MICHIGAN STATE UNIVERSITY

National Superconducting Cyclotron Laboratory

GENERALIZED ISOSCALING OF ISOTOPIC DISTRIBUTIONS

R. SHOMIN, M.B. TSANG, O. BJARKI, C.K. GELBKE, G.J. KUNDE, R.C. LEMMON, W.G. LYNCH, D. MAGESTRO, R. POPESCU, A.M. VANDER MOLEN, G. VERDE, G.D. WESTFALL, H.F. XI, W.A. FRIEDMAN, G. IMME, V. MADDALENA, C. NOCIFORO, G. RACITI, G. RICCOBENE, F.P. ROMANO, A. SAIJA, C. SFIENTI, S. FRITZ, C. GROSS, T. ODEH, C. SCHWARTZ, A. NADASEN, D. SISAN, K.A.G. RAO

CERN LIBRARIES, GENEVA



CM-P00040633

#2353667



JUNE 2002

Generalized Isoscaling of Isotopic Distributions

R.Shomin, M.B. Tsang, O. Bjarki, C.K. Gelbke, G.J.Kunde¹, R.C. Lemmon², W.G. Lynch, D. Magestro, R. Popescu³, A.M. Vandermolen, G. Verde, G.D. Westfall, H.F. Xi

Department of Physics and Astronomy and National Superconducting Cyclotron Laboratory, Michigan State University, East Lansing, MI 48824, USA,

W. A. Friedman

Department of Physics, University of Wisconsin, Madison, WI 53706, USA
G. Imme, V. Maddalena, C. Nociforo, G. Raciti, G. Riccobene, F.P. Romano, A. Saija, C. Sfienti.

Dipartimento di Fisica Università and I.N.F.N. Sezione and Laboratorio Nazionale del SUD - Catania, I95127, Catania, Italy

S. Fritz, C. Groß, T. Odeh, C. Schwarz,

Gesellschaft für Schwerionenforschung, D-64220 Darmstadt, Germany,

A. Nadasen, D. Sisan, K.A.G. Rao

Department of Natural Sciences, University of Michigan, Dearborn, MI 48128, USA

Abstract

Generalized isoscaling relationships are proposed that may permit one to relate the isotopic distributions of systems that may not be at the same temperature. The proposed relationships are applied to multifragmention excitation functions for central Kr+Nb and Ar+Sc collisions. The generalized scaling functions provide a more reasonable description of central collisions for the larger Kr+Nb system than for the smaller Ar+Sc system. For the latter case, the simple thermal approximations at the core of the generalized isoscaling relationships become increasingly inaccurate at higher incident energies.

Present address: Brookhaven National Laboratory, Upton, NY 11973, U.S.A.

¹ Present address: Yale University, New Haven CT 06520, U.S.A.

² Present address: CCLRC, Daresbury Laboratory, Daresbury, Cheshire WA44AD, U.K.

At incident energies in excess of about E/A=30 MeV, a rapid collective expansion of the combined system may occur during the later stages of a central collision between heavy nuclei [1,2]. At densities less than about 1/3 saturation density such systems disassemble into a mixture of fragments and light particles; the duration of fragment emission is of the order of 100 fm/c [3,4]. Even though the emission time is short, statistical models such as the bulk multifragmentation models which assume equilibrium at a single breakup density and temperature, have been used successfully to describe many experimental observables such as the fragment multiplicities, charge distributions, and the energy spectra of the emitted fragments [2,5-7]. These descriptions require careful, though not necessarily obvious, choices for the source size, excitation energy and collective velocity of expansion [2,6,7,8]. Many of these statistical models display a phase transition in nuclear matter with subsaturation density [9,10]; such models have been employed to extract the caloric curve, i.e. the relationship between excitation energy and temperature for the nuclear liquidgas phase transition [5,11-17] and to address whether finite nuclear systems may display negative heat capacities in analogy to those deduced for finite metallic clusters [18].

The success of thermal models raises the fundamental question of whether local thermal equilibrium is achieved in such collisions. It is important to note that there are problems with determining both the excitation energy [2,5,7,19] and the temperature [12-15,20-26] of multifragmenting systems. After correction for collective expansion, for example, calculated excitation energies for peripheral collisions at high energies must be further reduced by roughly 30% to reproduce experimental data and larger corrections are estimated for energetic central collisions [2,7,19,27]. Collective motion, pre-equilibrium emission and Coulomb barrier fluctuations increase significantly the temperatures deduced from kinetic energy spectra [1-3,12,19] while secondary decay modifies the temperatures deduced from excited state populations and isotope ratios [21,22]. Resolving these problems is an important priority.

Regardless of whether thermal equilibrium is actually achieved in such collisions, thermal models will remain extremely useful because they suggest a simple

and transparent description of the data. Here, we focus upon the description of isotopic data where thermal models have suggested an isotopic thermometer obtained from a double ratio of isotopic yields and an isoscaling relationship obtained from a single ratio of isotopic yields. The isoscaling relationship provides a remarkably accurate way to relate data for systems of different isotopic composition but with approximately the same excitation energy per nucleon or temperature [8,28-30].

In this paper, we begin by introducing these basic observables. We then generalize the isoscaling relationship to allow comparisons of systems at different temperatures and examine the accuracy of this generalized isoscaling relationship. We use the isotopic thermometer to provide input regarding the temperature difference required by the generalized isoscaling relationship. We study the empirical comparison between the isotopic temperatures and generalized isoscaling parameters obtained for a heavier Kr+Nb and those obtained for a lighter Ar+Sc system as a function of the incident energy per nucleon. These comparisons suggest that the lighter Ar+Sc system to be less consistent with thermal descriptions than the heavier Kr+Nb system.

Isotopic thermometers and isoscaling parameterizations

Due to the ease of measuring isotope cross-sections, the most widely used experimental method to measure temperatures in the caloric curves is to determine the relative isotopic abundances of two pairs of isotopes with large binding energy differences, B [11-17,20-23,25,26]. Most experimental isotope temperature determinations have used the following expression [23]:

$$\mathbf{T_{iso}} = \frac{\mathbf{B}}{\ln(\mathbf{a} \cdot \mathbf{R})} \tag{1}$$

where R = (Y(3)/Y(4))/(Y(1)/Y(2)) is fragment yield ratio of the ground states for isotope pairs (3,4) and (1,2), a is a ground state spin factor. Information on the four thermometers studied in the present work is listed in Table I.

Eq. (1) assumes that the excited systems are at statistical equilibrium and that the systems can be approximated by grand canonical ensembles i.e. finite size effects

and effects of sequential decays on the isotope yields are negligible. However, the sequential decay effects can be significant and at high temperatures dependent on the contributions of very short-lived unbound resonances [16,17,22,25,31-35]. These contributions can be calculated subject to certain model dependent assumptions about the continuum contributions and determined by direct measurements of the decays of these unbound states

It has been found empirically that isotope ratios from two statistical processes with same temperature exhibit isoscaling [28,29], i.e. the isotope ratios depends exponentially on the neutron number, N, and proton number, Z, of the isotope (N, Z)

$$R_{21}(N,Z)=Y_2(N,Z)/Y_1(N,Z)=Cexp(\alpha N+\beta Z)$$
 (2)

where α and β can be treated as empirical fitting parameters and C is the overall normalization factor. Eqs. (1) and (2) can be derived from the simple Grand Canonical model expression for the primary fragment yield for ith fragment in its kth state before secondary decay:

$$Y_{i,primary} \approx V \frac{A_i^{3/2}}{\lambda_T^3} (2J_{ik} + 1) \exp[(Z_i \mu_p + N_i \mu_n + B_{ik})/T]$$
 (3)

where μ_p and μ_n are the proton and neutron chemical potentials, $\lambda_T = h/\sqrt{2mT}$, B_{ik} and J_{ik} are the binding energy and spin of the fragment in the k^{th} state, and V is the free (unoccupied) volume of the system. The insertion of the ground state yields predicted by Eq. (3) into Eq. 2 results in the cancellations of binding energy terms provided the temperatures of the two reactions are equal. Similarly, the insertion of the ground state yields predicted by Eq. (3) into Eq. (1) results in the cancellation of the chemical potential terms; the spin and mass number terms contribute to the factor a in Eq. (1).

What is measured in an experiment, however, are the secondary yields after sequential decay. Calculations of the yields of secondary fragments after sequential decay require an accurate accounting for feeding from the particle decay of highly excited heavier nuclei [25,32,33]. Such calculations are doable but are somewhat non-transparent and subject to uncertainties regarding the levels that can be excited and the structure effects that govern their decay [25,26,32,33]. To construct simple

thermal expression, we adopt instead the thermal expressions in Eqs. 1 and 2 as rough empirical guides to the possible relationships between the temperature and the charge and mass distributions and explore the extent to which they can be used to describe experimental observations. A similar approach has been taken with Eq. (2) in refs. [28,29] and justified therein by statistical model calculations [30], which suggests that secondary decay corrections largely cancel when the two systems are at the same temperature. Likewise, this approach has also been taken with the isotope thermometric expression in Eq. (1); discussions of the modifications of Eq. (1) can be found in refs. [25,26,32,33].

We take this approach in order to see whether the isoscaling relationship can be extended to consider two systems at different temperatures. In general, the binding energy factors in Eq. (3) are not cancelled by the ratio in Eq. (2) if the two systems have different temperatures. However, one may try to extrapolate the isoscaling behavior to systems with different temperatures by multiplying $R_{21}(N,Z)$ by a binding energy dependent term:

$$R_{21}(N,Z)\exp(k_{21}\cdot BE(N,Z)) = C'\exp(\alpha'N + \beta'Z), \tag{4}$$

where $k_{21}=1/T_1-1/T_2$ is a temperature dependent correction factor. Because the two systems are at different temperatures, the scaling relationship of Eq. (4) may be more sensitive than that of Eq. (2) to the temperature dependent secondary decay corrections to the isotopic yields. In the following, we will use measured isotope ratio temperatures in Eq. (3) to test whether empirical isotope temperatures and the generalized isoscaling relationship in Eq. (4) can describe the evolution of the isotope distributions with excitation energy. We note that it might be possible to invert Eq. (4) and obtain a temperature for one system if the temperature of the other is known.

Experimental Analyses

The experiment was performed by bombarding ⁴⁵Sc targets of 3 mg/cm² areal density with ³⁶Ar beams at E/A=50, 100, 150 MeV and ⁹¹Nb targets of 3 mg/cm² areal density with ⁸⁶Kr beams at E/A=35, 50, 100, and 120 MeV at the National

Superconducting Cyclotron Laboratory at Michigan State University (MSU) [16]. Impact parameters were selected by assuming that the average multiplicity of identified charged particles detected at polar angles of 7° - 157° with 215 plastic ΔE -E phoswich of the MSU 4π array decreases monotonically with impact parameter [36,37]. Central collisions were defined by the requirement that the multiplicity N of identified charged particles lies within the highest 20% of the multiplicity distribution for N \geq 3. If b_{max} denotes the impact parameter corresponding to <N> = 3, this centrality requirement corresponds to values of the reduced impact parameter $\hat{b} = b/b_{max}$ of $\hat{b} < 0.45$.

We replaced two hexagonal modules of the MSU 4π array, located at 37° and 79° by 96 telescopes that covered approximate polar and azimuth angular ranges of 43° and 40°, respectively, in the laboratory [16]. To provide good coverage for light charged particles emitted at center-of-mass angles of 90°, where contributions from the decay of projectile-like and target-like fragments are minimal, the central angle of the hodoscope was placed at 47.9°, 42.6°, and 40.6° at incident energies of E/A=50, 100, 150 MeV, respectively. Each of these telescopes subtended a solid angle of 1.83 msr and consist of a 300 um thick silicon detector followed by 6 cm thick CsI(Tl) scintillation detector. The centers of neighboring telescopes were separated by relative angles of 3.3°.

The silicon detectors of these telescopes were calibrated to an accuracy of 2% using a precision pulser and alpha particles emitted from a ²¹²Po source. The CsI(Tl) scintillators were calibrated to an accuracy of 3% with recoil protons elastically scattered from a CH₂ target by ⁸⁶Kr ions at E/A=35 MeV and ⁴He ions at E/A=22 and 40 MeV. The accuracies (≈3%) of the overall calibrations are largely governed by the accuracy of the CsI(Tl) calibrations. With these telescopes, the spectra of isotopically resolved light particles with Z less than five were measured at angles between 20° and 70° in the laboratory frame.

The experimental apparatus at each incident energy samples somewhat different kinematic regions of the reaction. To compare similar kinematic regions for

the various incident energies and targer-projectile combinations and to minimize contaminations from the projectile- and target-like spectators, we have extracted the isotope yields for elements Z=1-4 at $\theta_{\text{CM}}=80^{\circ}$ -110° with center of mass energy threshold of 5 MeV per nucleon for all isotopes. The extracted yields lie within the acceptance of the hodoscope; they were obtained by fitting the experimental data and using the fits to predict the corresponding center of mass yields. The uncertainties in the fitted yields were determined by varying the fits. The energy thresholds are necessary because the experimental set up does not provide measurements at very forward angles. The low energy thresholds also minimize contributions from the evaporation of the residues at the lowest incident energies.

Fig. 1 shows the excitation function of the isotopic temperature measurements for the two systems, Kr+Nb (left panel) and Ar+Sc reactions (right panel). Different symbols represent different thermometers as specified in the left panel. The lines are drawn to guide the eyes. The increase of the ratio temperatures with incident energy seems to be more systematic for the Kr+Nb system. At the incident energy of 100A MeV where both systems have measurements, the isotope ratio temperatures obtained in the Ar+Sc are consistently higher than the temperatures obtained in the Kr+Nb system even though the uncertainties are large. We note that a dependence of isotopic temperatures on the system size has been reported previously [14].

Earlier studies have attributed the differences between the apparent temperatures for the different isotope pairs to the influence of the secondary decays of the heavier excited fragments formed in the early stages of the collisions. Over a moderate temperature range, one can reduce the differences between the thermometers by using the empirical relationship [15,20,21,24,]

$$1/\text{To} = 1/\text{Tapp} - \ln \kappa/B \tag{5}$$

where the values of lnk/B depends on specific isotope pairs used. Following this empirical approach, we applied the correction factors from Refs. [20,21] as listed in Table I to all of the isotopic temperatures in Fig.1. The resulting values for T₀, shown in Fig. 2 (kr_ar_to), display an increase with incident energy, but the variations

between different thermometers are much smaller than for the original temperature T_{app} . The variations in the corrected temperatures, T_0 , are larger for the Ar+Sc system (right panel) than for the Kr+Nb system (left panel).

To examine whether a generalized isoscaling can be applied to these reactions, we construct the isotope ratios, R₂₁(N,Z) from measurements on the same system at two different incident energies. The top panels of Fig. 3 show the isotope ratios measured in Kr+Nb collisions and the top panels in Fig. 4 show the isotope ratios measured in Ar+Sc collisions for Z=1, (open circles), Z=2 (closed circles), Z=3 (open squares), and Z=4 (closed squares) isotopes and different combinations of incident energies. The different incident energies involved in each ratio are labeled in each panel; e.g. the notation "70/35" in the upper left panel in Fig. 3 denotes the ratio of isotopic yields measured at E/A=75 MeV in Kr+Nb collisions to the corresponding yields measured at E/A=35 MeV. For simplicity, we adopt the convention that isotope yields from the higher energy collision are placed in the numerator. Clearly, the raw isotopic ratios in the upper panels of these figures don't show the systematic trend predicted by Eq. 2. Instead, the ratios fluctuate from isotope to isotope by a factor of two.

To determine whether these fluctuations are consistent with the binding energy term that results from a difference between the temperatures T_1 and T_2 for the two reactions measured at incident energies of E_1 and E_2 , we compensated approximately for the temperature difference using Eq. 2. For k_{12} , we used the average value $\langle k_{app} \rangle$ where:

$$\left\langle k_{app} \right\rangle = \left\langle 1/T_{1,app} - 1/T_{2,app} \right\rangle \tag{6}$$

where $T_{1,app}$ and $T_{2,app}$ are the measured isotopic temperature for a specific isotopic thermometer plotted in Fig. 1. The average is taken over all of the isotopic thermometers. These corresponding mean values $\langle k_{app} \rangle$ given in Table 2 and used as labels for the lower panels of Fig. 3 where the adjusted isotope ratios, $R_{21}(N,Z)\exp(\langle k_{app} \rangle \cdot BE(N,Z))$ are shown as the open and closed points. The degree to

which this procedure removes the fluctuations in isotopic ratios in Fig. 3 is remarkable. Alternatively, one can extract the k values by fitting the R_{21} data in the top panels of Fig. 3 with Eq. (3). These best fit values, given in the column in Table 2 labeled $k_{\rm fit}$, are statistically consistent with the mean values of $\langle k_{app} \rangle$. The values for α and β that describe the dependence in Eqs. 3 and 4 upon neutron and proton number are also given in the table.

When one performs the same procedure for Ar+Sc collisions, mean values of $\langle k_{app} \rangle = -0.039 \, \mathrm{and} - 0.028$ are obtained for the pairs of incident energies involved in the left and right panels of Fig. 4. The consistency between $\langle k_{app} \rangle$ and the corresponding best fit values for k_{21} is not as good as that obtained for the Kr+Nb system, see Table II. Furthermore, if the values for $\langle k_{app} \rangle$ are used to adjust the isotope ratios as shown in the lower panels of Fig. 4, the adjusted ratios do not follow parallel lines on a semilog plot. The same non-scaling results are obtained if the best fit values for k_{21} are used. Thus, the generalized isoscaling relationship is not as well satisfied by the Ar+Sc data as for the Kr+Nb data.

Dynamical stochastic mean field calculations suggest that the yields of excited fragments produced by dynamical models are not as consistent with isoscaling relations as the final yields after secondary decay [38]. Thus, the consistency of the Kr+Nb data with the generalized scaling relationships could be due to a higher degree of equilibration in the heavier system or to a greater abundance of heavier fragments that sequentially decay to the observed ones. In any case it is clear that measurements of other systems would be useful to establish the validity of generalized isoscaling more clearly. The generalized scaling procedure described here can be used to investigate if the equilibrium conditions are established in the measurements of the caloric curves.

In summary, evidence for the validity of generalized isoscaling relations that allow one to relate systems of different isotopic composition and at different excitation energies. The accuracy of these generalized scaling relationships is much greater for

the heavy system than for the light system, indicating that such equilibrium concepts may provide a more reasonable approximation to the final state for the collisions of heavy systems collisions.

Reference:

- [1] S.C. Jeong, N. Herrmann, J. Randrup, J.P. Alard, Z. Basrak, N. Bastid, I.M. Belaev, M. Bini, T. Blaich, A. Buta, R. Caplar, C. Cerruti, N. Cindro, J.P. Coffin, R. Dona, P. Dupieux, J. Ero, Z.G. Fan, P. Fintz, Z. Fodor, R. Freifelder, L. Fraysse, S. Frolov, A. Gobbi, Y. Grigorian, G. Guillaume, K.D. Hildenbrand, S. Holbling, A. Houari, M. Jorio, F. Jundt, J. Kecskemeti, P. Koncz, Y. Korchagin, R. Kotte, M. Kramer, C. Kuhn, I. Legrand, A. Lebedev, C., Phys. Rev. Lett. 72, 3468 (1994).
- [2] C. Williams, W.G. Lynch, C. Schwarz, M.B. Tsang, M.J. Huang, H.C. Hsi, D.R. Bowman, C.K. Gelbke, M.A. Lisa, G.F. Peaslee, L. Phair, D.O. Handzy, J. Dinius, M-C. Lemaire, S.R. Souza, G. Van Buren, R.J. Charity, L.G. Sobotka, G.J. Kunde, U. Lynen, J. Pochodzalla, H. Sann, W. Trautmann, D.Fox, R.T. de Souza, G. Peilert, A. Botvina, and N. Carlin; Phys. Rev. C <u>55</u>, (1997) R2132.
- [3] D.R. Bowman, G.F. Peaslee, N. Carlin, R.T. de Souza, C.K. Gelbke, W.G. Gong, Y.D. Kim, M.A. Lisa, W.C. Lynch, L. Phair, M.B. Tsang, C. Williams, N. Colonna, K. Hanold, M.A. McMahan, G.J. Wozniak, and L.G. Moretto, Phys. Rev. Lett. 70, 3534 (1993).
- [4] E. Bauge, A. Elmaani, R.A. Lacey, J. Lauret, N.N. Ajitanand, D. Craig, M. Cronqvist, E. Gualtieril, S. Hannuschke, T. Li, B. Llope, T.Reposeur, A. Vander Molen, G.D. Westfall, J.S. Winfield, J. Yee, S. Yennello, A. Nadasen, R.S. Tickle, and E. Norbeck, Phys. Rev. Lett. 70, 3705 (1993).
- [5] S. Das Gupta, A.Z. Mekjian and M.B. Tsang, Adv. Nucl. Phys. <u>26</u>, 91 (2001) and references therein.
- [6] M. D'Agostino, A.S. Botvina, P.M. Milazzo, M. Bruno, G.J. Kunde, D.R. Bowman, L. Celano, N. Colonna, J.D. Dinius, A. Ferrero, M.L. Fiandri, C.K. Gelbke, T. Glasmacher, F. Gramegna, D.O. Handzy, D. Horn, W.C. Hsi, M. Huang, I. Iori, M.A. Lisa, W.G. Lynch, L. Manduci, G.V. Margagliotti, P.F. Mastinu, I.N. Mishustin, C.P. Montoya, A. Moroni, G.F. Peaslee, F. Petruzzelli, L. Phair, R. Rui, C. Schwarz, M.B. Tsang, G. Vannini, and C. Williams, Phys. Lett. B 371, (1996) 175.
- [7] A.S. Botvina, I.N. Mishustin, M. Begemann-Blaich, J. Hubele, G. Imme, I. Iori, P. Kreutz, G.J. Kunde, W.D. Kunze, V. Lindenstruth, U. Lynen, A. Moroni, W.F.J. Muller, C.A. Ogilvie, J. Pochodzalla, G. Raciti, Th. Rubehn, H. Sann, A. Schuttauf, W. Seidel, W. Trautmann, and A. Worner, Nucl. Phys. A584, 737 (1995).
- [8] W.P. Tan, B-A. Li, R. Donangelo, C.K. Gelbke, M-J. van Goethem, X.D. Liu, W.G. Lynch, S. Souza, M.B. Tsang, G. Verde, A. Wagner and N.S. Xu, Phys. Rev. C 64, 051901 (2001) and the references therein.
- [9] J.P. Bondorf, A.S. Botvina, A.S. Iljinov, I.N, Mishustin, and K. Sneppen, Phys. Rep. 257, 133 (1995) and refs. therein.
- [10] D.H.E. Gross, Phys. Rep. 279, 119 (1997).
- [11] J. Pochodzalla, T. Mohlenkamp, T. Rubehn, A. Schuttauf, A. Worner, E. Zude, M. Begemann-Blaich, Th. Blaich, H. Emling, A. Ferrero, C. Gross, G. Imme, I. Iori, G. J. Kunde, W. D. Kunze, V. Lindenstruth, U. Lynen, A. Moroni, W. F. J. Muller, B. Ocker, G. Raciti, H. Sann, C.

- Schwarz, W. Seidel, V. Serfling, J. Stroth, W. Trautmann, A. Trzcinski, A. Tucholski, G. Verde, B. Zwieglinski, Phys. Rev. Lett. 75,1040 (1995).
- [12] J. A. Hauger, B. K. Srivastava, S. Albergo, F. Bieser, F. P. Brady, Z. Caccia, D. A. Cebra, A. D. Chacon, J. L. Chance, Y. Choi, S. Costa, J. B. Elliott, M. L. Gilkes, A. S. Hirsch, E. L. Hjort, A. Insolia, M. Justice, D. Keane, J. C. Kintner, V. Lindenstruth, M. A. Lisa, H. S. Matis, M. McMahan, C. McParland, W. F. J. Müller, D. L. Olson, M. D. Partlan, N. T. Porile, R. Potenza, G. Rai, J. Rasmussen, H. G. Ritter, J. Romanski, J. L. Romero, G. V. Russo, H. Sann, R. P. Scharenberg, A. Scott, Y. Shao, T. J. M. Symons, M. Tincknell, C. Tuvé, S. Wang, P. Warren, H. H. Wieman, T. Wienold, and K. Wolf, Phys. Rev. C 62, 024616 (2000)
- [13] V. E. Viola and K. Kwiatkowski, and W. A. Friedman, Phys. Rev. C 59, 2660 (1999).
- [14] J.B. Natowitz, R. Wada, K. Hagel, T. Keutgen, M. Murray, A. Makeev, L. Qin, P. Smith, C. Hamilton, Phys. Rev. C 65, 034618 (2002).
- [15] A. Ruangma, R. Laforest, E. Martin, E. Ramakrishnan, D. J. Rowland, M. Veselsky, E. M. Winchester, S.J. Yennello, L. Beaulieu, W.-c. Hsi, K. Kwiatkowski, T. Lefort, V.E. Viola, A. Botvina, R. G. Korteling, L. Pienkowski, H. Breuer, S. Gushue, L. P. Remsberg, nucl-ex/0110004
- [16] H.F. Xi, G.J. Kunde, O. Bjarki, C.K. Gelbke, RC. Lemmon, W.G. Lynch, D. Magestro, R. Popescu, R. Shomin, M.B. Tsang, A.M. Vandermolen, G.D. Westfall, G. Imme, V. Maddalena, C. Nociforo, G. Raciti, G. Riccobene, F.P. Romano, A. Saija, C. Sfienti, S. Fritz, C. Groß, T. Odeh, C. Schwarz, A. Nadasen, D. Sisan, K.A.G. Rao, Phys. Rev. C 58, (1998) R2636.
- [17] V. Serfling, C. Schwarz, R. Bassini, M. Begemann-Blaich, S. Fritz, S.J. Gaff, C. Groß, G. Immé, I. Iori, U. Kleinevoß, G.J. Kunde, W.D. Kunze, U. Lynen, V. Maddalena, M. Mahi, T. Möhlenkamp, A. Moroni, W.F.J. Müller, C. Nociforo, B. Ocker, T. Odeh, F. Petruzzelli, J. Pochodzalla, G. Raciti, G. Riccobene, F.P. Romano, A. Saija, M. Schnittker, A. Schüttauf, W. Seidel, C. Sfienti, W. Trautmann, A. Trzcinski, G. Verde, A. Wörner, Hongfei Xi, and B. Zwieglinski, Phys. Rev. Lett. 80, 3928 (1998).
- [18] M.D'Agostino et al., Phys. Lett. B473 (2000) 219. negative heat capacity; reference on negative heat capacity in clusters.
- [10] W. Reisdorf, D. Best, A. Gobbi, N. Herrmann, K.D. Hildenbrand, B. Hong, S.C. Jeong, Y. Leifels, C. Pinkenburg, J.L. Ritman, D. Schull, U. Sodan, K. Teh, G.S. Wang, J.P. Wessels, T. Wienold, J.P. Alard, V. Amouroux, Z. Basrak, N. Bastid, I. Belyaev, L. Berger, J. Biegansky, M. Bini, S. Boussange, A. Buta, R. Caplar, N. Cindro, J.P. Coffin, P. Crochet, R. Dona, P. Dupieux, M. Dzelalija, J. Ero, M. Eskef, P. Fintz, Z. Fodor, L. Fraysse, A. Genoux-Lubain, G. Goebels, G. Guillaume, Y. Grigorian, E. Hafele, S. Holbling, A. Houari, M. Ibnouzahir, M. Joriot, F. Jundt, J. Kecskemeti, M. Kirejczyk, P. Koncz, Y. Korchagin, M. Korolija, R. Kotte, C. Kuhn, D. Lambrecht, A. Lebedev, A. Lebedev, I. Legrand, C. Maazouzi, V. Manko, T. Matulewicz, P.R. Maurenzig, H. Merlitz, G. Mgebrishvili, J. Mosner, S. Mohren, D. Moisa, G. Montarou, I. Montbel, P. Morel, W. Neubert, A. Olmi, G. Pasquali, D. Pelte, M. Petrovici, G. Poggi, P. Pras, F. Rami, V. Ramillien, C. Roy, A. Sadchikov, Z. Seres, B. Sikora, V. Simion, K. Siwek-Wilczynska, V. Smolyankin, N. Taccetti, R. Tezkratt, L. Tizniti, M. Trzaska, M.A. Vasiliev, P. Wagner, K. Wisniewski, D. Wohlfarth, and A. Zhilin, Nucl. Phys. A612, 493 (1997).
- [20] M.B. Tsang, W.G. Lynch, H. Xi and W.A. Friedman, Phys. Rev. Lett. <u>78</u>, 3836 (1997)
- [21] H. Xi, M.B. Tsang, M.J. Huang, W.G. Lynch, J.D. Dinius, S.G. Gaff, C.K. Gelbke, T. Glasmacher, G.J. Kunde, L. Martin, C.P. Montoya, M. Azzano, G.V. Margagliotti, P.M. Milazzo, R. Rui, G. Vannini, L. Celano, N. Colonna, G. Tagliente, M. D'Agnosino, M. Bruno,

- M.L. Fiandri, A. Ferrero, I. Iori, A. Moroni, F. Petruzzelli and P.F. Mastinu, Phys. Lett. <u>B431</u>, 8 (1998).
- [22] M.J. Huang, H. Xi, W.G. Lynch, M.B. Tsang, J.D. Dinius, S.J. Gaff, C.K. Gelbke, T.Glasmacher, G.J. Kunde, L. Martin, C.P. Montoya, E. Scannapieco, P.M. Milazzo, M. Azzano, G.V. Margagliotti, R. Rui, G. Vannini, N. Colonna, L. Celano, G. Tagliente, M. D'Agostino, M. Bruno, M.L. Fiandri, F. Gramegna, A. Ferrero, I. Iori, A. Moroni, F. Petruzzelli, P.F. Mastinu, Phys. Rev. Lett. 78 (1997) 1648.
- [23] S. Albergo, S. Costa, E. Costanzo, A. Rubbino, Nuovo Cimento, A89, 1 (1985).
- [24] K. Kwiatkowski, A.S. Botvina, D.S. Bracken, E. Renshaw Foxford, W.A. Friedman, R.G. Korteling, K.B. Morley, E.C. Pollacco, V.E. Viola, and C. Volant Phys. Lett. B423, 21 (1998).
- [25] H. Xi, W. G. Lynch, M. B. Tsang, W. A. Friedman and D. Durand, Phys. Rev <u>C59</u>, 1567 1999).
- [26] A. Kolomiets, E. Ramakrishnan, H. Johnston, F. Gimeno-Nogues, B. Hurst, D. O'Kelly, D. J. Rowland, S. Shlomo, T. White, J. Winger*, and S. J. Yennello., Phys. Rev. C54, R472 (1996).
- [27] A. Schuttauf, W.D. Kunze, A. Worner, M. Begemann-Blaich, Th. Blaich, D.R. Bowman, R.J. Charity, A. Cosmo, A. Ferrero, C.K. Gelbke, C. Gross, W.C. Hsi, J. Hubele, G. Imme, I. Iori, J. Kempter, P. Kreutz, G.J. Kunde, V. Lindenstruth, M.A. Lisa, W.G. Lynch, U. Lynen, M. Mang, T. Mohlenkamp, A. Moroni, W.F.J. Muller, M. Neumann, B. Ocker, C.A. Ogilvie, G.F. Peaslee, J. Pochodzalla, G. Raciti, F. Rosenberger, Th. Rubehn, H. Sann, C. Schwarz, W. Seidel, V. Serfling, L.G. Sobotka, J. Stroth, L. Stuttge, S. Tomasevic, W. Trautmann, A. Trzcinski, M.B. Tsang, A. Tucholski, G. Verde, C.W. Williams, E. Zude, and B. Zwieglinski, Nucl. Phys. A607, 457 (1996).
- [28] H.S. Xu, M.B. Tsang, T.X. Liu, X.D. Liu, W.G. Lynch, W.P.Tan and G. Verde, L. Beaulieu, B. Davin, Y. Larochelle, T. Lefort, R.T. de Souza, R. Yanez, V.E. Viola, R.J. Charity, and L.G. Sobotka, Phys. Rev. Lett. 85, 716 (2000).
- [29] M.B. Tsang, W.A. Friedman, C.K. Gelbke, W.G. Lynch, G. Verde, H. Xu, Phys. Rev. Lett. 86, 5023 (2001).
- [30] M.B. Tsang, C.K. Gelbke, X.D. Liu, W.G. Lynch, W.P. Tan, G. Verde, H.S. Xu, W. A. Friedman, R. Donangelo, S. R. Souza, C.B. Das, S. Das Gupta, and D. Zhabinsky, Phys. Rev. C 64, 054615 (2001)
- [31] C. Schwarz, W.G. Gong, N. Carlin, C.K. Gelbke, Y.D. Kim, W.G. Lynch, T. Murakami, G. Poggi, R.T. de Souza, M.B. Tsang, and H.M. Xu, D.E. Fields, K. Kwiatkowski, V.E. Viola, Jr., and S.J. Yennello, Phys. Rev. <u>C48</u>, 676 (1993).
- [32] Z. Chen and C.K. Gelbke, Phys. Rev. <u>C38</u>, 2630 (1988).
- [33] T.K. Nayak, T. Murakami, W.G. Lynch, K. Swartz, D.J. Fields, C.K. Gelbke, Y.D. Kim, J. Pochodzalla, M.B. Tsang, H.M. Xu, F. Zhu, and K. Kwiatkowski, Phys. Rev. <u>C45</u>, 132 (1992).
- [34] F. Zhu, W.G. Lynch, D.R. Bowman, R.T. de Souza, C.K. Gelbke, Y.D. Kim, L. Phair, M.B. Tsang, C. Williams, H.M. Xu, and J. Dinius; Phys. Lett. <u>B282</u> (1992) 299.
- [35] F. Ajzenberg-Selove, Nucl. Phys. <u>A392</u> (1983) 1; <u>A413</u> (1984) 1; <u>A433</u> (1985) 1; <u>A449</u> (1985) 1; <u>A460</u> (1986) 1.
- [36] C. Cavata et al, Phys. Rev. C 42, 1760 (1990); Y.D. Kim et al, Phys. Rev. C 45, 338 (1992).
- [37] L. Phair, D.R. Bowman, C.K. Gelbke, W.G. Gong, Y.D. Kim, M.A. Lisa, W.G. Lynch, G.F. Peaslee, R.T. de Souza, M.B. Tsang, and F. Zhu, Nucl. Phys. A548, 489 (1992).
- [38] T.X. Liu, in preparation.

Figures Captions:

Fig. 1: Apparent temperatures extracted from the four isotope ratio thermometers listed in Table 1 as a function of incident energy for the Kr+Nb system (left panel) and Ar+Sc system (right panel). The lines are drawn to guide the eye.

Fig. 2. Corrected temperatures, T_o, for the Kr+Nb system (left panel) and Ar+Sc system (right panel).

Fig. 3. Top panel: Relative isotope yield ratios for Z=1 (open circles), Z=2 (closed circles), Z=3 (open squares), Z=4 (closed squares), for the Kr+Nb reactions. The ratios are obtained using the isotope yields from two different incident energies. The energies involved are labeled in the top panels as E_2/E_1 . See text for details. Bottom Panels: Relative isotope yield ratios corrected for temperature differences using Eq. 3 are shown as the open and closed points. The lines are drawn to guide the eye.

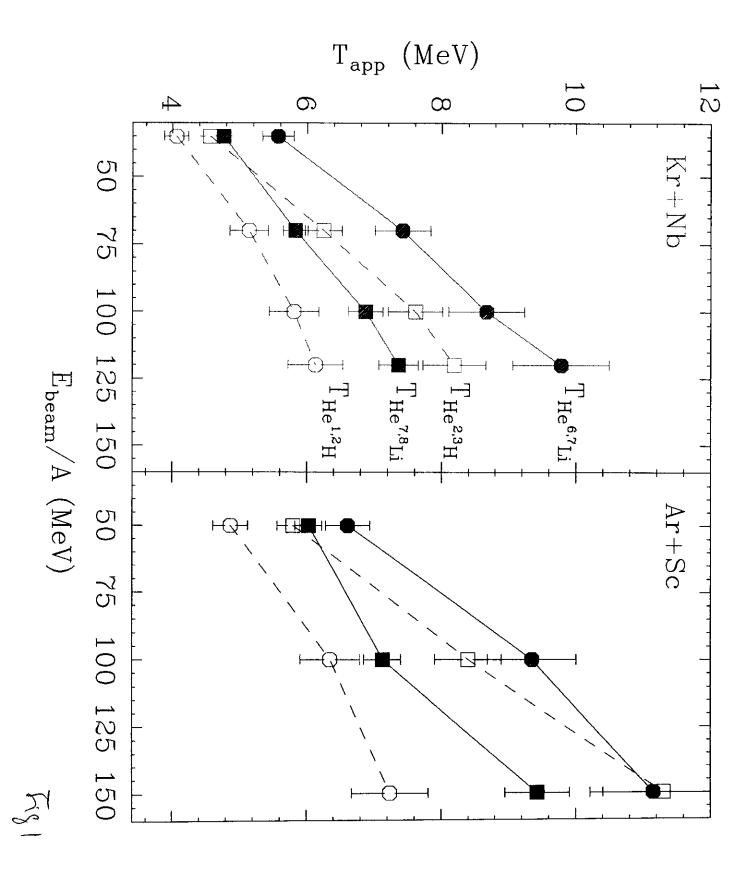
Fig. 4: Same as Fig. 3 for the Ar+Sc reactions. The lines are drawn to guide the eye.

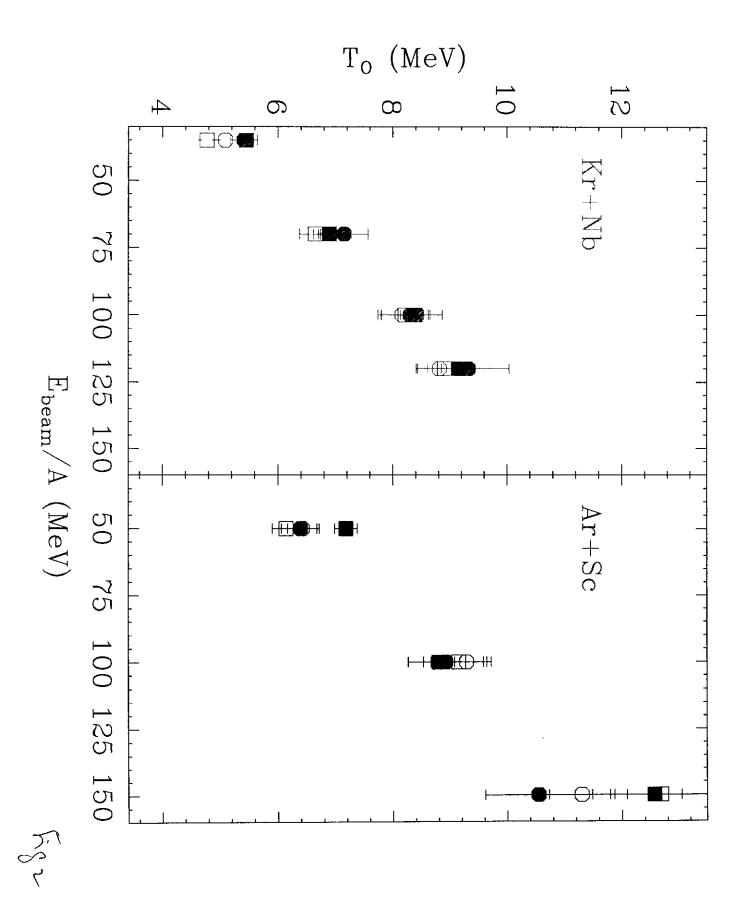
Table I: Four thermometers and their relevant parameters used in this article

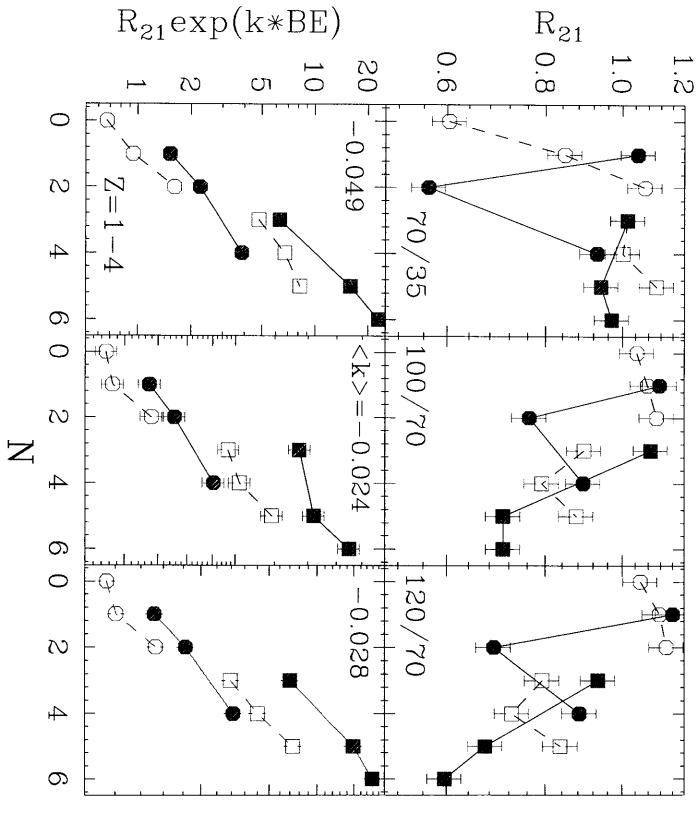
Isotope Ratio	a	В	(ln x/B)
	l	(MeV)	(MeV ⁻¹)
(^{6,7} Li, ^{3,4} He)	2.18	13.32	-0.0051
^{2,3} H, ^{3,4} He)	1.59	14.29	0.0097
$(^{1,2}H, ^{3,4}He)$	5.60	18.4	0.0496
(^{7,8} Li, ^{3,4} He)	1.98	18.54	0.0265

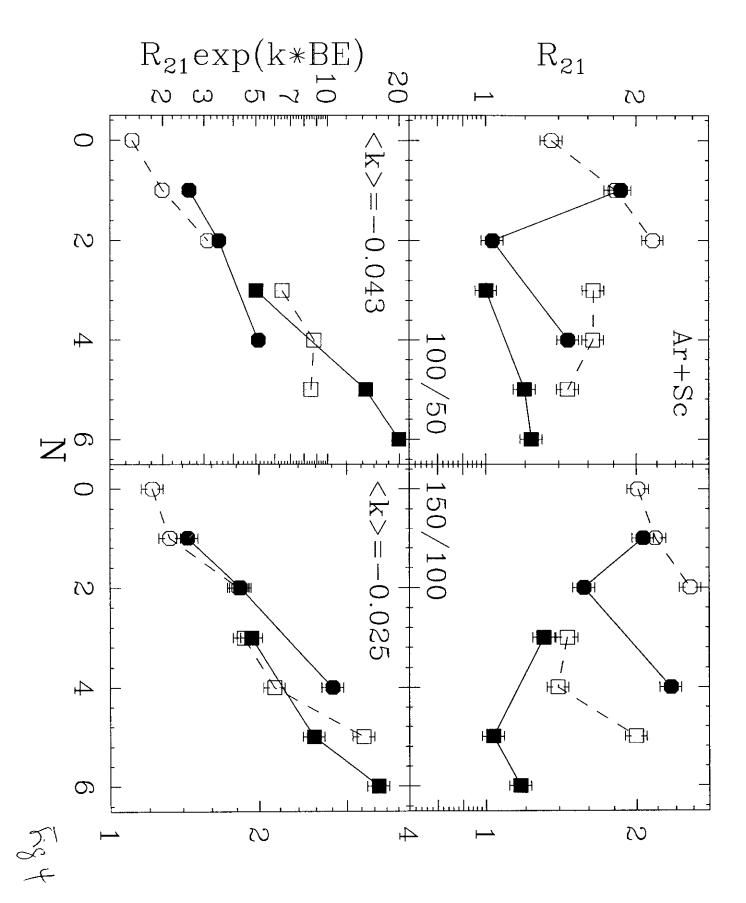
Table II. Generalized scaling parameters. The third and fourth columns list the values for $\langle k_{app} \rangle$ obtained from averaging the data in Fig. 1 and from the best fit, respectively. The values for α and β are weighted average of the scaling parameters from fits using $\langle k_{app} \rangle$ and $k_{\rm fit}$ to describe the temperature dependence.

Collision	$E_2(MeV)/E_1(MeV)$	$\langle k_{app} \rangle$	k _{fit}	α	β
Kr+Nb	70/35	-0.047±0.005	-0.040±0.005	0.3489	0.4034
Kr+Nb	100/70	-0.025±0.005	-0.024±0.004	0.0885	0.2073
Kr+Nb	120/70	-0.035±0.005	-0.028±0.003	0.1561	0.2340
Ar+Sc	100/50	-0.039±0.005	-0.028±0.004	0.2962	0.1461
Ar+Sc	150/100	-0.028±0.005	-0.025±0.003	0.2303	-0.051









		•