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### **Magnesio-ferri-hornblende, $\square\text{Ca}_2(\text{Mg}_4\text{Fe}^{3+})[(\text{Si}_7\text{Al})\text{O}_{22}](\text{OH})_2$ , a new member of the amphibole supergroup**

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## ABSTRACT

Magnesio-ferri-hornblende, ideally  $\square\text{Ca}_2(\text{Mg}_4\text{Fe}^{3+})[(\text{Si}_7\text{Al})\text{O}_{22}](\text{OH})_2$ , is a new mineral of the amphibole supergroup from the Husite granitic complex related to skarn-type Fe-Cu mineralization in the Western Tianshan, Xinjiang, northwestern China. The new species and the new name have been approved by the IMA-CNMNC (2021-100). Magnesio-ferri-hornblende is dark green to green-black with a vitreous lustre and a pale gray-green to gray-white streak. It occurs mostly as subhedral-columnar crystals with a length of 0.5 to 3 mm and shows well-developed {110}

21 cleavage. It has a Mohs hardness of ~5 and a Vickers microhardness of 389-448  
22 kg/mm<sup>2</sup> (VHN load in 100g) and is brittle with a conchoidal fracture. The measured  
23 and calculated densities are 3.275(6) and 3.204 g/cm<sup>3</sup>, respectively. In transmitted  
24 plane-polarized light, magnesio-ferri-hornblende is strongly pleochroic, *X* = pale  
25 yellow, *Y* = yellowish brown, *Z* = dark yellowish green. It is biaxial (-),  $\alpha = 1.651(2)$ ,  
26  $\beta = 1.658(2)$ ,  $\gamma = 1.662(2)$ ,  $2V$  (meas.) = 73 (1)° to 82 (1)° and  $2V$  (calc.) = 73.9 (1)°,  
27 dispersion is  $r > v$ , medium to strong. The orientation is:  $Y // b$ ,  $X \wedge a = 31.5^\circ$  ( $\beta$  obtuse),  
28  $Z \wedge c = 16.5^\circ$  ( $\beta$  acute).

29 Magnesio-ferri-hornblende is monoclinic, space group  $C2/m$ ,  $a = 9.8620(3)$ ,  $b =$   
30  $18.1060(5)$ ,  $c = 5.30810(10)$  Å,  $\beta = 104.8480(10)^\circ$ ,  $V = 916.17(4)$  Å<sup>3</sup>,  $Z = 2$ . The  
31 strongest seven lines in the powder X-ray diffraction pattern are [ $d$  in Å( $I$ )( $hkl$ )]:  
32 8.397(52)(110), 3.383(41)(150), 2.717(100)(151), 2.597(84)(061), 2.545(61)(20 $\bar{2}$ ),  
33 1.854(49)(1 $\bar{7}$  $\bar{2}$ ), and 1.519(62)(6 $\bar{2}$  $\bar{2}$ ). Analysis by a combination of electron  
34 microprobe and Mössbauer spectroscopy gave SiO<sub>2</sub> 47.37, TiO<sub>2</sub> 1.51, Al<sub>2</sub>O<sub>3</sub> 7.07,  
35 Fe<sub>2</sub>O<sub>3</sub> 3.86, FeO 11.62, MgO 12.77, CaO 11.22, SrO 0.15, MnO 0.39, Na<sub>2</sub>O 1.54,  
36 K<sub>2</sub>O 0.78, Cl 0.15, F 0, H<sub>2</sub>O<sub>calc</sub> 2.01, Cl≡O -0.03, sum 100.41 wt%. The empirical  
37 formula calculated on the basis of 24 (O + OH + F + Cl) with (OH + F + Cl) = 2 apfu  
38 is

39  $A_{(\square_{0.62}Na_{0.23}K_{0.15})\Sigma 1.00} B_{(Ca_{1.76}Na_{0.21}Mn_{0.02}Sr_{0.01})\Sigma 2.00} C_{(Mg_{2.79}Fe^{2+}_{1.42}Fe^{3+}_{0.43}Ti_{0.17}Al_{0.16}$   
40  $Mn_{0.03})\Sigma 5.00} T_{(Si_{6.94}Al_{1.06})\Sigma 8.00} O_{22} W_{(OH_{1.96}Cl_{0.04})\Sigma 2.00}$ . The crystal structure of magnesio-  
41 ferri-hornblende was refined to an  $R_1$  of 3.95% using 2185 data ( $>2\sigma$ ) collected with  
42 MoK $\alpha$  X-radiation. The *A* site is dominantly occupied by  $\square$  where  $A(Na + K + 2Ca) \leq$

43 0.5.  $T$ Al is ordered at the  $T(1)$  site.  $M(1)$  and  $M(3)$  are dominantly occupied by  $Mg^{2+}$ ,  
44 and  $M(2)$  is occupied by both  $Mg^{2+}$  and high-charged cations. The new mineral occurs  
45 most commonly in the porphyry-skarn Fe-Cu-Mo-Au- and hydrothermal Au-  
46 mineralized granitoids with high oxygen fugacity, but is rare or absent in barren  
47 intrusions. Its finding has important significance for magma fertility discrimination  
48 and can potentially be used in regional exploration for porphyry-skarn ore system.

49 **KEY WORDS:** magnesio-ferri-hornblende, new amphibole, optical properties, electron-  
50 microprobe analysis, crystal-structure refinement, magma fertility, Western Tianshan,  
51 China.

52

## INTRODUCTION

53 Amphiboles are common rock-forming minerals in a wide variety of rocks. The  
54 amphibole supergroup minerals are prevalent in igneous and metamorphic rocks  
55 formed under a broad temperature and pressure span, thus have a variety of chemical  
56 compositions (Leake, 1997; Hawthorne et al., 2007, 2012). Amphiboles are useful in  
57 understanding petrogenetic conditions and process, e.g., the Al-in-hornblende  
58 geobarometer was consistently explored since 1980s (Hammarstrom and Zen, 1986;  
59 Mutch et al., 2016 and references therein), and a series of amphibole  
60 thermobarometric formulations can give precious information ( $P$ ,  $T$ ,  $fO_2$ ,  $H_2O_{melt}$ ) on  
61 magmatic plumbing systems (Ridolfi et al., 2010; Ridolfi and Renzulli, 2012). The  
62 general chemical formula of the amphibole supergroup can be written as  
63  $AB_2C_5T_8O_{22}W_2$ , where  $A = \square, Na, K, Ca, Pb, Li$ ;  $B = Na, Ca, Mn^{2+}, Fe^{2+}, Mg, Li$ ;  $C =$   
64  $Mg, Fe^{2+}, Mn^{2+}, Al, Fe^{3+}, Mn^{3+}, Cr^{3+}, Ti^{4+}, Li$ ;  $T = Si, Al, Ti^{4+}, Be$ ;  $W = (OH), F, Cl,$

65  $O^{2-}$ . According to the current amphibole nomenclature scheme proposed by  
66 Hawthorne et al. (2012), the supergroup is divided into two groups of  $^W(OH, F, Cl)$ -  
67 dominant amphiboles and  $^WO$ -dominant amphiboles (oxo-amphiboles). The new  
68 mineral “magnesio-ferri-hornblende” reported in this study belongs to the calcium  
69 amphibole subgroup of the  $^W(OH, F, Cl)$ -dominant group.

70 In this contribution, the discovered amphibole has now been characterized as a new  
71 mineral species named magnesio-ferri-hornblende, where the prefix “magnesio-”  
72 indicates Mg dominance over  $C^{2+}$  cations and the prefix “ferri-” means  $Fe^{3+}$  is  
73 dominant over  $C^{3+}$  cations. The new species and the new name have been approved by  
74 the International Mineralogical Association Commission on New Minerals,  
75 Nomenclature and Classification (IMA-CNMNC, 2021-100). Holotype material is  
76 deposited in the Geological Museum of China, No. 16, Yangrou Hutong, Xisi, Beijing  
77 100031, People’s Republic of China, with the catalogue number M16131 (Zhang et al.,  
78 2022). Significantly, magnesio-ferri-hornblende occurs most commonly in fertilized  
79 felsic igneous rocks associated with porphyry-skarn Fe-Cu-Mo-Au mineralization, but  
80 is far less common or absent in barren intrusions. Therefore, it could be served as a  
81 potential mineral indicator for identifying fertile magmatic systems from barren felsic  
82 igneous rocks.

83

### OCCURRENCE

84 The new mineral magnesio-ferri-hornblende was firstly found in the Late  
85 Devonian Husite granitic complex in the Boluokenu island arc belt of the Western  
86 Tianshan, Xinjiang, NW-China ( $82^{\circ}43'26''$  E,  $44^{\circ}5'25''$  N). Situated in the western

87 section of the Central Asia Orogenic Belt (Fig. 1a), the Boluokenu island arc belt is  
88 one of the important metallogenic belts in the Western Tianshan, with a number of  
89 skarn and porphyry Fe-Cu-Mo polymetallic deposits associated with granitoids that  
90 formed during the Devonian to Early Carboniferous (Fig. 1b-c; Gu et al., 2014; Zhang  
91 et al., 2016; Wang et al., 2018). Since its initial discovery in the Husite granitic  
92 complex that is related to the Kekesala and Aimusidaiyi Fe-Cu skarn deposits,  
93 magnesio-ferri-hornblende has been subsequently found in several other intrusions in  
94 the belt, such as the Dawabulake granodiorite associated with the skarn-type Halegati  
95 Fe-Cu and Muzuke Pb-Zn deposits, the porphyry Mo ore-related Lailisigaoer  
96 plagiogranite porphyry, and the Nailenggele monzogranite (Fig. 1c).

97 The Husite intrusion elongates in NW direction, covering about 170 km<sup>2</sup> and  
98 intruding into the Ordovician limestone (Fig. 1c). It is composed of monzonitic  
99 granite, granodiorite, syenogranite, granitic aplite, dioritic porphyrite, gabbro and  
100 mafic microgranular enclaves (MMEs). Magnesio-ferri-hornblende dominantly occurs  
101 as subhedral-columnar crystals in the medium- to coarse-grained granodiorite  
102 lithofacies (Fig. 2). The grain size ranges from 0.5 to 3 mm, and the aggregate size  
103 varies from 0.5 to 1 cm. It is closely associated with plagioclase, quartz, K-feldspar,  
104 biotite, magnesio-hastingsite, as well as small amounts of zircon, magnetite, apatite,  
105 and sphene.

106

## MINERAL CHARACTERIZATION

### 107 **Physical and optical properties**

108 Magnesian-ferri-hornblende is dark green to green-black with a vitreous luster. It has  
109 a pale gray-green to gray-white streak and shows no fluorescence under long- or  
110 short-wave ultraviolet light. It is brittle and shows a conchoidal fracture. The Vickers  
111 microhardness is measured to be 389-448 kg/mm<sup>2</sup> (VHN load in 100g), and the  
112 converted Mohs hardness is ~5. It has the characteristic perfect {110} cleavage of  
113 amphiboles, intersecting at ~56°. The measured density is 3.275(6) g/cm<sup>3</sup> by  
114 hydrostatic weighing method using the ALFA MIRAGE SD-200L electronic density  
115 balance. The calculated density according to the empirical formula and unit cell  
116 volume refined from single crystal X-ray diffraction (XRD) data is 3.204 g/cm<sup>3</sup>.

117 In transmitted plane-polarized white light, magnesian-ferri-hornblende has strong  
118 pleochroism, with  $X$  = pale yellow,  $Y$  = yellowish brown, and  $Z$  = dark yellowish  
119 green. The orientation is  $Y \parallel b$ ,  $X \wedge a = 31.5^\circ$ , and  $Z \wedge c = 16.5^\circ$ . Magnesian-ferri-  
120 hornblende is biaxial negative, with indices of refraction  $\alpha = 1.651$  (2),  $\beta = 1.658$  (2),  
121  $\gamma = 1.662$  (2);  $2V$  (meas.) = 73 (1)° to 82 (1)° and  $2V$  (calc.) = 73.9 (1)°, under white  
122 light. Dispersion is medium to strong,  $r > v$ . Gladstone-Dale parameters yielded a  
123 compatibility index  $(1 - K_p/K_c)$  of 0.024, which is considered to be excellent  
124 (Mandarino, 1981).

### 125 **Mössbauer, infrared and Raman spectroscopy**

126 The Mössbauer spectrum of magnesian-ferri-hornblende was measured using a

127 constant acceleration in transmission mode (Wissel MVT-1000) with a  $^{57}\text{Co/Rh}$   
128 source at room temperature (20°C). The velocity was calibrated by 25  $\mu\text{m}$   $\alpha\text{-Fe}$  foil  
129 and the isomer shift relative to the center of  $\alpha\text{-Fe}$  at room temperature. The Mössbauer  
130 parameters are listed in Table 1, and the Mössbauer spectrum of the mineral is shown  
131 in Fig. 3.

132 The infrared spectrum of magnesio-ferri-hornblende powder pressed in a KBr pellet  
133 was analysed with a Perkin-Elmer Fourier transform infrared spectroscopy (FT-IR)  
134 in the 400-4000  $\text{cm}^{-1}$  wavelength range at a resolution of 4  $\text{cm}^{-1}$  upon 16 scanning. A  
135 pure KBr pellet was used as the reference sample. In the infrared spectrum of  
136 magnesio-ferri-hornblende (Fig. 4), Si-O stretching vibrations are identified by two  
137 absorption regions with bands at 1092  $\text{cm}^{-1}$ , 1044  $\text{cm}^{-1}$ , 987  $\text{cm}^{-1}$ , and 951  $\text{cm}^{-1}$ , and at  
138 753  $\text{cm}^{-1}$ , 688  $\text{cm}^{-1}$ , 660  $\text{cm}^{-1}$ , and 603  $\text{cm}^{-1}$ , respectively. In the low frequency range  
139 between 600  $\text{cm}^{-1}$  and 400  $\text{cm}^{-1}$ , the bands at 508  $\text{cm}^{-1}$  and 460  $\text{cm}^{-1}$  should be  
140 attributed to Si-O bending vibration and M-O vibration, respectively. The high-energy  
141 OH stretching vibrations exhibit bands for the OH groups coordinated by cationic  
142 triads: ( $\text{MgMgMg/MgMgFe}^{2+}\text{-OH-A}$ , 3688  $\text{cm}^{-1}$ ), ( $\text{MgMgMg-OH-}\square$ , 3672  $\text{cm}^{-1}$ ),  
143 and ( $\text{MgFe}^{2+}\text{Fe}^{2+}\text{-OH-}\square$ , 3645  $\text{cm}^{-1}$ ). The bands were assigned in accordance with  
144 Chukanov and Chervonnyi (2016), which further proved that the new mineral is Mg  
145 and  $\text{Fe}^{2+}$  at the  $M(1, 3)$  sites in composition. The bands at 3434  $\text{cm}^{-1}$  and 1632  $\text{cm}^{-1}$  of  
146  $\text{H}_2\text{O}$  appear in both sample KBr pellet and pure KBr pellet with similar absorbance  
147 extent, indicating that the  $\text{H}_2\text{O}$  comes from environments due to the hygroscopicity of  
148 KBr under current experimental conditions.

149 Raman spectra of magnesio-ferri-hornblende were obtained using a Renishaw  
150 Invia Reflex-type confocal Laser Raman microspectrometry with a laser excitation  
151 wavelength of 514.5 nm (20 mW, 1 $\mu$ m). The most distinct Raman peak detected in the  
152 magnesio-ferri-hornblende is around 666-725 cm<sup>-1</sup> (Fig. 5), which is assigned to the  
153 symmetric stretching vibrations of the Si-O-Si bridges (Apopei and Buzgar, 2010).  
154 The OH stretching region is found in the spectral domain of 3644 cm<sup>-1</sup>, 3659 cm<sup>-1</sup> and  
155 3673 cm<sup>-1</sup>, consistent with other amphiboles studied by various authors (e.g., Huang,  
156 2003; Su et al., 2009). The spectral bands around 532 cm<sup>-1</sup> were assigned to the OH<sup>-</sup>  
157 group (OH<sup>-</sup> translation, Kloprogge et al., 2001). The bands between 200 cm<sup>-1</sup> to 300  
158 cm<sup>-1</sup> may be assigned to lattice vibration (Huang, 2003) and the O-H-O group  
159 (Rinaudo et al., 2004).

#### 160 **Chemical composition**

161 Magnesio-ferri-hornblende was analyzed by electron microprobe JXA-8100  
162 operating in wavelength-dispersive mode with acceleration voltage 20 kV, beam  
163 current 20 nA, and beam diameter 2  $\mu$ m. The following standards and crystals were  
164 used for K $\alpha$  X-ray lines: Si, Na: albite, Ti: rutile; Al: sanidine; Fe: hematite; Mg, Ca:  
165 diopside; Sr: celestite; Mn: pyrophanite; K: biotite; Cl: tugtupite; OH: fluorapatite.  
166 The average of 18 analyses on three crystals, under consideration of the Mössbauer  
167 results, is given in Table 2.

168 The empirical chemical formula, calculated on the basis of 24 (O+OH+F) with

169 OH+F+Cl = 2 apfu is

170  $^A(\square_{0.62}\text{Na}_{0.23}\text{K}_{0.15})_{\Sigma 1.00}{}^B(\text{Ca}_{1.76}\text{Na}_{0.21}\text{Mn}_{0.02}\text{Sr}_{0.01})_{\Sigma 2.00}{}^C(\text{Mg}_{2.79}\text{Fe}^{2+}_{1.42}\text{Fe}^{3+}_{0.43}\text{Ti}_{0.17}\text{Al}_{0.16}$



171  $\text{Mn}_{0.03}\Sigma_{5.00}^{\text{T}}(\text{Si}_{6.94}\text{Al}_{1.06})\Sigma_{8.00}\text{O}_{22}^{\text{W}}(\text{OH}_{1.96}\text{Cl}_{0.04})\Sigma_{2.00}$ . The ideal formula is  
172  $\square\text{Ca}_2(\text{Mg}_4\text{Fe}^{3+})(\text{Si}_7\text{Al})\text{O}_{22}(\text{OH})_2$ , which requires  $\text{SiO}_2$  49.90,  $\text{Al}_2\text{O}_3$  6.05,  $\text{CaO}$  13.31,  
173  $\text{MgO}$  19.13,  $\text{Fe}_2\text{O}_3$  9.47,  $\text{H}_2\text{O}$  2.14, total 100 wt%.

#### 174 **X-ray diffraction and structure refinement**

175 X-ray powder diffraction data were collected with a two-dimensional PHOTON  
176 100 CMOS detector with Bruker D8 QUEST, using monochromated  $\text{MoK}\alpha$  radiation  
177 (0.71073 Å) from a sealed microfocus tube. Generator settings were 50 kV and 1 mA.  
178 The recorded pattern is reported in Table 3. The strongest seven diffraction lines are  
179 [ $d$  in Å( $hkl$ )]: 8.397(52)(110), 3.383(41)(150), 2.717(100)(151), 2.597(84)(061),  
180 2.545(61)(202), 1.854(49)( $\overline{172}$ ), and 1.519(62)( $\overline{622}$ ). The unit-cell parameters refined  
181 from the powder-XRD data (space group  $C2/m$ ) are:  $a = 9.889$  (5) Å;  $b = 18.153$  (6) Å;  
182  $c = 5.313$  (3) Å;  $\beta = 104.57$  (3)°;  $V = 923.1$  (7) Å<sup>3</sup>;  $Z = 2$ . The  $a:b:c$  ratio calculated  
183 from unit-cell parameters is 0.5447 : 1 : 0.2932.

184 A crystal of 0.058×0.078×0.110 mm size was examined for the single crystal XRD  
185 experiment, using a Bruker D8 QUEST diffractometer (I $\mu$ S Tube, 100W,  $\text{MoK}\alpha$   $\lambda =$   
186 0.71073 Å) and a PHOTON 100 CMOS detector. The frames were integrated with the  
187 Bruker SAINT Software Package using a narrow-frame algorithm. The integration of  
188 the data using a monoclinic unit cell yielded a total of 26665 reflections to a  
189 maximum  $\theta$  angle of 40.25° (0.55 Å resolution), of which 2953 were independent  
190 (average redundancy 9.030, completeness = 100.0%,  $R_{\text{int}} = 7.26\%$ ,  $R_{\text{sig}} = 3.51\%$ ) and  
191 2185 (73.99%) were greater than  $2\sigma$  ( $F^2$ ). The final cell constants of  $a = 9.8620$ (3) Å,  
192  $b = 18.1060$ (5) Å,  $c = 5.30810$ (10) Å,  $\beta = 104.8480$ (10)°, and  $V = 916.17$ (4) Å<sup>3</sup> are

193 based upon the refinement of the *XYZ*-centroids of 242 reflections above  $20 \sigma(I)$  with  
194  $4.485^\circ < 2\theta < 45.66^\circ$ . Data were corrected for absorption effects using the Multi-Scan  
195 method (SADABS). The ratio of minimum to maximum apparent transmission was  
196 0.918. Data collection information and structure-refinement parameters for magnesio-  
197 ferri-hornblende are given in Appendix Table 1.

198 The structure was solved and refined using the Bruker SHELXTL Software  
199 Package, using the space group *C2/m*, with  $Z = 2$  yielding the formula unit,  
200  $\text{K}_{0.07}\text{Na}_{0.72}\text{Ca}_{1.82}\text{Mg}_{3.06}\text{Fe}_{1.94}\text{Si}_{6.94}\text{Al}_{1.06}\text{O}_{24}\text{Cl}_{0.11}$ . The final anisotropic full-matrix  
201 least-squares free refinement on  $F^2$  with 116 variables converged at  $R_1 = 0.0395$  and  
202  $wR_2 = 0.0775$  using 2185 data ( $>2\sigma$ ). The goodness-of-fit was 1.058. On the basis of  
203 the final model, the calculated density was  $3.213 \text{ g/cm}^3$  and  $F(000)$ ,  $875 e^-$ . Atom  
204 coordinates, site occupancy factors and site-scattering factors, and anisotropic atomic  
205 displacement parameters are given in Appendix Table 2, and selected interatomic  
206 distances are listed in Table 4. Refined site-scattering values are given in Table 5. The  
207 Crystallographic Information File (CIF) is available as Supplemental Material.

## 208 **Derivation of site populations**

209 The occupancy of  $A(m)$  and  $A(2)$  was fixed based on the EPMA results. There are no  
210 restraints in  $B$ ,  $M(1)$ ,  $M(2)$  and  $M(3)$  sites during structure refinement. The results of  
211 structure refinement clearly show that  $A$  site is dominantly occupied by vacant,  $M(1)$   
212 and  $M(3)$  are dominantly occupied by  $\text{Mg}^{2+}$ , and  $M(2)$  is occupied by both  $\text{Mg}^{2+}$  and  
213 high-charged cations.

214 Site populations were obtained considering the experimental mean bond lengths

215 and site-scattering values for the individual sites (Hawthorne et al., 1995; Oberti et al.,  
216 2007).  $T$ Al is ordered at the  $T(1)$  site, the occupancy of  $T(1)$ Al is following the  
217 equations  $T(1)$ Al (apfu) = [ $\langle T(1)-O \rangle - 1.6193$ ]  $\times$  34.2199, and  $\langle T(1)-O \rangle$  is 1.650 Å.  
218 We get  $T(1)$ Al (apfu) = 1.061 (Oberti et al., 2007). It is in excellent agreement with the  
219 chemical analyses. So we fix the occupancy of Si (0.735) and Al (0.265) at  $T(1)$  site.  
220 The refined values of site-scattering obtained by structure refinement are in good  
221 agreement with those calculated based on the proposed site populations (Hawthorne et  
222 al., 1995). Calculation of the site populations (Table 5) was based on the  $\langle \text{cat}-O \rangle$   
223 values of Mg: 2.078 Å, Al: 1.929 Å, Ti: 1.960 Å,  $\text{Mn}^{2+}$ : 2.173 Å,  $\text{Fe}^{2+}$ : 2.125 Å, and  
224  $\text{Fe}^{3+}$ : 2.025 Å (Oberti et al., 2016b).

#### 225 **RELATION TO OTHER SPECIES**

226 The name “magnesio-ferri-hornblende” has been used in the literature  
227 (Zarubina et al., 2016), with no previous approval by the IMA-CNMNC. A certain  
228 amount of hornblende EPMA data published in literatures should belong to magnesio-  
229 ferri-hornblende, but the authors did not fully identify it and named it magnesio-  
230 hornblende (e.g., He et al., 2018; Gong et al., 2018; and references in Appendix Table  
231 3). Before the approval of magnesio-ferri-hornblende, there were four mineral species  
232 approved by the IMA-CNMNC with hornblende root name, e.g., magnesio-  
233 hornblende, ferro-hornblende, magnesio-ferri-fluoro-hornblende, and ferro-ferri-  
234 hornblende. The differences between magnesio-ferri-hornblende and the known  
235 species with hornblende root name are shown in Table 6.

236 According to “the IMA-CNMNC dominant-constituent rule revisited and

237 extended” (Hatert and Burke, 2008; Nickel and Grice, 1998) and “Nomenclature of  
238 the amphibole supergroup” (Hawthorne et al., 2012), magnesio-ferri-hornblende  
239 should be regarded as a new member of calcium amphibole subgroup. Magnesio-ferri-  
240 hornblende is the  $\text{Fe}^{3+}$  dominant at  $\text{C}^{3+}$  analogue of magnesio-hornblende.

## 241 **DISCUSSION**

### 242 **Potential mineral indicator of fertile porphyry systems**

243 Recalculating the published amphibole data according to Hawthorne et al. (2012),  
244 preliminary statistics show that magnesio-ferri-hornblende is distributed in various  
245 types of rocks from more than 40 locations worldwide (Ghent et al., 1990; El-Shazly  
246 et al., 1997; Mogessie et al., 2000; Zarubina et al., 2016; Horváth et al., 2019;  
247 Stepanov et al., 2021; Gil et al., 2022; and references in Appendix Table 3).  
248 Magnesio-ferri-hornblende occurs either in metamorphic rocks such as blueshist  
249 (Ghent et al., 1990), epidote amphibolite (El-Shazly et al., 1997), dolomite (Zarubina  
250 et al., 2016), and amphibolitized eclogite (Gil et al., 2022), or in mafic-ultramafic  
251 igneous rocks (Liu et al., 2012; Mogessie et al., 2000; Zhang et al., 2021) and more  
252 broadly in intermediate-acid intrusive rocks (for references see Appendix Table 3). Of  
253 particular note is that the occurrence frequency of magnesio-ferri-hornblende in  
254 fertilized felsic intrusions related to porphyry-skarn type Fe-Cu-Mo-Au polymetallic  
255 and magmatic hydrothermal Au mineralization is significantly higher than in barren  
256 felsic intrusions (Fig. 6), based on statistics of ~320 amphibole compositions  
257 documented in the literature (Appendix Table 3).

258 Magnesian-ferri-hornblende in fertilized felsic intrusions accounts for ~70% of  
259 the amphiboles analysed, and is followed by magnesian-hastingsite (22%) and small  
260 amounts of actinolite, Ti-rich magnesian-hastingsite, and pargasite (Fig. 6a). In barren  
261 intrusions, the proportion of magnesian-ferri-hornblende decreases significantly to  
262 ~24%, whereas the proportion of magnesian-hastingsite increases to ~47%. Hastingsite  
263 and ferro-ferri-hornblende occur only in barren intrusions and account for 20% and  
264 8.6% of the amphibole population, respectively (Fig. 6b). Such difference and  
265 variation in types and proportions of amphibole essentially reflect the change in  
266 physical-chemical conditions and the compositional evolution of magmatic systems.  
267 According to the thermobarometric formulations retrieved by Ridolfi et al. (2010), we  
268 calculated crystallization physical-chemical parameters including pressure ( $P$ ),  
269 temperature ( $T$ ), oxygen fugacity ( $f_{O_2}$ ), and  $H_2O_{melt}$  based on amphibole compositions  
270 (Fig. 7). In the  $P$ - $T$  diagram (Fig. 7a), magnesian-hastingsite and hastingsite  
271 crystallized at relatively high temperatures ( $>800^\circ\text{C}$ ) and pressures (from 110 to 520  
272 MPa), whereas magnesian-ferri-hornblende and actinolite formed under relatively low  
273 temperatures and pressures. Oxygen fugacity calculated for amphiboles from barren  
274 intrusions (FMQ to NNO+1) is generally lower by 1-2 orders of magnitude than that  
275 calculated for amphiboles from fertilized intrusions (typically from NNO+1 to  
276 NNO+3) (Fig. 7b). Figure 7c shows that  $H_2O_{melt}$  is indistinguishable between  
277 amphiboles from fertilized and barren intrusions. Therefore, barren intrusions  
278 characterized by dominant magnesian-hastingsite and distinctive hastingsite are  
279 inferred to have crystallized in high  $P$ - $T$  and low  $f_{O_2}$  environments, while fertilized

280 intrusions characterized by predominant magnesio-ferri-hornblende with subordinate  
281 magnesio-hastingsite and actinolite have formed under relatively low  $P$ - $T$  but high  $f_{O_2}$   
282 conditions.

283 A broad consensus is that porphyry-skarn polymetallic deposits are closely  
284 associated with hydrous and oxidized magmas (e.g., Ballard et al., 2002; Richards,  
285 2003; Sillitoe, 2010; An et al., 2015; Lu et al., 2016). In magmatic systems with a  
286 high oxidation state ( $\Delta FMQ > 1.5-2$ ), sulfur is dominantly present as sulfate, with the  
287 result that chalcophile elements such as Cu and Au behave as incompatible elements  
288 and are retained in the melt, which is conducive to the polymetallic mineralization in  
289 the upper crust (e.g., Carroll and Rutherford, 1985; Richards, 2003; Mustard et al.,  
290 2006; Sun et al., 2015; Duan and Jiang, 2017). Possible explanations for the high  
291 oxidation state include the transfer of fluids and the effect of later crustal processes  
292 including protracted fractionation and degassing of volatile-rich magmas (Lee et al.,  
293 2010; Humphreys et al. 2015). Water, as the most abundant volatile in magmas, is  
294 considered the key factor driving melt oxidation during degassing, however, recent  
295 studies suggest that water degassing has no effect on oxidation (Crabtree and Lange,  
296 2012; Waters and Lange, 2016). Given that  $f_{O_2}$  of fertilized intrusions is significantly  
297 higher than barren intrusions, while  $H_2O$  content of the two intrusion types is  
298 basically overlapped as shown in this study (Fig. 7c),  $H_2O$  degassing might have little  
299 effect on the oxidation state of magmatic liquids.

300 Recent study by Zhang et al. (2022) shows that bulk  $Fe^{3+}/\Sigma Fe$  ratio of cumulates  
301 from the Kohistan arc decreases from 0.4 to 0.2 with decreasing Mg#, as  $Fe^{3+}/\Sigma Fe$

302 ratios in amphibole decrease from 0.35 to 0.2, leading to increasing  $\text{Fe}^{3+}$  abundance in  
303 residual melts by multistage amphibole fractionation. Their modeling suggests that  
304 amphibole fractionation could elevate the oxidation state of evolved magmas in  
305 subduction zones, which is probably a reliable mechanism for triggering porphyry  
306 copper mineralization in subduction settings. In our study, the  $\text{Fe}^{3+}/\Sigma\text{Fe}$  ratios in  
307 amphibole of barren intrusions are mostly less than 0.3 (from 0.21 to 0.32, Fig. 8a),  
308 while the  $\text{Fe}^{3+}/\Sigma\text{Fe}$  ratios in amphibole of fertilized intrusions have a wide range and  
309 relatively high values (from 0.18 to 0.55). In terms of element occupation of  
310 amphibole, both  $\text{Fe}^{2+}$  and  $\text{Fe}^{3+}$  at C site generally decrease from barren intrusions to  
311 fertilized intrusions, with increasing of Mg at C site (Fig. 8b-c). Figure 8d shows a  
312 positive linear correlation between Mg# and  $\Delta\text{NNO}$  values, which is consistent with  
313 other previous studies (Anderson and Smith, 1995; Scaillet and Evans, 1999; Ridolfi  
314 et al., 2010; Shane and Smith, 2013), suggesting that Mg# values are redox-sensitive  
315 and can be used to evaluate the oxidation state and fertility of porphyry magmas.

316 In summary, magnesio-ferri-hornblende that occurs as a dominant amphibole  
317 species with relatively high Mg# values in mineralized intrusions is indicative of a  
318 high  $f\text{O}_2$  crystallization condition of its parent magma. Given that the more oxidized  
319 felsic magma has a better sulfur-carrying capacity and is favorable for porphyry-skarn  
320 mineralization, the extensive occurrence of magnesio-ferri-hornblende could be used  
321 as a potential indicator for identifying fertile porphyry systems.

### 322 **Linkage between magma mixing and porphyry system mineralization**

323 Mafic magmatic enclaves (MMEs) occur commonly in felsic intrusions. The

324 mixing/mingling process between injected mafic magma and crustal magma is  
325 considered to be the most likely formation mechanism (e.g., Barbarin, 2005; Kumar  
326 and Rino, 2006; Chen et al., 2009; Clemens and Stevens, 2012; Cao et al., 2022). For  
327 the felsic intrusions containing MMEs without mineralization clue at present, the  
328 proportion of amphiboles ( $n=183$ ) is similar to fertilized intrusions, but significantly  
329 different from barren intrusions (Fig. 6). Amphiboles in the felsic intrusions with  
330 magma mixing are predominated by magnesio-ferri-hornblende with a proportion of  
331 86.7%, followed by actinolite (7.2%), magnesio-hastingsite (3.9%), magnesio-  
332 hornblende (1.7%), and pargasite (0.6%). The occurrence of pargasite that represents  
333 high- $P$  mantle amphiboles (Ridolfi et al., 2010) in both felsic intrusions of magma  
334 mixing origin and fertilized intrusions indicates a contribution of mantle-derived  
335 magmas. In contrast, hastingsite and ferro-ferri-hornblende with low Mg# values  
336 ( $<0.5$ ) occur only in barren intrusions, indicating characteristics of crustal amphiboles  
337 (Ridolfi et al., 2010). Significantly, MMEs were also observed in about half of  
338 statistical fertilized intrusions (e.g., Kekeshala-Aimsduyi and Halegati-Muzuke skarn  
339 type Fe-Cu deposits, Xilicuduke, Baogutu, Machagnqing, Jono and Pulang porphyry  
340 Cu-Au-Mo deposits, Appendix Table 3 and references therein). The predominance of  
341 the new mineral magnesio-ferri-hornblende in both fertilized magmas and unfertilized  
342 magma mixing intrusions suggests that the recharge of mafic magma marked by  
343 MMEs should play an important role in porphyry system mineralization.

344 Keith et al. (1997) proposed that the volcanic rocks from the Bingham and Tintic  
345 districts (Utah) were formed by mixing of an S-rich mafic alkaline magmas with more



346 silicic magmas, indicating a large budget of mantle-derived S, Cu, Ag, and Au was  
347 contributed to the process of the formation of porphyry and mesothermal vein world-  
348 class deposits. Richards (2003) suggested that mixing between crustal- and mantle-  
349 derived magmas yields evolved (andesitic to dacitic), volatile-rich, metalliferous,  
350 hybrid magmas, which are of sufficiently low density to rise through the crust.  
351 Mungall et al. (2015) used laboratory experiments to show that droplets of sulfide  
352 melt in mafic magmas (which subsequently mix with more felsic magmas) can attach  
353 to vapor bubbles to form compound drops that float, which is a key step in the  
354 formation of porphyry deposits. Cao et al. (2022) identified two groups of amphibole  
355 crystals with different mineral chemistry and suggested that MMEs and host granites  
356 at the giant Pulang porphyry Cu-Au deposit in Tibet are sourced from mafic and felsic  
357 magma reservoirs, respectively. Together with partitioning models, they proposed that  
358 repeated mafic magma recharge may have supplied significant amounts of S and some  
359 Cl, thus aiding formation of the large, long-lived magma reservoir that produced the  
360 porphyry deposits. Indeed, recharge of mafic magma may contribute significant  
361 amount of sulfur-, metal-, and/or chlorine-rich magmatic fluids to the felsic magma  
362 (Hattori and Keith, 2001; Audétat and Simon, 2012; Zhang et al., 2019; Yang et al.,  
363 2020; Cao et al., 2022) or remobilize precipitated magmatic sulfides (Wilkinson,  
364 2013). Consequently, magma mixing recorded by MMEs and/or contemporaneous  
365 mafic intrusions as well as massive magnesio-ferri-hornblende could be used to  
366 identify felsic magma fertility that is particularly predisposed to ore formation.  
367 However, not all felsic melts mixed with mafic magmas can eventually evolve into

368 fertile magmas required to generate ore deposits, because formation of porphyry ore  
369 systems is controlled by the interplay of several processes, including cyclical  
370 enrichment of magmas with metals and volatiles in the deep crust, sulfide saturation  
371 of the magmas, efficient transfer of metals into hydrothermal fluids exsolved from the  
372 magmas, and efficient precipitation of ore minerals at the deposit trap sites (Wilkinson,  
373 2013).

#### 374 **IMPLICATIONS**

375 Magnesio-ferri-hornblende, the new mineral found in felsic magmas, is ubiquitous  
376 in both fertilized magmas related to porphyry ore system and granitoids containing  
377 MMEs of magma mixing origin, but rare in barren intrusions. Our findings indicate  
378 that amphibole compositions and proportions, especially massive occurrences of  
379 magnesio-ferri-hornblende in felsic magmas may represent a potential application as a  
380 useful tool to identify ore-bearing and barren intrusions within porphyry-skarn and  
381 hydrothermal polymetallic deposits. Amphiboles in barren intrusions are dominantly  
382 characterized by magnesio-hastingsite and hastingsite crystallized in high  $P$ - $T$  and low  
383  $f_{O_2}$  environments. The distinctive hastingsite and ferro-ferri-hornblende observed only  
384 in barren intrusions are mainly crustal amphiboles (high  $Fe^{2+}$  at  $C$  sites and  $Mg\# < 0.5$ ).  
385 Felsic magmas with ore-potential record evidence of lower  $P$ - $T$  and high  $f_{O_2}$   
386 conditions, characterized by predominant magnesio-ferri-hornblende and subordinate  
387 magnesio-hastingsite that have low  $Fe^{2+}$  at  $C$  site and high  $Mg\#$  values. Importantly,  
388 amphibole compositions could be an effective exploration indicator, based on an  
389 empirical relationship between magmatic  $f_{O_2}$  and  $Mg\#$  (Ridolfi et al., 2010). Magma

390 mixing process has been genetically linked to porphyry system mineralization.  
391 However, there are also felsic rocks containing MMEs that have not shown clues of  
392 mineralization yet, which should be paid attention in the further target mineralization  
393 exploration.

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634 **FIGURE CAPTIONS**

635 **FIGURE 1.** (a) Sketch tectonic map of the Central Asian Orogenic Belt (modified  
636 after Long et al., 2011); (b) Geological map of the Western Tianshan (NW  
637 China), showing the location of major skarn and porphyry deposits (modified  
638 after Wang et al., 2018, and references therein); (c) Regional geologic map of the  
639 Boluokenu area in the Western Tianshan (modified after Gu et al., 2014 and  
640 Zhang et al., 2016). Magnesio-ferri-hornblende was found in the granitoids  
641 marked by yellow stars.

642 **FIGURE 2.** Photographs showing occurrence of magnesio-ferri-hornblende. (a) Hand  
643 specimen photo of granodiorite from the Husite complex that contains magnesio-  
644 ferri-hornblende. (b) Thin section of granodiorite containing magnesio-ferri-  
645 hornblende (mosaics of multiple images taken in plane-polarized light). (c and d)  
646 Magnesio-ferri-hornblende is associated with plagioclase, quartz, K-feldspar, and  
647 biotite (viewed in plane-polarized light). Mfhbl = magnesio-ferri-hornblende; Pl  
648 = plagioclase; Qtz = quartz; Kf = K-feldspar; Bt = biotite.

649 **FIGURE 3.** The Mössbauer spectrum of magnesio-ferri-hornblende

650 **FIGURE 4.** The infrared spectrum of magnesio-ferri-hornblende

651 **FIGURE 5.** The Raman spectrum of magnesio-ferri-hornblende

652 **FIGURE 6.** Pie graphs showing proportion of amphiboles under different geological

653 backgrounds. (a) Amphiboles in felsic magmas related to porphyry-skarn Fe-Cu-  
654 Mo-Au and hydrothermal Au mineralization; (b) Amphiboles from barren felsic  
655 magmas in the same ore belt of fertilized magmas; (c) Amphiboles in felsic  
656 magmas containing mafic magmatic enclaves (MMEs) of magma mixing origin,  
657 where no mineralization clue has yet been found.

658 **FIGURE 7.**  $P$ - $T$  (a),  $\log f_{O_2}$ - $T$  (b), and  $T$ - $H_2O_{melt}$  (c) diagrams for several types of  
659 amphiboles from fertile and barren felsic intrusions. The curves FMQ, NNO and  
660 MH are after O'Neill (1987a, 1987b) and Schwab and Küstner (1981). Owing to  
661 the applicability of amphibole thermobarometry (Ridolfi et al., 2010),  $T$ - $P$ -  
662  $H_2O_{melt}$ - $f_{O_2}$  parameters were partially determined for certain amphiboles. Mfhbl =  
663 Magnesio-ferri-hornblende, Mhst = Magnesio-hastingsite, Ti-Mhst = Ti-rich  
664 Magnesio-hastingsite, Act = Actinolite, Hst = Hastingsite.

665 **FIGURE 8.** (a)  $Fe^{3+}/\Sigma Fe$  ratios versus Mg# diagram for amphiboles; (b and c)  ${}^C Fe^{2+}$   
666 and  ${}^C Fe^{3+}$  versus  ${}^C Mg$  diagrams for amphiboles; (d)  $\Delta NNO$  values versus Mg#  
667 diagram for amphiboles. Mg# = atomic  $[Mg/(Mg+Fe_T)]$ , where  $Fe_T = Fe^{2+} + Fe^{3+}$ .  
668 Mfhbl = Magnesio-ferri-hornblende, Mhst = Magnesio-hastingsite, Ti-Mhst = Ti-  
669 rich Magnesio-hastingsite, Act = Actinolite, Prg = Pargasite, Hst = Hastingsite,  
670 Ffhbl = Ferro-ferri-hornblende.

671 **TABLE CAPTIONS**

672 **TABLE 1.**  ${}^{57}Fe$  hyperfine parameters for magnesio-ferri-hornblende

673 **TABLE 2.** Chemical composition (wt%) and unit formula (apfu) for magnesio-ferri-  
674 hornblende

- 675 **TABLE 3.** X-ray powder diffraction data ( $d$  in Å) for magnesio-ferri-hornblende
- 676 **TABLE 4.** Selected interatomic distances (Å) for magnesio-ferri-hornblende
- 677 **TABLE 5.** Site populations (apfu) for magnesio-ferri-hornblende
- 678 **TABLE 6.** Predominant-constituents in the structural sites of magnesio-ferri-
- 679 hornblende in the hornblende species approved so far
- 680 **SUPPLEMENTAL MATERIALS**
- 681 **APPENDIX TABLE 1.** Miscellaneous information for magnesio-ferri-hornblende.
- 682 **APPENDIX TABLE 2.** Atomic coordinates, equivalent isotropic and anisotropic atomic
- 683 displacement parameters (Å<sup>2</sup>) for magnesio-ferri-hornblende.
- 684 **APPENDIX TABLE 3.** Major elements (wt%) from electron microprobe analysis and
- 685 calculated physical-chemical conditions of amphiboles from felsic fertilized
- 686 intrusions (a), barren intrusions (b) and intrusions of magma mixing origin (c).
- 687 **ONLINE MATERIAL.** The Crystallographic Information File (CIF) for magnesio-ferri-
- 688 hornblende.



TABLE 1.  $^{57}\text{Fe}$  hyperfine parameters for magnesio-ferri-hornblende

Component	Isomer shift(mm/s)	Quadrupole splitting (mm/s)	Line width (mm/s)	Area (%)
$\text{Fe}^{2+}$	0.964±0.003	2.431±0.007	0.571±0.011	77.62
$\text{Fe}^{3+}$	0.046±0.007	0.935±0.014	0.384±0.020	22.38

TABLE 2. Chemical composition (wt%) and unit formula (apfu) for

magnesio-ferri-hornblende

Constituent	Mean	Range	SD* ( $\sigma$ )		apfu
$\text{SiO}_2$	47.37	46.76-48.29	0.489	$\text{Si}^{4+}$	6.94
$\text{TiO}_2$	1.51	1.28-1.78	0.166	$\text{Al}^{3+}$	1.06
$\text{Al}_2\text{O}_3$	7.07	6.4-7.6	0.363	$\Sigma\text{T}$	8.00
$\text{FeO}_{\text{total}}$	15.09	13.78-16.07	0.700	$\text{Al}^{3+}$	0.16
$\text{Fe}_2\text{O}_3^{**}$	3.86			$\text{Ti}^{4+}$	0.17
$\text{FeO}^{**}$	11.62			$\text{Fe}^{3+}$	0.43
$\text{MgO}$	12.77	12.05-13.94	0.587	$\text{Fe}^{2+}$	1.42
$\text{CaO}$	11.22	10.66-11.55	0.217	$\text{Mg}^{2+}$	2.79
$\text{SrO}$	0.15	0.09-0.21	0.030	$\text{Mn}^{2+}$	0.05
$\text{MnO}$	0.39	0.26-0.48	0.062	$\text{Sr}^{2+}$	0.01
$\text{Na}_2\text{O}$	1.54	1.30-1.95	0.141	$\text{Ca}^{2+}$	1.76
$\text{K}_2\text{O}$	0.78	0.58-0.93	0.084	$\text{Na}^+$	0.21
$\text{Cl}$	0.15	0.06-0.20	0.030	$\Sigma\text{B+C}$	7.00
$\text{Cl}\equiv\text{O}$	-0.03			$\text{Na}^+$	0.23
$\text{H}_2\text{O}^{***}$	2.01			$\text{K}^+$	0.15
$\text{F}$	0			$\Sigma\text{A}$	0.38
Total	100.41			$\text{Cl}^-$	0.04
				$(\text{OH})^-$	1.96
				$\Sigma\text{W}$	2.00

Note: \*\*Conversion of  $\text{FeO}_{\text{total}}$  by Mössbauer spectrum data;  
 \*\*\*Stoichiometric calculation on the basis of structural refinement; SD\*  
 Standard deviation, quoted at  $1\sigma$  level

TABLE 3. X-ray powder diffraction data ( $d$  in Å) for magnesio-ferri-hornblende

$l_{\text{meas}}$	$d_{\text{meas}}$	$d_{\text{calc}}$	$h$	$k$	$l$
<b>52</b>	<b>8.397</b>	<b>8.466</b>	<b>1</b>	<b>1</b>	<b>0</b>
7	5.082	5.115	1	3	0
26	4.890	4.906	1	$\bar{1}$	$\bar{1}$
21	4.526	4.538	0	4	0
3	4.222	4.233	2	2	0
26	3.882	3.898	1	$\bar{3}$	$\bar{1}$
<b>41</b>	<b>3.383</b>	<b>3.395</b>	<b>1</b>	<b>5</b>	<b>0</b>
27	3.279	3.293	2	4	0
38	3.134	3.142	3	1	0
22	2.947	2.957	1	$\bar{5}$	$\bar{1}$
6	2.803	2.822	3	3	0
<b>100</b>	<b>2.717</b>	<b>2.723</b>	<b>1</b>	<b>5</b>	<b>1</b>
<b>84</b>	<b>2.597</b>	<b>2.608</b>	<b>0</b>	<b>6</b>	<b>1</b>
<b>61</b>	<b>2.545</b>	<b>2.548</b>	<b>2</b>	<b>0</b>	<b><math>\bar{2}</math></b>
26	2.341	2.349	3	$\bar{5}$	$\bar{1}$
26	2.298	2.311	1	$\bar{7}$	$\bar{1}$
9	2.218	2.222	2	$\bar{4}$	$\bar{2}$
37	2.167	2.176	2	6	1
20	2.131	2.131	4	$\bar{4}$	$\bar{1}$
18	2.022	2.024	4	0	$\bar{2}$
<b>49</b>	<b>1.854</b>	<b>1.855</b>	<b>1</b>	<b><math>\bar{7}</math></b>	<b><math>\bar{2}</math></b>
5	1.813	1.811	1	9	1
2	1.750	1.754	5	$\bar{1}$	$\bar{2}$
33	1.691	1.691	3	$\bar{1}$	$\bar{3}$
16	1.653	1.653	4	$\bar{8}$	$\bar{1}$
21	1.623	1.626	1	11	0
39	1.585	1.585	5	$\bar{5}$	$\bar{2}$
<b>62</b>	<b>1.519</b>	<b>1.518</b>	<b>6</b>	<b><math>\bar{2}</math></b>	<b><math>\bar{2}</math></b>
14	1.476	1.475	3	7	2
12	1.460	1.460	1	$\bar{7}$	$\bar{3}$
28	1.341	1.340	3	9	2
37	1.299	1.297	0	14	0

Note: The intensities of the seven strongest lines are in **bold**.

TABLE 4. Selected interatomic distances (Å) for magnesio-ferri-hornblende

A(m)	-O5×2	2.9840	A(2)	-O5×2	2.6200
	-O6×2	2.8280		-O6×2	2.8300
	-O7	2.5710		-O7×2	2.5390
	-O7	2.4920		<A(2)-O>	2.6630
	<A(m)-O>	2.7812			
M(4)	-O2×2	2.4124	M(4')	-O2×2	2.1400
	-O4×2	2.3251		-O4×2	2.2740
	-O5×2	2.7210		-O5×2	2.9360
	-O6×2	2.5395		-O6×2	2.8090
	<M(4)-O>	2.4995		<M(4')-O>	2.5398
M(1)	-O1×2	2.0610	M(2)	-O1×2	2.1112
	-O2×2	2.1192		-O2×2	2.0862
	-O3×2	2.0949		-O4×2	1.9933
	<M(1)-O>	2.0917		<M(2)-O>	2.0636
M(3)	-O1×4	2.0961	T(1)	-O1	1.6382
	-O3×2	2.0834		-O6	1.6604
	<M(3)-O>	2.0919		-O7	1.6398
T(2)	-O4	1.5990		-O5	1.6613
	-O5	1.6491		<T(1)-O>	1.6499
	-O2	1.6260			
	-O6	1.6630			
	<T(2)-O>	1.6343			

TABLE 5. Site populations (apfu) for magnesio-ferri-hornblende

Site	Site population(apfu)	Site scattering (epfu)		Bond distance (Å)	
		Refined	Calculated	Refined	Calculated
T(1)	2.94Si+1.06Al	54.94	54.94		
T(2)	4Si	56.00	56.00		
M(1)	1.3Mg+0.7Fe <sup>2+</sup>	34.86	33.80	2.093	2.092
M(2)	0.94Mg+0.3Fe <sup>2+</sup> +0.43Fe <sup>3+</sup> +0.16Al+0.17Ti	34.16	36.08	2.063	2.064
M(3)	0.55Mg+0.42Fe <sup>2+</sup> +0.03Mn	18.09	18.30	2.092	2.092
ΣC cations		87.12	88.18		
B cations	1.76Ca+0.21Na+0.02Mn+0.01Sr	39.72	38.39		
A cations	0.23Na+0.15K	6.03	5.38		
W cations	1.93OH+0.07Cl	19.84	18.56		

TABLE 6. Predominant-constituents in the structural sites of magnesio-ferri-hornblende  
 in the hornblende species approved so far

Mineral name	Site occupation						Reference
	<i>A</i>	<i>B</i>	<i>C</i> <sup>2+</sup>	<i>C</i> <sup>3+</sup>	<i>T</i>	<i>W</i>	
Magnesio-ferri-hornblende	□	Ca	Mg	Fe <sup>3+</sup>	Si <sub>7</sub> Al	OH	This work
Magnesio-hornblende	□	Ca	Mg	Al	Si <sub>7</sub> Al	OH	Oberti et al. (2018)
Ferro-hornblende	□	Ca	Fe	Al	Si <sub>7</sub> Al	OH	Hawthorne et al. (2012)
Magnesio-ferri-fluoro-hornblende	□	Ca	Mg	Fe <sup>3+</sup>	Si <sub>7</sub> Al	F	Oberti et al. (2016a)
Ferro-ferri-hornblende	□	Ca	Fe <sup>2+</sup>	Fe <sup>3+</sup>	Si <sub>7</sub> Al	OH	Oberti et al. (2016b)

Figure 1, Zhang et al., AM

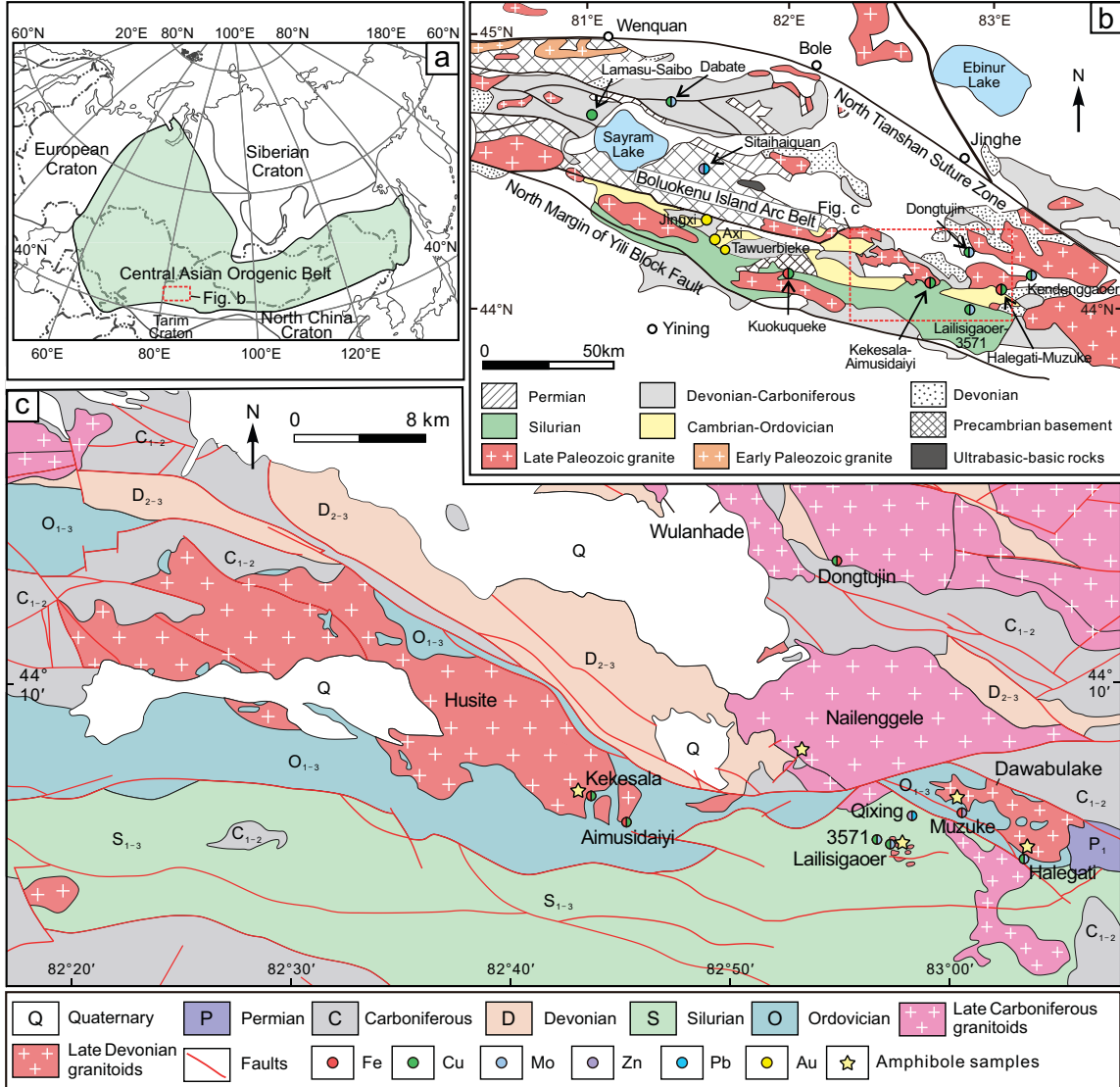


Figure 2, Zhang et al., AM

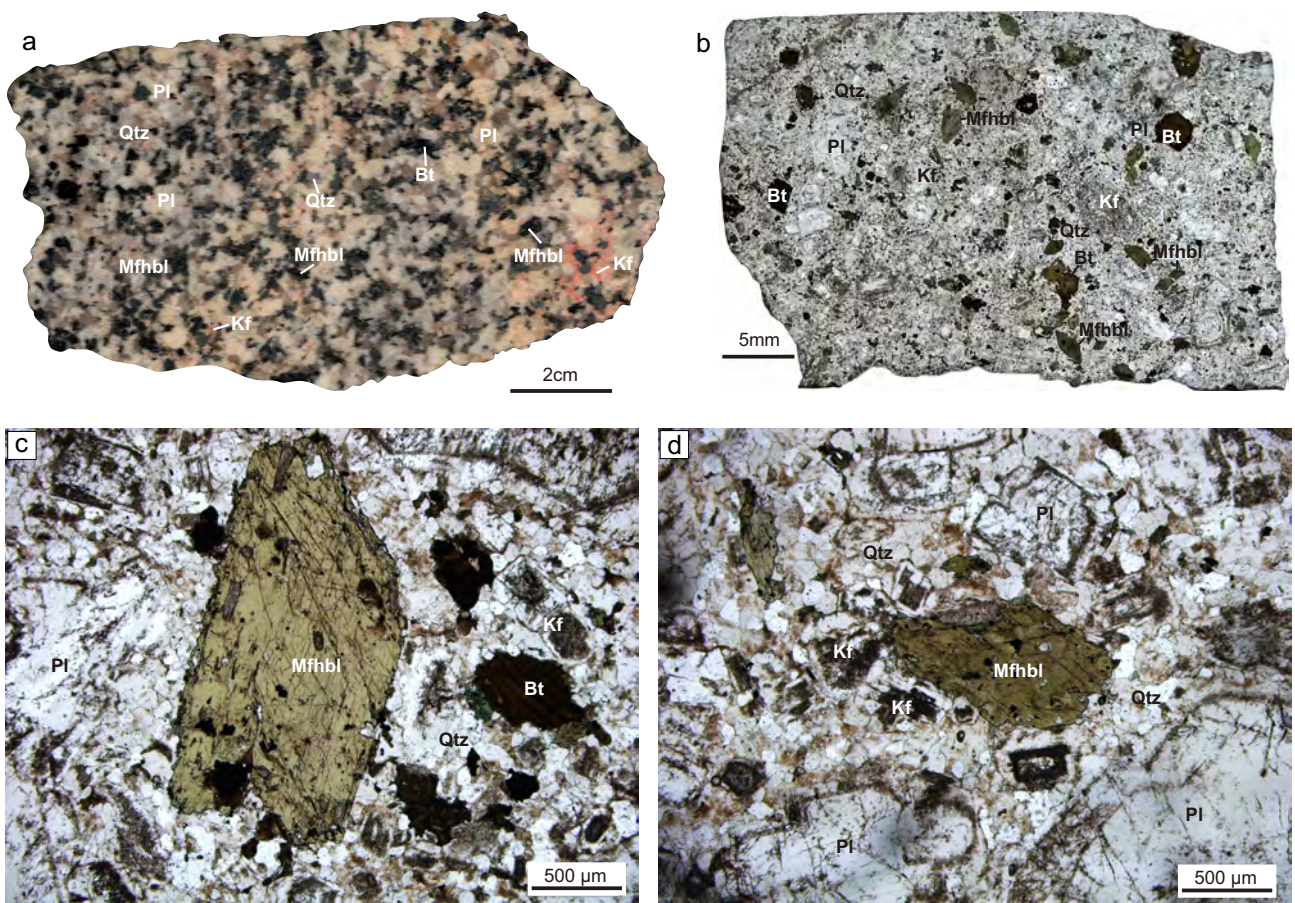


Figure 3, Zhang et al., *AM*

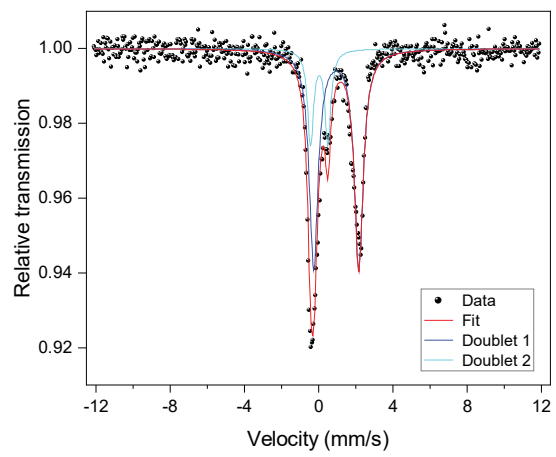


Figure 4, Zhang et al., AM

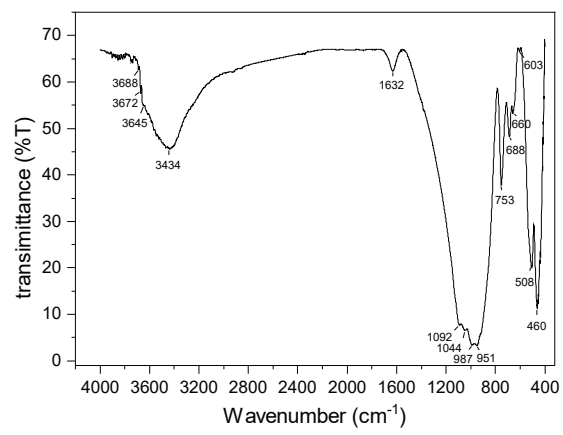




Figure 5, Zhang et al., *AM*

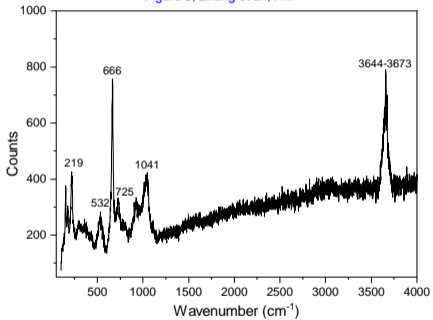


Figure 6, Zhang et al., AM

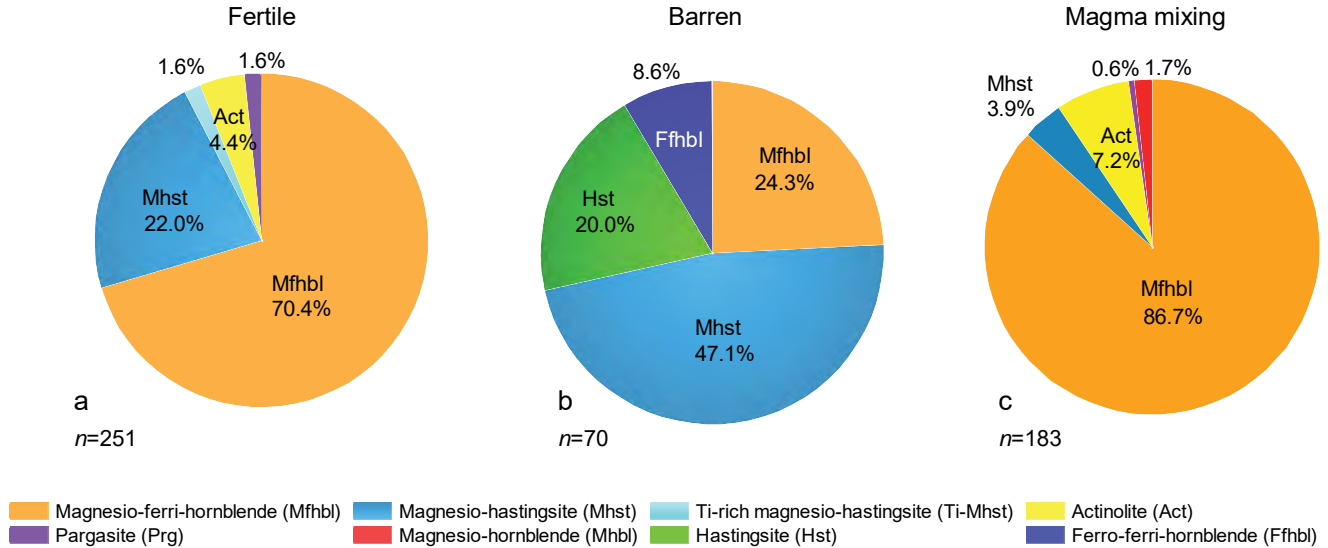


Figure 7, Zhang et al., AM

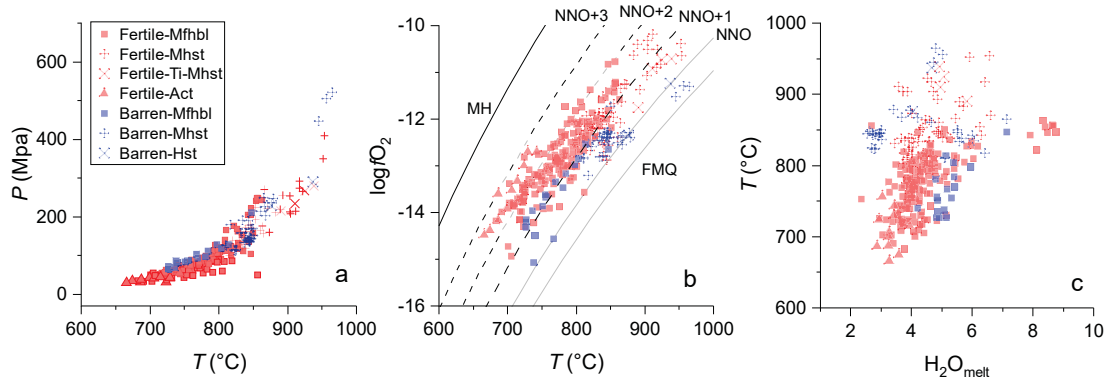


Figure 8, Zhang et al., *AM*

