



Simulation of IBS (and cooling)

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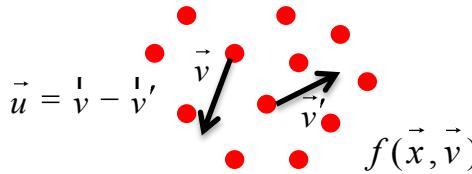
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‘Disclaimer:’ I will report about IBS and cooling simulation schemes developed and validated for the HESR (15 GeV pbars), SIS-300 (35 GeV/u U^{92+}) and RHIC at BNL.

Multiple Coulomb Collisions in a Plasma

Fokker-Planck equation

$$\frac{df(\vec{v})}{dt} = -\sum_j \frac{\partial}{\partial v_i} (f K_j) + \frac{1}{2} \sum_{i,j} \frac{\partial^2}{\partial v_i \partial v_j} (f D_{i,j})$$



friction vector: $\vec{K}(v) = 8\pi \left(\frac{q^2}{4\pi \bar{U}_0 m} \right)^2 L_c \int f(v') \frac{\vec{r}}{u^3} d^3 v'$

diffusion tensor: $D_{i,j}(\vec{v}) = 4\pi \left(\frac{q^2}{4\pi \bar{U}_0 m} \right) L_c \int f(v') \frac{u^2 \delta_{i,j} - u_i u_j}{u^3} d^3 v'$

Coulomb logarithm: $L_c = \ln \left(\frac{r_{beam}}{b_{90^\circ}} \right) \approx 10 - 20$

(small angle collisions dominate)

Diffusion rate: $\tau_{j,j}^{-1} \approx \left\langle \frac{D_{j,j}}{v_j^2} \right\rangle$

Longitudinal diffusion coefficient: $D_{\parallel} \approx n \left(\frac{q^2}{4\pi \bar{U}_0 m} \right)^2 \frac{L_c}{\Delta_{\perp}}$

Longitudinal diffusion rate: $\tau_{\parallel}^{-1} \approx n \left(\frac{q^2}{4\pi \bar{U}_0 m} \right)^2 \frac{L_c}{\Delta_{\perp} \Delta_{\parallel}}$

Gaussian velocity distribution:

$$f(\vec{v}, t) = \frac{n}{\pi \sqrt{2\pi} \Delta_{\parallel} \Delta_{\perp}} \exp \left(-\frac{v_{\perp}^2}{\Delta_{\perp}^2} \right) \exp \left(-\frac{v_{\parallel}^2}{2\Delta_{\parallel}^2} \right)$$

Anisotropic velocity distribution:

$$\Delta_{\perp} \ll \Delta_{\parallel}$$

IBS rates for high energy beams

Ratio of longitudinal/transverse velocities in the beam frame:

$$\frac{v_{\parallel}}{v_{\perp}} = \frac{\hat{\beta}_{\perp} \delta}{\gamma^2 \delta} = 1 \quad \text{with} \quad \tilde{\delta} = \left\langle \frac{\Delta p}{p} \right\rangle$$

(rms momentum spread)

Longitudinal ‘plasma’ diffusion in the lab frame: $D_{\parallel}^{ibs} = \frac{r_i^2 c N L_c}{\pi R \beta_0^3 \gamma_0^3 \left\langle \hat{\beta}_{\perp} \right\rangle^{1/2} \delta^{3/2}}$

Longitudinal ‘plasma’ IBS heating rate:

$$(\tau_{\parallel}^{-1})^{ibs} = \frac{1}{\delta} \frac{d\delta}{dt} = \frac{D_p^{IBS}}{\delta} = \frac{\Lambda_{\parallel}^{IBS}}{\delta^{3/2}} \quad \text{with} \quad \Lambda_{\parallel}^{IBS} = \frac{N c r_i^2 L_c}{8 \sqrt{\pi} R \beta_0^3 \gamma_0^3 \left\langle \hat{\beta}_{\perp}^{1/2} \right\rangle}$$

(corresponds to the Bjorken-Mtingwa result for high energies)

J.D. Bjorken, S.K. Mtingwa, Part. Accel. 13 (1983)

RHIC: A.V. Fedotov et al., PAC 2005

HESR: O. Boine-Frankenheim et al., NIM 2006

Touschek loss rate: $(\tau_{\square, loss}^{-1})^{ibs} = \frac{D_p^{ibs}}{L_c \delta_{max}^2}$

Transverse IBS heating rates:

$$(\tau_h^{-1})^{ibs} \approx (\tau_{\parallel}^{-1})^{ibs} \frac{\tilde{\delta}^2}{\epsilon_{\perp}} \left\langle \frac{D_x^2 + \tilde{D}_x^2}{\hat{\beta}_x} \right\rangle$$

with $\tilde{D}_x = D_x \hat{\alpha}_x + D'_x \hat{\beta}_x$

Cooling equilibrium for a constant cooling time :

$$\tau_c = \tau_{\parallel}^{IBS} : \quad \tilde{\delta}^2 = \frac{\tau_c \Lambda_{\parallel}^{ibs}}{\epsilon_{\perp}^{3/2}} \propto N^{2/5}$$

$$\tau_c = \tau_{\perp}^{IBS} : \quad \epsilon_{\perp} = \left(\tau_c \Lambda_{\parallel}^{ibs} \left\langle \frac{D_x^2 + \tilde{D}_x^2}{\hat{\beta}_x} \right\rangle \right)^{2/5}$$



Why/When do we need kinetic IBS simulation ?

1. Ultra-cold beams, strong correlations ($L_c < 1$).

2. Interplay of IBS with:

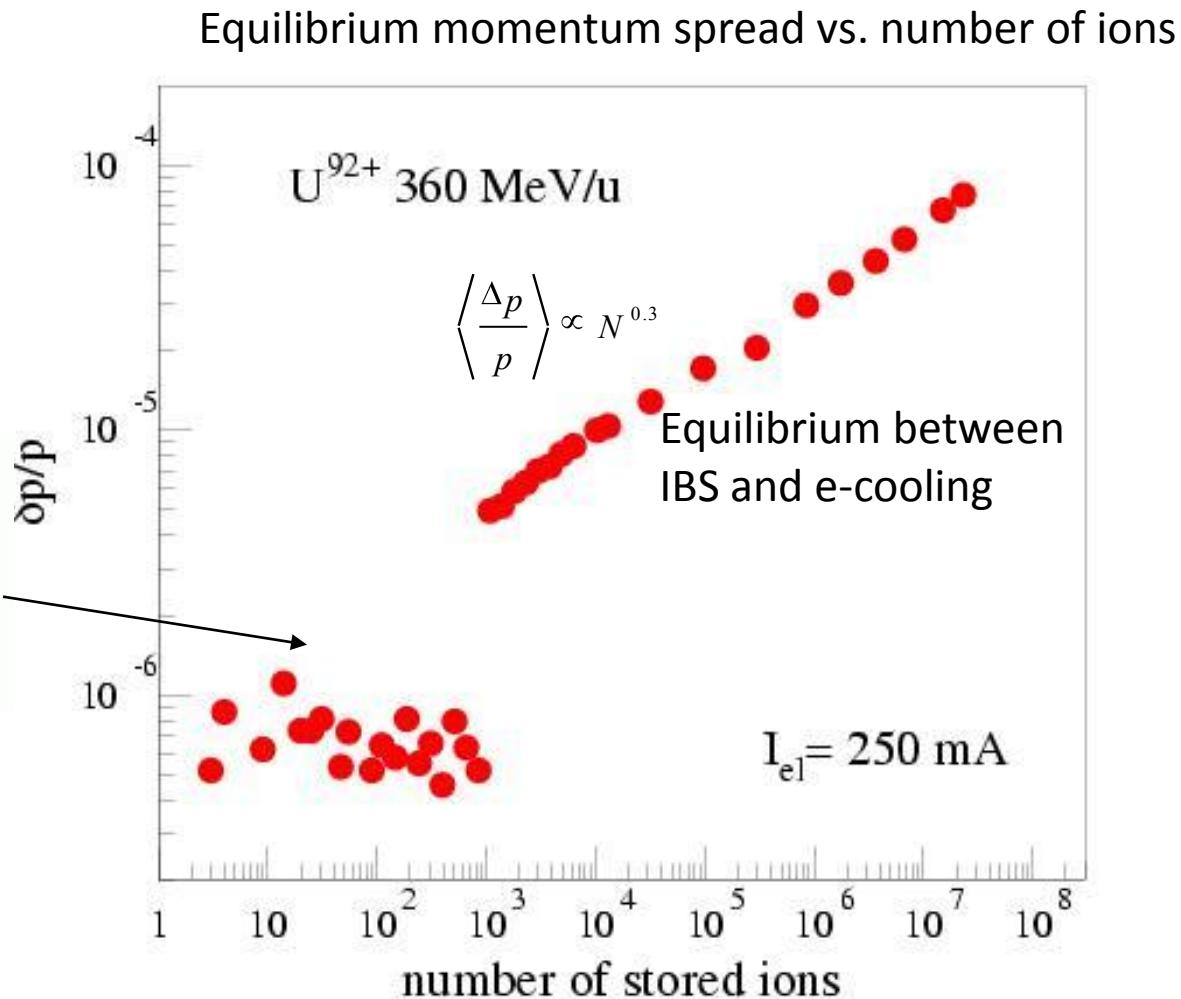
- nonlinear resonances (beam-beam, e-clouds, space charge)
 - impedances and wakes
 - internal targets
 - (nonlinear) cooling force and particle losses
- > non-Gaussian distribution functions**

Coulomb strings in coasting ion beams

$L_c \leq 1 :$

- Molecular dynamics simulations
- Tree codes

Ultra-cold beams



Experiment: M. Steck et al., Phys. Rev. Lett. 1996

Theory/Simulations: R. Hasse Phys. Rev. Lett. 1999, 2001, 2003

Why/When do we need kinetic IBS simulation ?

1. **Ultra-cold beams**, strong correlations ($L_c < 1$).

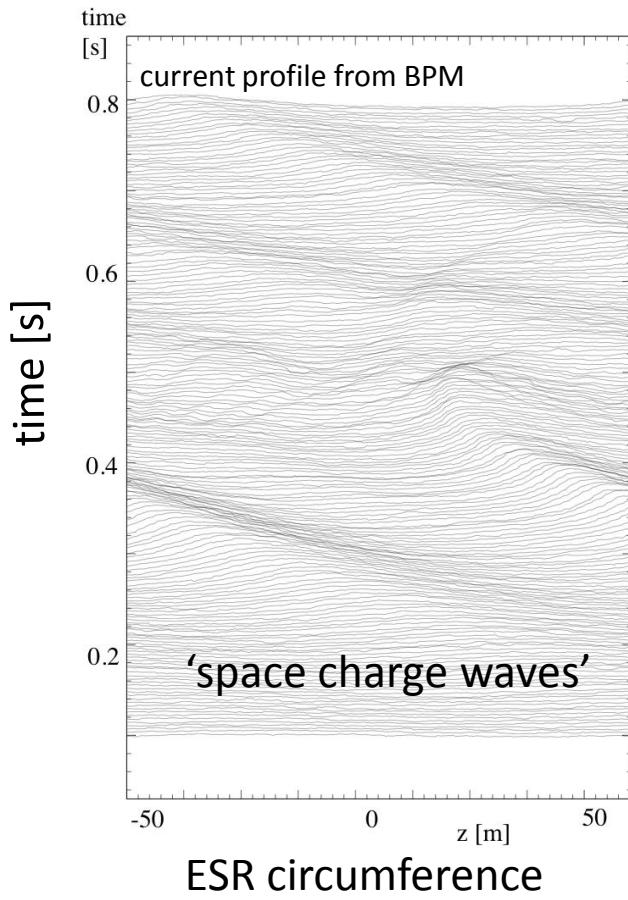
2. **Interplay of IBS with:**

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Interplay of IBS, cooling and impedances

1D coasting beam example

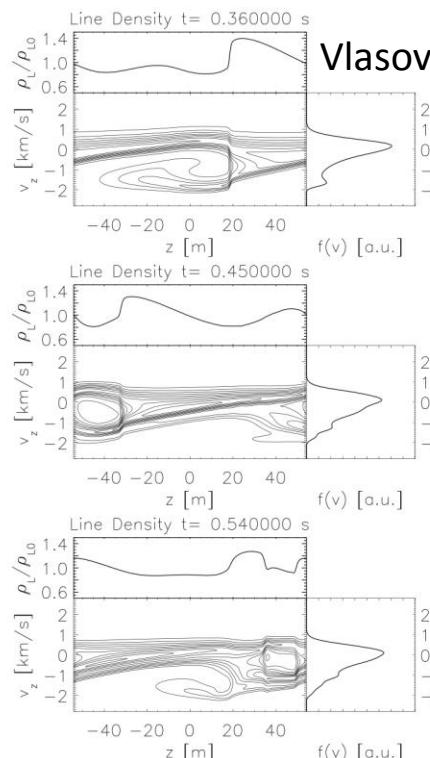
Self-bunching in a coasting beam (observed in the ESR) due to the rf cavity impedance:



1D Vlasov-Fokker-Planck equation for $f(z, \delta, t)$

$$\frac{\partial f}{\partial t} - \eta_0 v_0 \delta \frac{\partial f}{\partial z} + \frac{q V(z, t)}{p_0} \frac{\partial f}{\partial \delta} = - \frac{\partial}{\partial \delta} (F_e(\delta) f) + D_{IBS}^{IBS} \frac{\partial^2 f}{\partial \delta^2}$$

Impedances (harmonic n): $V_n(t) = (Z_n^{sc} + Z_n^{cav}) I_n(t)$



Vlasov-F-P simulation results

Direct numerical solution
of the V-F-P equation on a
grid in (z, δ) phase space.

Simplified 3D Fokker-Planck approach

Assumptions:

- Gaussian distribution
- Constant diffusion coefficients $D_{i,j}$
- No vertical dispersion

Averaging of the FP coefficients over the field and test particles:

$$\left\langle K_i \right\rangle = \frac{A_0 L_c}{2} \left\langle \frac{u_i^2}{u^3} \right\rangle \quad \left\langle D_{i,j} \right\rangle = A_0 L_c \left\langle \frac{u^2 \delta_{i,j} - u_i u_j}{u^3} \right\rangle$$

Diffusion tensor:

$$D_{i,j} = \begin{pmatrix} D_{x,x} & 0 & 0 \\ 0 & D_{y,y} & 0 \\ D_{z,x} & 0 & D_{z,z} \end{pmatrix}$$

Calculation of the coefficients using the B-M formalism:

$$D_{i,j} = A_N (\delta_{i,j} \sum_{i=1}^3 INT_{i,i} - INT_{i,j})$$

with $INT_{z,z} = 4\pi \int_0^\infty \frac{\sqrt{\lambda} (a_2 + \lambda) d\lambda}{[(a_1 + \lambda)(a_2 + \lambda) - \alpha^2]^{3/2} (a_3 + \lambda)^{1/2}}$

$$INT_{z,x} = -8\pi\alpha \int_0^\infty \frac{\sqrt{\lambda} d\lambda}{[(a_1 + \lambda)(a_2 + \lambda) - \alpha^2]^{3/2} (a_3 + \lambda)^{1/2}}$$

$$INT_{y,y} = 4\pi \int_0^\infty \frac{\sqrt{\lambda} d\lambda}{[(a_1 + \lambda)(a_2 + \lambda) - \alpha^2]^{1/2} (a_3 + \lambda)^{3/2}}$$

$$INT_{x,x} = 4\pi \int_0^\infty \frac{\sqrt{\lambda} (a_1 + \lambda) d\lambda}{[(a_1 + \lambda)(a_2 + \lambda) - \alpha^2]^{3/2} (a_3 + \lambda)^{1/2}}$$

and $\frac{a_1}{2} = \frac{1}{\varepsilon_z} + \frac{\gamma^2 (D_x^2 + \tilde{D}_x^2)}{\beta_x \varepsilon_x}, \quad \frac{a_2}{2} = \frac{\beta_x}{\varepsilon_x}, \quad \frac{a_3}{2} = \frac{2\beta_y}{\varepsilon_y}$

$$\alpha = \frac{2\gamma \tilde{D}_x}{\varepsilon_x}, \quad \tilde{D}_x = \ddot{D}_x' \beta_x + \ddot{D}_x \alpha_x$$

3D numerical solution of the Fokker-Planck equation: Langevin equations

$$\vec{P} = \begin{pmatrix} x' \\ y' \\ \frac{1}{\gamma} \frac{\Delta p}{p} \end{pmatrix} \quad \text{Langevin equation:}$$

$$P_i(t + \Delta t) = P_i(t) - K_i P_i(t) \Delta t + \sqrt{\Delta t} \sum_{j=1}^3 C_{i,j} \xi_j$$

ξ_i : Random numbers with Gaussian distribution

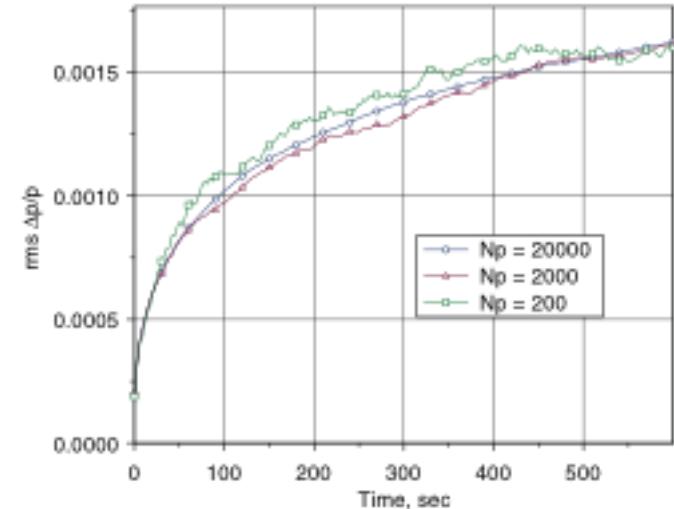
Relation between Langevin
and diffusion coefficients: $\sum_{k=1}^3 C_{i,k} C_{j,k} = D_{i,j}$

(e.g. H. Risken, *The Fokker Planck equation*, 1984)

$$D_{i,j} = \begin{pmatrix} D_{x,x} & 0 & 0 \\ 0 & D_{y,y} & 0 \\ D_{z,x} & 0 & D_{z,z} \end{pmatrix} \Rightarrow \begin{aligned} C_{11} &= \sqrt{D_{x,x} - C_{13}^2} \\ C_{22} &= \sqrt{D_{y,y}} \\ C_{33} &= \sqrt{D_{z,z}} \end{aligned}$$

$$C_{13} = \frac{1}{C_{33}}$$

$$C_{31} = 0$$



Choice of C_{ij} gives correct growth
of the emittances (B-M theory)

Outline of the algorithm:

1. Calculation of $C_{i,j}$ at every lattice elements
2. Three random numbers ξ for each macro-particle
3. Apply Langevin kick.
4. Transport particles through the lattice element

P. Zenkevich, O. Boine-Frankenheim, A. Bolshakov, NIM A (2006)

General diffusion tensor: Meshkov et al., BNL Report, 2007, **BETACOOL code**

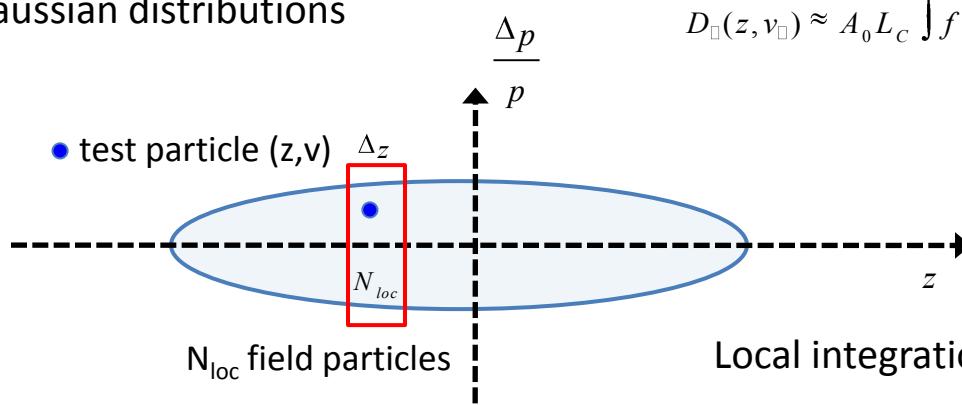


Local IBS model

a 1D illustration

Motivation:

-accurately treat tail/halo particles
in non-Gaussian distributions



Longitudinal diffusion function:

$$D_{\parallel}(z, v_{\parallel}) \approx A_0 L_C \int f(z, v'_P) \frac{\langle u_{\perp}^2 \rangle - (v_P - v'_P)^2}{\langle u_{\perp}^3 \rangle} dv'_P \quad \text{with} \quad \vec{u} = \vec{v} - \vec{v}'$$

Local integration over field particles (index m'):

$$D_{\parallel}^{\Delta z}(v_P) \approx A_0 L_C \frac{1}{N_{loc}} \sum_{m'=0}^{N_{loc}} \frac{\langle u_{\perp}^2 \rangle - (v_P - v'_{P,m'})^2}{\langle u_{\perp}^3 \rangle}$$

Langevin equation for macro-particles (index m):

$$v_{\parallel,m}(z, v, t + \Delta t) = v_{P,m}(z, v, t) - \xi_m \sqrt{\Delta t D_{\parallel}(z, v_P)}$$

Outline/problems of the algorithm:

- Diffusion coefficient has to be calculated every time step for every macro-particle
- N_{loc} large enough to reduce **numerical noise**

Plasma: Manheimer, Lampe, Joyce, J. of Comput. Phys. (1997)

RHIC: Meshkov et al., BNL Report, 2007; Fedotov, Proc. of ICFA-HB2010

IBS and internal targets

Momentum kick: $\Delta \delta_j = -F_P^e(\delta_i) + Q_j(t)\sqrt{D_P^{ibs}\Delta t} - \frac{\Delta \dot{U}_j(t)}{\beta_0^2 E_0}$

e-cooling force IBS diffusion (Langevin force) target

Target energy kick:

Integrated probability:

$$P(\Delta\epsilon_j) = \frac{\Delta t}{T_0} \int_I^{\Delta\epsilon} w(\epsilon) d\epsilon$$

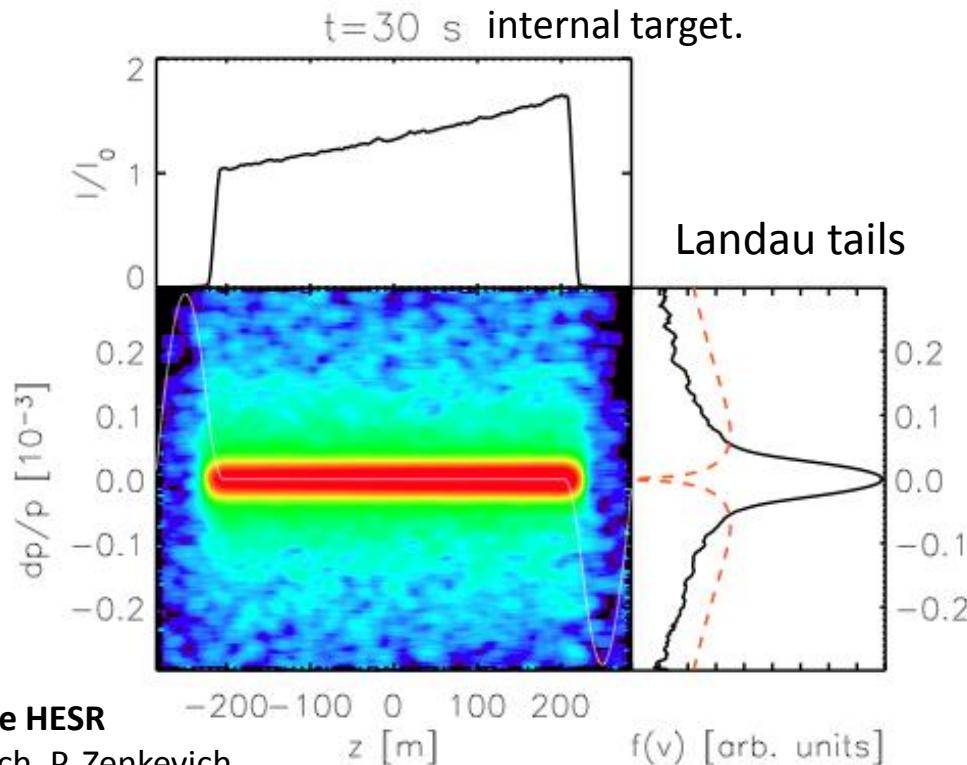
Random number:

$$P(\Delta\epsilon_j) = \xi_j(t)$$

'Thin target' requirement:

$$\Delta t \ll T_0 \frac{I}{\xi}$$

Simulation of the interaction of an electron cooled bunch in a barrier bucket with an internal target.



Beam equilibrium and beam loss with internal targets in the HESR

O. Boine-Frankenheim, R. Hasse, F. Hinterberger, A. Lehrach, P. Zenkevich
 Nucl. Inst. and Methods A 560 (2006) 245.



Validating IBS simulation modules

Global IBS: IBS emittance growth rates for Gaussian beams

IBS for non-Gaussian distributions: Integration of the stationary F-P equation

$$\frac{\partial}{\partial v} (K_{\perp}^{cool}(v) f_0(v)) + \frac{\partial^2}{\partial v^2} (D_{\perp}^{ibs}(v) f_0(v)) = 0 \Rightarrow f_0(v)$$

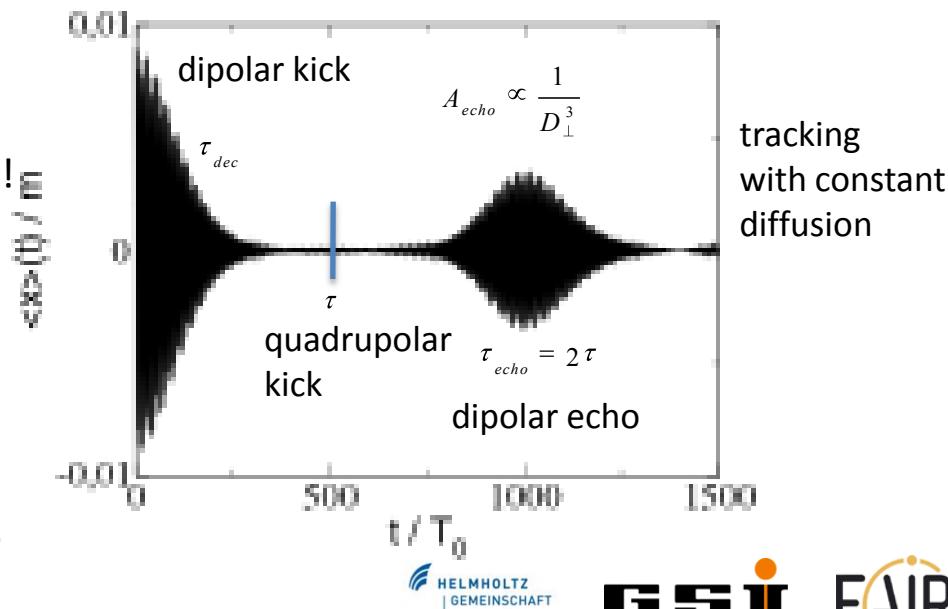
Problem in beam dynamics simulations with IBS modules:

Artificial numerical diffusion vs. IBS diffusion in (detailed) IBS simulations

Test case: (transverse) beam echo amplitudes

Echo amplitude depends very sensitive on diffusion !

$$A_{echo,rel} = \frac{\langle x \rangle_{echo,max}}{\langle x \rangle_0} = \frac{q}{\alpha^3} \frac{\tau}{\tau_{dec}} \quad \alpha = \frac{2}{3} \frac{\tau}{\tau_{diff}} \left(\frac{\tau}{\tau_{dec}} \right)^2 + 1 \quad D_{\perp} = \frac{\epsilon_{\perp}}{\tau_{diff}}$$



Simulation of beam echoes and code validation:

Sorge, Boine-Frankenheim, W. Fischer, Proc. ICAP 2006

Al-khateeb, Boine-Frankenheim, Hasse, PRST-AB 2003

Conclusions

IBS simulation schemes based on the Fokker-Planck equation have been presented.

- a. Direct solution of the Vlasov-Fokker-Planck equation in 1D
- b. Langevin equations in 1D-3D:
 - Constant diffusion coefficients from B-M theory
 - Local coefficients for non-Gaussian distributions

Code validation for IBS effect is an important issue (numerical noise vs. IBS) !

For most applications the Langevin equations with constant diffusion coefficients might be the method of choice.