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## LONGITUDINAL RESISTIVE WALL INSTABILITIES OF AN INTENSE COASTING BEAM

## FOR AXIALLY ASSYMMETRIC GEOMETRY

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## Introduction

The longitudinal resistive wall instability for azimuthally uniform beam moving in a straight vacuum chamber has been investigated for different arrangements of beam and pipe but the beam was always assumed to be centered in the pipe.

The case of a round and rectangular geometry is treated in Ref. [1] whereas the elliptic geometry is used in Ref. [2].

Here we will study the case of a beam placed asymmetrically in a round pipe.

## 2. Solution of Maxwell Equations

The approach is the same as in Ref.  $\begin{bmatrix} 2 \end{bmatrix}$ . We assume a perturbation in the charge distribution of the form

$$\rho = \rho_1 e^{i(kz - \omega t)},$$

$$k = R/n$$

 $2\pi R$  - length of the closed orbit n - harmonic number.

We use the following parameters:

r - radius of the round beam

b - radius of the pipe

d - distance between the axis of the two parallel
 cylinders.

Further we use Bi-cylinder coordinates  $\eta, \Theta, z$  (Ref. [3]). The coordinate surfaces in the cartesian frame may be written

a) 
$$\eta = \text{const.}, -\infty \leq \eta \leq +\infty$$
,

$$x^{2} + y^{2} + a^{2} = 2ax \operatorname{cotgh} \eta,$$
 (1)

this is a cylinder with the axis at  $x = a \operatorname{cotgh} \eta$ , y = 0.

b)  $\Theta$  = const. (this is not the azimuthal variable along the closed orbit),  $-\pi \in \Theta \leq \pi$ 

$$x^{2} + (y - a \cot g \theta)^{2} = \frac{a^{2}}{\sin^{2} \theta}$$

c) z = const.

The pipe is represented by the cylinder  $\eta = \eta_0 > 0$ , the beam by  $\eta = \eta_1 > 0$ .  $\eta = +\infty$  is identical with x = a which is chosen to be near the axis of the beam  $x = a \operatorname{cotgh} \eta_1$ , y = 0. Using (1) we get for the radii

$$r = a/senh \eta_1$$
  
 $b = a/senh \eta_0$ 

and

$$\operatorname{cotgh} \eta_0 - \operatorname{cotgh} \eta_1 = d/a$$
.

As r, b and a are known for a particular problem we figure out a,  $\eta_0$ ,  $\eta_1$ . For CESAR with b = 5 cm and r = 0,15 cm we get the following table.

		TABLE I	•• 2	
d	a	η <sub>l</sub>	n <mark>o</mark> 	q/n
l cm	12,00 cm	5,1	1,6	7,0 · 10 <sup>-3</sup>
~2 <b>c</b> m	5,25 cm	4,2	0,92	3,1
3 cm	2,67 cm	3,6	0,52	1 <b>,</b> 5
4 cm	1,12 cm	2,7	0,22	0,7
4,8 cm .	0,14 cm	0,84	0,028	0 <b>,09</b>

For the source of the electromagnetic field we take the following: the charge density  $\rho = \rho_1 e^{i(kz - \omega t)} \text{ for } \eta > \eta_1$ 

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$$j_{\eta} = j_{\varphi} = 0, j_{z} = \omega \rho/k$$

From Maxwell equations we get the Helmotz' equation for the longitudinal field  $E_z$ . If we use the following expression for  $E_z$ .

$$E_{z} = f(\eta, \theta) e^{i(kz - \omega t)}$$

we can write

· ;

$$\frac{\partial^2 \mathbf{f}}{\partial \eta^2} + \frac{\partial^2 \mathbf{f}}{\partial \theta^2} - \frac{\mathbf{k}^2}{\mathbf{\gamma}_{W}^2} g_{11} \mathbf{f} = 4\pi \mathbf{i} \frac{\mathbf{k}^2}{\gamma_{W}^2} g_{11}$$
(2)

where  $g_{11}$  is the length element of the bi-polar coordinates

$$g_{11} = \frac{a^2}{(\cosh \eta - \cos \theta)^2}$$
and
$$\gamma_w^2 = 1/(1-\beta_w^2)$$

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$$g_{11} = \frac{a^2}{(\cosh \eta - \cos \theta)^2}$$

In the region outside the beam,  $\eta < \eta_1$ , we have to solve the homo-

$$\frac{\partial^2 f}{\partial \eta^2} + \frac{\partial^2 f}{\partial \theta^2} - \frac{k^2}{\gamma_w^2} \frac{a^2 f}{(\cosh \eta - \cos \theta)^2} = 0, \quad (3)$$

In order to make equation (3) separable we introduce an approximation. The smallest value of  $\cosh \eta$  occurs at the surface of the pipe and is always bigger than one. If we put in equation (3)

$$(\cosh \eta - \cos \theta) \sim \cosh \eta$$
 (5)

which is strictly valid for all  $\eta$  at  $\mathcal{O} = \frac{+}{2} \frac{\pi}{2}$  and which is a good approximation for  $\mathcal{O} \neq \pm \frac{\pi}{2}$  and  $\eta$  large, we obtain

$$\frac{\partial^2 \mathbf{f}}{\partial \eta^2} + \frac{\partial^2 \mathbf{f}}{\partial \theta^2} + \frac{\mathbf{k}^2}{\gamma_w^2} \frac{\mathbf{a}^2 \mathbf{f}}{\cosh^2 \eta} = 0.$$
 (6)

Equ. (6) now <u>separates.</u> If we put

$$f = H(\eta) \Theta(\theta)$$

we get the following two equations

$$\frac{d^{2} H}{d\eta^{2}} - \left(\frac{k^{2} a^{2}}{\gamma_{w}^{2} \cosh^{2} \eta} + c_{1}\right) H = 0$$
(7a)

$$\frac{d^2 \Theta}{d \theta^2} + c_1 \theta = 0 .$$
 (7b)

Because  $E_z$  is a periodic function of  $\Theta$   $c_1$  cannot be negative. Further as  $E_z$  cannot change its sign for symmetry reasons we get  $c_1 = 0$ and  $\Theta$  is a constant. (7a) yields

$$\frac{d^{2}H}{d\eta^{2}} - \frac{q^{2}H}{\cosh^{2}\eta} = 0$$
(7c)

with

$$q = ka/\gamma_{\rm W} \tag{8}$$

In CESAR it is  $\gamma_{_W}$  = 4.5 and R = 380; the quantity q/n is tabulated in table I.

Equ. (7c) can be solved by the series method as shown in Appendix A. For large value of  $\eta$  a first integral is given approximatively by

$$H_{1} = 1 + \frac{q^{2}}{4 \cosh^{2} \eta} + \dots$$
 (9)

Because q is very small  $H_1$  is a function always near one. The second integral of (7a)  $H_2$ , dan be found by the Euler method.

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## 3. Boundary conditions and expressions for the longitudinal field.

The longitudinal field  $\mathop{\rm E_Z}_Z$  outside the beam,  $\eta < \eta_4$  is represented by

 $E_{z} = \left[A_{1} H_{1}(\eta) + A_{2} H_{2}(\eta)\right] e^{i(kz - \omega t)}$ (13)

As in the beam,  $\eta > \eta_1$ ,  $E_z$  must remain finite we have to drop H<sub>2</sub>( $\eta$ ) in the homogeneous solution and we obtain the following general solution for  $\eta > \eta_1$ .

$$E_{z} = B_{1} H_{1} (\eta) e^{i(kz - \omega t)} - 4\pi i \frac{\rho}{k}.$$
 (14)

The last term in the right of the equ. (14) is the particular integral of the inhomogeneous equation (3), valid for  $\eta \nearrow \eta_1$ .

From the following Maxwell equations

 $\operatorname{rot}_{\mathcal{O}} \stackrel{\rightarrow}{\operatorname{E}} = \operatorname{i} \frac{\omega}{c} \stackrel{\operatorname{H}}{\operatorname{H}}_{\mathcal{O}}$  $\operatorname{rot}_{\eta} \stackrel{\rightarrow}{\operatorname{H}} = -\operatorname{i} \frac{\omega}{c} \stackrel{\operatorname{E}}{\operatorname{H}}_{\eta}$ 

we get a relation between  $H_{\rho}$  and  $E_{z}$  taking into account that for reasons of symmetry  $H_{z}$  is zero:

$$H_{\theta} = -\frac{i\beta_{w}\gamma_{w}^{2}}{ak} \quad (\cosh \eta + \cos \theta) \quad \frac{\partial E_{z}}{\partial \eta} \quad . \tag{15}$$

The boundary conditions which have to be satisfied are

- a) continuity of  $E_z$  at  $\eta = \eta_1 \cdot \eta_2 \cdot \eta_3$
- b) continuity of  $H_{\beta}$  at  $\eta = \eta_1$ , and
- c) normalization between  $E_z$  and  $H_z$  at  $\eta = \eta$ , which means

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$$\mathbf{E}_{z} = (1 - i) \widehat{\mathbf{R}} \mathbf{H}_{\widehat{\mathbf{V}}}$$
(16)

where

$$\mathcal{R} = \sqrt{\frac{\omega}{8\pi \, \odot}}$$

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 $\sigma$  - conductivity of the walls of vacuum chamber. Condition (16) cannot be satisfied completely because there is the dependence on  $\cos \Theta$  in the expression (15) for H<sub>Q</sub>. But in the following we shall apply the approximation (5) again. This means that the boundary condition at  $\eta = \eta_0$  is satisfied only in the regions of the surface for which  $\cos \Theta$  is near zero.

In Appendix B it is shown how the three constants  $A_1$ ,  $A_2$  and  $B_1$  have been evaluated from the boundary conditions a), b) and c). The results for the longitudinal field  $E_z$  evaluated on the polar axis  $\eta = \infty$ , which is very near to the beam axis are

$$E_{z}(\infty) = 4\pi i \frac{9}{k} \left[ H_{2}'(\eta_{1}) - 1 + H_{1}'(\eta_{1}) \frac{H_{2}(\eta_{0}) - T H_{2}'(\eta_{0})}{H_{1}(\eta_{0}) - T H_{1}'(\eta_{0})} \right]$$
(17)

with

$$T = -(1+i)\frac{\hat{\mathcal{R}}\beta_{w}\gamma_{w}^{2}}{ak}\cosh\eta_{o} = -(1+i)T_{1}. \qquad (18)$$

4. Discussion of the dispersion relation and results for CESAR.

Following Ref. [1] we write  $\omega = n\omega_0$ . For CESAR we have

$$\omega_{0} = 7,5 \cdot 10^{7} \text{ cs}^{-1}$$

$$o^{-} = 1,2 \cdot 10^{16} \text{ s}^{-1}$$

this results in

$$T_{1} = \frac{2.4 \cdot 10^{-2}}{\sqrt{n}} \operatorname{cotgh} \eta_{0}$$
.

As one can see from table I  $T_1 \leq 1$  and  $\eta \gg \frac{q^2}{4}$  is fulfilled. Introducing in (17) the results of Appendix A and (18) we obtain

$$E_{z}(\infty) = -i \frac{k}{\gamma_{w}^{2}} \lambda \operatorname{senh}^{2} \eta_{1} \left[ \epsilon(\eta_{1}) - (\eta_{1} - \eta_{0}) \epsilon'(\eta_{1}) + i\epsilon'(\eta_{1}) T_{1} \right]$$
(19)

 $\epsilon$  (n) is a function described in Appendix A.  $\lambda$  is the charge per unit length.

$$\lambda = \lambda_1 e^{i(kz - \omega t)}.$$

In Ref. [1] the Vlasov equation for the distribution function  $\Psi(w, \theta, t)$  in the beam is solved and a dispersion relation is obtained.

With the same procedure as applied in Ref. [1] and Ref. [2] we get for the complex quantity (U - iV)

$$U = Ne^{2} \frac{k}{\gamma_{w}^{2}} tgh^{2} \eta_{1} \left[1 + 2(\eta_{1} - \eta_{0}) tgh \eta_{1}\right]$$
(20a)

$$V = 2 \operatorname{Ne}^{2} \operatorname{R} \operatorname{\beta}_{W} \operatorname{tgh}^{3} \eta_{1} \operatorname{cotgh} \eta_{0} / b .$$
 (20b)

The results in Ref. [1] for a round beam placed in the centre of a round pipe are

$$U_{c} = Ne^{2} \frac{k}{\gamma_{W}^{2}} \left[1 + 2 \ln (b/r)\right]$$
 (21a)

 $\nabla_{c} = 2 N e^{2} R / \frac{3}{W} / b$  (21b)

d zero. As

$$\eta_1 - \eta_0 \rightarrow \ln \frac{b}{r}$$
 for  $\eta_1, \eta_0 \rightarrow \infty$ 

is obvious, we find that (21a), (21b) is just a special case of (20a), (20b).

When the beam is moved away from the centre, the quantity U decreases monotonically and V increases as shown in Fig. 1.

We can write also

$$U = U_{c} \frac{\operatorname{tgh}^{2} \eta_{1} \left[ 1 + 2 (\eta_{1} - \eta_{o}) \operatorname{tgh} \eta_{1} \right]}{1 + 2 \ln (b/r)}$$
(22a)

and

$$V = V_{c} \operatorname{tgh}^{3} \eta_{1} \operatorname{cotgh} \eta_{0}.$$
 (22b)

For CESAR, b = 5 cm, r = 0,15 cm, (22a) and (22b) give the following values:

			TABLE TT	
d.		v/v <sub>c</sub>	U/U <sub>c</sub>	$V/U \cdot \sqrt{n} \cdot 10^3$
0	cm	1.0	1.00	5.6
1	cm	1.1	1.00	6.2
2	cm	1.4	0.95	8,3
3	cm	2.1	0.90	13
4	cm	4.5	0.75	34
4,8	cm	<b>19</b> ()	0.19	560

For a small value of d we get  $V \ll U$ . Above transition growth time and stability criterion against longitudinal instability can be written in the form

$$\mathcal{L} = \left[ n \left| k_{0} \right| \ U \right]^{-1/2}$$

$$S > \left[ U/n \left| k_{0} \right| \right]^{1/2}$$
(23a)
(23b)

as shown in Ref.  $\begin{bmatrix} 1 \end{bmatrix}$ :

2 - total frequency spread in the beam,

 $k_0$  - constant related to the momentum compaction factor.

For a large d, V becomes larger than U and (23a), (23b) are no longer valid. For V >> U the equations for growth time and stability become (see Ref. [4])

$$\overline{L} = \left[ 2/n |k_0| \overline{V} \right]^{1/2}$$
(24a)  
$$\delta > \left[ V/n |k_0| \right]^{1/2}$$
(24b)

## 5. <u>Conclusions</u>

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Taking into account (23) and (24) we can infer from Fig. 1 or table II that above a certain value of d the beam is less stable than for the case d = 0. The growth time is plotted against d in Fig. 2.

This could explain qualitatively why the beam in CESAR is more unstable at injection orbit which is very near to the plate of the pulsed inflector. than at central orbit.

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The functions H (n), H (n)

We consider the equation

$$\frac{d^2H}{d\eta^2} - \frac{q^2 H}{\cosh^2 \eta} = 0$$
 (A1)

The quantity  $q^2$  is very small. For CESAR and for n = 10 we get a  $q^2$  quantity of the order of  $10^{-4}$  for d = 3cm. If  $\eta$  increases, the second term goes to zero. Thus we can omit the second term for  $\eta \rightarrow \emptyset$ and write  $\frac{d^2 H}{d\eta^2} = 0$ . (A2)

Two particular solutions are and the second second

 $H_1 = 1 \text{ and } H_2 = \eta$  . (A3)

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We are interested in the particular integrals of (Al) of which (A3) are the assymptotic expressions.

The differential equation (Al) can be integrated by the method of series. We look for an integral of the type

$$H_{1}(\eta) = \sum_{0}^{\infty} i \frac{\alpha' i}{\cosh^{1} \eta} \qquad (A4)$$

that, if we take  $\alpha_{\eta} = 1$ , for  $\eta \rightarrow$  tends to unit.

Substitution of (A4) in (A1) yields  $\alpha_1 = 0$  and the following correlation between  $\alpha_i$  and  $\alpha_{i+1}$ 

$$i(i+1)\alpha_{i} + q^{2}\alpha_{i} - (i+2)^{2}\alpha_{i+2} = 0$$
, (A5)

If  $q \ll 1$  and i > 0 is fulfilled we obtain from (A5) with good

approximation  

$$\begin{aligned}
\mathcal{O}_{i+2} &= \frac{i(i+1)}{(i+2)^2} \mathcal{O}_{i} \quad (A6) \\
\text{For odd i it is } \mathcal{O}_{i} &= 0 \quad \mathbf{a}_{2} &= -\frac{q^{2}}{4} \text{ results from (A5) for } i = 0. \\
\text{It is easily verified that (A6) and even i give} \\
\alpha_{i} &= \frac{(i-1)!}{\left[i - (i-2)(i-4)(i-6) \dots 4\right]^{2}} \alpha_{2} \quad (A7) \\
\text{then if we put} \\
\delta_{i} &= \frac{(i-1)!}{\left[i (i-2)(i-4) \dots 4\right]^{2}} \quad (A8) \\
\text{we get for } H_{1}(\eta) \\
\end{aligned}$$

$$H_{1}(\eta) = 1 + \frac{q^{2}}{4} \frac{\sum_{j=1}^{\infty} i \frac{S_{2i}}{\cosh^{2i} \eta}}{\cosh^{2i} \eta}$$
(A9)

with

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$$S = \frac{\sum_{i=1}^{\infty} i \delta_{2i}}{1} \delta_{2i}$$

is convergent which was checked numerically and S is equal to 2;8. This implies that the series (A9) or (A4) are always convergent. Further the convergence becomes very fast for  $\eta \rightarrow \infty$ .

For all values of  $\eta$ , the second term on the right-hand side in (A9) is always small in comparison to one. Therefore we can write

$$H_{1}(\eta) = 1 + \frac{q^{2}}{4} \epsilon(\eta)$$
 (A10)

• . . . The maximum value of  $\epsilon$  ( $\eta$ ) is 2,8 and the function decreases rapidly to zero by increasing  $\eta_{\bullet}$ 

Derivation of  $H_1$  ( $\eta$ ) with respect to  $\eta$  is

$$H_1'(\eta) = -\frac{q^2}{4} \epsilon'(\eta)$$
 (A11)

For  $\eta \rightarrow C0^{-}$  H<sup>+</sup> ( $\eta$ ) tends to 0. Note again that  $\epsilon'(\eta)$  is a negative quantity. The second integral H<sub>2</sub> ( $\eta$ ) of (Al) can be derived by the Euler method.

$$H_{2}(\eta) = H_{1}(\eta) \int_{0}^{\eta} \frac{d\eta}{H_{1}^{2}(\eta)}$$
 (A12)

From (AlO) and with  $q \ll 1$  it results in

$$H_{1}^{2}(\eta) = 1 + \frac{q^{2}}{2} \epsilon(\eta)$$

and

(. <sub>2</sub>.)

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$$\frac{1}{H_1^2(\eta)} = 1 - \frac{q^2}{2} \epsilon(\eta)$$

then we get

$$\int_{0}^{\eta} \frac{d\eta}{H_{1}^{2}} = \eta - \frac{q^{2}}{2} g(\eta)$$
(A13)

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wherein

$$g(\eta) = \int_{0}^{1+\alpha} \varepsilon(\eta) d\eta ,$$

Inserting (A13) in (A12) we get a constraint of a line with a synthetic set of the end of the end

(114) 
$$H_{2}(\eta) = \eta - \frac{q^{2}}{2} \left[ g(\eta) - \frac{1}{2} \eta \varepsilon (\eta) \right].$$
(A14)

whereas for  $\eta = 0$  the same function is zero. Thus we can write

$$\lim_{\eta \to \infty} H_2 H_1' = 0$$

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Finally we get from (A14)

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$$H_{2}'(\eta) = 1 - \frac{q^2}{4} \left[ \varepsilon(\eta) - \eta \varepsilon'(\eta) \right] ; \qquad (A15)$$

this quantity is always near to unity and the quantity in brackets is always positive.

From the relation above we can infer that the Wronskian W is a constant and can be evaluated for  $\eta = 00$ . It turns out that W is equal to one.

Evaluation of the constant  $\underline{A_1, A_2, B_1}$ . From the continuity of the longitudinal field  $\underline{B_2}$  (13) (14) at  $\eta = \eta_1$ , the beam boundary, we get (1)  $A_1 H_1 (\eta_1) + A_2 H_2 (\eta_1) = B_1 H_1 (\eta_1) - 4\pi i \frac{\rho_1}{k}$ . (B1)

The continuity of  $H_{\Omega}$  (15) at  $\eta = \eta_1$  demands

$$A_{1} H_{1}' (\eta_{1}) + A_{2} H_{2}' (\eta_{1}) = B_{1} H_{1}' (\eta_{1})$$
(B2)

where the prime denotes derivation with respect to  $\eta_{\bullet}$ 

From (B1) and B2) we obtain the constants  $A_1$  and  $A_2$  as functions of the  $B_1$ . Now we bear in mind that the wronskian of the two functions  $H_1$  and  $H_2$  is one as shown in App. A. This yields

$$A_1 = B_1 - 4\pi i \frac{\rho_1}{k} H_2'(\eta_1)$$
 (B3)

$$A_2 = 4\pi i \frac{\beta_1}{k} H_1'(\eta_1)$$
 (B4)

<sup>B</sup><sub>1</sub> can now be determined applying the boundary condition (16) at  $\eta = \eta_0$ . From (13) and (15) it results

$$A_{1} H_{1} (\eta_{0}) + A_{2} H_{2} (\eta_{0}) =$$

$$= - (1 + i) \left( \int_{a}^{b} \frac{\eta_{w}^{2} \cosh \eta_{0}}{a k} \left[ A_{1} H_{1}' (\eta_{0}) + A_{2} H_{2}' (\eta_{0}) \right] \right). \quad (B5)$$

Deriving (B5) we applied the approximation (5) for the reasons mentioned above. Putting (B3) and (B4) in (B5) we get finally

$$\mathbf{A}_{1} \left[ \mathbf{H}_{1} \left( \mathbf{\eta}_{0} \right) - \mathbf{T}\mathbf{H}_{1} \left( \mathbf{\eta}_{0} \right) \right] + \mathbf{A}_{2} \left[ \mathbf{H}_{2} \left( \mathbf{\eta}_{0} \right) - \mathbf{T}\mathbf{H}_{2} \left( \mathbf{\eta}_{0} \right) \right] = 0$$

or

$$B_{1} = 4\pi i \frac{\rho_{1}}{k} \left[ H_{2}'(\eta_{1}) - H_{1}'(\eta_{1}) \frac{H_{2}'(\eta_{0}) - TH_{2}'(\eta_{0})}{H_{1}(\eta_{0}) - TH_{1}'(\eta_{0})} \right].$$
(B6)

As  $H_1(\eta)$  at  $\eta = \infty$  is one, using (B6) we get the expression (17) for the longitudinal field  $E_z$  near the beam centre.

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<u>References</u>,

- V. K. Neil and A.M. Sessler Longitudinal Resistive <u>Instabilities</u> of Intense Coasting Beams in Particle Accelerator Rev. Sci. Instr. 36, 429, (1965)
- A.G. Ruggiero
   Longitudinal Resistive Instabilities of Intense Coasting Beams
   for Elliptic Geometry.
   ISR-SRM /67-13,
- 3) P. Moon and D.E. Spencer Field Theory Handbook Springer Verlag, 1961
- 4) V.G. Vaccaro

Longitudinal Instability of a coasting Beam above transition, due to the action of lumped Discontinuities. ISR-RF/66-35

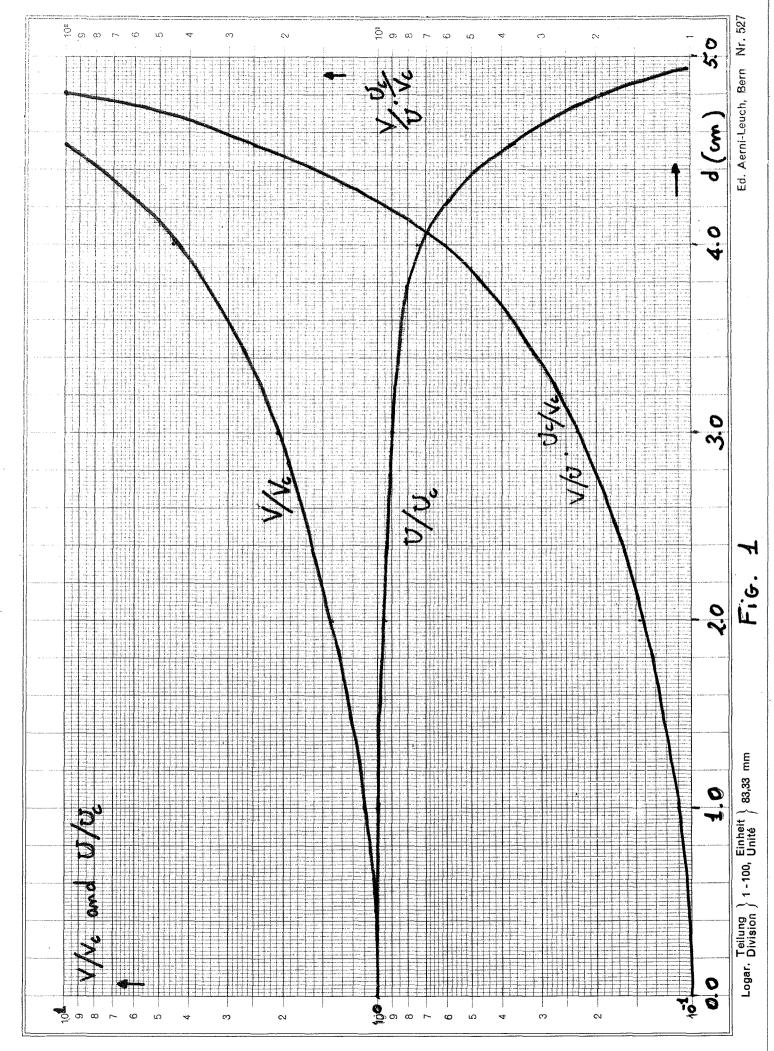
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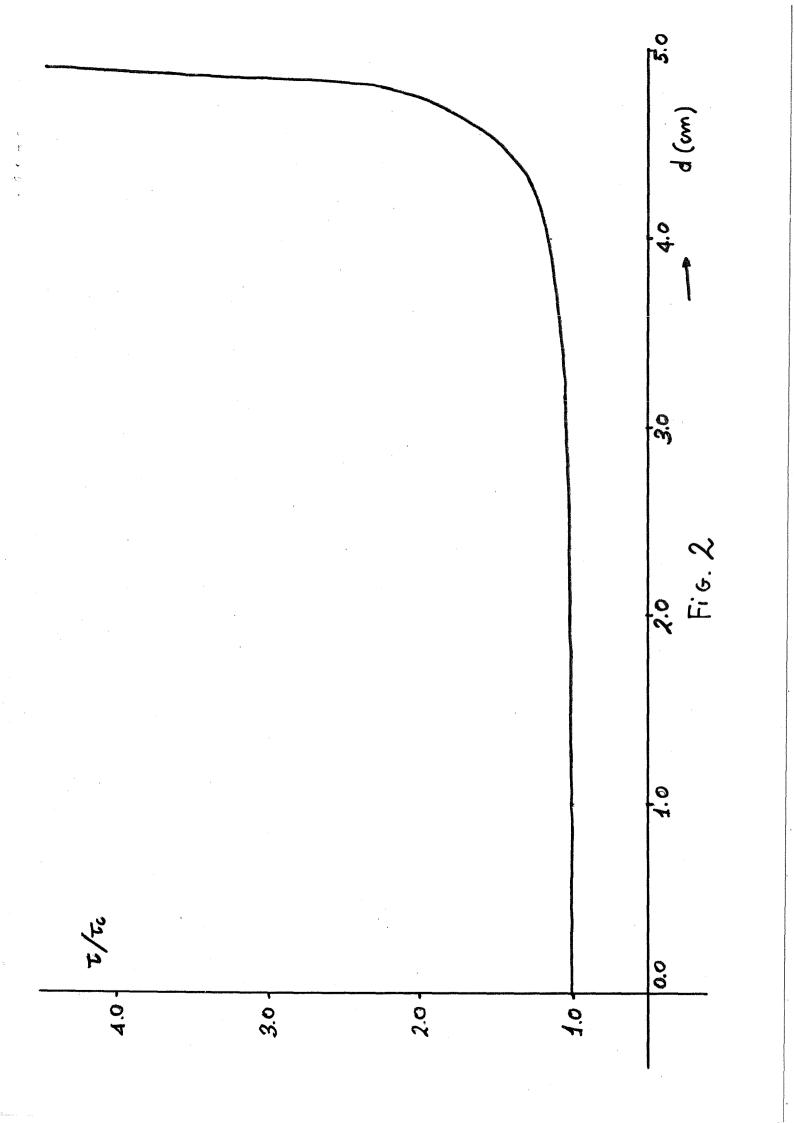
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