

CERN-EP-2024-324
04 December 2024

First detection of a tagged neutrino in the NA62 experiment

The NA62 Collaboration ¹

Abstract

The NA62 experiment at the CERN SPS reports the first detection of a tagged neutrino candidate based on the data collected in 2022. The candidate consists of a $K^+ \rightarrow \mu^+ \nu_\mu$ decay where the charged particles are reconstructed and the neutrino is detected through a charged-current interaction in a liquid krypton calorimeter.

To be submitted to Physics Letters B

¹Corresponding authors: Bianca De Martino, Mathieu Perrin-Terrin
email : bianca.de.martino@cern.ch, mathieu.perrin-terrin@cern.ch

The NA62 Collaboration

Université Catholique de Louvain, Louvain-La-Neuve, Belgium

E. Cortina Gil¹, J. Jerhot¹, N. Lurkin

TRIUMF, Vancouver, British Columbia, Canada

T. Numao, B. Velghe, V. W. S. Wong

University of British Columbia, Vancouver, British Columbia, Canada

D. Bryman²

Charles University, Prague, Czech Republic

Z. Hives, T. Husek³, K. Kampf, M. Kolesar, M. Koval

Aix Marseille University, CNRS/IN2P3, CPPM, Marseille, France

B. De Martino*, M. Perrin-Terrin*

Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), Garching, Germany

B. Döbrich, S. Lezki, J. Schubert⁴

Institut für Physik and PRISMA Cluster of Excellence, Universität Mainz, Mainz, Germany

A. T. Akmete, R. Aliberti⁵, M. Ceoletta⁶, L. Di Lella, N. Doble, L. Peruzzo, S. Schuchmann, H. Wahl, R. Wanke

Dipartimento di Fisica e Scienze della Terra dell'Università e INFN, Sezione di Ferrara, Ferrara, Italy

P. Dalpiaz, R. Negrello, I. Neri, F. Petrucci, M. Soldani

INFN, Sezione di Ferrara, Ferrara, Italy

L. Bandiera, N. Canale, A. Cotta Ramusino, A. Gianoli, M. Romagnoni, A. Sytov

Dipartimento di Fisica e Astronomia dell'Università e INFN, Sezione di Firenze, Sesto Fiorentino, Italy

M. Lenti, P. Lo Chiatto, I. Panichi, G. Ruggiero

INFN, Sezione di Firenze, Sesto Fiorentino, Italy

A. Bizzeti⁷, F. Bucci


Laboratori Nazionali di Frascati, Frascati, Italy

A. Antonelli, V. Kozhuharov⁸, G. Lanfranchi, S. Martellotti, M. Moulson, T. Spadaro, G. Tinti







Dipartimento di Fisica "Ettore Pancini" e INFN, Sezione di Napoli, Napoli, Italy

F. Ambrosino, M. D'Errico, R. Fiorenza⁹, M. Francesconi, R. Giordano, P. Massarotti, M. Mirra, M. Napolitano, I. Rosa, G. Saracino

Dipartimento di Fisica e Geologia dell'Università e INFN, Sezione di Perugia, Perugia, Italy

G. Anzivino 

INFN, Sezione di Perugia, Perugia, Italy

P. Cenci , V. Duk , R. Lollini , P. Lubrano , M. Pepe , M. Piccini 

Dipartimento di Fisica dell'Università e INFN, Sezione di Pisa, Pisa, Italy

F. Costantini , M. Giorgi , S. Giudici , G. Lamanna , E. Lari , E. Pedreschi , J. Pinzino , M. Sozzi 

INFN, Sezione di Pisa, Pisa, Italy

R. Fantechi , F. Spinella 

Scuola Normale Superiore e INFN, Sezione di Pisa, Pisa, Italy

I. Mannelli 





Dipartimento di Fisica, Sapienza Università di Roma e INFN, Sezione di Roma I, Roma, Italy

M. Raggi 





INFN, Sezione di Roma I, Roma, Italy

A. Biagioni , P. Cretaro , O. Frezza , A. Lonardo , M. Turisini , P. Vicini 

INFN, Sezione di Roma Tor Vergata, Roma, Italy

R. Ammendola , V. Bonaiuto ¹⁰ , A. Fucci, A. Salamon , F. Sargeni ¹¹ 





Dipartimento di Fisica dell'Università e INFN, Sezione di Torino, Torino, Italy

R. Arcidiacono ¹² , B. Bloch-Devaux ^{3, 13} , E. Menichetti , E. Migliore 







INFN, Sezione di Torino, Torino, Italy

C. Biino ¹⁴ , A. Filippi , F. Marchetto , D. Soldi 



Institute of Nuclear Physics, Almaty, Kazakhstan

Y. Mukhamejanov , A. Mukhamejanova ¹⁵ , N. Saduyev , S. Sakhiyev 

Instituto de Física, Universidad Autónoma de San Luis Potosí, San Luis Potosí, Mexico

A. Briano Olvera , J. Engelfried , N. Estrada-Tristan ¹⁶ , R. Piandani , M. A. Reyes Santos ¹⁶ , K. A. Rodriguez Rivera 

Horia Hulubei National Institute for R&D in Physics and Nuclear Engineering, Bucharest-Magurele, Romania

P. Boboc , A. M. Bragadireanu, S. A. Ghinescu , O. E. Hutanu

Faculty of Mathematics, Physics and Informatics, Comenius University, Bratislava, Slovakia

T. Blazek , V. Cerny , T. Velas , R. Volpe ¹⁷ 

CERN, European Organization for Nuclear Research, Geneva, Switzerland

J. Bernhard¹⁸, L. Bician¹⁸, M. Boretto, F. Brizioli¹⁹, A. Ceccucci, M. Corvino, H. Danielsson, F. Duval, L. Federici, E. Gamberini, R. Guida, E. B. Holzer, B. Jenninger, Z. Kucerova, G. Lehmann Miotto, P. Lichard, K. Massri²⁰, E. Minucci²¹, M. Noy, V. Ryjov, J. Swallow²², M. Zamkovsky

Ecole Polytechnique Fédérale Lausanne, Lausanne, Switzerland

X. Chang, A. Kleimenova, R. Marchevski

School of Physics and Astronomy, University of Birmingham, Birmingham, United Kingdom

J. R. Fry, F. Gonnella, E. Goudzovski, J. Henshaw, C. Kenworthy, C. Lazzeroni, C. Parkinson, A. Romano, J. Sanders, A. Sergi²³, A. Shaikhiev²⁰, A. Tomczak

School of Physics, University of Bristol, Bristol, United Kingdom

H. Heath

School of Physics and Astronomy, University of Glasgow, Glasgow, United Kingdom

D. Britton, A. Norton, D. Protopopescu

Physics Department, University of Lancaster, Lancaster, United Kingdom

J. B. Dainton, L. Gatignon, R. W. L. Jones

Department of Physics and Astronomy, George Mason University, Fairfax, Virginia, USA

P. Cooper, D. Coward²⁴, P. Rubin

Authors affiliated with an international laboratory covered by a cooperation agreement with CERN

A. Baeva, D. Baigarashev²⁵, V. Bautin, D. Emelyanov, T. Enik, V. Falaleev¹⁷, V. Kekelidze, D. Kereibay, A. Korotkova, L. Litov⁸, D. Madigozhin, M. Misheva²⁶, N. Molokanova, I. Polenkevich, Yu. Potrebenikov, K. Salamatin, S. Shkarovskiy

Authors affiliated with an Institute formerly covered by a cooperation agreement with CERN

S. Fedotov, K. Gorshanov, E. Gushchin, S. Kholodenko²⁷, A. Khotyantsev, Y. Kudenko, V. Kurochka, V. Kurshetsov, A. Mefodev, V. Obraztsov, A. Okhotnikov, A. Sadovskiy, V. Sugonyaev, O. Yushchenko

* Corresponding authors: B. De Martino, M. Perrin-Terrin,
email: bianca.de.martino@cern.ch, mathieu.perrin-terrin@cern.ch

¹Present address: Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), D-85748 Garching, Germany

²Also at TRIUMF, Vancouver, British Columbia, V6T 2A3, Canada

³Also at School of Physics and Astronomy, University of Birmingham, Birmingham, B15 2TT, UK

⁴Also at Department of Physics, Technical University of Munich, München, D-80333, Germany

⁵Present address: Institut für Kernphysik and Helmholtz Institute Mainz, Universität Mainz, Mainz, D-55099, Germany

- ⁶Also at CERN, European Organization for Nuclear Research, CH-1211 Geneva 23, Switzerland
- ⁷Also at Dipartimento di Scienze Fisiche, Informatiche e Matematiche, Università di Modena e Reggio Emilia, I-41125 Modena, Italy
- ⁸Also at Faculty of Physics, University of Sofia, BG-1164 Sofia, Bulgaria
- ⁹Present address: Scuola Superiore Meridionale e INFN, Sezione di Napoli, I-80138 Napoli, Italy
- ¹⁰Also at Department of Industrial Engineering, University of Roma Tor Vergata, I-00173 Roma, Italy
- ¹¹Also at Department of Electronic Engineering, University of Roma Tor Vergata, I-00173 Roma, Italy
- ¹²Also at Università degli Studi del Piemonte Orientale, I-13100 Vercelli, Italy
- ¹³Present address: Université Catholique de Louvain, B-1348 Louvain-La-Neuve, Belgium
- ¹⁴Also at Gran Sasso Science Institute, I-67100 L'Aquila, Italy
- ¹⁵Also at al-Farabi Kazakh National University, 050040 Almaty, Kazakhstan
- ¹⁶Also at Universidad de Guanajuato, 36000 Guanajuato, Mexico
- ¹⁷Present address: INFN, Sezione di Perugia, I-06100 Perugia, Italy
- ¹⁸Present address: Charles University, 116 36 Prague 1, Czech Republic
- ¹⁹Also at INFN, Sezione di Perugia, I-06100 Perugia, Italy
- ²⁰Present address: Physics Department, University of Lancaster, Lancaster, LA1 4YB, UK
- ²¹Present address: Syracuse University, Syracuse, NY 13244, USA
- ²²Present address: Laboratori Nazionali di Frascati, I-00044 Frascati, Italy
- ²³Present address: Dipartimento di Fisica dell'Università e INFN, Sezione di Genova, I-16146 Genova, Italy
- ²⁴Also at SLAC National Accelerator Laboratory, Stanford University, Menlo Park, CA 94025, USA
- ²⁵Also at L. N. Gumilyov Eurasian National University, 010000 Nur-Sultan, Kazakhstan
- ²⁶Present address: Institute of Nuclear Research and Nuclear Energy of Bulgarian Academy of Science (INRNE-BAS), BG-1784 Sofia, Bulgaria
- ²⁷Present address: INFN, Sezione di Pisa, I-56100 Pisa, Italy

Introduction

The next generation of long-baseline neutrino experiments [1, 2] will collect datasets more than ten times larger than their predecessors. Controlling systematic uncertainties in neutrino oscillation studies in these experiments becomes crucial [3], and will be vital for the precision measurements required to probe the three-neutrino paradigm [4]. These experiments require precise knowledge of neutrino fluxes, cross-sections, interaction models, and energy estimators with well-controlled biases [2, 4–7].

A tagged-neutrino experiment [8–12] would provide improved systematic uncertainties and neutrino energy estimates. In the setup proposed recently [12], beam neutrinos from $\pi^\pm \rightarrow \mu^\pm \bar{\nu}_\mu$ and $K^\pm \rightarrow \mu^\pm \bar{\nu}_\mu$ decays are kinematically reconstructed from the incoming and outgoing charged particles, which are detected using instrumentation installed in the beamline. Hence, the neutrino fluxes from $\pi^\pm \rightarrow \mu^\pm \bar{\nu}_\mu$ and $K^\pm \rightarrow \mu^\pm \bar{\nu}_\mu$ decays, and from all other K^\pm decays, particularly $K^\pm \rightarrow \pi^0 e^\pm \bar{\nu}_e$, are precisely determined. Moreover, the detected neutrinos can be individually associated with their decay parents, thus obtaining a precise estimate of their energy from the decay kinematics. Compared to energy reconstruction algorithms based on the measured neutrino interaction final state, which typically yield 10–30% precision with significant biases [2, 13], the kinematic reconstruction provides sub-percent precision and negligible bias [12]. In short-baseline experiments, neutrino tagging would allow precise measurements of the total and differential interaction cross-sections. In long-baseline experiments, neutrino tagging would improve the study of neutrino oscillations due to the excellent energy resolution, reduced backgrounds from non-oscillated neutrinos, and improved flavour and interaction identification.

The development of the neutrino tagging technique has been hindered so far by the prohibitive particle rates in a neutrino beamline. These obstacles are being overcome by the development of neutrino beamlines operating with slow extraction [14, 15] and beam spectrometers capable of operating in a high rate environment [16].

The NA62 experiment at the CERN SPS allows the possibility of a feasibility study of the neutrino tagging technique. To measure the properties of the very rare $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ decay [17], NA62 uses a high-intensity K^+ beam which copiously produces neutrinos, most importantly via $K^+ \rightarrow \mu^+ \nu_\mu$ decays. The experiment employs a state-of-the-art beam spectrometer [18] and a liquid krypton calorimeter which can be used as an instrumented target to detect neutrino interactions. Associating neutrino interactions with the K^+ and the μ^+ from their parent decays would compellingly demonstrate the feasibility of the neutrino tagging technique.

In the following, we report the first tagged neutrino detection based on data collected by NA62 in 2022. Details about the detector are given in Section 1 and event selection is described in Section 2. Section 3 reports the background assessments and Section 4 the signal expectation. The properties of the tagged neutrino candidate are given in Section 5.

1 Beam, detector and data sample

The layout of the NA62 beamline and detector [19] is shown schematically in Figure 1. An unseparated secondary beam of π^+ (70%), protons (23%) and K^+ (6%) is created by directing 400 GeV protons extracted from the CERN SPS onto a beryllium target in spills of 4.8 s duration. The target position defines the origin of the NA62 reference system: the beam travels along the Z axis in the positive direction (downstream), the Y axis points vertically up, and the X axis is directed to form a right-handed coordinate system. The central beam momentum is 75 GeV/ c , with a momentum spread of 1% (rms).

Beam kaons are tagged with a time resolution of 70 ps by a differential Cherenkov counter (KTAG), which uses nitrogen gas at 1.75 bar pressure contained in a 5 m long vessel as radiator. Beam particle momentum, direction and time are measured with 0.15 GeV/ c , 16 μ rad and 100 ps

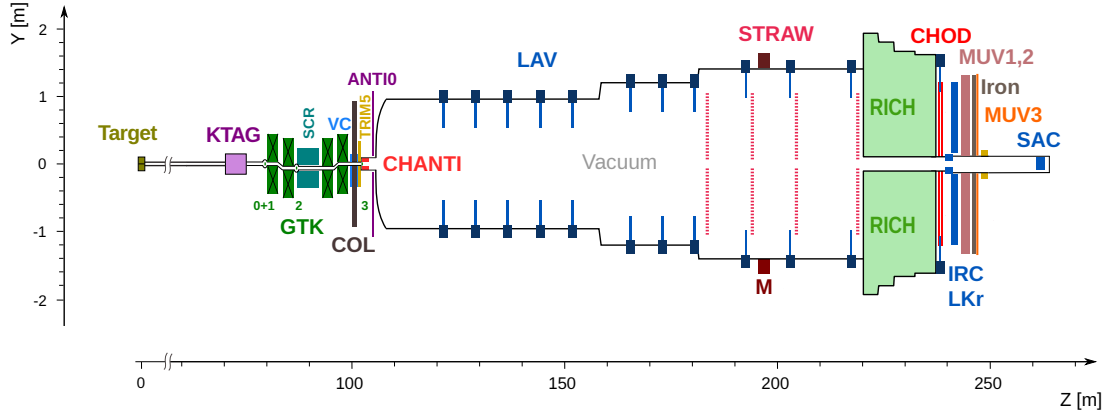


Figure 1: Schematic side view of the NA62 beamline and detector used in 2022. Information from VC and ANTI0 counters is not used in this analysis.

resolutions by a silicon pixel spectrometer consisting of four stations (GTK0,1,2,3) and four dipole magnets. A toroidal muon sweeper, called scraper (SCR), is installed downstream of GTK2. A 1.2 m thick steel collimator (COL) with a $76 \times 40 \text{ mm}^2$ central aperture and $1.7 \times 1.8 \text{ m}^2$ outer dimensions is placed upstream of GTK3 to absorb hadrons from upstream K^+ decays. Inelastic interactions of beam particles in GTK3 are detected by an array of scintillator bars (CHANTI). A dipole magnet (TRIM5) providing a $90 \text{ MeV}/c$ horizontal momentum kick is located in front of GTK3. The beam is delivered into a vacuum tank evacuated to a pressure of 10^{-6} mbar, which contains a 75 m long fiducial volume (FV) starting 2.6 m downstream of GTK3. The beam angular spread at the FV entrance is 0.11 mrad (rms) in both horizontal and vertical planes. Downstream of the FV, undecayed beam particles continue their path in vacuum.

Momenta of charged particles produced in K^+ decays are measured by a magnetic spectrometer (STRAW) located in the vacuum tank downstream of the FV. The spectrometer consists of four tracking chambers made of straw tubes, and a large aperture dipole magnet (M), located between the second and third chambers, that provides a horizontal momentum kick of $270 \text{ MeV}/c$ directed along $-X$ ($+X$) for positively (negatively) charged particles. The momentum resolution is $\sigma_p/p = (0.30 \oplus 0.005 \cdot p)\%$, with the momentum p expressed in GeV/c .

A ring-imaging Cherenkov detector (RICH) consisting of a 17.5 m long vessel filled with neon at atmospheric pressure (with a Cherenkov threshold of $12.5 \text{ GeV}/c$ for pions) provides particle identification, charged particle time measurements with a typical resolution of 70 ps and the trigger time. Two scintillator hodoscopes (CHOD), which include a matrix of tiles and two planes of slabs arranged in four quadrants located downstream of the RICH, provide trigger signals and time measurements. The tile matrix hodoscope has a time resolution of 1 ns, while the slab hodoscope measures time with 200 ps precision.

A 27 radiation length thick quasi-homogeneous liquid krypton (LKr) electromagnetic calorimeter is used as an active neutrino target in this analysis. The calorimeter has a thickness of 127 cm and an active volume of 7 m^3 segmented in the transverse direction into 13248 projective cells of $2 \times 2 \text{ cm}^2$ size, and provides an energy resolution $\sigma_E/E = (4.8/\sqrt{E} \oplus 11/E \oplus 0.9)\%$, with E expressed in GeV. To achieve hermetic acceptance for photons emitted in K^+ decays in the FV at angles up to 50 mrad from the beam axis, the LKr calorimeter is supplemented by annular lead glass detectors (LAV) installed in 12 positions inside and downstream of the vacuum tank, and two lead/scintillator sampling calorimeters (IRC, SAC) located close to the beam axis. An iron/scintillator-strip sampling hadronic calorimeter formed of two modules (MUV1,2) and a muon detector (MUV3), consisting of 148 scintillator tiles located behind an 80 cm thick iron wall, are used for particle identification. The eight smaller tiles of MUV3 adjacent to the beam

Table 1: L0 trigger conditions relevant to this study.

L0 Condition	Requirements
RICH	at least two signals in RICH
RICH16	signals in at least 16 groups of eight adjacent PMTs in the RICH
Q1	at least one signal in the tile CHOD
Q2	at least one signal in each of two different quadrants of the tile CHOD
UTMC	fewer than five signals in the tile CHOD
E5	at least 5 GeV deposited in the LKr
MOQX	at least one signal in each of two diagonally-opposite quadrants of MUV3

pipe are referred to as the *inner tiles*, while the remaining 140 regular tiles are called the *outer tiles*. The four quadrants of MUV3 include only the outer tiles.

The data sample is obtained from 3.2×10^5 SPS spills collected in 2022, with a typical beam intensity of 3.0×10^{12} protons per spill. This value corresponds on average to a 600 MHz instantaneous beam particle rate at the FV entrance, and a 4.5 MHz K^+ decay rate in the FV.

A two-level trigger system is used, comprising a hardware low level trigger (L0) and a software high level trigger (L1) [20]. The main trigger line is dedicated to the $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ study. Other trigger lines, based on combinations of individual detector conditions, operate concurrently. The L0 trigger conditions relevant to this study are described in Table 1.

The signal sample is collected using the Neutrino trigger line. The L0 conditions are $\text{RICH} \cdot \overline{\text{RICH16}} \cdot \text{UTMC} \cdot \overline{\text{Q2}} \cdot \text{E5} \cdot \text{MOQX}$. The L1 conditions require K^+ identification in KTAG, a positively charged track originating in the FV with momentum below 65 GeV/ c and no activity in LAV. This trigger line selects $K^+ \rightarrow \mu^+ \nu_\mu$ decays followed by a charged-current neutrino interaction in the LKr calorimeter which produces a hadronic shower and a μ^- .

A normalisation sample of $K^+ \rightarrow \mu^+ \nu_\mu$ events is used to determine the number of kaon decays and estimate the corresponding number of expected signal events. This sample is collected using the Minimum bias trigger line ($\text{RICH} \cdot \text{Q1}$) downscaled by a factor of 600; no conditions are applied at L1.

Monte Carlo (MC) simulations of particle interactions with the detector and its response are performed using a software package based on the GEANT4 [21] and GENIE [22–25] toolkits.

2 Event selection

The event selection consists of two sets of criteria. The *common selection* is applied to both signal and normalisation samples, and is designed to select $K^+ \rightarrow \mu^+ \nu_\mu$ decays with a single positively charged track originating in the FV and identified as a muon. The *interaction selection* is designed to select the signal sample comprising $K^+ \rightarrow \mu^+ \nu_\mu$ decays followed by a neutrino interaction in the LKr calorimeter. The signal region defined in the following is kept masked until the completion of the background evaluation.

Common selection

Trigger condition Events are required to satisfy all RICH and CHOD L0 conditions used in the Neutrino and Minimum bias trigger lines ($\text{RICH} \cdot \overline{\text{RICH16}} \cdot \text{Q1} \cdot \overline{\text{Q2}} \cdot \text{UTMC}$).

Positively charged muon A positively charged track must be reconstructed in the STRAW spectrometer with momentum below 65 GeV/ c . Its extrapolated positions at the CHOD, RICH, LKr, MUV1, MUV2 and MUV3 front planes should be within the respective geometrical acceptances. Spatial association of in-time signals in these detectors is required. The RICH time defines the track time. Muon identification is performed as follows:

- both *track-seeded ring* and *single ring* algorithms exploiting RICH signals and track properties [26] must favour the muon hypothesis;
- a MUV3 signal must be associated with the extrapolated track position.

Parent kaon A K^+ candidate is defined as a beam track reconstructed in GTK in time with a KTAG signal. The K^+ candidate with the KTAG time closest to the μ^+ track time defines the parent kaon. Inelastic kaon interactions in GTK3 are excluded by applying the following criteria: no CHANTI signal in time with the μ^+ track; no other K^+ candidate forming a vertex with the μ^+ track at a Z coordinate consistent with the GTK3 position (100 – 105 m).

Decay vertex The mid-point of the smallest segment between the μ^+ and K^+ tracks defines the decay vertex. The closest distance of approach of these tracks must be smaller than 4 mm. The Z coordinate of the vertex must be in the FV range 105 – 170 m.

Neutrino candidate The neutrino candidate is defined as the neutral particle originating from the decay vertex with 4-momentum equal to $P_{K^+} - P_{\mu^+}$, where P_{K^+} and P_{μ^+} are the K^+ and μ^+ 4-momenta under the assumption of K^+ and μ^+ masses, respectively. The squared missing mass, $m_{\text{miss}}^2 = (P_{K^+} - P_{\mu^+})^2$ must be consistent with zero, $|m_{\text{miss}}^2| < 0.006 \text{ GeV}^2/c^4$, as shown in Figure 2-left. The extrapolated position of the neutrino candidate at the LKr front plane must be within the LKr geometrical acceptance and at least 250 mm from the beam axis. The extrapolated position at the MUV3 front plane must be at least 300 mm from the beam axis. The distance between the extrapolated positions of the neutrino and the μ^+ candidates should be larger than 200 mm at the LKr front plane.

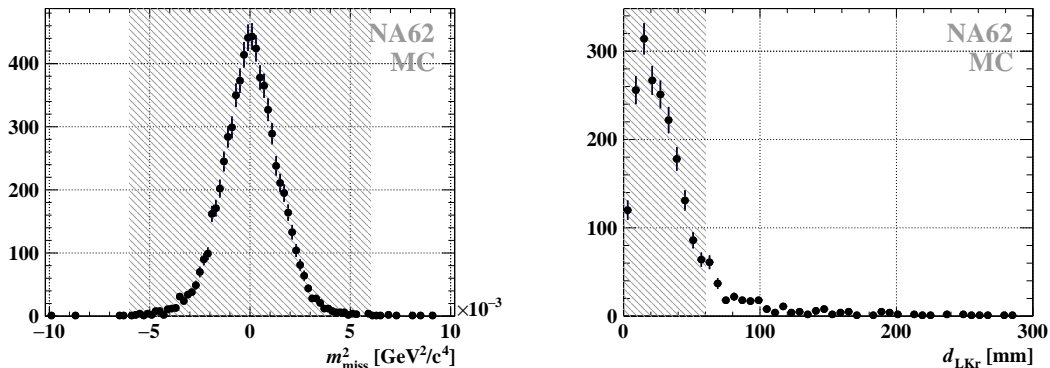


Figure 2: Left: distribution of the squared missing mass m_{miss}^2 for simulated $K^+ \rightarrow \mu^+ \nu_\mu$ decays. Right: distribution of the distance d_{LKr} between the extrapolated position of the neutrino candidate at LKr front plane and the associated interaction position. The shaded areas represent the signal region.

Extra activity In-time activity related to other beam tracks or K^+ decays is excluded by requiring:

- no more than three beam kaons based on KTAG signals detected within 8 ns of the RICH time;
- fewer than five beam tracks in GTK within 0.6 ns of the RICH time;
- no partially reconstructed STRAW tracks pointing to the decay vertex;
- no activity in LAV downstream of the vertex Z position within 3 ns of the RICH time;

- no more than one signal associated to the μ^+ in each CHOD hodoscope;
- no signals in LKr, MUV1, and MUV2 within 250 mm, 200 mm and 300 mm of the beam axis, respectively;
- no signals in SAC, IRC.

Interaction selection

Neutrino interaction Neutrino interaction signals should be reconstructed in LKr, MUV1, MUV2 and MUV3, matching in time signals from the μ^+ track from the K^+ decay. The distance, d_{LKr} , between the extrapolated position of the neutrino candidate at the LKr front plane and the associated LKr signal position, should be less than 60 mm, as shown in Figure 2-right. For kinematic reasons, neutrino interactions are not expected in the LKr inner region centered on the Z axis with a diameter 500 mm, extended on the $X > 0$ side to a half-ellipse of 500 mm (900 mm) minor (major) axes along $Y(X)$. Except for the signals associated with the neutrino candidate and the μ^+ , no in-time LKr signal should be reconstructed. The MUV1 and MUV2 signals should be spatially consistent with the neutrino candidate interaction in LKr. The MUV3 signals associated with the neutrino candidate and the μ^+ should be separated by at least 440 mm and located in diagonally opposite quadrants.

Interaction energy The LKr energy deposit associated with the neutrino candidate should be larger than 5 GeV, to match the E5 condition at L0, and between 5% and 50% of E_ν , the neutrino energy reconstructed kinematically from the $K^+ \rightarrow \mu^+ \nu_\mu$ decay. The total in-time energy in MUV1 and MUV2 should be larger than 8 GeV. The sum of the energy deposits associated with the interaction in LKr, MUV1 and MUV2 should be smaller than E_ν . As the hadronic shower is expected to be contained in LKr and MUV1, the energy deposit associated with the interaction in MUV2 is required to be lower than 10% of the sum of the associated energy deposits in LKr, MUV1 and MUV2.

The production angle of the μ^- increases as a function of the neutrino energy fraction transferred to the hadronic shower. The distance, d_{MUV3} , between the extrapolated position of the neutrino candidate at MUV3 front plane and the MUV3 signal position is expected to decrease linearly with the inelasticity $y = (E_\nu - E_{\text{LKr}} - E_{\text{MUV1}})/E_\nu$, where E_{LKr} and E_{MUV1} are the energy deposits associated with the neutrino candidate in LKr and MUV1. As a result, d_{MUV3} is required to satisfy the condition $880 \text{ mm} < d_{\text{MUV3}} + 880 \text{ mm} \times y < 1300 \text{ mm}$.

Interaction topology The neutrino and its interaction products, the hadronic shower and the μ^- , should be coplanar. Using a cylindrical coordinate system with the neutrino candidate direction defining the longitudinal axis, the azimuthal angle, φ , between the LKr signal associated with the neutrino candidate and the MUV3 signal associated with the μ^- must be compatible with π radians, and is required to satisfy $|\varphi - \pi| < 1.44 \text{ rad}$.

3 Background estimate

The background events originate from (A) $K^+ \rightarrow \mu^+ \nu_\mu$ decays with unrelated in-time activity mimicking a neutrino interaction, and (B) mis-reconstructed non- $K^+ \rightarrow \mu^+ \nu_\mu$ decays.

A data-driven method is used to estimate the contribution of background events by modelling their observed distribution in the regions $|m_{\text{miss}}^2| \geq 0.006 \text{ GeV}^2/c^4$ or $d_{\text{LKr}} \geq 60 \text{ mm}$ and integrating the model in the signal region.

Background A is studied using a sample satisfying the signal selection but inverting the d_{LKr} condition and removing the requirements on φ and on d_{MUV3} to increase the sample size. Assuming that the accidental activity is uniform over the LKr front plane, the number of

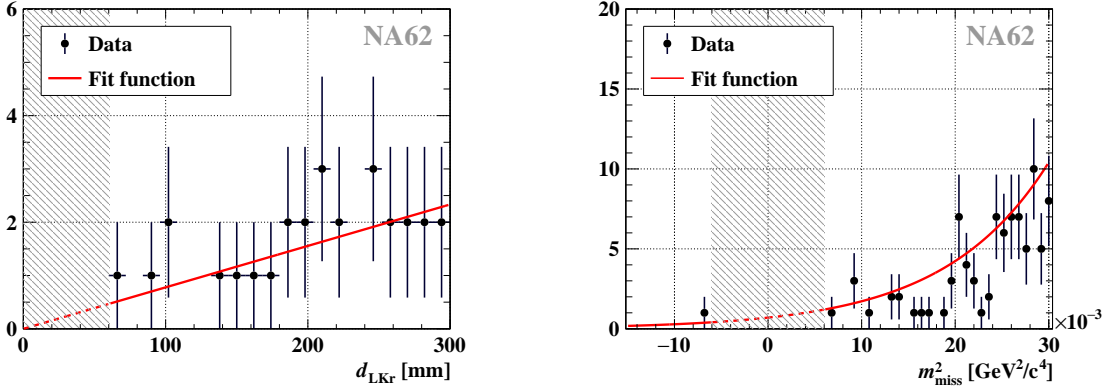


Figure 3: Left: d_{LKr} distribution of the selected events in the background A sample, fitted with a linear function. Right: m_{miss}^2 distribution of the selected events in the background B sample, fitted with an exponential function. The shaded areas represent the masked signal region.

background events per unit of d_{LKr} is proportional to d_{LKr} . The d_{LKr} distribution of the selected sample is fitted with a linear function from 60 mm to 300 mm as shown in Figure 3-left. The expected number of events in the signal region is obtained as the integral of the fit function in this region, scaled by the efficiency of the removed selection criteria, which is measured in the sidebands. The contribution is found to be $0.030^{+0.041}_{-0.023}|_{\text{stat}} \pm 0.004_{\text{syst}}$. The statistical uncertainties account for those of the fit parameter and the scale factor. The systematic uncertainty accounts for the choice of the fit range, and is computed as the difference between the scaled integrals using the fit ranges [60, 300] mm and [60, 180] mm.

Background B is similarly studied by inverting the m_{miss}^2 requirement and removing the interaction energy and topology conditions to increase the sample size. The m_{miss}^2 distribution of the selected sample is fitted from -0.015 to $0.03 \text{ GeV}^2/c^4$ with an exponential function, excluding the signal region $|m_{\text{miss}}^2| < 0.006 \text{ GeV}^2/c^4$, as shown in Figure 3-right. The integration of the fit function in the signal region, after scaling by the measured efficiency of the removed criteria, gives an estimated contribution of $0.0039^{+0.0060}_{-0.0034}|_{\text{stat}} \pm 0.0007_{\text{syst}}$. The statistical uncertainties account for those of the fit parameters and scale factor. The systematic uncertainty accounts for the choice of the fit function and is computed as the difference between the scaled integrals using exponential and quadratic functions.

The total number of expected background events in the signal region, dominated by background A, is $0.034^{+0.041}_{-0.023}|_{\text{stat}} \pm 0.004_{\text{syst}}$.

4 Expected signal

The expected number of signal events, N_{sig} , is obtained from the number of observed normalisation events, $N_{\text{norm}} = 2.22 \times 10^8$:

$$N_{\text{sig}} = N_{\text{norm}} \cdot D \cdot P^{\text{int}} \cdot \frac{\varepsilon_{\text{sig}}}{\varepsilon_{\text{norm}}}, \quad (1)$$

where $D = 600$ is the downscaling factor of the L0 Minimum bias trigger line used to collect the normalisation events, P^{int} is the average neutrino interaction probability for the selected signal sample, and ε_{sig} ($\varepsilon_{\text{norm}}$) is the selection and trigger efficiency of the signal (normalisation) sample.

Given the charged-current interaction cross section ($0.68 \times 10^{-38} \text{ cm}^{-2} \text{ GeV}^{-1}$ [27]), the mean neutrino energy (40 GeV) and the thickness of the LKr calorimeter (127 cm), the interaction probability, P^{int} , is expected to be a few 10^{-11} . A more precise value is obtained using simulations by GENIE including elaborated models of cross sections. The neutrino from each simulated

$K^+ \rightarrow \mu^+ \nu_\mu$ decay is transferred to GENIE which calculates the neutrino interaction probability in the LKr calorimeter, generates the interaction position and the final state particles, and returns this information to the NA62 simulation software. The average interaction probability for the selected signal events is found to be $P^{\text{int}} = (5.6 \pm 0.1_{\text{syst}}) \times 10^{-11}$, where the uncertainty is due to the precision of the neutrino cross-section measurements.

Removing the trigger components common to the signal and normalisation selections, the efficiency ratio becomes

$$\frac{\varepsilon_{\text{sig}}}{\varepsilon_{\text{norm}}} = \varepsilon^{\text{int}} \cdot \varepsilon^{\text{MOQX}} \cdot \varepsilon^{\text{E5}} \cdot \varepsilon^{\text{L1}}, \quad (2)$$

where ε^{int} , $\varepsilon^{\text{MOQX}}$, ε^{E5} , and ε^{L1} are the efficiencies of the interaction selection and trigger conditions applied only in the signal selection.

The interaction selection efficiency is $\varepsilon^{\text{int}} = 0.0367 \pm 0.0022_{\text{stat}} \pm 0.0013_{\text{syst}}$, as evaluated using simulations and correcting for the random loss induced by non-simulated extra activity in the detectors. This loss is estimated from data as the fraction of events rejected when requiring in-time activity in LKr and MUV3 in the normalisation sample.

The MOQX trigger efficiency is measured using a sample of $K^+ \rightarrow \pi^+ \pi^+ \pi^-$ decays where two pions of opposite charge decay in flight as $\pi^\pm \rightarrow \mu^\pm \bar{\nu}_\mu$. The sample is collected with a trigger line based only on RICH and CHOD signals and therefore independent of MUV3. Events are required to have two reconstructed tracks associated with MUV3 signals located in diagonally opposite outer quadrants. The MOQX efficiency is computed as a function of the MUV3 tile pair. The efficiency $\varepsilon^{\text{MOQX}} = 0.976 \pm 0.007_{\text{stat}}$ is obtained by weighting the tile-pair efficiencies by the tile-pair distribution obtained from signal simulation.

The E5 trigger efficiency is evaluated using a sample of $K^+ \rightarrow \pi^+ \pi^0$ decays where the π^+ produces a hadronic shower in LKr and the photons from the π^0 decay are detected in the LAV. The sample comprises events collected by trigger lines that do not rely on LKr signals. Events are required to have no activity in the LKr inner region, IRC and SAC and exactly one LKr signal in time with the π^+ track, ensuring a purely hadronic energy deposit in LKr. The E5 efficiency is measured as a function of the in-time energy deposit in the LKr. The efficiency $\varepsilon^{\text{E5}} = 0.847 \pm 0.008_{\text{stat}} \pm 0.003_{\text{syst}}$ is obtained by weighting the measurements in individual energy bins by the energy distribution of selected signal events obtained from simulations. The systematic uncertainty is obtained by varying by 10% the neutrino energy used in the GENIE simulation of the interaction.

The L1 efficiency is found to be $\varepsilon^{\text{L1}} = 0.918 \pm 0.002_{\text{stat}}$, using a downscaled dataset recorded without L1 conditions applied [20].

The resulting number of expected signal events is

$$N_{\text{sig}} = 0.208 \pm 0.013_{\text{stat}} \pm 0.009_{\text{syst}},$$

leading to a signal over background ratio of 6/1.

5 Results and prospects

After unmasking the signal region, one event is observed, the main properties of which are reported in Table 2. A display of the activity in the downstream detectors is shown in Figure 4. Given the total number of expected events ($0.242^{+0.043}_{-0.027} |_{\text{stat}} \pm 0.010_{\text{syst}}$), the probability of observing one event in the signal region is 19% while the probability of observing one background event candidate is 3%.

The observed event is the first tagged neutrino candidate and demonstrates the feasibility of the neutrino tagging technique. The obtained result is an important step toward dedicated tagged neutrino experiments, allowing the collection of large samples of tagged neutrinos, which would open new perspectives for accelerator-based neutrino physics. The data already collected by NA62 in 2023–2024 can be used to consolidate the result of the study presented here.

Table 2: Properties of the tagged neutrino candidate event. Its energy evaluated from kinematics has a relative uncertainty of 0.34% which is significantly better than the 10-30% typically achievable in neutrino detectors.

Property	Value
K^+ momentum	77.34 GeV/c
μ^+ momentum	25.26 GeV/c
Decay vertex position z_{vtx}	161.2 m
Neutrino energy $E_\nu = E_{K^+} - E_{\mu^+}$	52.09 GeV
Distance of the LKr signal to the Z axis	334.9 mm
LKr energy associated with the neutrino E_{LKr}	12.54 GeV
MUV1 energy associated with the neutrino E_{MUV1}	13.75 GeV
MUV2 energy associated with the neutrino E_{MUV2}	10.74 GeV
Inelasticity y	0.69
Azimuthal angle between the LKr and MUV3 signals φ	3.28 rad
d_{LKr}	31.4 mm
d_{MUV3}	567.9 mm
m_{miss}^2	-0.00086 GeV ² /c ⁴

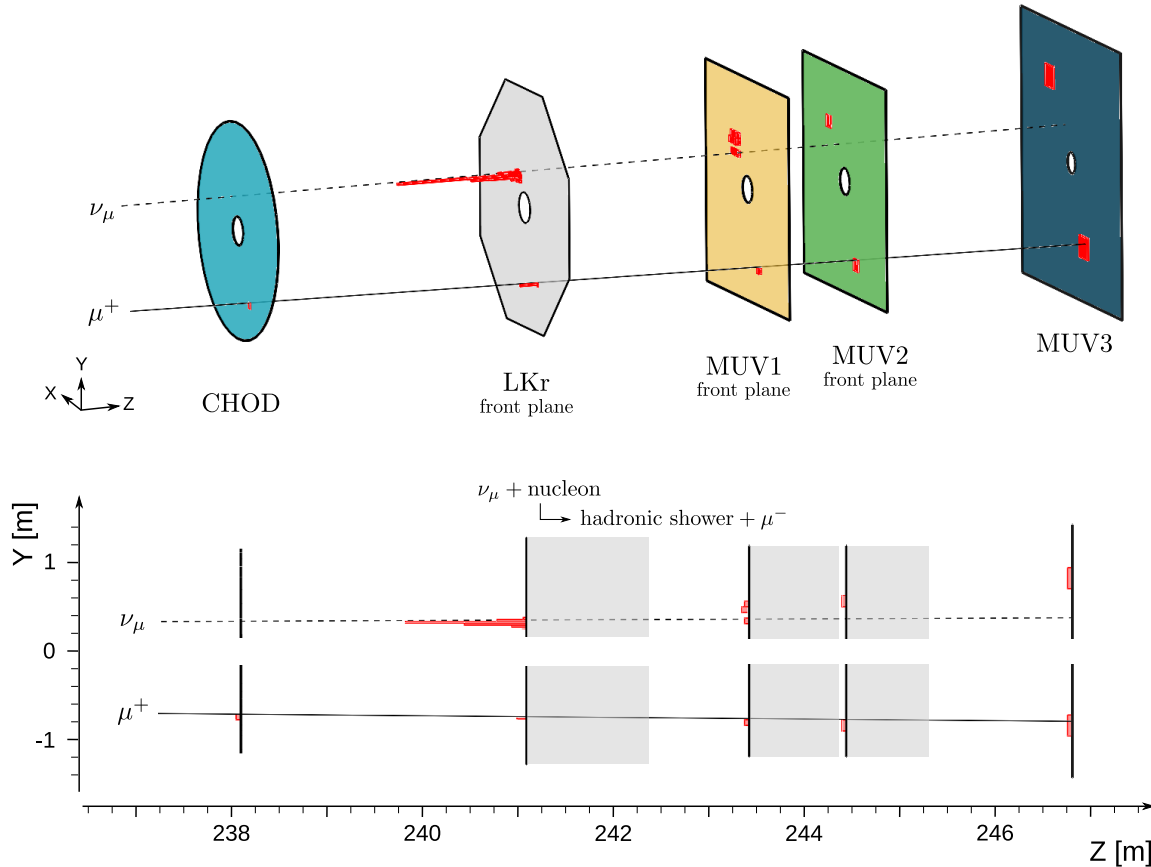


Figure 4: Display of the activity recorded in CHOD, LKr, MUV1, MUV2 and MUV3 for the signal candidate. No extra in-time activity is recorded in these detectors. The red histograms represent the activity in each detector projected at the front plane. For LKr, MUV1 and MUV2 the heights correspond to the fraction of the total energy deposited in each cell. The solid line represents the μ^+ trajectory extrapolated from the STRAW track; the dashed line represents the neutrino trajectory using the K^+ and μ^+ momenta. The shaded areas in the projection represent the detector volumes.

Acknowledgments

The authors are grateful to Marco Roda and Costas Andreopoulos from the GENIE Collaboration for their assistance in implementing the neutrino interaction simulation in the NA62 software. It is a pleasure to express our appreciation to the staff of the CERN laboratory and the technical staff of the participating laboratories and universities for their efforts in the operation of the experiment and data processing.

The cost of the experiment and its auxiliary systems was supported by the funding agencies of the Collaboration Institutes. We are particularly indebted to: F.R.S.-FNRS (Fonds de la Recherche Scientifique - FNRS), under Grants No. 4.4512.10, 1.B.258.20, Belgium; CECI (Consortium des Equipements de Calcul Intensif), funded by the Fonds de la Recherche Scientifique de Belgique (F.R.S.-FNRS) under Grant No. 2.5020.11 and by the Walloon Region, Belgium; NSERC (Natural Sciences and Engineering Research Council), funding SAPPJ-2018-0017, Canada; MEYS (Ministry of Education, Youth and Sports) funding LM 2018104, Czech Republic; BMBF (Bundesministerium für Bildung und Forschung) contracts 05H12UM5, 05H15UMCNA, 05H18UMCNA, 05H21UMCNA, and 05H24UMCNA, Germany; INFN (Istituto Nazionale di Fisica Nucleare), Italy; MIUR (Ministero dell’Istruzione, dell’Università e della Ricerca), Italy; CONACyT (Consejo Nacional de Ciencia y Tecnología), Mexico; IFA (Institute of Atomic Physics) Romanian CERN-RO Nr. 06/03.01.2022 and Nucleus Programme PN 19 06 01 04, Romania; MESRS (Ministry of Education, Science, Research and Sport), Slovakia; CERN (European Organization for Nuclear Research), Switzerland; STFC (Science and Technology Facilities Council), United Kingdom; NSF (National Science Foundation) Award Numbers 1506088 and 1806430, U.S.A.; ERC (European Research Council) “UniversaLepto” advanced grant 268062, “KaonLepton” starting grant 336581, Europe.

Individuals have received support from: Charles University (Research Center UNCE/SCI/013, grant PRIMUS 23/SCI/025), Czech Republic; Czech Science Foundation (grant 23-06770S); Agence Nationale de la Recherche (grant ANR-19-CE31-0009), France; Ministero dell’Istruzione, dell’Università e della Ricerca (MIUR “Futuro in ricerca 2012” grant RBFR12JF2Z, Project GAP), Italy; the Royal Society (grants UF100308, UF0758946), United Kingdom; STFC (Rutherford fellowships ST/J00412X/1, ST/M005798/1), United Kingdom; ERC (grants 268062, 336581 and starting grant 802836 “AxScale”); EU Horizon 2020 (Marie Skłodowska-Curie grants 701386, 754496, 842407, 893101, 101023808).

References

- [1] B. Abi, et al., Long-baseline neutrino oscillation physics potential of the DUNE experiment, *Eur. Phys. J. C* 80 (10) (2020) 978. [arXiv:2006.16043](https://arxiv.org/abs/2006.16043), [doi:10.1140/epjc/s10052-020-08456-z](https://doi.org/10.1140/epjc/s10052-020-08456-z).
- [2] K. Abe, et al., Hyper-Kamiokande Design Report, <https://doi.org/10.48550/arXiv.1805.04163> (5 2018). [arXiv:1805.04163](https://arxiv.org/abs/1805.04163), [doi:10.48550/arxiv.1805.04163](https://doi.org/10.48550/arxiv.1805.04163).
- [3] 2020 Update of the European Strategy for Particle Physics (Brochure), <http://cds.cern.ch/record/2721370> (2020). [doi:10.17181/CERN.JSC6.W89E](https://doi.org/10.17181/CERN.JSC6.W89E).
- [4] P. Huber, et al., Snowmass Neutrino Frontier Report, <https://doi.org/10.48550/arXiv.2211.08641> (Nov. 2022). [doi:10.48550/ARXIV.2211.08641](https://doi.org/10.48550/ARXIV.2211.08641).
- [5] P. Huber, et al., On the impact of systematical uncertainties for the CP violation measurement in superbeam experiments, *JHEP* 03 (2008) 021. [doi:10.1088/1126-6708/2008/03/021](https://doi.org/10.1088/1126-6708/2008/03/021).

- [6] A. Branca, G. Brunetti, A. Longhin, M. Martini, F. Pupilli, F. Terranova, A New Generation of Neutrino Cross Section Experiments: Challenges and Opportunities, *Symmetry* 13 (9) (2021) 1625. doi:10.3390/sym13091625.
- [7] A. Abed Abud, et al., Deep Underground Neutrino Experiment (DUNE) Near Detector Conceptual Design Report, *Instruments* 5 (4) (2021) 31. doi:10.3390/instruments5040031.
- [8] B. Pontecorvo, Tagging direct neutrinos: a first step to neutrino tagging, *Lett. Nuovo Cim.* 25 (1979) 257–259. doi:10.1007/BF02813638.
- [9] I. P. Nedyalkov, Single spectrometer station for neutrino tagging, <http://inis.jinr.ru/s1/NTBLIB/JINR-E1-84-515.pdf> (7 1984).
URL <http://inis.jinr.ru/s1/NTBLIB/JINR-E1-84-515.pdf>
- [10] G. Bohm, Project of a tagged neutrino facility at Serpukhov, in: *Physics at Future Accelerators. Proceedings, 10th Warsaw Symposium on Elementary Particle Physics, Kazimierz, Poland, 1987, 1987.*
URL <https://inis.iaea.org/search/21064823>
- [11] R. Bernstein, Neutrino physics at Fermilab in the 1990's, *Nuclear Physics B - Proceedings Supplements* 31 (1993) 255–261. doi:10.1016/0920-5632(93)90143-t.
- [12] M. Perrin-Terrin, Neutrino tagging: a new tool for accelerator based neutrino experiments, *Eur. Phys. J. C* 82 (5) (2022) 465. doi:10.1140/epjc/s10052-022-10397-8.
- [13] A. Friedland, S. W. Li, Understanding the energy resolution of liquid argon neutrino detectors, *Phys. Rev. D* 99 (3) (2019) 036009. arXiv:1811.06159, doi:10.1103/PhysRevD.99.036009.
- [14] F. Acerbi, et al., Design and performance of the ENUBET monitored neutrino beam, *The European Physical Journal C* 83 (10) (Oct. 2023). doi:10.1140/epjc/s10052-023-12116-3.
- [15] A. Baratto-Roldán, M. Perrin-Terrin, E. Parozzi, M. Jebramcik, N. Charitonidis, Nutag: a proof-of-concept study for a long-baseline neutrino beam, *Eur. Phys. J. C* 84 (10 2024). doi:10.1140/epjc/s10052-024-13324-1.
- [16] F. Borgato, et al., Charged-particle timing with 10 ps accuracy using TimeSPOT 3D trench-type silicon pixels, *Frontiers in Physics* 11 (2023).
URL <https://www.frontiersin.org/articles/10.3389/fphy.2023.1117575>
- [17] E. Cortina Gil, et al., Measurement of the very rare $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ decay, *JHEP* 06 (2021) 93. arXiv:2103.15389, doi:10.1007/JHEP06(2021)093.
- [18] G. Aglieri Rinella, et al., The NA62 GigaTracKer: a low mass high intensity beam 4D tracker with 65 ps time resolution on tracks, *JINST* 14 (2019) P07010. arXiv:1904.12837, doi:10.1088/1748-0221/14/07/P07010.
- [19] E. Cortina Gil, et al., The Beam and detector of the NA62 experiment at CERN, *JINST* 12 (05) (2017) P05025. arXiv:1703.08501, doi:10.1088/1748-0221/12/05/P05025.
- [20] E. Cortina Gil, et al., Performance of the NA62 trigger system, *JHEP* 03 (2023) 122. doi:10.1007/JHEP03(2023)122.
- [21] J. Allison, et al., Recent developments in Geant4, *Nucl. Instrum. Meth. A* 835 (2016) 186–225. doi:10.1016/j.nima.2016.06.125.

- [22] C. Andreopoulos, et al., The GENIE Neutrino Monte Carlo Generator, Nucl. Instrum. Meth. A 614 (2010) 87–104. doi:10.1016/j.nima.2009.12.009.
- [23] J. Tena-Vidal, et al., Neutrino-nucleon cross-section model tuning in GENIE v3, Phys. Rev. D 104 (7) (2021) 072009. doi:10.1103/PhysRevD.104.072009.
- [24] J. Tena-Vidal, et al., Neutrino-nucleus $CC0\pi$ cross-section tuning in GENIE v3, Phys. Rev. D 106 (11) (2022) 112001. doi:10.1103/PhysRevD.106.112001.
- [25] J. Tena-Vidal, et al., Hadronization model tuning in GENIE v3, Phys. Rev. D 105 (1) (2022) 012009. doi:10.1103/PhysRevD.105.012009.
- [26] E. Cortina Gil, et al., An investigation of the very rare $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ decay, JHEP 11 (2020) 042. arXiv:2007.08218, doi:10.1007/JHEP11(2020)042.
- [27] J. A. Formaggio, G. P. Zeller, From eV to EeV: Neutrino Cross Sections Across Energy Scales, Rev. Mod. Phys. 84 (2012) 1307–1341. doi:10.1103/RevModPhys.84.1307.