIX_ 2μ | PHENIX_e |
|---|---------|--------|-----------|-----------|-----------|----------------|----------|
| Ì | W^+ | 64,600 | 71,500 | 78,400 | 4,650 | 9,300 | 14,900 |
| 1 | W^- | 15,000 | 20,600 | 26,200 | 5,050 | 10,100 | 2,600 |
| I | Z^{0} | 2,700 | 4,200 | 6,200 | 25 | 310 | 120 |

Acceptances and Event Rates

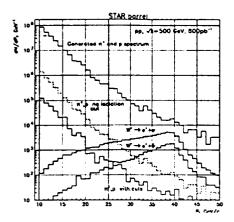


Figure 3: Charged hadron background vs P_T before and after applying rejection criteria along with e^{\pm} from W^{\pm} decays.

The contamination of the W^{\pm} event samples by high- P_T charged hadrons misidentified as electrons is expected to be the most serious source of background. A number of the STAR detector features can be used to separate charged hadrons from electrons in order to get a necessary rejection power of $\sim 10^3$.

At the lowest trigger level a threshold ~ 20 - 25 GeV/c should be applied to the $P_T^{e^{\pm}}$. Since hadrons mostly deposit only a fraction of their energy in the EM calorimeter, a single hadron will effectively be "seen" in the EMC as a particle with a lower P_T than it actually is. Thus, due to the rapid drop of the hadron P_T spectrum when P_T increases, it effectively provides a hadron P_T spectrum, measured in the EMC, lying well below the actual one. The effective hadron suppression power of this mechanism varies from about 50 to 150 in the P_T region of 10 - 50 GeV/c. Some other selection criteria, common for the practice of other collider experiments, also have been studied (isolation cut; limit to the shower width, defined with the fine-graned Shower Maximum Detector placed at the depth $5 \cdot X_0$ in the EMC).

The results of simulations are shown in Fig. 3, providing that the W^+ signal can be extracted in STAR_b at the background level of $\sim 3-7\%$, while the detection efficiency, due to applying cuts, drops by $\sim 15-25\%$. Background to W^- can be rejected to the level of $\sim 10-30\%$, depending on the selection criteria.

Conclusion

The estimated event rates prove the possibility for STAR and PHENIX to carry out the substantial spin physics with Drell-Yan pairs, W^{\pm} and Z^{0} at the statistical error level of a few percents, provided the polarization of colliding protons is ~50-70%. The background to the W^{\pm} signal in STAR is expected to be at the acceptable level.

Acknowledgments

We are pleased to express our appreciation to G. Bunce, D. Grosnick, S. Heppellmann, Yo. Makdisi, S. Nurushev, J. Soffer, H. Spinka, M. Tannenbaum, A. Vasiliev and D. Underwood for the useful and encouraging discussions.

References

- 1. Bourrely, C. and Soffer, J., Parton Distributions and Parity-Violating Asymmetries in W[±] and Z⁰ Production at RHIC, CPT-93/P.2865, CNRS Luminy (France).
- 2. Tannenbaum, M. J., Polarized Protons at RHIC, Proc. of the Polarized Collider Workshop, University Park, PA (1990), AIP conf. proc. No 223, AIP, New York (1991), p. 201.
- 3. Doncheski, M. A., et al., Hadronic W production and the Gottfried Sum Rule, Preprint MAD/PH/744, June 1993.
- Cheng, H.-Y. and Lai, S.-N., Phys. Rev., D41, 91 (1991);
 Jaffe, R. L. and Ji, X., Phys. Rev. Lett., 67, 552 (1991);
 Ji, X., Phys. Lett., B284, 137 (1992).
- 5. RHIC Spin Collaboration, Proposal on Spin Physics Using the RHIC Polarized Collider, August 14, 1992; Proposal Update, September, 1993, and references therein.
- 6. Derevschikov, A. A. and Rykov, V. L., Notes on the Drell-Yan Pairs, Z⁰ and W[±] in STAR and PHENIX at RHIC, Internal RSC Report RSC-BNL/IHEP-4, August 1992 (unpublished).
- 7. Rykov, V. L. and Shestermanov, K. E., W[±] and Z⁰ Event Rates and Background Estimates for the STAR Detector at RHIC in pp Collisions, ANL-HEP-TR-93-89, Argonne, 1993 (unpublished).
- 8. Eichten, E., et al., Rev. Mod. Phys., 56 579 (1984) and Erratum 58, 1065 (1986).
- 9. The STAR collaboration, Conceptual Design Report for the Solenoidal Tracker at RHIC (STAR), June 15, 1992.
- 10. The PHENIX collaboration, PHENIX Conceptual Design Report, January 29, 1993.