CERN/TC/NBC 65-4 17.6.1965 G. KELLNER

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Geometrical Reconstruction for 150 cm BNHBC

 \sim (results valid for experiments after Jan.1965 only) a service to the world of the

The aim of this note is to describe the necessary modifications in the geometry program for the reconstruction of events in the \sim British Hydrogen Bubble Chamber. Since lens distortions and tilt of the film plane with respect to the reference plane of the chamber can not be neglected, an additional treatment of IEP measurements has to be included.

All calculations were done with the program PYTHON, written by J. Zoll (see T.C.Program Library, Section : UTILITY-PYTHON). Copious reference is given to this manual to avoid a too long description.

The following description is grouped into several sections :

- A) Calculation of Chamber Constants
- B) Results
C) Applica
- C) Application in Geometry Program
D) Discussion
- Discussion
- E) Appendix I : Grid Plate
- F) Appendix II : Grid calibration
- G) Appendix III : Fiducial marks on chamber windows

It should be stressed that these results are only valid for the experiments after January 1965 since the camera plate has been and some in modified at this time.

A) Calculation of Chamber Constants

<u>ing Alexandria</u> Two methods have been used to determine these parameters : a) a grid plate, and b) actual chamber fiducials.

a) A grid plate (glass plate with lines engraved on the face towards the cameras, see Appendix I) has been photographed in 3 positions, corresponding approximately to hydrogen surface of front glass, centre of chamber and hydrogen surface of back glass. Only 8 horizontal and 23 vertical lines (184 intersection points) could be seen on all photographs.

Measurements were done such that points on a line were measured in between two intersecting lines. Any two noighbouring points were used to calculate the equation of the line.
Intersection of the two lines through $\begin{array}{c} x \to \text{each point give then the best possible} \end{array}$

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values for the coordinates of these intersection points on the film $(see$ PYTHON, INMESH).

Various combinations of two positions were used to calculate camera positions, film tilt and lens distortions.

The influence of film stretch was removed by measuring the 4 camera based fiducials (fiducials in the film gate) which are printed onto every photograph.

The transformation - from measured camera based fiducials on the film to fiducials in space - is found by a least-squares fit (see PYTHON, SHAPE).

To get the position of these marks with high accuracy, they were photographed separately with orthochromatic plates (Gevaert 23D56) and measured on a precision microscope.

After alignment of two grid positions in space (telescope readings, see Appendix II) an iteration cycle was entered for the calculation of. camera positions. Parameters which were fixed during the iteration cycle were : camera coordinates (the chamber was shifted relative to the cameras instead), inter-lens spacing of cameras, distance between two grid positions and distances between fiducial marks. After 10 - 20 iterations tho restriction of fixed inter-lens spacing was removed. Usually the results are slightly improved.

fit to the following expression (see PYTHON, CAMPOS, DISTIP) : Film tilt and lens distortions were calculated by a least-squares

$$
\sum_{i=1}^{N} w ((x^{i} - x^{T})^{2} + (y^{i} - y^{T})^{2}) = Min.
$$

x, y T_{xx} T $\mathtt{x}^{\mathtt{-}},\ \mathtt{y}^{\mathtt{-}}$ w measured coordinates on film (transformed into the optic axis system) true coordinates, obtained by projection of the chamber fiducials (grid intersection points) onto an ideal film plane $weight$

distance film-lens (then the α_1 are dimensionless and independent of the arrangement)

Note: The correct formula for fitting would be

 $(\frac{x}{y})' = (\frac{x}{y}) (1 + \beta_1 x + \beta_2 y + \beta_3 x^2 + \beta_4 x^4)$ $r^2 = x^2 + y^2$

 β_1 , β_2 coefficients for. film tilt

 β_7 , β_4 rotational symmetric lens distortions

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 α_3 , α_4 , α_5 were introduced to study effects of non-rotational arrangement and other, not obvious, contributions.

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b) Chamber fiducials: The distortion coefficients resulting from the previous calculations with t' ; grid plate are now used as fixed parameters in the iteration cycle for the camera positions. (Other fixed parameters are : camera coordinates, fiducial mark distance, depth of chamber, hydrogen refractive index, glass thickness and refractive index). Again, linear distortions of the film are removed by measuring the camera based fiducials before the iteration cycle. starts. Since calculation of camera position is a standard, wellknown, procedure no further description is necessary.

The fiducial mark position in space as well as depth of the chamber, were taken from a report of P. Williams (CERN/TC/NBC 65-1). All 34 fiducials have been used though not all of them are visible on all 3 views (see Appendix III). Thos'e fiducials hnve been given a small weight and did not affect the least squares fitting.

The chamber working conditions are 70 P.S.I. at 27° K. This corresponds to a hydrogen density of 0.0605 g/cm², and a refractive index of 1.0945 (calculated from Progress in Cryogenics, vol.2, p.112, 1960)

B) Results

1) Camera based fiducial marks

Relative position of fiducials on film gate.

Film gate (face to chamber)

Distances are given in mm. with an error of 0,002 mm. (Note: A black spot in fiducial 2 of camera 3 might be useful for orientation of measurements).

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B) 2) Analysis of film stretch

A typical example is given for the transformation

 $(x+) = (A)(x - x_0)$

found between measured camera based fiducials and the positions in space mentioned above :

$$
\left(\begin{array}{c}\nx \\
y\n\end{array}\right)' = 2.545333 \left(\begin{array}{c}\n0.999674 & 0.008175 \\
-0.008274 & 1.000257\n\end{array}\right) \left(\begin{array}{c}\nx - 16699.05 \\
y - 6672.72\n\end{array}\right)
$$

This transformation (see PYTHON, B 505, TRANAL) can be decomposed into:

> i) a rotation of the coordinate system into the axis of the stretch

$$
\alpha = -1.650817
$$

ii) a diagonal matrix which removes the stretch

2. 546086 *(* 1. 000000 0. 000000 ')

' 0.000000 0.999408

iii) a rotation and translation of the coordinate system to coincide with the (arbitrarily chosen) system, in which the camera based fiducials are given in space

$$
\beta = 1.659042, x_0 = 16699.05, y_0 = 6672.72
$$

The differences between true and transformed (measured) pattern are small (few microns). Other views and measurements give similar results for the film stretch.

3) grid measurements

The results for the near-centre and the near-far positions of the grid will be discussed,

The calculations were done in 3 steps :

- i) 2 iterations without taking into account distortions. This results in more accurate values for the rough input data and avoids meaningless distortion coefficients due to wrong setup.
- ii) 20 iterations with calculation of distortions. The inter-lens distances are fixed in this (and the previous) step to ensure that no unphysical deformation of the camera base is introduced.
- iii)lO iterations with calculation of distortions without fixing inter-lens distances. Since the actual position of the optical axis of the camera cannot be given precisely the above restriction would not be meaningful.

The easiest way to check the results is a comparison of X^2 - values after the various steps. (Table 1). Corrections are calculated for plane 1 and plane 2 separately (both planes are used for calculation of canera positions, but only one plane is used for calculation of distortions at a time)

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The distortion coefficients, together with errors and X^2 of the fit, are shown separately for the plane 1 and plane 2 after step iii) (Table 2).

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The errors (see PYTHON, DISTIP) are
 $\Lambda \alpha = \left[\begin{array}{ccc|ccc} 1 & \gamma^{2-\beta} & -1 & 1/2 \end{array}\right]$ $\Delta \alpha_{\mathbf{k}} = \left| \frac{\mathbf{N} - 7 - 1}{N - 1} \right| \cdot \left| \frac{\mathbf{A}}{\mathbf{A}} \cdot \mathbf{A}_{\mathbf{k},k} \right|$ $X^{2} = \frac{N}{i} w_{i} \left\{ (x_{i}^{N} - x_{i}^{N})^{2} + (y_{i}^{N} - y_{i}^{N})^{2} \right\}$ ^N*=* number of points determined by with $w = weight = 1/((v_1 \delta_1)^2 + (v_2 \delta_2)^2)$ ^v*=* relative error of point 6 *==* global error on points $A_{k, k}^{-1} =$ diagonal elements of the inverted matrix.

The measurement errors, which are important for the value of x^2 , were taken as

 δ_1 ^{ϵ} = 0.0013 cm. for the error on the grid intersection points (quoted in Appendix I)

 δ_{α} = 0.0005 cm. (2 fringes) for the error on the calculated intersection points on the film

Corroction for distortions gives a drastical improvement of the film measurements. In Fig. 1 the differences between true and measured (transformed) intersection points are plotted, without and with correction for distortions.

After correction the mean error $(\Delta x, \Delta y)$ on a measured point is about 100μ in space.

4) Chamber fiducial marks

A similar procedure as mentioned under 3) above has been applied for these calculations, except that the distortion coefficients remained fixed during the iteration cycle. Various sets of coefficients were used.

Parameters which remained unchanged for all calculations are shown in Table 3.

The results for the camera positions, calculated with the distortion coefficients found in the near-centre and near-far positions of the grid, are shown in Table 4 and Table 5. Also given are the X^2 valuos for the front and back fiducial plane.

The fiducial marks on the chamber windows as well as on the reference fiducial plane (see section C for definition) are given (again for all 4 combinations, as above) in Table 6 and Table 7.

The mean error on reconstructed fiducial marks is about 100 μ for Δx , Δy and 300 μ for Δz (see PYTHON, NEWFID, RECON)

C) . Application in Geometry Program

The definitions used in the description are shown in the following sketch :

The reference fiducial plane, shown above, carries the image' of the fiducial marks on the chamber windows, projected along the optic axis through the distorting lens onto the tilted film plane and scaled to chamber units (usually cm.). If no lens distortions are present and the film plane is parallel to the chamber reference plane $(z = 0)$ the reference fiducial plane is identical to the apparent fiducial plane.

The following steps have to be executed in the reconstruction for each view :

- 1) Transfer reference fiducials into optic axis system for each camera. This is necessary only if the fiducials are given in a unique system attached to the chamber
- 2) Find transformation between fiducials on the measured film and the reference fiducial plane (least-squares fit).
- 3) Apply this transformation to any measured point (fiducial mark, point, point of track). This results in :
	- a) transformation into optic axis system on the film,
	- b) elimination of linear distortions due to film stretch, obliquity of IEP axes, etc.

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- c) scaling of film measurements to chamber units.
- 4) Correct for film tilt and lens distortion (points and points of tracks only). The same formula which was used for the determination of the coefficients (see section A) is used. Only d (z-coordinate \sim of camera) has to be taken instead of f (distance film lens) in this formula.

Correspondence between chamber and film system (to first order optics for simplicity) is given by :

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 $x = X$. $\frac{f}{d}$ in film units

Application of the transformation in 3) scales to chamber units $x' = \frac{d}{f}$. x

Division by d gives now:

 \underline{x} ' = \underline{x} d f

(For chambers where no lens distortions and film tilt are present stop 4) will just be skipped)

After last view : Reconstruction in space using available views.

The necessary modifications have been included in the new CERN Fortran/Fortran IV versions of THRESH

D) Discussion

The results quoted above show clearly that reconstruction of events without correcting for distortions will not give reliable geometry output. It is therefore advisable to use one of the given 4 sets of data (near-centre with corrections for plane near or centre, near-far with corrections for plane near or far). Any of these sets can be used but, at least for high energy tracks which lie usually close together in the middle of the chamber, it is probably best to use the correction coefficients found for the centre position of the grid (near-centre, corrections for plane centre).

About 1000 events of the 10 GeV/c K^- experiment have been passed through THRESH and GRIND and the results seem to be correct. The expected percentage of fits is observed and probability distributions for secondary and primary vertex look reasonable. Further analysis of the data is still to be done.

One interesting fact is the significant difference between the distortion coefficients $\alpha_4(x^2)$ and $\alpha_5(y^2)$, which should be about equal, since rotational lens distortions are proportional to $r^2 = x^2 + y^2$. There exist several explanations, but no one of them is really convincing. This result will be studied in more detail.

As a final remark it should be stated that points in space can be resonstructed with about the same precision as in the 80 cm chamber. The 75 μ (for Δx , Δy) quoted for this chamber should be compared with the 100 μ found above, but due to the larger demagnification (15 instead of 10 for the front glass) this gives the same precision for the reconstruction from film measurements.

Tables and Figures :

- Table 1. x^2 -values (179 degrees of freedom) for the least-squares fit of the film measurements to the grid intersection points in space. Results for the fit to plane 1 and plane 2, resp., are given.
- Table 2. Distortion coefficients, their errors and X^2 -values of the fit (separately for the fit to plane 1 and plane 2).
- Table 3. Chamber parameters fixed during the iteration cycle for camera positions.
- Table 4. Camera coordinates and X^2 -values after least-squares fit for fiducial marks, with near-centre position of grid used for calculation of distortions, Eor comparison results are shown if no distortion corrections are introduced.
- Table 5. Ditto, with near-far position of grid.
- Table 6. Fiducial marks on the chamber and on the reference plane in a unique coordinate system attached to the chamber with near-centre position of grid used for calculation of distortions (corrections for near and centre).
- Table 7. Ditto, with near-far position of grid. (corrections for near and far)
- Figure 1. Results of least-squares fit of the film measurements to the grid intersection points in space. View 1 of the near-centre positions of the grid (fit for plane 1) has been chosen as an example. The differences Δx and Δy after step i) and step iii) are shown for the two planes.

Figure 2. Grid Plate

Figure 3. Visible and not visible fiducial marks on the 150 cm BNHBC. Numbering of fiducials according to CERN/TC/NBC 65-1.

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APPENDIX I. $(Fig. 2)$

Grid plate for testing lens distortions

The figure shows the general layout of the rulings on the plate, the lines being ruled at 2 inch intervals on a jig borer at 20° C. The rulings are in two sections A and B, the relation between those two being given below. In the A section, (left hand section), the coordinate axes are roughly as indicated, the exact position of the origin being such that the point P has coordinates in inches $(32,000, 14,000)$ and the grid rulings are parallel to the axes of coordinates.¹⁸ Let the grid intersections be denoted by pairs of numbers (m,n) which are nominally equal to the (x,y) coordinates, and actuallv equal to them in the A soction. Thus the point Q in the B section is labelled (44, 28) although these are not exactly its coordinates.

Then the actual coorodinates of grid points in the B section, referred to the axes in the A section, are given as follows :

 $x = m + 0.643 \times 10^{-4} n + 0.0003$ $y = n - 0.643 \times 10^{-4}$ m $- 0.0005$ if $m \geqslant 38$

 $(N.B.$ In fact this transformation implies the following:

The coordinates of the point $(46, 28)$ are $(46, 0021, 27.9965)$ and the B system is rotated anticlockwise 0.643×10^{-4} radian with respect to the A system).

The special crosses 1, 2, $\frac{7}{1}$ are in line with the nominal positions of the axes of cameras 1, 2, 3; the special crosses A, B, C are in line with the alignment telescope holes in the lens plate.

The grid plate is actually concave on its ruled surface with a sperical curvature of about 1.7×10^{-5} inch⁻¹; the ruling plane was parallel to the tangent plane at the point symmetrically situated between the camera axes.

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APPENDIX II.

Camera Calibration, 14th January, 1965

Before this calibration the film gates were replaced; the new gates incorporate an improved suction system which should hold the film flat with 100% reliability. The "race-track" camera fiduciary has-been replaced by a camera fiduciary system of four X crosses at the coruers of a rectanglo of approximate dimensions 28 by 84 mm. Also the reference surface on the lens plate has now been re-adjusted so that it is parallel to the film plane within about a milliradian.

The grid plate, described in the accompanying note, was photographed simultaneously by the three ameras; three positions of the grid plate were taken, corresponding approximately to the inside of the camera side window (near), the mid-plane of the chamber (middle) and the inside of the illumination side window (far). As usual the cameras are numbered 1, 2 , 3 in order from the top, i.e. the top camera is No.l. The exact positions of the grid plate are given below.

In order to define the positions of the grid plate unambiguously we set up right-handed coordinate axes $0xyz$, with positive z direction normal to the datum surface on the lens plate, along the axis of alignment telescope hole A, and pointing from the lens plate towards the grid plate; the positive y direction is downwards and parallel to the line joining the centres of the alignment telescope holes B and C in the lens plate. Thus positive x is roughly parallel to the beam direction. The grid plate is set so that, as seen from the cameras, its x and y axes are roughly parallel to these. The actual positions of the grid plate are then specified by the following data :

> L, M, the x and y direction cosines of the normal to the plate in the region near the mark A. Δx , Δy , the coordinates of mark A. h, the distance which mark B is below the axis of the alignment telescope hole B. (This gives the rotation of the grid plate in its own plane). Δz , the displacement of mark A along the z ayis relative to tho middle position of the grid plate. The numerical values are as follows :

Mark A was $60.4"({+0.1})$ from the datum plane, in the middle position.

The temperature was 20 $^{\circ}$ C for the near and middle positions, 18° C for the far position. The thermal expansion coefficient of the grid plate is 8 x 10^{-6} . PS/4944/dmh W.T. Welford.

APPENDIX III.

Fiducial marks on chamber

On the attached Fig. 3 all visible and not visible fiducial marks (outside the dashed lines) are drawn. The numbering convention is the same as indicated in CERN/TC/NBC 65-1, that means front glass fiducials are numbered 21 through 37, back glass fiducials 1 through 17.

Table 1.

a) Near-centre

b) Near-far

 \mathbb{R}^2

 $\left\langle \left\langle \cdot , \cdot \right\rangle \right\rangle$

 \mathbb{R}^2

 $\bar{\gamma}$

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a) Near-centre

 ζ

b) Near-far

Table 4

-)

 \overline{z}

a) distortion corrections for plane 1 (near)

b) distortion corrections for plane 2 (centre)

c) distortions ignored

a) distortion corrections for plane 1 (near)

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b) distortion corrections for plane 2 (far)

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 $\mathcal{L}_{\mathbf{X}}$

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 $\mathcal{L}_{\mathcal{A}}$

 \bar{z}

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 \bar{z}

Table 6

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 \bar{z}

 $\bar{\mathcal{A}}$

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 $\hat{\mathcal{A}}$

Table 6 cont.

a) with distortion corrections for plane 1 (near) - back fiducial plane

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 $**Table 6 - cont.**$ </u>

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- 1)

b) with distortion corrections for plane 2 (centre)

 $\begin{picture}(220,20) \put(0,0){\line(1,0){10}} \put(15,0){\line(1,0){10}} \put(15,0){\line($

 \rightarrow

a) with distortion corrections for plane 1 (near)

 $\ddot{}$

b) with distortion corrections for plane 2 (far)

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