

## Minutes of the QCD Meeting on 28. Sept. 1989

Present: Günter Hauser, Eberhard Lange, Thomas Lohse, Wolfgang Männer, Scott Menary, Andreas Roth, Gerald Rudolph, Richard St. Denis, Heinz Sander, Michael Schmelling, Hans-Jörg Schmidt, Ron Settles, Terry Sloan, Makoto Takashima, Werner Wiedenmann, Weimin Wu, Dezhi Xu.

1. The distributions for the first QCD paper were discussed.
2. Brian Webber gave us a talk on his model, "HERWIG".
3. - Contribution by Gerald Rudolph: Several investigations of the parameter fitting method have been done.

I. Fit 4 parameters of LUND 2<sup>nd</sup> order to LUND shower at 92 GeV:

A comparison of my fit results (2<sup>nd</sup> iteration) with Hans-Joerg Schmidt's shows

- Agreement in  $\sigma$  and B;
- Some discrepancy in  $y_{min}$  and  $\Lambda$ ;
- H.J.'s errors are 5 - 10 times bigger than mine;
- Overall, the  $\chi^2$  of my fit is slightly better, but H.J.'s set fits thrust better.

Since the  $\chi^2$  is anyway bad for most of the distributions, the result depends on the set of distributions chosen (matching the 2 models is not possible).

The deviations in  $\Lambda$  and  $y_{min}$  reported earlier when fitting test samples generated at a point on the hypersphere have been studied once more. Fitting only 1 parameter at a time (get rid of correlation) shows slight systematic deviations in each of the parameters, probably due to the parabolic parametrisation.

II. Fit 5 parameters of LUND shower to LUND shower at 92 GeV:

Since this model fits lower energy data best, it should be studied in more detail.

Fitting the 5 parameters  $\Lambda$ ,  $M_{min}$  (or  $Q_0$ ),  $\sigma$ , A and B to several test event samples generated at the center of the sphere (where the method should be most

reliable) and using the distributions  $x$ ,  $S$ ,  $A$  gives the correct result on average but large fluctuations occur sometimes.

The  $\chi^2$  of the fit is excellent, as expected. Correlations between parameters are very high (around 0.9). They don't go away if 1 or 2 parameters are held fixed. Results are more stable if only 1 parameter is fitted at a time, but still same fluctuations are larger than the  $1-\sigma$  errors from MINUIT.

III. A fit of LUND shower to 35 GeV TASSO data gives an overall good fit and

$$\Lambda = 0.380 \pm .008 \text{ GeV}$$

$$M_{min} = 1.640 \pm .080 \text{ GeV}$$

$$\sigma = 0.375 \pm .004 \text{ GeV}$$

$$A = 0.628 \pm .030 \text{ GeV}$$

$$B = 0.939 \pm .040 \text{ GeV}^{-2},$$

quite close to the defaults, except perhaps  $M_{min}$ . Therefore, I don't confirm the rather strange values for  $\Lambda$ ,  $A$  and  $B$  obtained by Phil Burrows and published in the TASSO paper. Phil's set gives a worse overall  $\chi^2$ .

Correlations are not as strong as in the case above.

1. Fit Lund  $0(d_s^2)$  to Lund parton shower, 92 GeV

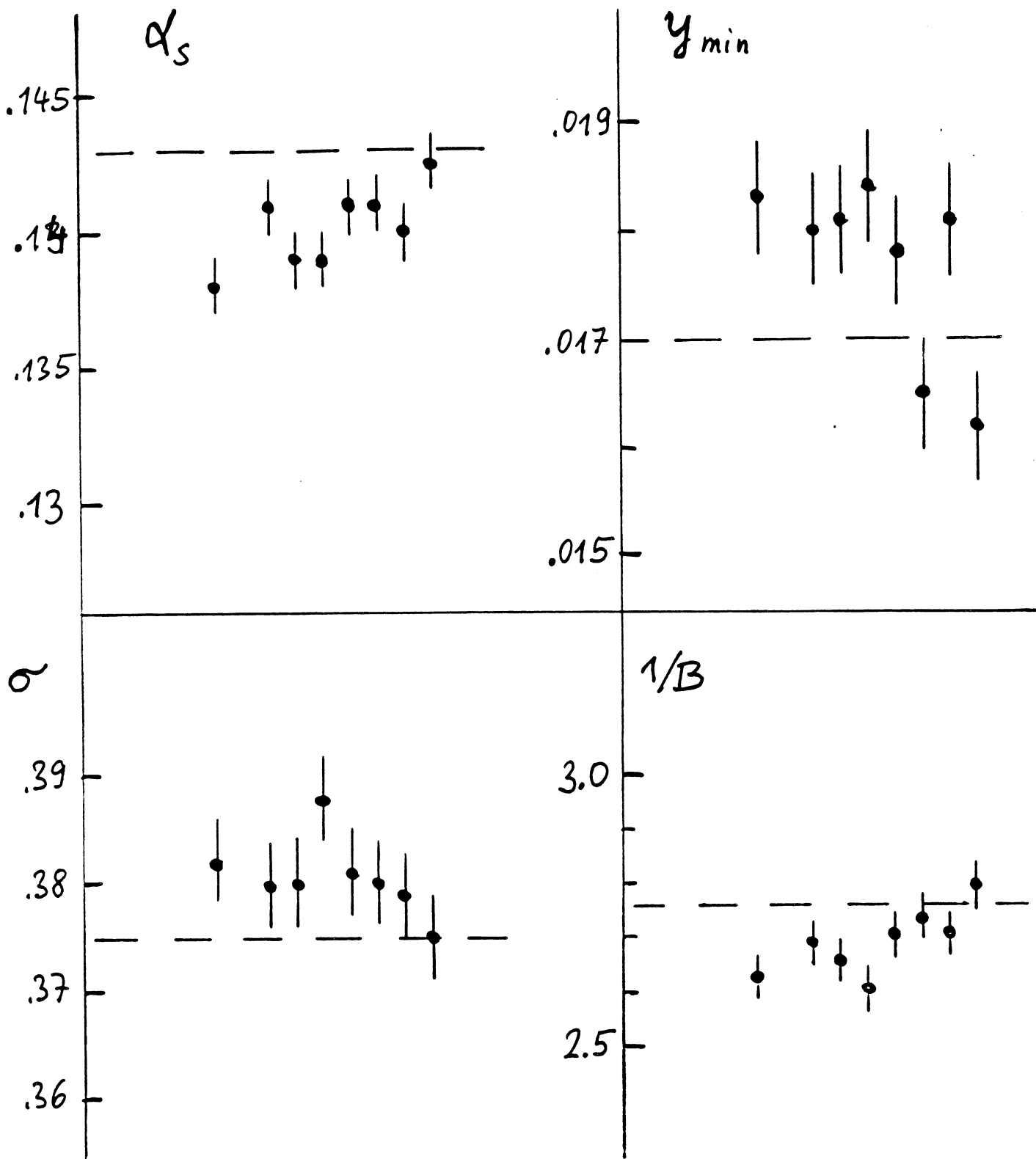
parameter	default (PETRA/PEP)	compare fit results :	
		GERALD	HANS-JÖRG
$\Lambda$ [GeV]	.50	$.815 \pm .055$	$1.04 \pm .60$
$y_{\min}$	.02	$.0174 \pm .0005$	$.029 \pm .006$
$\sigma$ [GeV]	.40	$.497 \pm .005$	$.490 \pm .030$
$B$ [GeV $^{-2}$ ]	.70	$.291 \pm .008$	$.287 \pm .030$

Comments : good agreement in  $B$  and  $\sigma$ .  
agreement in  $\Lambda$  within error.  
discrepancy in cut-off parameter  $y_{\min}$ .

Observable	# intervals	$\chi^2$ - value (15000 ev. Lund $0(d_s^2)$ - 3000 ev. Lund p.s.)	
$N_{ch}$	13	69	64
$X_p$	26	89	188 $\leftarrow$ too many high - x too few soft
$P_{Tout}$	9	85	138
$P_{Tin}$	9	30	63
$Q_1$	6	155	159
$Q_2$	7	13	118 $\leftarrow$ too much peaked
$T$	9	260 $\leftarrow$ miserable!	43 $\leftarrow$ much better!
$Min$	7	285	198
$Obl$	8	4	95
Sum $\chi^2 =$		990	1066

My fit used  $N_{ch} \times P_{T0} P_{T1} Q_1 Q_2 Min Obl$  as observables  
Correlation coefficients (global):

$\alpha_s$	$y_{\min}$	$\sigma$	$\frac{1}{B}$
.45	.66	.66	.84



Fit of Lund  $O(\alpha_s^2)$  to Lund  $O(d_s^2)$  test events  
 at 92 GeV. (8 samples, 3000 ev. each,

Testpoint lies on boundary of parameter region (sphere).

Fit each parameter at a time.  
 (avoid correlations)

2. Fit Lund p.s. to Lund p.s., at 92 GeV  
 (as a test of the method)

Parameter ranges:

$$\Lambda_{LLA} : .20 - .40 - .60 \text{ GeV}$$

$$Q_0 : 1.0 - 2.0 - 3.0 \text{ GeV}$$

$$\sigma : .25 - .35 - .45 \text{ GeV}$$

$$A : .20 - .50 - .80$$

$$B : .50 - .90 - 1.3 \text{ GeV}^2$$

↑  
center of sphere (= default values)

MC statistics generated :  $1/2 \cdot 10^6$

Fit observables  $X, S_c, A_c$

Adjust 5 parameters simultaneously:

large fluctuations in some cases

prob. due to large degree of correlation:

$\Lambda$	$Q_0$	$\sigma$	A	B
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.90	.70	.90	.92	.95
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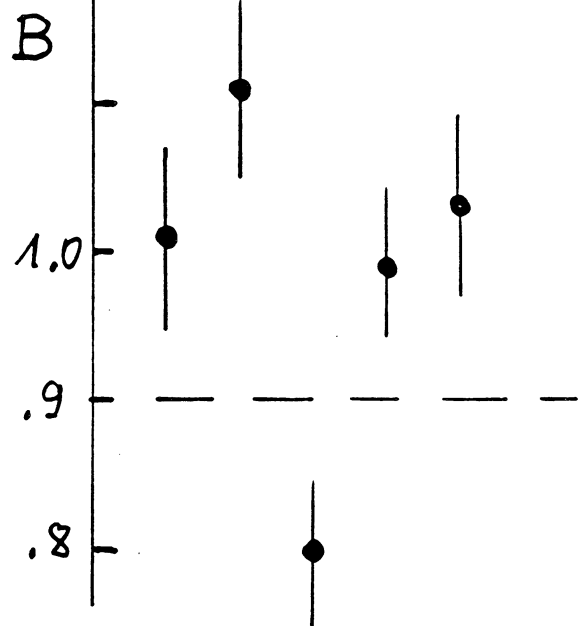
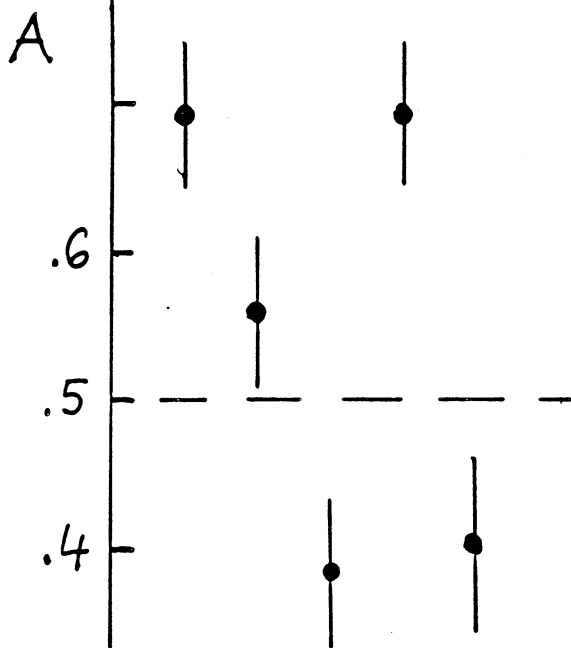
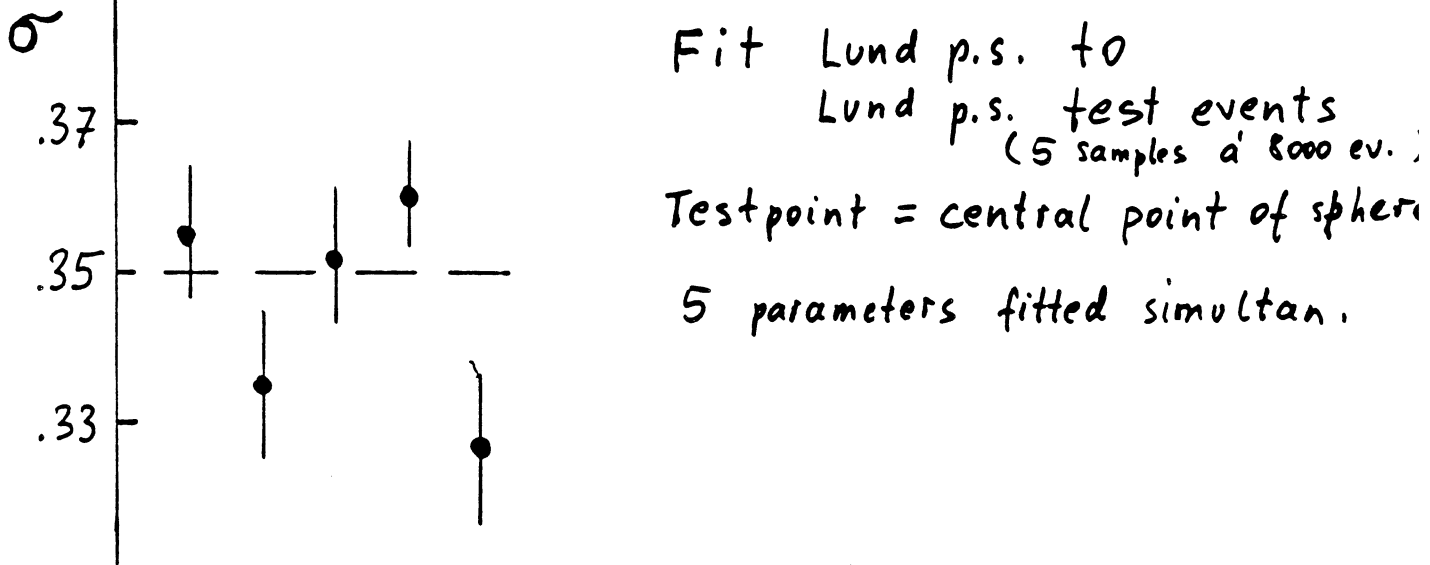
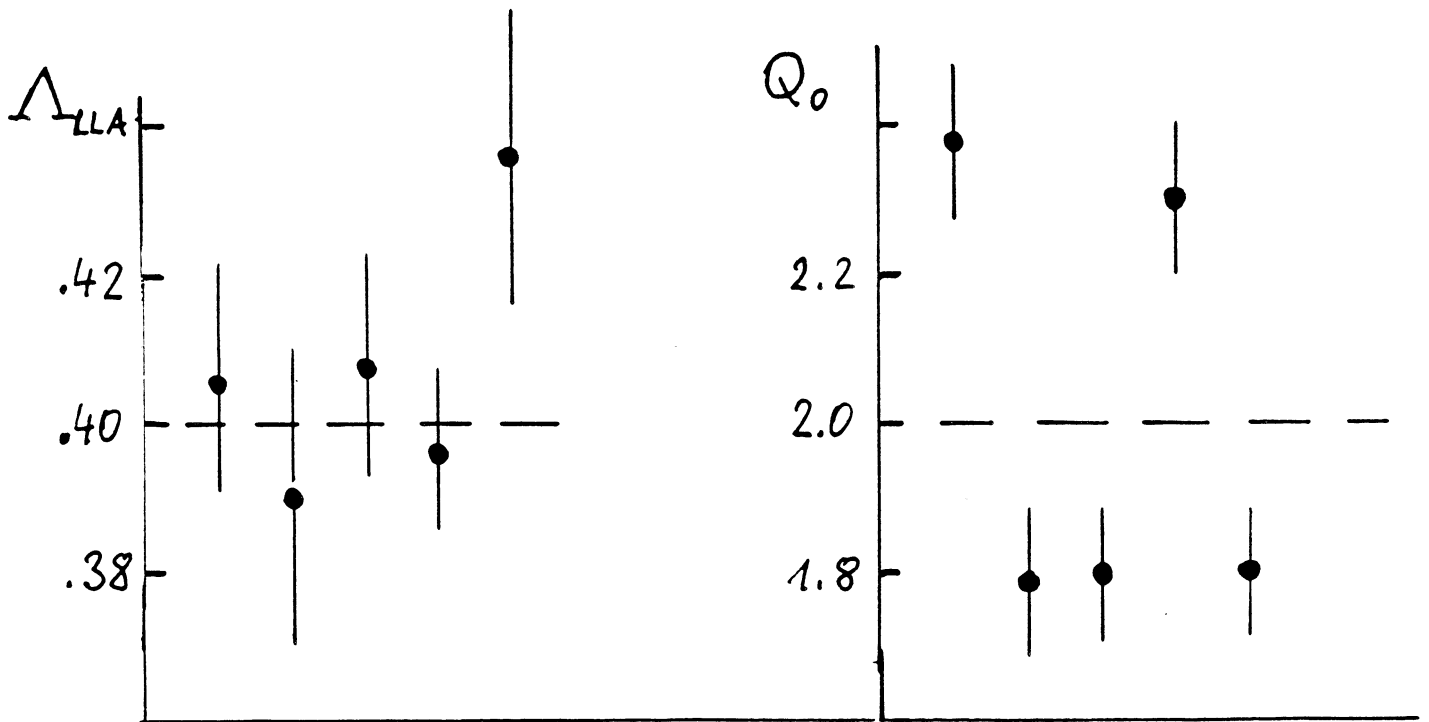
"global" correlation coefficients  
given by MINUIT

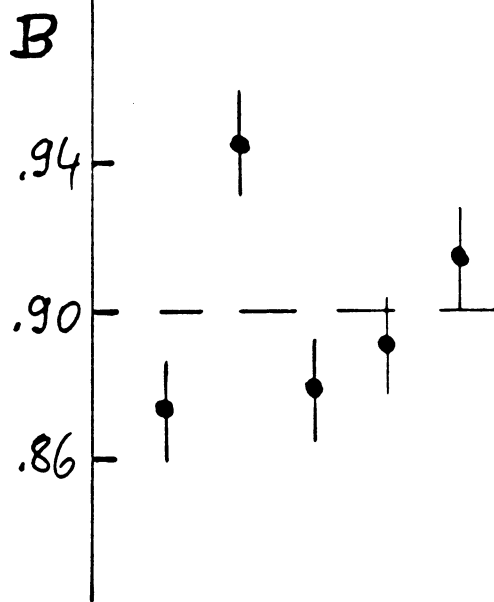
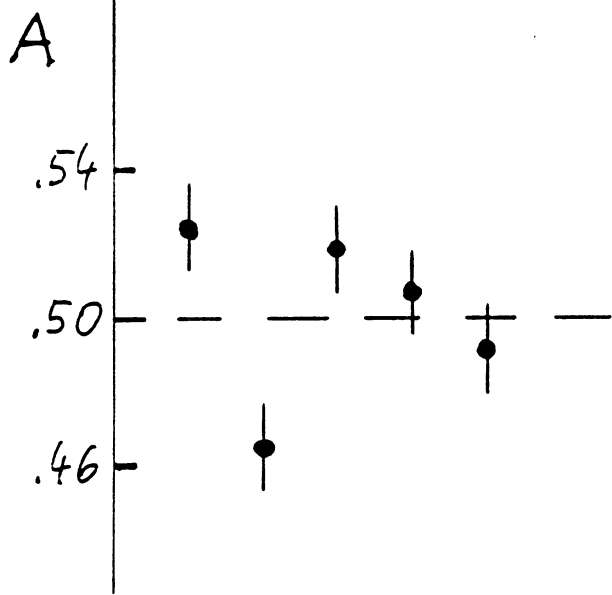
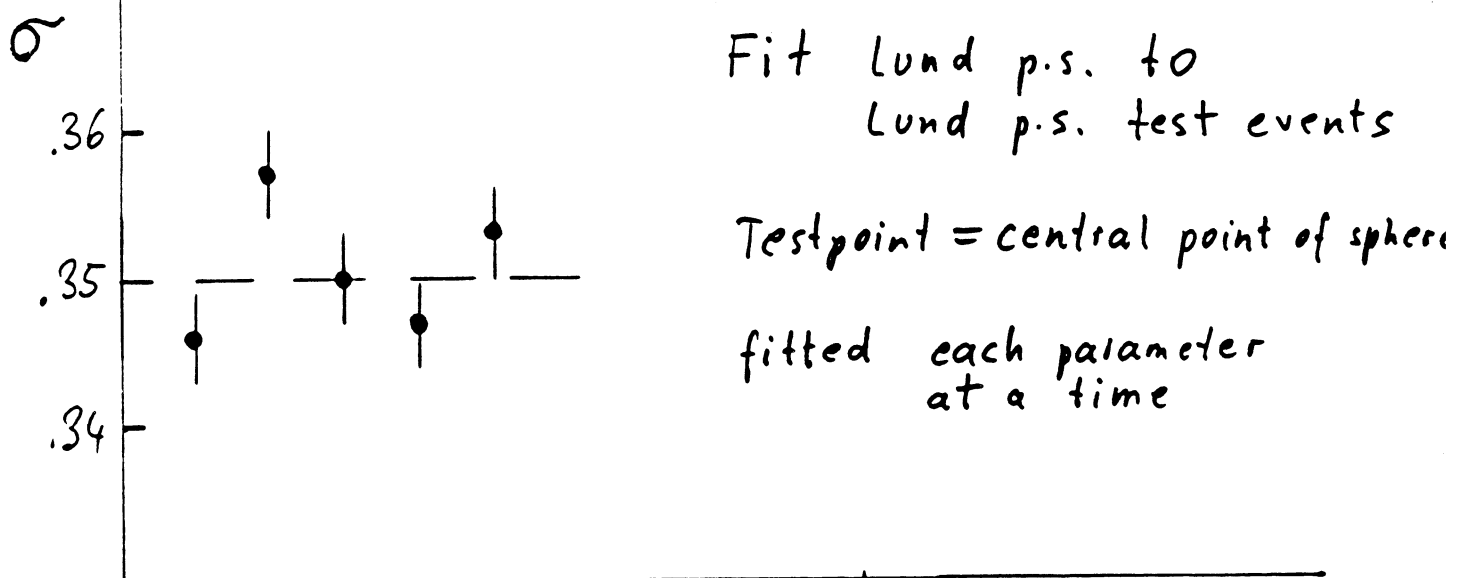
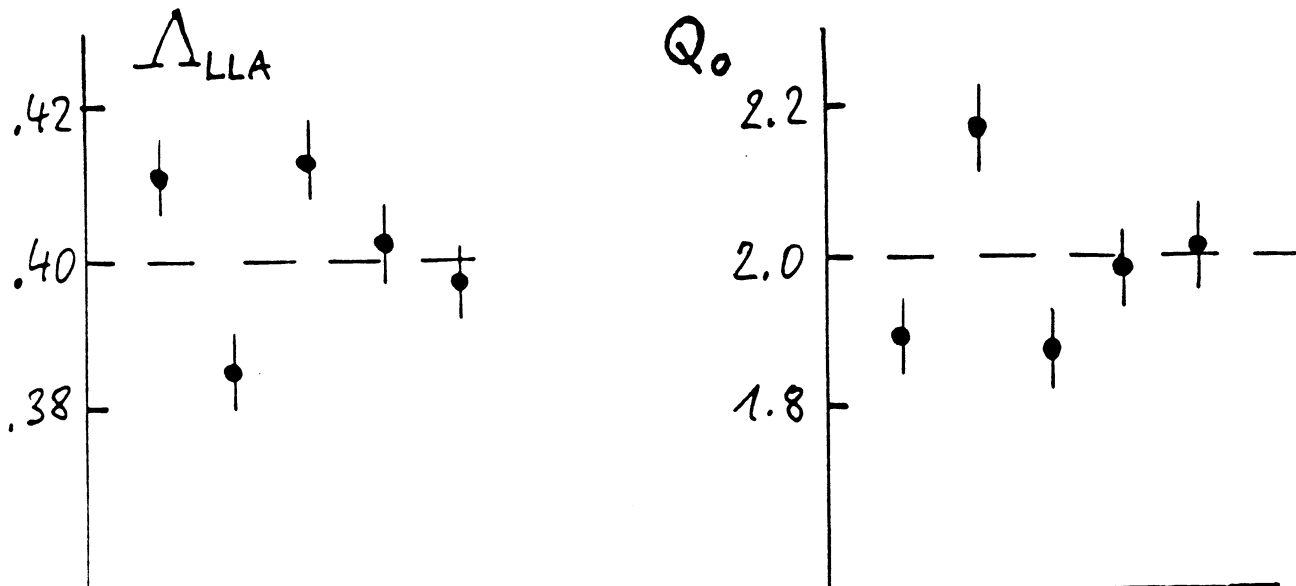
Adjust one parameter at a time (fixing the other 4 at the values they should have):

good results

no systematic shifts,

errors given by MINUIT seem too small, however





### 3. Fit Lund p.s. to TASSO Data, 35 GeV

Used same parameter region than at 92 GeV

Statistics generated  $1/2 \cdot 10^6$  events

global fit to observables  $X$   $P_{T0}$   $P_{Ti}$   $\langle P_{T0}^2 \rangle$   $\langle P_{Ti}^2 \rangle$  S A T

Parameter	def.	MARK II FIT 29 GeV	My fit to MARK II data	TASSO FIT 35 GeV	My fit to TASSO data
$\Lambda$	.40	.40	$.43 \pm .004$	.26	$.38 \pm .008$
$Q_0$	1.0	1.0	$1.1 \pm .05$	[1.0]	$1.64 \pm .08$
$\sigma$	.35	.354	$.34 \pm .003$	.39	$.375 \pm .004$
A	.50	.50	$.45 \pm .01$	.18	$.628 \pm .03$
B	.90	.90	$.77 \pm .02$	.34	$.939 \pm .04$

$$\frac{\chi^2}{ndf} = \frac{209}{132} = 1.6$$

Result close to defaults.

correlation coeff. (global):

$\Lambda$	$Q_0$	$\sigma$	A	B
.60	.65	.70	.70	.80

TASSO published parameter set  
gives somewhat worse description

$$\frac{\chi^2}{ndf} = \frac{252}{132} = 1.9$$

for above set of distrib's.  
( $p_T$  tails worse,  
but Aplanarity better)