



Improving topological cluster reconstruction using calorimeter cell timing in ATLAS

The ATLAS Collaboration

Clusters of topologically connected calorimeter cells around cells with large absolute signal-to-noise ratio (*topo-clusters*) are the basis for calorimeter signal reconstruction in the ATLAS experiment. Topological cell clustering has proven performant in LHC Runs 1 and 2. It is, however, susceptible to *out-of-time pile-up* of signals from soft collisions outside the 25 ns proton-bunch-crossing window associated with the event's hard collision. To reduce this effect, a calorimeter-cell timing criterion was added to the signal-to-noise ratio requirement in the clustering algorithm. Multiple versions of this criterion were tested by reconstructing hadronic signals in simulated events and Run 2 ATLAS data. The preferred version is found to reduce the out-of-time pile-up jet multiplicity by $\sim 50\%$ for jet $p_T \sim 20$ GeV and by $\sim 80\%$ for jet $p_T \gtrsim 50$ GeV, while not disrupting the reconstruction of hadronic signals of interest, and improving the jet energy resolution by up to 5% for $20 < p_T < 30$ GeV. Pile-up is also suppressed for other physics objects based on *topo-clusters* (electrons, photons, τ -leptons), reducing the overall event size on disk by about 6% in early Run 3 pile-up conditions. Offline reconstruction for Run 3 includes the timing requirement.

Contents

| | | |
|----------|---|-----------|
| 1 | Introduction | 2 |
| 2 | ATLAS detector | 3 |
| 2.1 | Calorimeter timing measurement | 4 |
| 3 | Data and Monte Carlo samples | 6 |
| 4 | The topo-cluster reconstruction algorithm | 7 |
| 5 | The cell time cut | 8 |
| 6 | Performance of the time cut on MC simulation | 10 |
| 6.1 | Performance on topo-clusters | 10 |
| 6.2 | Performance on jets | 13 |
| 7 | Performance on ATLAS data | 21 |
| 7.1 | Checks for time cut inefficiencies | 29 |
| 7.2 | Impact on the ATLAS event size | 34 |
| 8 | Conclusion | 35 |

1 Introduction

The ATLAS experiment [1], one of the four major experiments at the LHC [2], relies on the precise measurement of particle showers as the starting point for physics object reconstruction. Calorimeter showers are reconstructed by clustering groups of topologically connected calorimeter cells, the first step in this process involving cells with high absolute energy relative to calorimeter noise [3]. The produced clusters are referred to as *topo-clusters* and provide the input to the reconstruction of a number of physics objects, including jets, electrons, τ -leptons and missing transverse momentum (E_T^{miss}). The topological cell clustering algorithm is well established and performed very well during the LHC Run 1 and Run 2 data-taking periods, from 2010 to 2012 and 2015 to 2018 respectively. However, further improvement is possible, especially in view of the increased luminosity foreseen in LHC Run 3 and beyond.

At the LHC, bunches of about 10^{11} protons collide every 25 ns, giving rise to multiple pp interactions in each *bunch-crossing*.¹ Signals from particles produced in additional soft pp collisions pile up on top of those from the pp *hard-scattering* that triggered the ATLAS data-recording process. If the additional particles are produced in the same bunch-crossing that produced the recorded event's pp *hard-scattering*, they are referred to as *in-time pile-up*, whereas if they are produced in the previous or next bunch-crossings they are called *out-of-time pile-up*. Calorimeter signals are sensitive to both in-time and out-of-time pile-up, which can result in the formation of additional topo-clusters, as well as spurious contributions to those originating from the hard scattering. Multiple pile-up suppression techniques have been studied in ATLAS, such as the grooming algorithms used for the reconstruction of boosted objects [5–7] or the Constituent Subtraction [8] and Soft Killer [9] methods used prior to jet reconstruction [10]. These methods, however,

¹ The average number of pp collisions per bunch-crossing was 25 in 2016, 38 in 2017 and 36 in 2018 data-taking [4].

are all applied after topo-cluster reconstruction, whereas in this paper a low-level mitigation technique is explored with the goal of reducing pile-up contributions at cell level while building the topo-clusters.

The ATLAS calorimeters provide a measurement of the signal time, in addition to the deposited energy. In this paper, the topo-clustering algorithm is modified by applying a calorimeter time-measurement selection criterion (the *time cut* in the following) to cells with large signal-to-noise ratio in order to reject contributions from out-of-time pile-up. Multiple versions of the time cut were compared and tested on both Monte Carlo (MC) simulated events and a sample of ATLAS data events. An Upper Limit, switching off the time cut for very high energy signals, was explored and it was adopted as a safety measure to avoid rejecting potential new-physics processes which might produce out-of-time signals. The performance of the new algorithm has been tested in the context of jet reconstruction, focusing on the suppression of pile-up-initiated jets as well as its effects on jet energy calibration and resolution.

The rest of this paper is organised as follows. Section 2 describes the detector. Section 3 lists the data and MC samples used for this analysis. Section 4 describes the current topo-cluster reconstruction algorithm and Section 5 introduces the time cut. Section 6 discusses the time cut’s performance on topo-cluster and jet kinematics as evaluated using MC samples. Section 7 discusses the time cut’s performance on ATLAS data taken during Run 2 of the LHC, including checks for possible time-cut inefficiency and the cut’s impact on the ATLAS event size. Conclusions are presented in Section 8.

2 ATLAS detector

The ATLAS detector [1] at the LHC covers nearly the entire solid angle around the collision point.² It consists of an inner tracking detector surrounded by a thin superconducting solenoid, electromagnetic and hadron calorimeters, and a muon spectrometer incorporating three large superconducting air-core toroidal magnets.

The inner-detector system is immersed in a 2 T axial magnetic field and provides charged-particle tracking in the range $|\eta| < 2.5$. The high-granularity silicon pixel detector covers the vertex region and typically provides four measurements per track, the first hit normally being in the insertable B-layer [11, 12] installed before Run 2. It is followed by the silicon microstrip tracker, which usually provides eight measurements per track. These silicon detectors are complemented by the transition radiation tracker (TRT), which enables radially extended track reconstruction up to $|\eta| = 2.0$. The TRT also provides electron identification information based on the fraction of hits (typically 30 in total) above a higher energy-deposit threshold corresponding to transition radiation.

The calorimeter system covers the pseudorapidity range $|\eta| < 4.9$. Within the region $|\eta| < 3.2$, electromagnetic (EM) calorimetry is provided by barrel and endcap high-granularity lead/liquid-argon (LAr) calorimeters, with an additional thin LAr presampler covering $|\eta| < 1.8$ to correct for energy loss in material upstream of the calorimeters. Hadron calorimetry is provided by the steel/scintillator-tile calorimeter (*Tile calorimeter* in the following), segmented into three barrel structures within $|\eta| < 1.7$,

² ATLAS uses a right-handed coordinate system with its origin at the nominal interaction point (IP) in the centre of the detector and the z -axis along the beam pipe. The x -axis points from the IP to the centre of the LHC ring, and the y -axis points upwards. Cylindrical coordinates (r, ϕ) are used in the transverse plane, ϕ being the azimuthal angle around the z -axis. The pseudorapidity is defined in terms of the polar angle θ as $\eta = -\ln \tan(\theta/2)$. Angular distance is measured in units of $\Delta R \equiv \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2}$. The rapidity $y = \frac{1}{2} \ln \left(\frac{E+p_z}{E-p_z} \right)$ can also be used for kinematic distributions instead of η , as it accounts for the particle’s mass.

and two copper/LAr hadronic endcap calorimeters. The solid angle coverage is completed with forward copper/LAr and tungsten/LAr calorimeter modules optimised for electromagnetic and hadronic energy measurements respectively. Electromagnetic and hadronic calorimeters based on LAr technology are collectively referred to as the *LAr calorimeter* in the following. Additional details about the timing measurements in the ATLAS calorimeters are provided in Section 2.1.

The muon spectrometer comprises separate trigger and high-precision tracking chambers measuring the deflection of muons in a magnetic field generated by the superconducting air-core toroidal magnets. The field integral of the toroids ranges between 2.0 and 6.0 T · m across most of the detector. Three layers of precision chambers, each consisting of layers of monitored drift tubes, cover the region $|\eta| < 2.7$, complemented by cathode-strip chambers in the forward region, where the background is highest. The muon trigger system covers the range $|\eta| < 2.4$ with resistive-plate chambers in the barrel, and thin-gap chambers in the endcap regions.

Interesting events are selected by the first-level (L1) trigger system implemented in custom hardware, followed by selections made by algorithms implemented in software in the high-level trigger (HLT) [13]. The first-level trigger accepts events from the 40 MHz bunch-crossings at a rate below 100 kHz, which the high-level trigger reduces in order to record events to disk at about 1 kHz. Triggers with acceptance rates that are too large are prescaled, i.e. only a fraction of the events satisfying the trigger are written to disk. The prescale factor is then applied to estimate the original rate. An extensive software suite [14] is used in data simulation, in the reconstruction and analysis of real and simulated data, in detector operations, and in the trigger and data acquisition systems of the experiment.

2.1 Calorimeter timing measurement

In addition to the energy deposited in each cell, both the LAr and Tile calorimeters can measure the time of arrival of the particle depositing the energy. LAr signals are read out using Front End Boards (FEB). They shape the signal and sample it at 40 MHz, four samples are then digitised if the event passes the L1 trigger. The shape of the LAr signal is optimised to minimise the noise contribution [15]. An example of the typical LAr calorimeter pulse shape is shown in Figure 1. The long negative tail of the shaped LAr pulse implies that out-of-time pile-up provides a negative energy contribution on average, while the contribution from in-time pile-up is positive on average. Because of this, the average pile-up energy per event is zero. While this feature is very useful in correcting for the average pile-up energy deposition, suppression methods like the one presented in this paper are still very useful in suppressing pile-up contributions to individual physics objects.

The signal amplitude A (proportional to its energy) and timing (t) are both reconstructed using an optimal filtering algorithm [16] applied to the digitised samples S_i :

$$A = \sum_{i=1}^{n_{\text{samples}}} a_i S_i, \quad t = \frac{1}{A} \sum_{i=1}^{n_{\text{samples}}} b_i S_i, \quad (1)$$

where the optimal filtering coefficients a_i and b_i are computed from the predicted pulse shape and measured noise. The cell time is only measured if the detected energy is above a certain configurable threshold. Typically, threshold values equal to three times the cell noise ($3\sigma_{\text{noise}}$) are used. If the reconstructed energy is below threshold, then the cell time is not computed and $t = 0$ is stored.

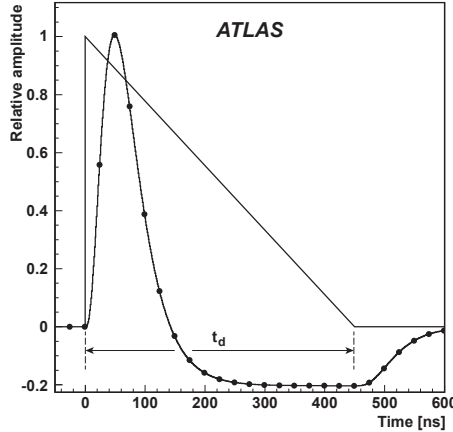


Figure 1: Example of the LAr (central EM barrel) calorimeter pulse shape from Ref. [3]. The unipolar triangular pulse is the one generated by fast ionising particles in the liquid argon. Its characteristic time is the drift time $t_d \simeq 450$ ns in the example shown. The shaped pulse is superimposed, with a characteristic duration of $t_{\text{signal}} \simeq 600$ ns. The full circles on the shaped pulse indicate the nominal bunch-crossings at 25 ns intervals.

The LAr time measurement is synchronised with the LHC clock and fine-tuned at the FEB level [17]. Time alignment corrections are recalculated using beam-splash and early collision data [18]. The measured time is also monitored as part of the data quality assessment [19] in order to ensure its stability. The LAr time resolution is typically parameterised as the sum in quadrature of a constant term and a $\sim 1/E$ noise term:

$$\sigma_t = \frac{p_1}{E} \oplus p_0.$$

In Run 2 and after applying calibration constants obtained from data, the constant term p_0 was found to reach ~ 200 ps, while the noise term p_1 was $\mathcal{O}(1 \text{ GeV ns})$ [18].

The Tile calorimeter signal is also shaped, amplified and sampled at 40 MHz by the front-end electronics. Unlike the LAr calorimeter, seven rather than four samples are digitised and the cell energy and time are reconstructed via an optimal filtering method as described by Eq. (1). If the signal amplitude is below 5 ADC counts,³ then the timing is not measured and a default value of $t = 0$ is stored. Since most Tile calorimeter cells are read out by two independent channels, the average of the two times is taken as the measured cell time [20]. The Tile calorimeter uses its Laser calibration system [21] to monitor the timing of all channels during physics runs. Laser events are recorded during empty bunch-crossings in physics data-taking and the stability of the timing is monitored as part of the data quality assessment, allowing corrections for possible misconfiguration to be made in the main data processing. Channels that exhibit a large number of timing jumps are flagged as *bad timing* and are not used for further physics object reconstruction [21]. Cells flagged as Tile *bad timing* are not considered when implementing the time cut. The Tile timing resolution is described with a functional form similar to the one commonly used for the calorimeter energy resolution:

$$\sigma_t = p_0 \oplus \frac{p_1}{E} \oplus \frac{p_2}{\sqrt{E}}.$$

³ In a few Tile cells affected by higher radiation (*gap scintillators* [20]), the minimum amplitude for time reconstruction is 15 ADC counts.

In Run 2, the constant term was found to be ~ 300 ps, while p_1 and p_2 are $O(1 \text{ GeV ns})$ and $O(1 \sqrt{\text{GeV}} \text{ ns})$, respectively, with variations depending on the gain [22].

3 Data and Monte Carlo samples

The time cut's performance has been studied on simulated multi-jet event samples as well as pp collision data from the ATLAS detector. The data sample used corresponds to one run, taken by ATLAS in 2017 during Run 2 of the LHC at a centre-of-mass energy $\sqrt{s} = 13 \text{ TeV}$, with an integrated luminosity of $\mathcal{L} = 76 \pm 2 \text{ pb}^{-1}$. Because the goal of this study is only to compare jet kinematic distributions, a larger data sample is not needed. In the selected run the LHC provided collisions from bunch trains with a nominal bunch-crossing spacing of 25 ns. The recorded average mean number of interactions per bunch-crossing $\langle \mu \rangle = 38.8$ is consistent with the average $\langle \mu \rangle$ recorded during the 2017 data-taking period. Data quality requirements are applied to ensure that all detector components were fully operational [23]. Data were collected using multiple single-jet triggers (see Section 7). Events selected by each trigger are scaled by the corresponding trigger prescale factor. In a few *Luminosity Blocks*⁴ at the beginning of the run, the prescale factors for the chosen single-jet triggers were found to be very large, resulting in an apparent loss of statistical precision due to recording few events with very large trigger weights. To avoid this issue, the Luminosity Blocks in question were excluded from the analysis. The loss of integrated luminosity due to this additional rejection amounts to 6.5%. The integrated luminosity after data quality requirements are applied is $\mathcal{L} = 71 \pm 2 \text{ pb}^{-1}$. The uncertainty in the 2017 integrated luminosity is assumed for this run; it amounts to 2.4% [4], obtained using the LUCID-2 detector [24] for the primary luminosity measurements.

Additional data samples were used to check for inefficiencies caused by time miscalibration and to test the ATLAS event size for reductions due to the time cut. For time-cut inefficiency checks, one Luminosity Block of data collected in 2018 was used, with an average number of interactions per bunch-crossing of $\langle \mu \rangle = 42.5$. The data used for the event size test also consists of one Luminosity Block, taken in 2022. The average pile-up for this run is $\langle \mu \rangle = 41.3$.

Multi-jet production was modelled using PYTHIA 8.230 [25] with leading-order matrix elements for dijet production which were matched to the parton shower. The renormalisation and factorisation scales were set to the geometric mean of the squared transverse masses of the two outgoing particles in the matrix element, $\sqrt{(p_{T,1}^2 + m_1^2)(p_{T,2}^2 + m_2^2)}$. The NNPDF2.3LO set of parton distribution functions (PDFs) [26] was used in the matrix element generation, the parton shower, and the simulation of the multi-parton interactions. The A14 [27] set of tuned parameters was used. The modelling of fragmentation and hadronisation was based on the Lund string model [28, 29]. The effect of both in-time and out-of-time pile-up was modelled by overlaying the simulated hard-scattering event with inelastic proton–proton (pp) events generated with PYTHIA 8.186 [30] using the NNPDF2.3LO PDF set [26] and the A3 set of tuned parameters [31].

The MC events have been reweighted to account for the difference between the simulated number of interactions per bunch-crossing (μ) profile and the one from the single run used for this paper. The correction factor is defined in bins of μ , as the ratio of the data to MC sample μ spectra, normalised to

⁴ A Luminosity Block (LB) is a period of time during which the instantaneous luminosity, detector and trigger configuration, and data quality conditions are considered constant. In general, one LB corresponds to a time period of 60 s, although LB duration is flexible and actions that might alter the run configuration or detector conditions trigger the start of a new LB before 60 s have elapsed. LB start and end timestamps are assigned in real time during data-taking by the ATLAS Central Trigger Processor [23].

unity and prior to any event selection. This reweighting is not applied when performing MC-only studies, to allow a wider μ range to be covered.

An additional MC sample was used to assess the ATLAS event size reduction due to the time cut. This MC sample describes $t\bar{t}$ production in the fully hadronic decay channel. The production of $t\bar{t}$ events was modelled using the POWHEG BOX v2 [32–35] generator at NLO with the NNPDF3.0_{NLO} [36] PDF set and the h_{damp} parameter set to $1.5 m_{\text{top}}$ [37]. The events were interfaced to PYTHIA 8.307 [25] to model the parton shower, hadronisation, and underlying event, with parameters set according to the A14 tune [27] and using the NNPDF2.3_{LO} set of PDFs [26]. The decays of bottom and charm hadrons were performed by EVTGEN 2.1.1 [38].

4 The topo-cluster reconstruction algorithm

The cell time cut is introduced into the existing ATLAS topo-cluster reconstruction algorithm, a detailed description is provided in Ref. [3]. In short, ATLAS topo-clustering is based on the cell signal significance $\zeta_{\text{cell}}^{\text{EM}}$, defined as the ratio of the cell signal energy to the average expected noise (Eq. (2)). Both are measured at the electromagnetic (EM) scale, defined by calibrating the calorimeter energy measurement to the energy deposited by an electron or photon, without any correction for the non-compensating response of the calorimeter.

$$\zeta_{\text{cell}}^{\text{EM}} = \frac{E_{\text{cell}}^{\text{EM}}}{\sigma_{\text{noise,cell}}^{\text{EM}}}. \quad (2)$$

Topo-clusters are formed by a growing-volume algorithm, configured by three threshold parameters $\{S, N, P\}$:

$$|\zeta_{\text{cell}}^{\text{EM}}| > S \quad (\text{primary seed threshold}), \quad (3)$$

$$|\zeta_{\text{cell}}^{\text{EM}}| > N \quad (\text{threshold for growth control}), \quad (4)$$

$$|\zeta_{\text{cell}}^{\text{EM}}| > P \quad (\text{principal cell filter}); \quad (5)$$

set to $\{S = 4, N = 2, P = 0\}$ in Runs 1, 2 and 3. The algorithm is seeded by cells (*seed* cells in the following) whose signal significance exceeds a threshold S (Eq. (3)) and which have not been marked as having either read-out or general signal extraction problems in the actual run conditions. Seed cells are then sorted from highest to lowest energy significance and topo-clusters are grown by adding all neighbouring cells that satisfy Eq. (5).

Two cells are considered to be neighbours if they are directly adjacent in a given sampling layer, or, if in adjacent layers, if they at least partially overlap in the plane. If a neighbouring cell has a signal significance above the threshold N in Eq. (4), then the procedure is iterated over its neighbours, while if it satisfies Eq. (5) but not Eq. (4), its neighbours are not included in the growing topo-cluster. This algorithm iterates over further neighbours until no more neighbouring cells satisfying Eq. (4) are found. Since the S, N and P thresholds are applied to the absolute significance, cells of either positive or negative energy can seed a cluster or be included in one. As discussed in Section 2.1, negative cell signals in the ATLAS calorimeters are often the result of out-of-time pile-up, since the residual signal trace produced by out-of-time particles is scaled by the negative tail of the LAr calorimeter signal-shaping function [15]. Only clusters with an overall positive energy are used as inputs for jet reconstruction. The algorithm described so far, however, has no limitations to how large a cluster can grow and it could potentially lead to very large, unphysical clusters. For this reason, clusters are then split around local energy maxima, as detailed in Ref. [3].

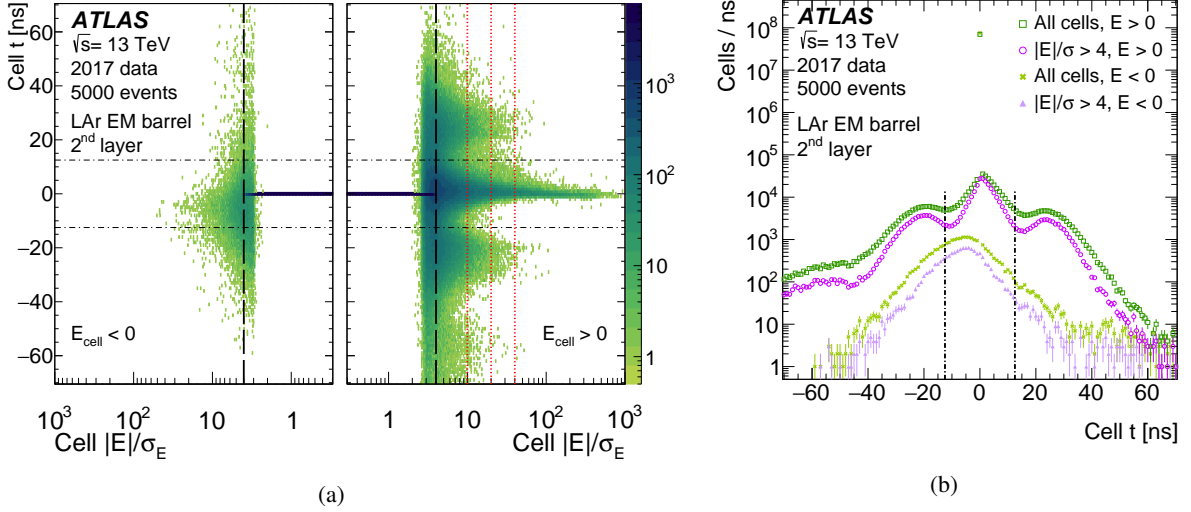


Figure 2: Time and energy significance for calorimeter cells in the second layer of the LAr EM barrel calorimeter. (a) shows the two-dimensional cell time vs significance spectrum. The cell time is computed if the cell energy exceeds a given threshold (typically $3 \cdot \sigma_{\text{noise,cell}}^{\text{EM}}$), below which a default value of $t = 0$ is stored. The dashed vertical line represents the seed candidate requirement $S = 4$ (Eq. (3)), while the horizontal lines represent the cell time rejection limits of ± 12.5 ns. The three dotted vertical lines represent the three possible values considered for the Upper Limit (see Section 5.0.3). (b) shows a comparison between the inclusive cell time spectrum and the one for seed candidates ($|E|/\sigma_E > 4$). The vertical lines represent the cell time rejection limits of ± 12.5 ns.

5 The cell time cut

As discussed in Section 2.1, both the ATLAS LAr and Tile calorimeters provide a time measurement for signals of sufficiently large energy. Figure 2 shows the distribution of the energy significance and time of calorimeter signals in real data. Calorimeter cells belonging to the LAr EM barrel region are picked as an example, but similar distributions occur in other calorimeter regions. At high energy significance, two secondary peaks at ± 25 ns are clearly visible, associated with out-of-time pile-up from the previous and next bunch-crossings. The asymmetry visible between positive and negative cell time is due to the LAr calorimeter being affected by a larger number of bunch-crossings before the current one than after it.

At lower energy significance the time resolution is poorer, and the secondary peaks cannot in general be distinguished. When limiting the study to candidate *seed* cells, i.e. those with energy significance greater than $S = 4$, the three peaks are separated well enough to justify a selection on the absolute cell time, requiring it to be within ± 12.5 ns of the midpoint between the secondary peaks. Multiple approaches to implementing a cell time cut in topo-cluster building were explored. The most relevant are detailed below.

5.0.1 The *Seed* time cut

The requirement for seeding a topo-cluster is modified. Seed cells must satisfy Eq. (6), i.e. have both an absolute cell signal significance larger than $S = 4$ and a cell time within ± 12.5 ns:

$$\begin{cases} |\zeta_{\text{cell}}^{\text{EM}}| > S & (\text{primary seed threshold, default } S = 4) \\ |t_{\text{cell}}| < 12.5 \text{ ns.} \end{cases} \quad (6)$$

The cluster growth and splitting stages of the topo-clustering algorithm are left untouched. Since this selection is exclusively applied to topo-cluster seeds, it is referred to as the *Seed cut* in the following.

5.0.2 The *Seed Extended* time cut

In the *Seed cut* implementation, a cell exceeding the signal significance threshold in Eq. (6), but failing the time cut, would be prevented from seeding a cluster. It could, however, be included in another cluster as a neighbouring cell if it falls in the vicinity of an in-time signal. Due to the high level of pile-up at the LHC, this phenomenon is not unlikely. To prevent their inclusion, a tighter version of the cut was defined, in which candidate seed cells that satisfy Eq. (3), but not Eq. (6), are vetoed from being included in growing topo-clusters. In the following, this version of the time cut is referred to as the *Seed Extended* cut.

5.0.3 The Upper Limit

Although this paper focuses on the reduction of out-of-time pile-up, delayed calorimeter signals with substantial energy can also arise from new physics, for example from the decays of slow-moving long-lived particles (LLPs) that are predicted by several theories for physics beyond the Standard Model [39, 40]. Optimising the time cut for LLP signals would require taking into account the many signatures considered for LLP searches and is outside the scope of this paper. A conservative approach was therefore adopted, where the time cut is not applied to signals with very large significance, thus leaving the decision of whether or not to reject such signals to a later stage of the reconstruction chain, where differentiation between standard and LLP-dedicated selections is possible.

The time cut was therefore modified by introducing an Upper Limit (UL). The time cut is turned off if the cell signal significance $\zeta_{\text{cell}}^{\text{EM}} = E_{\text{cell}}^{\text{EM}} / \sigma_{\text{noise,cell}}^{\text{EM}}$ is positive and larger than a given value X_{UL} , i.e. if:

$$\begin{cases} |\zeta_{\text{cell}}^{\text{EM}}| > S & (\text{primary seed threshold, default } S = 4) \\ |t_{\text{cell}}| < 12.5 \text{ ns} \end{cases} \quad \text{OR} \quad \zeta_{\text{cell}}^{\text{EM}} > X_{\text{UL}}.$$

A lower value of X_{UL} implies a larger phase space left untouched by the time cut. Three values $X_{\text{UL}} = 10, 20$ and 40 were tested in order to estimate how low the Upper Limit can be set without reintroducing too much pile-up into the event. While the Upper Limit could in principle be combined with either the *Seed cut* or the *Seed Extended* cut, in this paper it is only studied in combination with the *Seed Extended* cut, as the latter appeared more promising than the *Seed cut* (Section 6). The Upper Limits $X_{\text{UL}} = 10, 20$ and 40 are shown as dotted lines in Figure 2. As expected, $X_{\text{UL}} = 40$ affects a very limited number of out-of-time cells and it is expected to have a negligible effect, while $X_{\text{UL}} = 10$ is expected to have the largest impact.

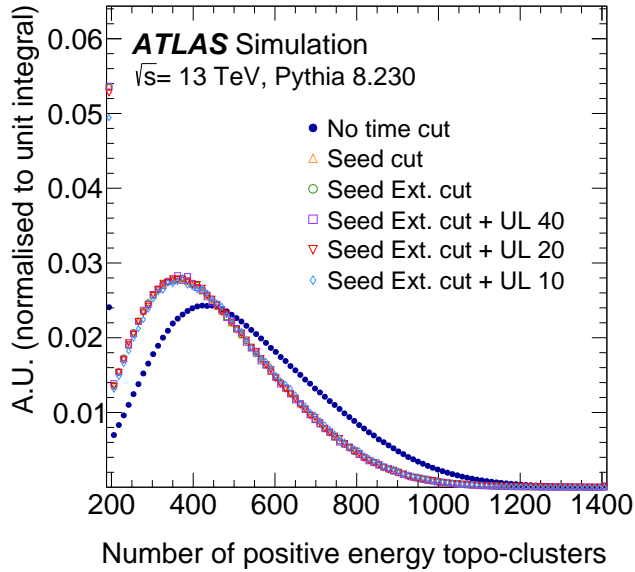


Figure 3: Number of positive energy ($E_{cl} > 0$) reconstructed topo-clusters per event: comparison between no time cut and multiple choices for the time cut. The first and last bins show the overflow and underflow.

6 Performance of the time cut on MC simulation

The performance of the time cut was evaluated using a PYTHIA 8.230 dijet MC sample, as described in Section 3. The event selection is kept as loose as possible in order to observe ample ranges of cluster energy and jet energy.

6.1 Performance on topo-clusters

Since the time cut is applied in the topo-clustering algorithm, the overall number of clusters is expected to be reduced, because the applied cut leads to a smaller number of cluster seeds. It should be noted, however, that the topo-clusters built with the time cut applied are not necessarily a subsample of those built using only the cell energy significance. Excluding out-of-time signals may change the number of local maxima, prompting additional cluster splitting. Figure 3 shows the number of clusters reconstructed per event; only positive-energy topo-clusters (i.e. eligible inputs to the jet reconstruction) are shown. The number of positive-energy clusters in each event is on average $\sim 15\%$ smaller when the time cut is applied, with very little difference between the different choices for the cut. The Seed and Seed Extended cuts differ only in their treatment of out-of-time signals with significant energy when these signals are included as neighbouring cells, and hence the two algorithms can be expected to differ more strongly in their effect on topo-cluster properties than on the number of reconstructed topo-clusters. Since the UL only truncates the relatively sparsely populated high-energy tail of the topo-cluster spectrum, its effect then becomes subdominant when considering the inclusive number of clusters.

Removing out-of-time contributions can also result in a change in the kinematic properties of the reconstructed topo-clusters. This effect is expected to be more significant when applying the Seed Extended time cut as opposed to the Seed time cut. Figure 4 shows the multiplicity of topo-clusters that have various

kinematic properties, comparing the various time cut options. Each kinematic moment of the topo-clusters is defined as the weighted average of the property of interest over the cells contributing to the cluster. A full description of the topo-clusters' properties is given in Ref. [3]. An additional calibration using the local cell weighting (LCW) scheme is applied⁵ to take into account the non-compensating response of the calorimeter, out-of-cluster energy deposits and energy deposited in the dead material within the detector [3]. Only the topo-clusters with overall positive energy in each event are shown. Bottom panels show the 'cut / no-cut' ratio plot. Uncertainties in the ratios must take into account that the distributions with and without the time cut are obtained from the same events, and hence are correlated in a non-trivial way. To properly account for the correlations, uncertainties are computed by splitting the available sample into subsamples of approximately 10 000 events each, and the cut / no-cut ratio is recomputed for each subsample. The ratio's uncertainty is estimated from the standard deviation of the resulting distribution. The estimated uncertainty is propagated to the inclusive ratio as a relative uncertainty. This method is used to compute the uncertainties in the cut / no-cut ratio plots throughout this paper.

Figure 4(a) shows the cluster time. When no time cut is applied, two shoulders are clearly visible near ± 25 ns, consistent with being due to out-of-time pile-up from the previous and next bunch-crossings. Once the time cut is applied, the number of clusters around ± 25 ns is reduced by more than 80% by the Seed Extended time cut, smoothing out the two shoulders, while the main peak at 0 ns appears unchanged. The Seed cut has a smaller impact, reducing the two peaks by $\sim 70\%$, while the impact of the Seed Extended plus Upper Limit $X_{UL} = 10$ cut lies between those of the Seed and Seed Extended cuts. Overall, topo-clusters with absolute time greater than 12.5 ns are not removed entirely, since out-of-time contributions can still occur due to predominantly low-energy cells being included in topo-clusters as neighbouring cells.

The relative effect of the time cut on the number of clusters is also found to vary with jet rapidity (Figure 4(b)), the largest effect being a $\sim 25\%$ reduction in the cluster multiplicity for $2 \lesssim |y| \lesssim 3.5$. The energy deposited by pile-up is expected to increase at larger absolute rapidity, but the time cut has only a $\sim 10\%$ effect above $|y| \sim 3.5$ because each forward calorimeter cell covers a larger pseudorapidity interval. This geometrical effect implies that larger numbers of both in-time particles and out-of-time particles deposit energy in the same cell, thus resulting in a less effective distinction between in-time and out-of-time signals. The effect of the time cut is also seen to vary with the cluster energy (Figure 4(c)). The number of clusters with energy between about 1 GeV and 100 GeV is reduced by $\sim 25\%$, with the reduction decreasing to $\sim 10\%$ at both very low and very high energies. Finally, the dependence on the cluster p_T (Figure 4(d)) provides the best way to observe the effect of the Upper Limit. As expected, $X_{UL} = 40$ has hardly any effect, while $X_{UL} = 20$ produces a small difference and it starts to diverge from the Seed Extended cut for clusters with $p_T \gtrsim 15$ GeV. However, $X_{UL} = 10$ has a larger impact, which starts to appear for clusters with $p_T \gtrsim 2$ GeV. At $p_T = 5$ GeV for instance, the Seed Extended cut reduces the number of clusters by $\sim 20\%$, while the Seed Extended plus $X_{UL} = 10$ cut reduces their number by only $\sim 15\%$.

Figure 5 shows two topo-cluster moments. The topo-cluster isolation (Figure 5(a)) is computed from the number of non-clustered cells on the outer perimeter of the topo-cluster [3]: an isolation value close to 1 indicates an isolated topo-cluster, while it tends to 0 when the topo-cluster is not isolated. Since the lower values of the isolation spectrum are affected the most by the time cut, it can be concluded that topo-clusters reconstructed with the time cut tend to be more isolated than those from the standard algorithm. This observation can be explained as a combination of two effects. Firstly, isolated topo-clusters are mostly produced by EM showers, and are less sensitive to pile-up because of their smaller size. Secondly, the pile-up reduction introduced by the timing cut is likely to produce more isolated topo-clusters. Specific

⁵ The LCW scheme is not propagated to the jets, in order to mimic the approach most commonly followed by ATLAS physics analyses.

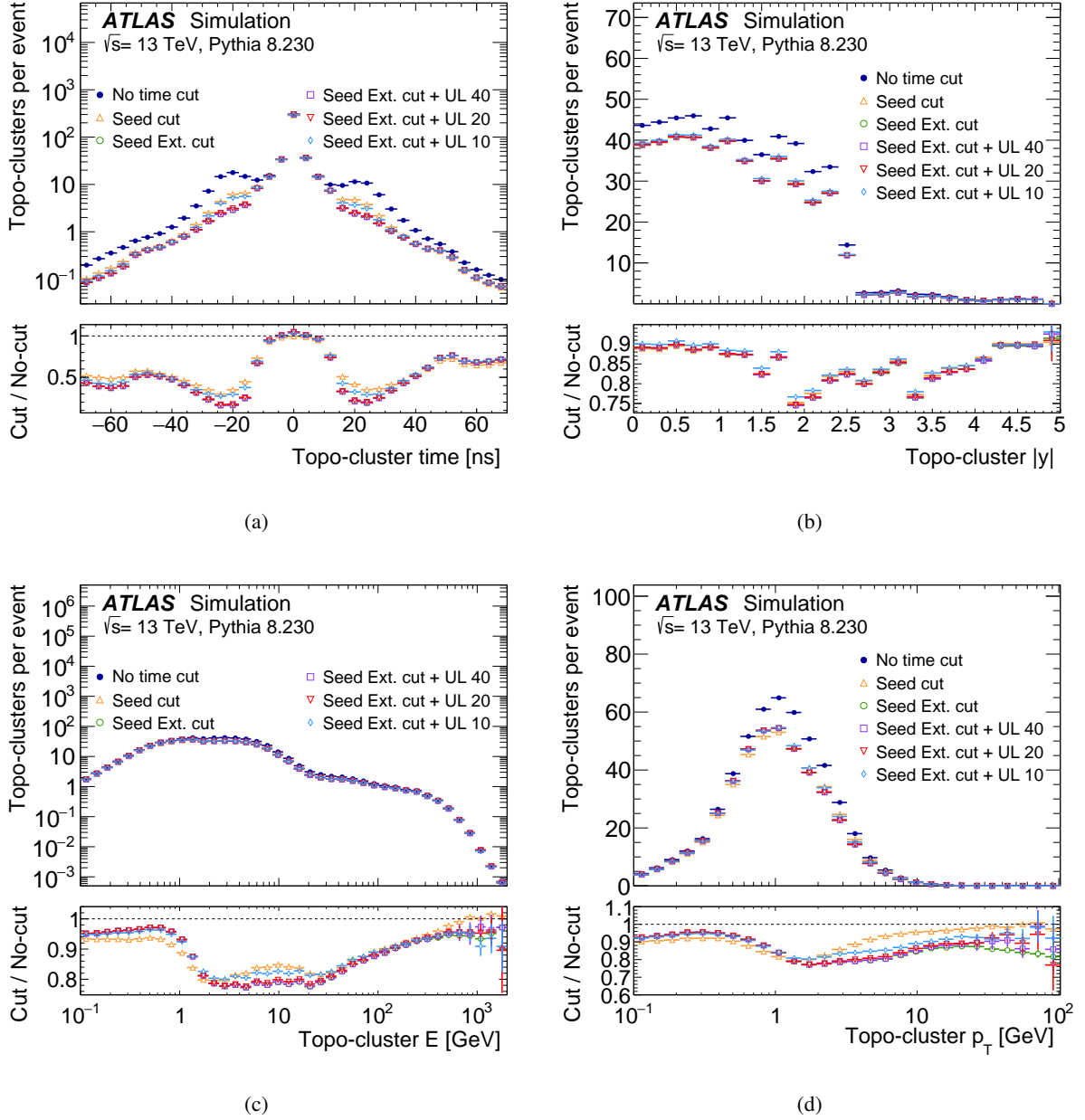


Figure 4: Kinematics of reconstructed topo-clusters, compared between no time cut, the Seed cut, the Seed Extended cut and the Seed Extended cut combined with multiple choices of Upper Limit: $X_{UL} = 40$, $X_{UL} = 20$ and $X_{UL} = 10$. Each kinematic moment of the topo-clusters is defined as the weighted average of the property of interest over the cells contributing to the cluster [3]. Topo-clusters with energy $E_{cl} > 0$ are shown. Plots are normalised to the total number of positive-energy topo-clusters per event. Uncertainties in the cut / no-cut ratios are computed by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty.

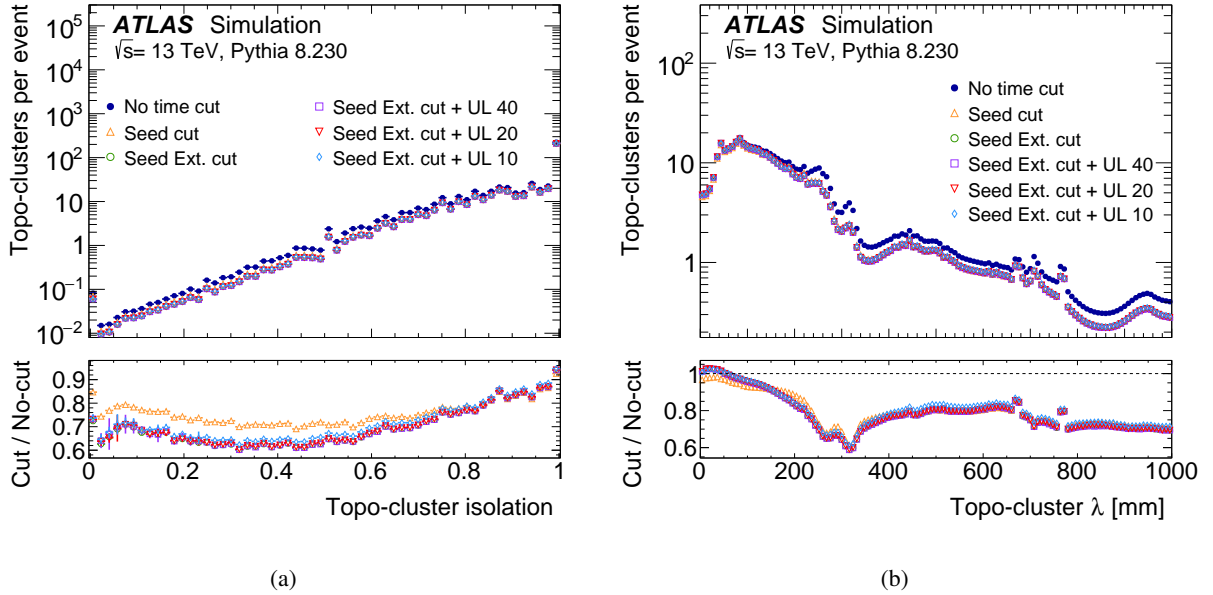


Figure 5: Topo-cluster (a) isolation and (b) distance λ from the calorimeter front face. A comparison between no time cut, the Seed cut, the Seed Extended cut and the Seed Extended cut combined with $X_{UL} = 40$, $X_{UL} = 20$ and $X_{UL} = 10$ is shown. Each kinematic moment of the topoclusters is defined as the weighted average of the property of interest over the cells contributing to the cluster [3]. Topoclusters with energy $E_{cl} > 0$ are shown. Plots are normalised to the total number of positive-energy topoclusters per event. Uncertainties in the cut / no-cut ratios are computed by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty.

studies are needed to further understand the interplay between these two effects. The distance λ of the cluster's centre of gravity from the calorimeter's front face [3] is shown in Figure 5(b). The region of $\lambda \lesssim 100$ mm is known to be dominated by EM clusters, while clusters located at $\lambda \gtrsim 400$ mm are predominantly hadronic [3]. In Figure 5(b) very little difference is visible in the region dominated by EM clusters while a relatively uniform reduction of about 20% is visible for predominantly hadronic clusters. The area between $\lambda \sim 100$ mm and $\lambda \sim 400$ mm corresponds to a mixture of EM and hadronic clusters. Here, peaks are visible in the spectrum of standard clusters, while they are almost entirely removed by the time cut. This effect hints that the timing cut may increase the separation between hadronic and EM showers in the presence of pile-up.

6.2 Performance on jets

Topoclusters serve as inputs to the reconstruction of many physics objects in ATLAS. This paper focuses on the effect of the time cut on jets reconstructed by applying the anti- k_r clustering algorithm with radius parameter $R = 0.4$ [41]. In ATLAS, jets are most often reconstructed from *particle-flow* objects [42, 43], which are built by combining topoclusters with matching inner-detector tracks. Despite it being the preferred method in ATLAS, particle flow adds a layer of complexity to the jet reconstruction: jets built directly from topoclusters provide a more straightforward way of evaluating the time cut's performance and are hence used in this paper. Preliminary studies show that the effect of the time cut on $R = 0.4$ particle-flow jets is about the same size as the effect on those built directly from topoclusters. Since the

ATLAS tracking is performed within the triggered bunch crossing, it is reasonable to expect that most out-of-time calorimeter signals would not match tracks, so the time cut can be expected to primarily affect objects classified as neutral by the particle-flow algorithm, which in turn consist only of topo-clusters, with no track-based correction.

6.2.1 Jet calibration

Jets are calibrated to correct for both the pile-up contribution and the detector response. A complete description of jet calibration in ATLAS can be found in Ref. [43]. The calibration procedure consists of three stages. First, a two-step pile-up correction is made. Then a MC-based calibration factor is applied to correct for the detector response (MC jet energy scale: *MC-JES*) and improve the jet resolution by reducing the calibration's dependence on additional jet properties (*Global Sequential Calibration*). Finally, an *in situ* correction is applied to account for discrepancies between data and MC events.

The first stage of the calibration procedure starts with an *area-based* correction in which the average pile-up contribution is subtracted from the jet's transverse momentum:

$$p_T^{\text{area}} = p_T - \rho \times A,$$

where p_T is the jet p_T , A is the jet catchment area and ρ is the median p_T density of the event. The value of ρ is calculated using jets reconstructed by applying the k_t algorithm [44, 45] to positive-energy topo-clusters with $|\eta| < 2$. A residual pile-up correction is then applied to account for residual dependence of the jet p_T on pile-up activity in the event. Two η -dependent correction terms are derived independently by fitting the dependence of the jet p_T on μ and the number of reconstructed primary vertices per event, N_{PV} . In general, μ provides an estimate of the amount of out-of-time pile-up in the event, while N_{PV} estimates the amount of in-time pile-up. The residual correction is then applied as:

$$p_T^{\text{residual}} = p_T^{\text{area}} - \alpha \times (N_{\text{PV}} - 1) - \beta \times \mu,$$

where α and β are the correction factors estimated from MC simulation.

The pile-up contribution to jets is expected to be affected by the time cut. Figure 6 shows the spectrum of the median p_T density ρ , as well as the dependence of ρ on μ . When the time cut is applied, ρ is found to be $\sim 20\%$ smaller overall, and also to increase more slowly as a function of μ . Since ρ is built from topo-clusters, the improved pile-up rejection of the time cut leads to smaller ρ values, decreasing the pile-up-dependent contribution to the jet energy. To account for possible remaining data–MC differences in the pile-up contribution to jets, not covered by the jet-area correction, the residual correction is recalculated after the application of the time cut. The standard procedure summarised above is repeated for both the Seed and Seed Extended cuts. Since the effect of the Upper Limit is smaller and limited to high-energy regions, the residual correction is not recalculated for any of the X_{UL} values, and the residual correction calculated for the Seed Extended cut is used.

6.2.2 Jet selection

Reconstructed jets are calibrated by applying the pile-up correction and the MC-JES calibration. They are then preselected, requiring them to lie within $|\eta| < 4.5$ and satisfy $p_T > 7$ GeV. Jets are also required to be isolated, i.e. there must be no other preselected jet satisfying $\Delta R(j_1, j_2) < 1.5 R_{\text{jet}}$, where j_1 and

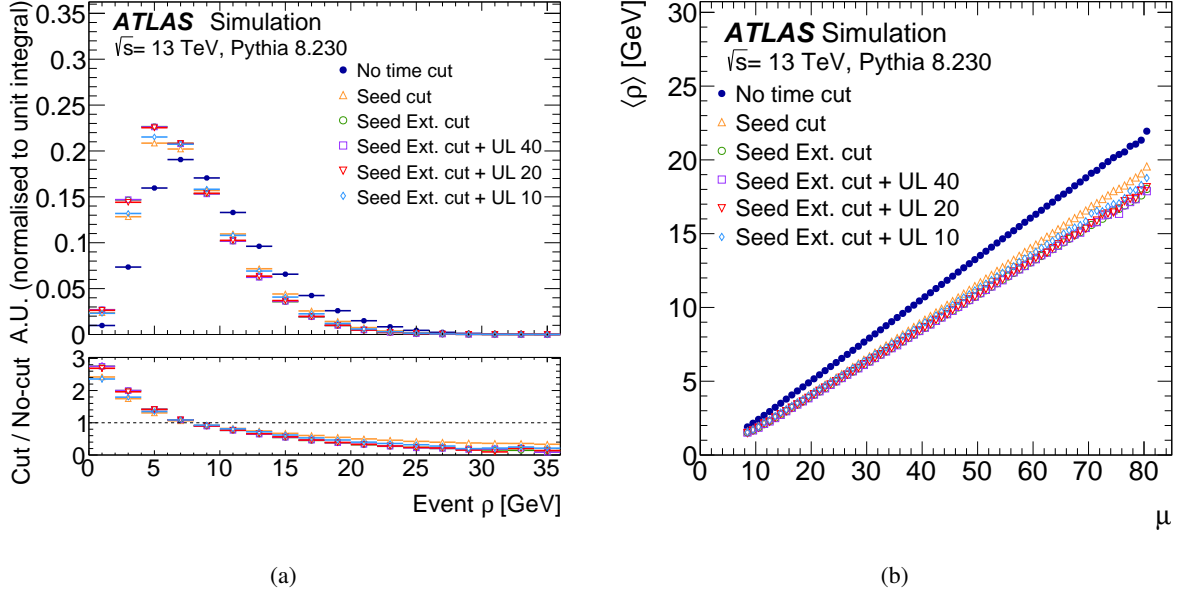


Figure 6: Topo-cluster median p_T density (ρ): comparison between the no time cut and the considered time cut options. (a): ρ spectrum. (b): average ρ in bins of the number of interactions per bunch-crossing (μ). Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty.

j_2 are any two reconstructed jets and R_{jet} is the jet radius parameter used in the anti- k_r algorithm (here $R_{\text{jet}} = 0.4$). Since this study focuses on both signal and pile-up jets, pile-up suppression requirements such as the one using the Jet Vertex Tagger [46] are not applied.

In order to distinguish jets produced by the hard-scattering from jets originating from pile-up, reconstructed jets are matched to *truth-level jets* in MC events. Truth-level jets are reconstructed by applying the anti- k_r algorithm ($R = 0.4$) to stable final-state particles, defined as having $c\tau > 10$ mm in the generator's event record, excluding muons and neutrinos. Three types of truth jets are distinguished. First, truth jets originating from the hard scattering (*HS-truth jet*) are obtained by clustering stable particles from the simulated hard-scattering event, excluding particles from pile-up interactions. Next, two types of pile-up truth jets are obtained by clustering stable particles from the minimum-bias events which are overlaid on the hard scattering in order to simulate pile-up [47]. Jets are clustered for each minimum-bias event individually and then grouped into two pile-up jet collections: jets from in-time overlay events are included in the in-time pile-up jet collection (*IT-truth jet* in the following), while those from out-of-time overlay events are included in the out-of-time pile-up jet collection (*OOT-truth jet* in the following). The latter are built from 32 bunch-crossings before the current one and 6 after [47].

Default cuts requiring $p_T > 10$ GeV and $p_T > 15$ GeV are applied to IT-truth jets and OOT-truth jets respectively. HS-truth jets are also preselected by requiring $p_T > 7$ GeV. All truth jets must satisfy $|\eta| < 5$. It should be noted that while a p_T match can be expected between HS-truth jets and reconstructed jets after the calibration is applied, this is less the case for OOT-truth jets: because of negative energy contributions, one might expect the reconstructed energy of an out-of-time pile-up jet to be lower than that of the truth jet. This is the reason for having a higher p_T selection for truth pile-up jets than for HS-truth jets. Truth-level

jets are also required to be isolated: there must be no other truth jet satisfying $\Delta R(j_1, j_2) < 2.5 R_{\text{jet}}$. Here j_1 and j_2 are any two truth jets, all three truth-jet categories being considered together when computing the isolation. Reconstructed jets are matched to truth jets via a geometrical matching requirement: a reconstructed jet j_r is said to match a truth jet j_t if their separation satisfies $\Delta R(j_r, j_t) < 0.3$. Reconstructed jets are first checked for matches with a HS-truth jet. If no such match is found, then matches with IT-truth and OOT-truth jets are attempted.

6.2.3 Jet kinematics

Figure 7 shows the calibrated transverse momentum distributions for the jet categories described in Section 6.2.2. Both the reconstructed jets matching HS-truth jets and those matching IT-truth jets show only a percent-level impact from the time cut. It should be pointed out that the time cut can affect the reconstructed jet kinematics, possibly resulting in bin-to-bin migrations in the spectra shown in this subsection. Energy (and p_T) variations can occur in either direction, since the cell energies included in the topo-clusters can be either positive or negative. Out-of-time cells often produce negative energy signals and this explains why the ratio plots in Figures 7(a) and 7(b) have values above unity in some places. The small p_T -dependent decrease in the fraction of jets matching HS-truth jets was found to be due to random matches. The fraction of jets that match OOT-truth jets, on the other hand, is drastically reduced by the Seed Extended cut, especially in the high p_T region. The Seed Extended cut reduces the multiplicity of out-of-time jets by $\sim 60\%$ at $p_T = 20$ GeV, and no out-of-time jets are found above ~ 45 GeV. The Seed cut, on the other hand, reduces the multiplicity of the out-of-time jets by only $\sim 30\%$ at 20 GeV and only $\sim 20\%$ at 50 GeV, indicating that the Seed Extended cut rejects out-of-time pile-up more effectively. Figure 7 also illustrates the effect of the different Upper Limit values. At $p_T \sim 45$ GeV, the reduction in out-of-time jet multiplicity is $\sim 90\%$ for an Upper Limit of 40, $\sim 70\%$ for an Upper Limit of 20, and $\sim 50\%$ for an Upper Limit of 10, confirming that lower Upper Limit values allow more pile-up in the event.

Figure 8 shows an analogous comparison for the jet rapidity. In order to limit the kinematics plots to the phase space of interest in physics analyses, an additional requirement of $p_T > 20$ GeV is applied to the jets. Overall, all cuts except the Seed cut and Seed Extended plus Upper Limit 10 cut provide a $\sim 40\%$ suppression of out-of-time jets within $|y| < 1.5$. The effect increases to a 50% – 60% reduction for $1.5 < |y| < 3.2$ and then decreases until it has negligible impact for $|y| > 4$. From these observations, it can be concluded that the Seed Extended time cut is preferable to the Seed cut because it is more effective in suppressing pile-up. It can also be seen that the Upper Limit $X_{\text{UL}} = 10$ allows too much pile-up to remain in the event, so the lowest practical Upper Limit among those tested is $X_{\text{UL}} = 20$.

6.2.4 Jet resolution

As discussed in Section 6.2.1, the time cut affects the jet calibration because of the different amount of pile-up remaining. To check for its impact on the performance, the jet energy resolution after calibration is compared for the different time cuts. The jet response is defined as the ratio of the calibrated energy of the reconstructed jet to the energy of the matching truth jet: $R = E_j^{\text{calibrated}} / E_j^{\text{truth}}$. Only those jets matched to a HS-truth jet are considered for this study. The response distribution is computed in bins of truth-jet p_T and rapidity ($|y|$), or truth-jet p_T and the event μ . For each bin, the response distribution is fitted with a Gaussian function. The jet resolution is then computed for each bin, as the ratio of the fitted standard deviation of the jet response R to its fitted mean value ($\sigma_R^{\text{fit}} / \langle R \rangle^{\text{fit}}$).

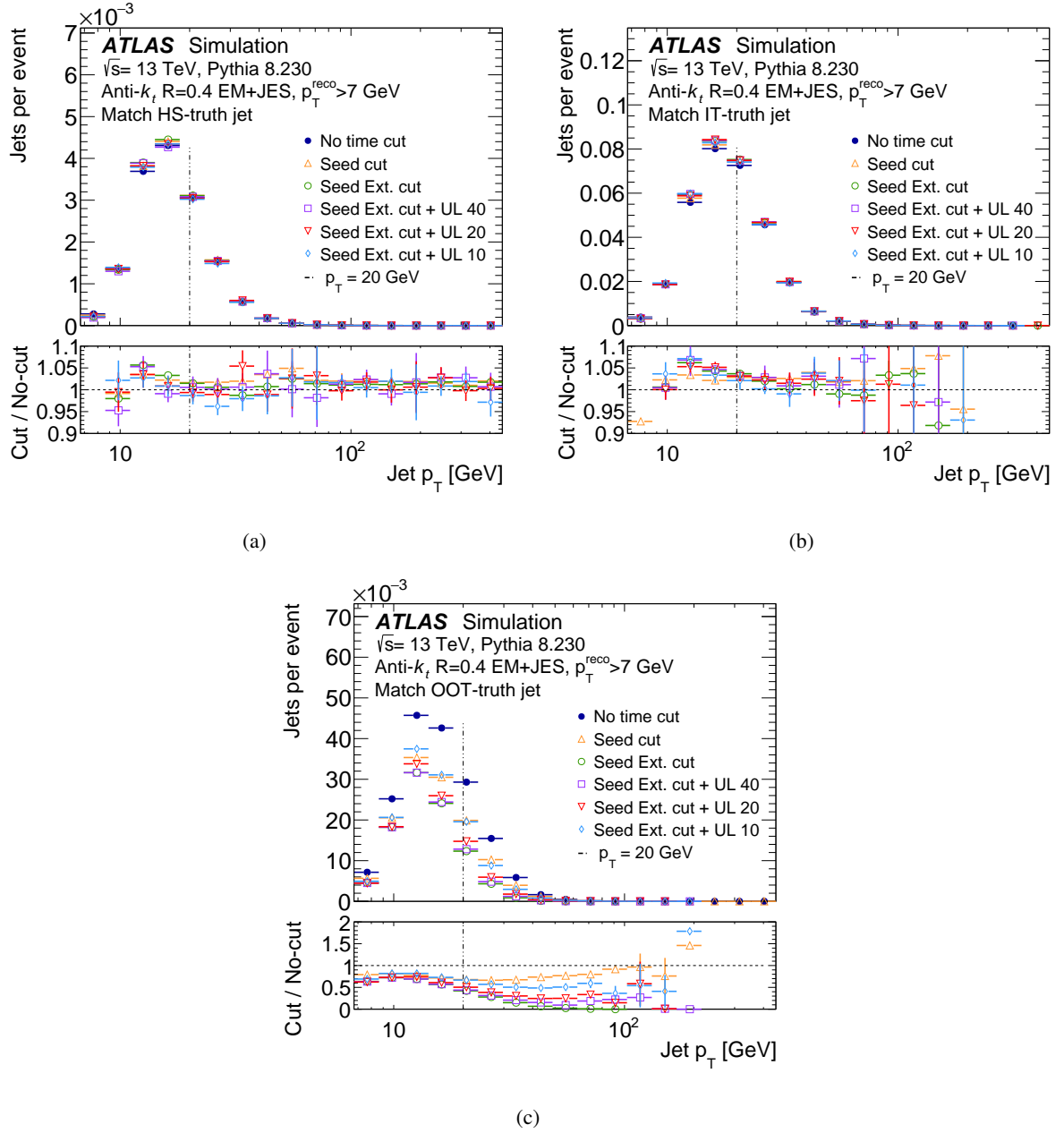
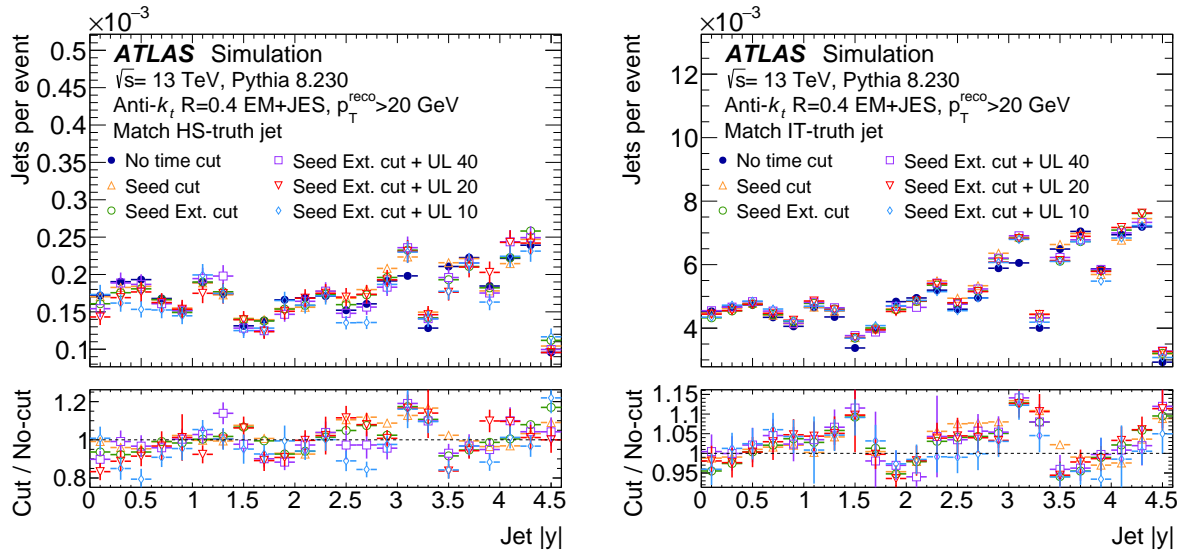
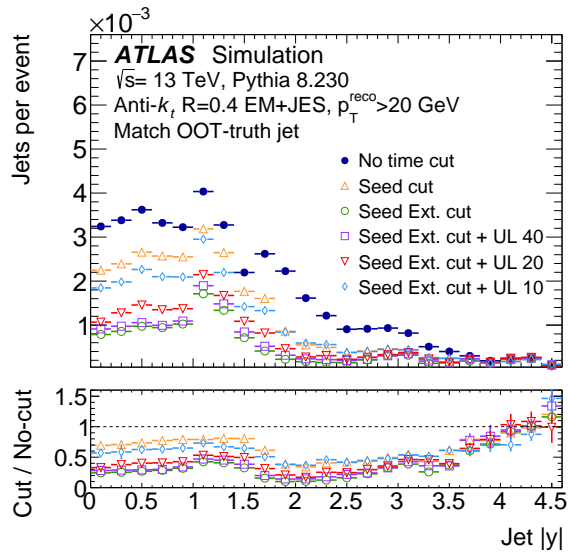


Figure 7: Comparison between jet p_T spectra obtained when the Seed, Seed Extended, or no time cut is used. The Seed Extended cut is also shown in combination with the Upper Limit for $X_{UL} = 40$, $X_{UL} = 20$, and $X_{UL} = 10$. (a): jets matching HS-truth jets. (b): jets matching IT-truth jets. (c): jets matching OOT-truth jets. Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty. The vertical dashed line represents the minimum jet p_T of 20 GeV normally required in ATLAS physics analyses.



(a)

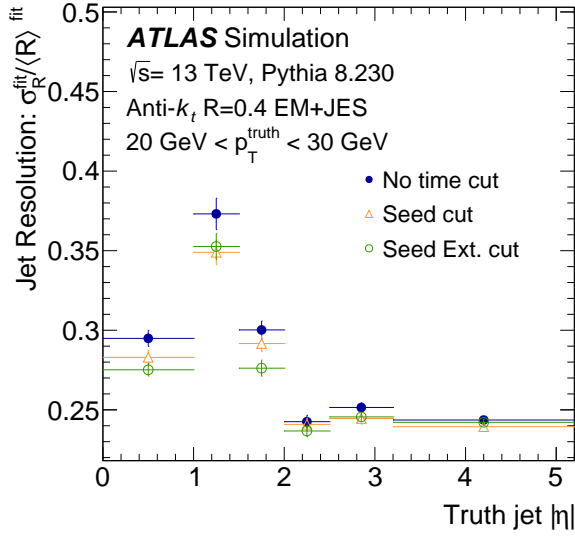
(b)



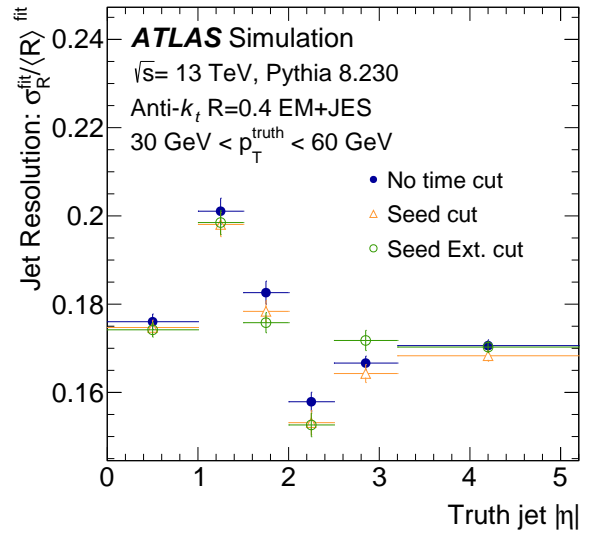
(c)

Figure 8: Comparison between jet rapidity spectra obtained when the Seed, Seed Extended, or no time cut is used. The Seed Extended cut is also shown in combination with the Upper Limit for $X_{UL} = 40$, $X_{UL} = 20$, and $X_{UL} = 10$. (a): jets matching HS-truth jets. (b): jets matching IT-truth jets. (c): jets matching OOT-truth jets. Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty.

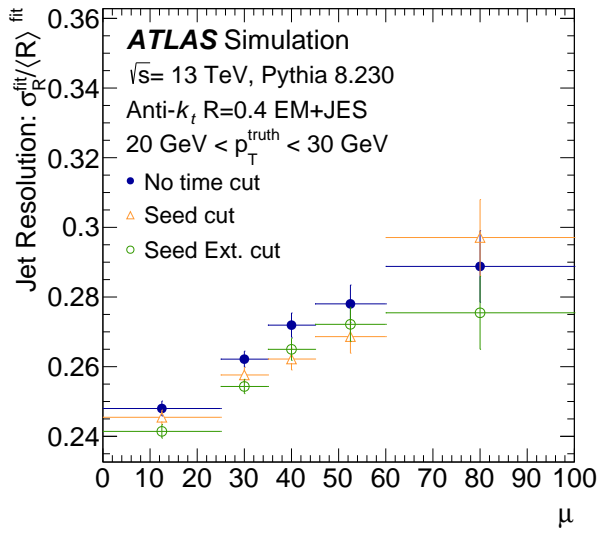
The estimated resolution is shown for two p_T bins in Figures 9(a) and 9(b) as a function of the jet rapidity and in Figures 9(c) and 9(d) as a function of μ . Some improvement in the jet resolution is visible, amounting to 5%–10% for $20 \text{ GeV} \leq p_T \leq 30 \text{ GeV}$, but no strong dependence on μ or $|y|$ is observed. A more accurate estimation of the jet resolution and calibration performance after the time cut is left for future studies, following the implementation of the algorithm in the standard ATLAS reconstruction chain.



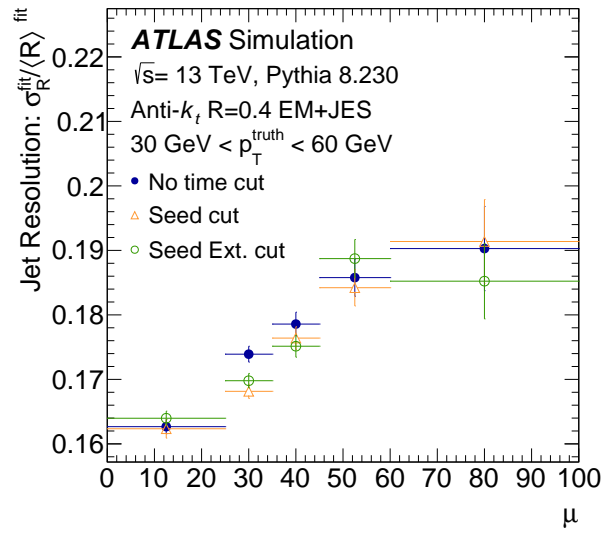
(a)



(b)



(c)



(d)

Figure 9: Jet energy resolution after calibration as a function of (a), (b) the jet $|\eta|$ and (c), (d) the number of interactions per bunch-crossing (μ). Results are shown in bins of the jet p_T : (a), (c) $20 \text{ GeV} < p_T^{\text{truth}} < 30 \text{ GeV}$; (b), (d) $30 \text{ GeV} < p_T^{\text{truth}} < 60 \text{ GeV}$. The resolution is defined as the ratio of the standard deviation and mean of a Gaussian function fitted to the response distribution.

7 Performance on ATLAS data

As a further probe of its performance, the time cut was tested on data collected by the ATLAS detector in 2017, corresponding to an integrated luminosity of $\mathcal{L} = 71 \pm 2 \text{ pb}^{-1}$ after data quality requirements [23]. Single-jet triggers were used to collect the data, requiring at least one jet with sufficiently large p_T . Multiple triggers, listed in Table 1 and implementing different minimum p_T requirements, were studied independently. Relatively low minimum p_T requirements are considered, as the time cut is known to mostly effect low-energy jets. Triggers with different L1 seeds are also considered: events passing triggers seeded by random L1 triggers are expected to have a larger contamination from out-of-time pile-up than those triggered by single jet at L1. Since all the triggers in Table 1 are prescaled, the data were reweighted event-by-event by the respective trigger prescale factor.

Table 1: Triggers used for data selection.

| Trigger name | L1 trigger seed | Trigger minimum p_T (≥ 1 jet with $p_T \geq X$) | Offline minimum p_T (≥ 1 jet with $p_T \geq X$) |
|----------------------|-------------------------------------|---|---|
| HLT_j15 | Random for filled bunches | 15 GeV | 20 GeV |
| HLT_j25 | Random for filled bunches | 25 GeV | 30 GeV |
| HLT_j45 | ≥ 1 jet with $p_T \geq 15$ GeV | 45 GeV | 50 GeV |
| HLT_j45_L1RD0_FILLED | Random for filled bunches | 45 GeV | 50 GeV |

Data and simulated events are required to contain at least one primary vertex; the one with the highest sum of squared track p_T is considered in this section. In addition, at least one jet must satisfy a minimum p_T requirement, as listed in Table 1. This requirement is needed to restrict selected events to a phase space where the trigger efficiency is within $\sim 20\%$ of its plateau value. A fully efficient trigger is not necessary for this study, as the time cut is not applied at trigger level and a smaller trigger efficiency would not affect the comparison between the cut and no-cut scenarios. In order to properly compare data and MC events, jets are calibrated using the full calibration chain, including the *in situ* correction.⁶ All jets are required to have $p_T > 20$ GeV. The Seed Extended cut and Seed Extended plus $X_{UL} = 20$ cut are compared with no-cut spectra for data and MC events. Figures 10, 11 and 12 show different jet kinematic properties, while topo-cluster properties are shown in Figures 13, 14 and 15. Each histogram shows the distributions for events selected by a different trigger. Uncertainties shown for MC distributions are a combination of the MC statistical uncertainty and the luminosity uncertainty. The latter is based on the 2.4% luminosity uncertainty in 2017 [4]. Uncertainties shown for data are statistical only.

The data and MC distributions are not in perfect agreement, and exhibit relatively uniform differences of about 30%–40%. This is to be expected as the MC sample used is a dijet simulation at leading order: even though the hadronisation and jet characteristics are well described by PYTHIA 8.230, the overall MC normalisation is likely to be too low due to the missing higher orders. Moreover, low- p_T events are dominated by soft QCD radiation, which is known to be difficult to model. A complete jet cross-section measurement would have to use an improved simulation and consider additional correction factors. Since this study only seeks to verify that the time cut’s behaviour in data reflects what is observed in MC simulation, this level of agreement is considered acceptable.

The effect of the time cut depends on the trigger. Events passing HLT_j45, which has both a higher p_T threshold and a non-random L1 seed, are not affected at all by the time cut. The numbers of events passing

⁶ As in Section 6, the residual pile-up correction is calculated after applying the time cut.

the other three triggers exhibit a consistent reduction of $\sim 20\%$ for values of the leading-jet p_T between 30 and 50 GeV (Figure 10). Effects of similar size can be seen throughout the tested kinematic spectra.

Except for the HLT_j45 trigger, the time cut tends to have a more pronounced effect on data than on MC events, the difference being of $\mathcal{O}(10\%)$ in most of the phase space. This is most likely due to the simulation not accounting for events consisting exclusively of pile-up. Such events are expected to be present in non-negligible numbers for the other three triggers, which have relatively low p_T thresholds and are seeded by random events. In addition, the MC simulation does not reproduce bunch-to-bunch luminosity fluctuations, which have a more significant effect on the reconstruction of out-of-time signals [3].

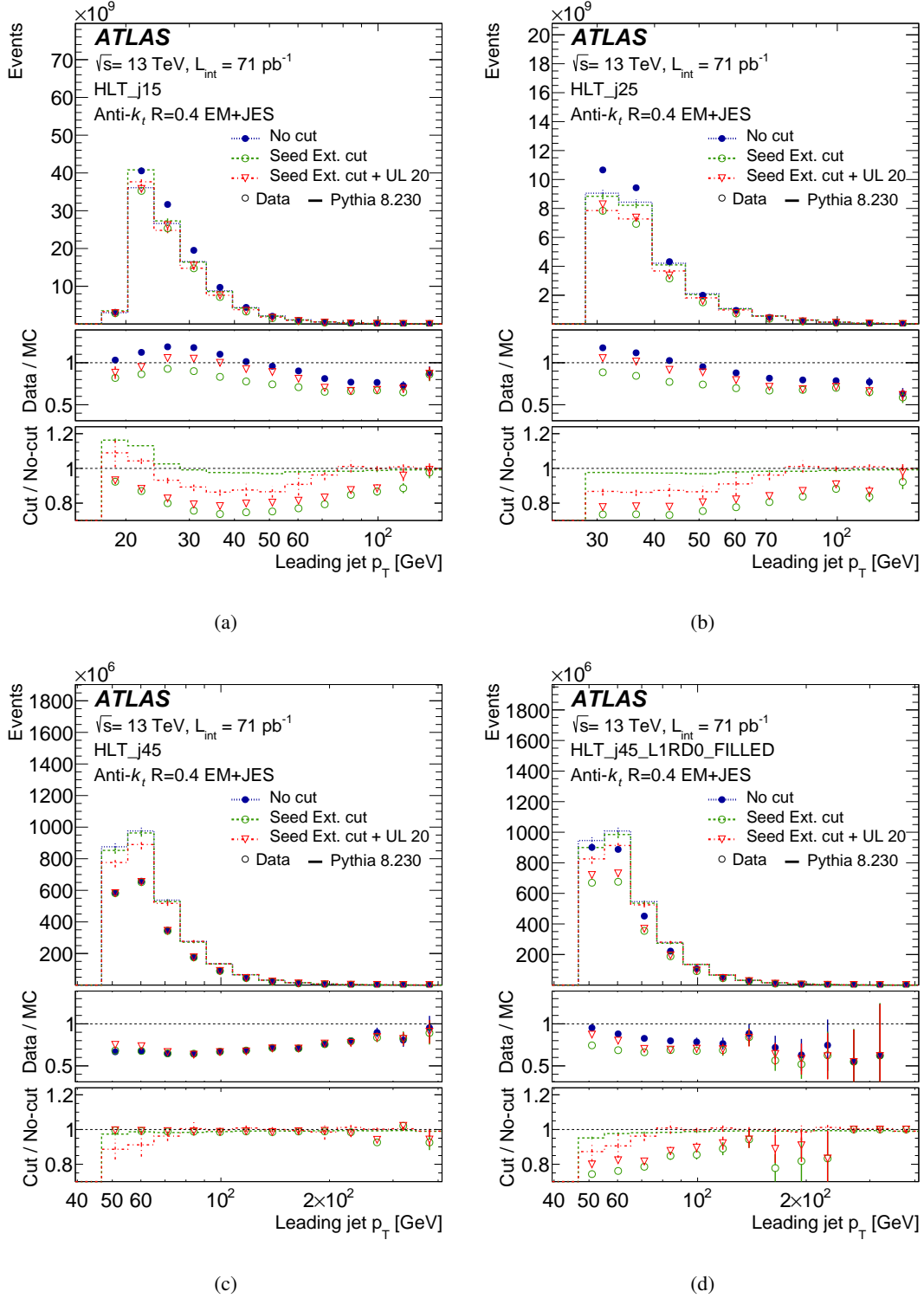


Figure 10: Leading jet p_T spectrum in data and MC selected multi-jet events. Four triggers are compared: (a) HLT_j15, (b) HLT_j25, (c) HLT_j45, and (d) HLT_j45_L1RD0_FILLED. The error bars convey the statistical and luminosity uncertainties. Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty. The luminosity uncertainty does not apply to the cut / no-cut ratio. In the plots, lines represent MC events and markers represent data, different colours and styles represent different cuts.

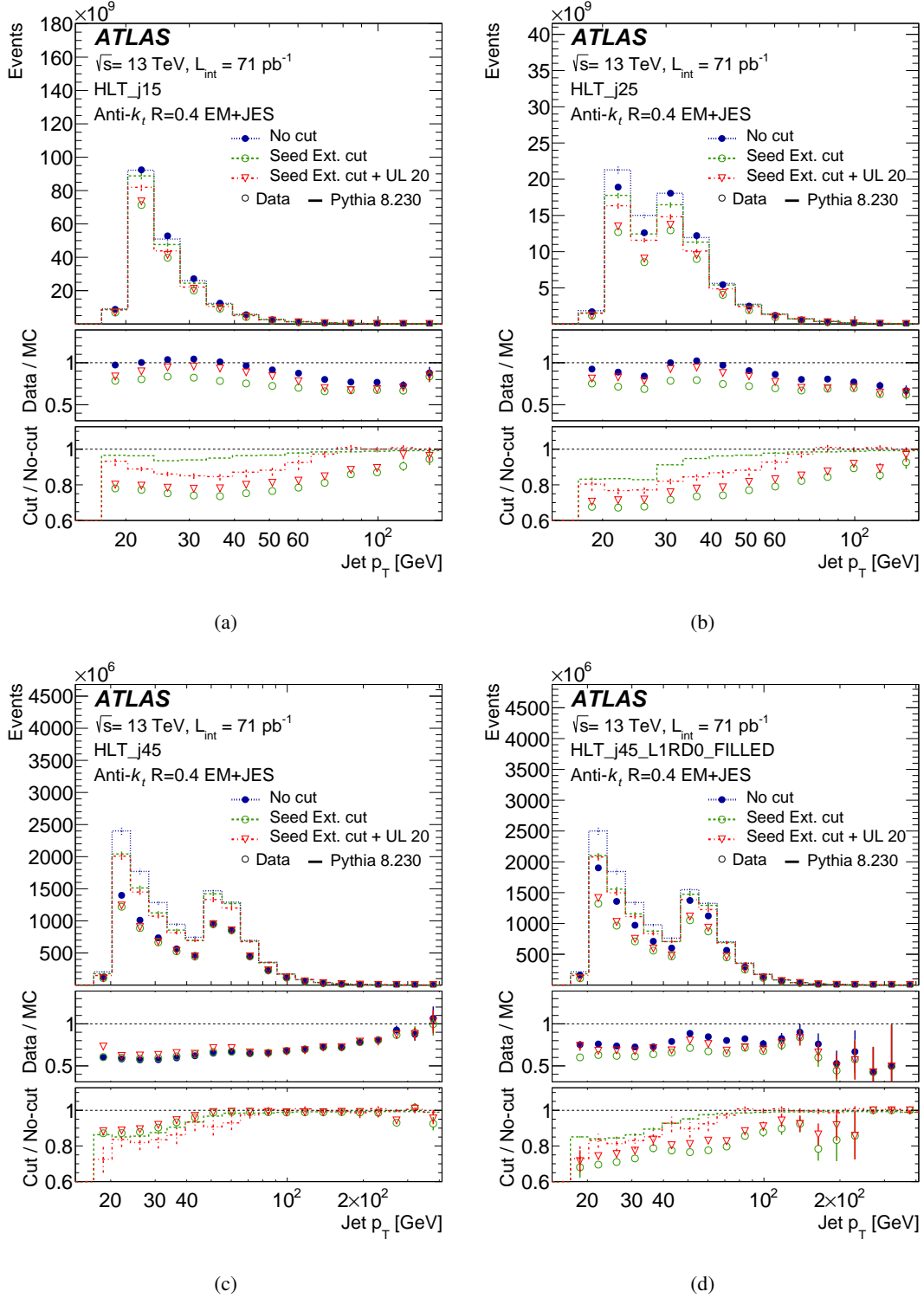


Figure 11: Jet p_T spectrum in data and MC selected multi-jet events. All jets passing the minimum p_T requirement ($p_T > 20 \text{ GeV}$) are shown. Four triggers are compared: (a) HLT_j15, (b) HLT_j25, (c) HLT_j45, and (d) HLT_j45_L1RD0_FILLED. The error bars convey the statistical and luminosity uncertainties. Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty. The luminosity uncertainty does not apply to the cut / no-cut ratio. In the plots, lines represent MC events and markers represent data, different colours and styles represent different cuts.

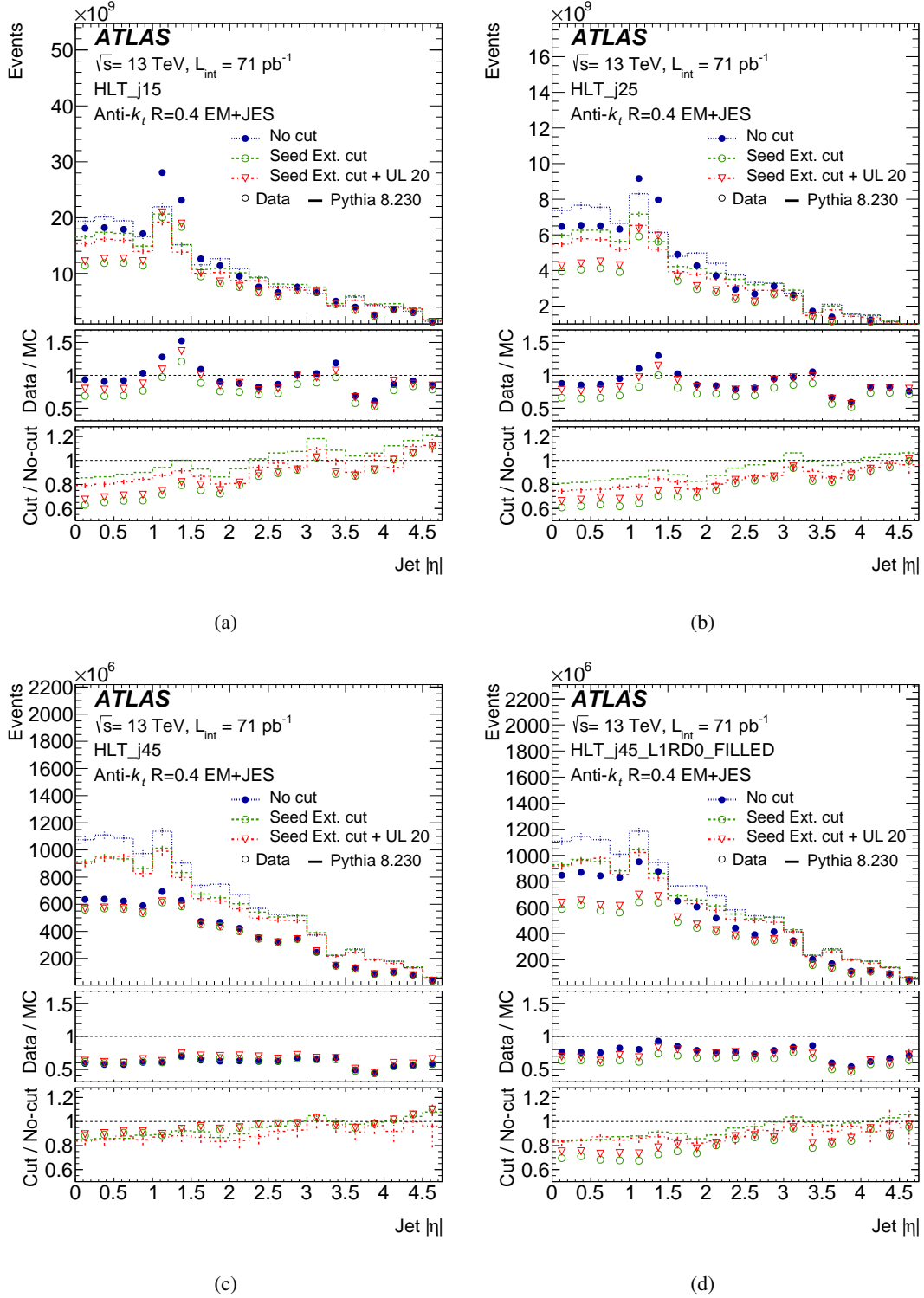


Figure 12: Jet $|\eta|$ spectrum in data and MC selected multi-jet events. All jets passing the minimum p_T requirement ($p_T > 20$ GeV) are shown. Four triggers are compared: (a) HLT_j15, (b) HLT_j25, (c) HLT_j45, and (d) HLT_j45_L1RD0_FILLED. The error bars convey the statistical and luminosity uncertainties. Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty. The luminosity uncertainty does not apply to the cut / no-cut ratio. In the plots, lines represent MC events and markers represent data, different colours and styles represent different cuts.

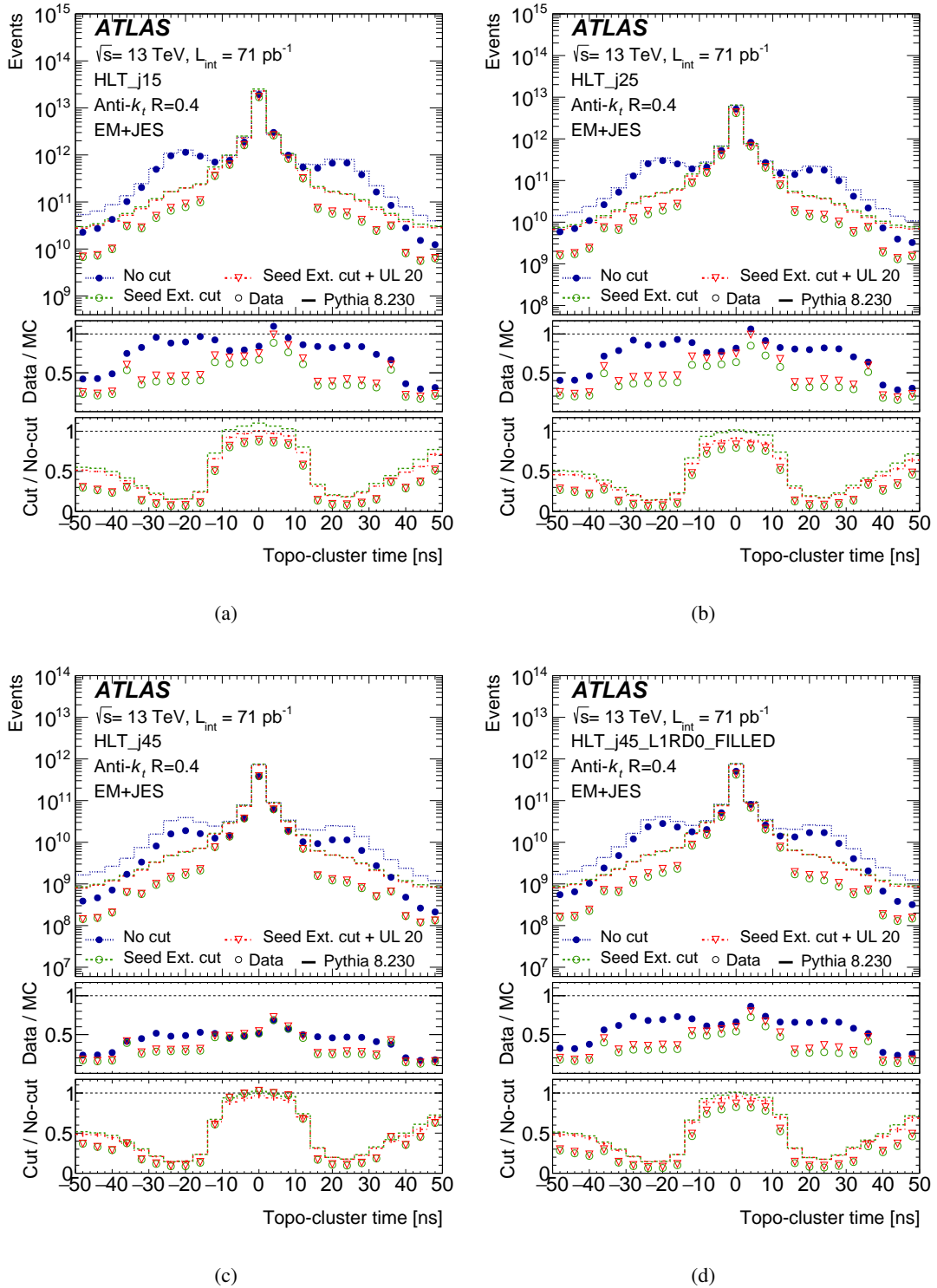


Figure 13: Topo-cluster time spectrum in data and MC selected multi-jet events. Four triggers are compared: (a) HLT_j15, (b) HLT_j25, (c) HLT_j45, and (d) HLT_j45_L1RD0_FILLED. The error bars convey the statistical and luminosity uncertainties. Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty. The luminosity uncertainty does not apply to the cut / no-cut ratio. In the plots, lines represent MC events and markers represent data, different colours and styles represent different cuts.

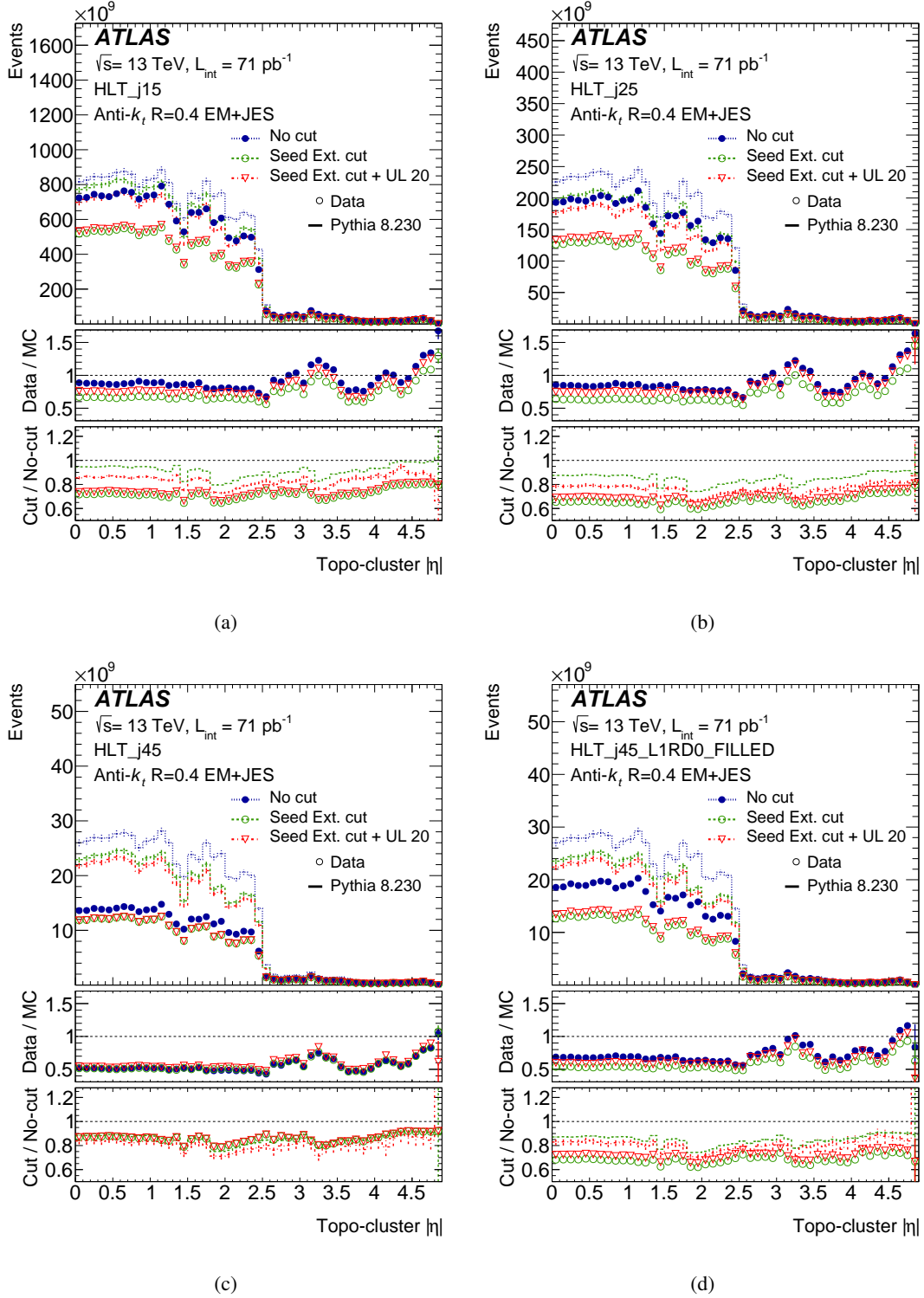


Figure 14: Topo-cluster $|\eta|$ spectrum in data and MC selected multi-jet events. Four triggers are compared: (a) HLT_j15, (b) HLT_j25, (c) HLT_j45, and (d) HLT_j45_L1RD0_FILLED. The error bars convey the statistical and luminosity uncertainties. Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and recomputing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty. The luminosity uncertainty does not apply to the cut / no-cut ratio. In the plots, lines represent MC events and markers represent data, different colours and styles represent different cuts.

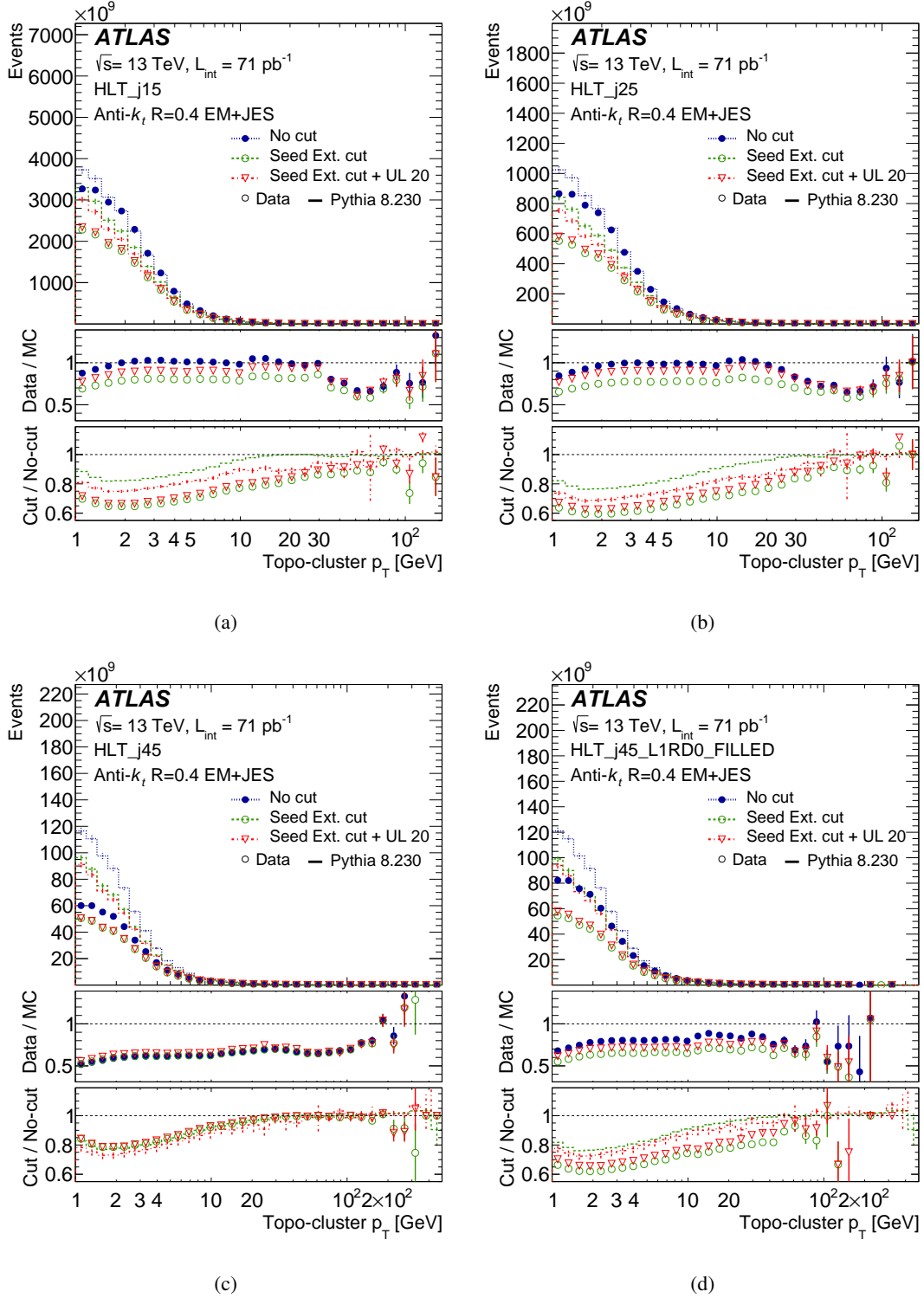


Figure 15: Topo-cluster p_T spectrum in data and MC selected multi-jet events. Four triggers are compared: (a) HLT_j15, (b) HLT_j25, (c) HLT_j45, and (d) HLT_j45_L1RDO_FILLED. The error bars convey the statistical and luminosity uncertainties. Uncertainties in the cut / no-cut ratios are obtained by splitting the available sample into subsamples and re-computing the cut / no-cut ratio for each subsample. The standard deviation of the distribution of the ratio is used to estimate the ratio's uncertainty. The luminosity uncertainty does not apply to the cut / no-cut ratio. In the plots, lines represent MC events and markers represent data, different colours and styles represent different cuts.

7.1 Checks for time cut inefficiencies

The time cut relies on the accurate measurement of signal timing provided by the calorimeters. The precision of the calorimeter timing measurement is guaranteed by periodic realignment and constant monitoring, as discussed in Section 2.1. As a further safety check, the robustness of the cut against a local time miscalibration was tested. The events selected for this check were those in which one calorimeter channel had been flagged as producing fake out-of-time signals due to cross-talk. The effect of the time cut on particles hitting the affected region was then compared with the effect of the time cut on particles hitting unaffected regions of the same subdetector.

The affected channel was located in the second layer of the LAr EM endcap calorimeter, on the C-side (along the negative z -axis). As an example, Figure 16 shows a combination of six events in which an electron from a $Z \rightarrow ee$ decay hit the affected area. Events were reconstructed with no cut, the Seed Extended cut and the Seed Extended plus $X_{UL} = 20$ cut. These events were identified first by applying a standard $Z \rightarrow ee$ event selection. Events are required to pass photon triggers analogous to those used in diphoton resonance searches [48, 49] and to contain at least two electrons with $p_T > 10$ GeV and $|\eta| < 2.47$ (excluding the transition region $1.37 < |\eta| < 1.52$ between the LAr EM barrel and endcap calorimeters), satisfying the Medium identification selection and FCLoose isolation criteria [50]. The reconstructed track matched to each electron candidate must be consistent with having originated from the primary vertex: its longitudinal impact parameter z_0 and transverse impact parameter d_0 must satisfy $|z_0 \cdot \sin \theta| < 0.5$ mm and $|d_0|/\sigma_{d_0} < 5$ mm respectively. Furthermore, the dielectron pair is required to have an invariant mass m_{ee} satisfying $68 \text{ GeV} < m_{ee} < 108 \text{ GeV}$ and a pseudorapidity separation $|\Delta\eta(e_1, e_2)| > 0.1$. Electrons hitting the affected channel were then flagged by requiring that the most energetic cell in the second layer of the EM calorimeter has high enough energy ($E_{\text{cell}}^{\text{max}} > 5$ GeV) and its measured time satisfies $|t_{\text{cell}}^{\text{max}}| > 12.5$ ns. Figure 16 shows that the Seed Extended time cut does reject the energy from one spot in the centre of the topo-cluster. Despite this, the overall topo-cluster is not lost. Moreover, the addition of the Upper Limit restores most of the lost energy, so that no cell is entirely missing from the electron topo-cluster.

To better quantify the effect, the average $\eta \times \phi$ location of flagged $Z \rightarrow ee$ electrons was used to define a 3×3 -cell *Test Region* containing the affected channel in the C-side LAr EM endcap’s second layer. The $\eta \times \phi$ ranges corresponding to this area are listed in Table 2. In order to compare the time cut’s effect in the Test Region with the normal time-cut behaviour, three *Control Regions* were also defined by inverting the sign of either η , ϕ or both in the Test Region’s definition. The Control Regions’ $\eta \times \phi$ boundaries are also listed in Table 2.

Table 2: Definition of Test and Control Regions.

| | η range | ϕ range |
|------------------|------------------|--------------------|
| Test Region | $[-1.9, -1.825]$ | $[-0.638, -0.565]$ |
| Control Region 1 | $[+1.9, +1.825]$ | $[-0.638, -0.565]$ |
| Control Region 2 | $[-1.9, -1.825]$ | $[+0.638, +0.565]$ |
| Control Region 3 | $[+1.9, +1.825]$ | $[+0.638, +0.565]$ |

Events from one Luminosity Block of 2018 data, during which the Test Region was known to be affected by cross-talk, were used to test the time cut’s behaviour. Rather than restricting the set of events to

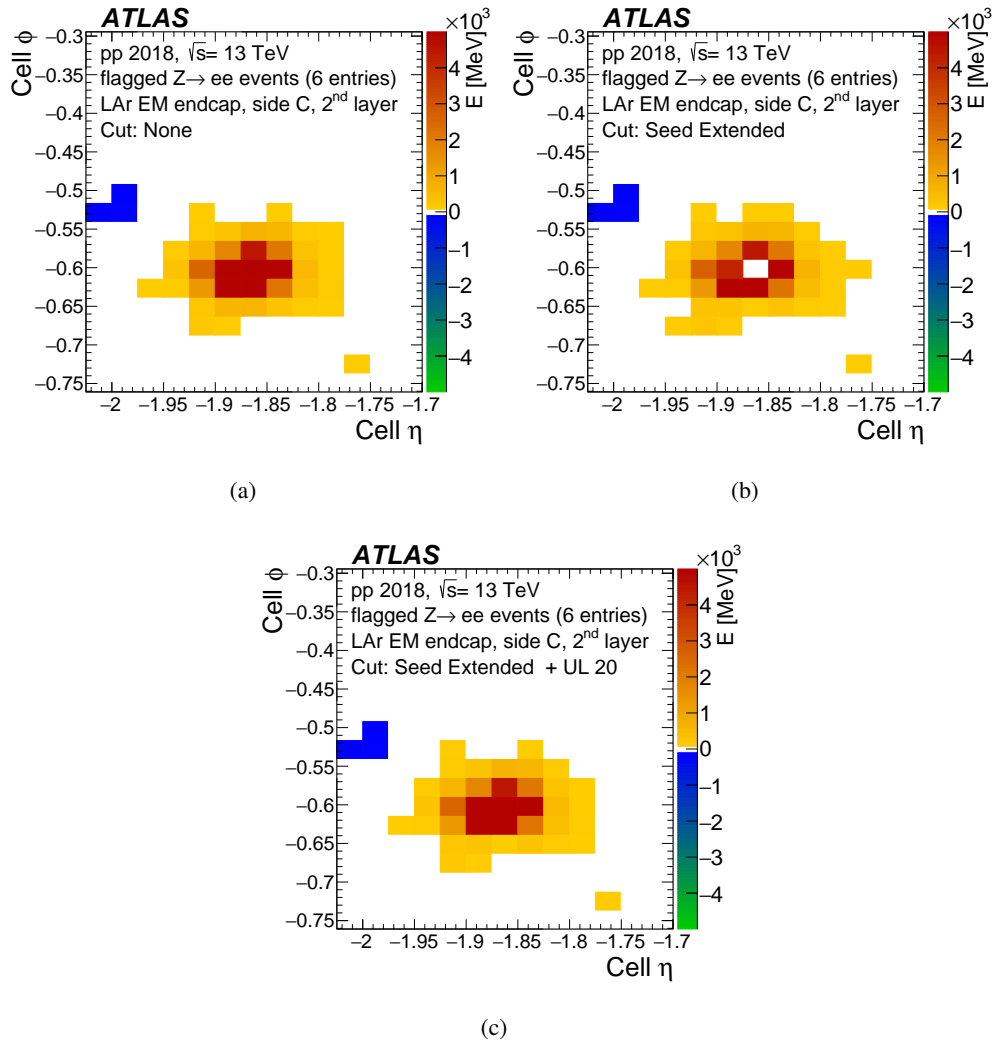


Figure 16: The $\eta \times \phi$ distribution of cells belonging to topo-clusters in the second layer of the LAr EM endcap calorimeter, side C. The colour scale represents the recorded energy. Six events, in which $Z \rightarrow ee$ electrons hit the affected area, are averaged. The three panels show results with (a) no time cut, (b) the Seed Extended cut and (c) the Seed Extended plus $X_{UL} = 20$ cut.

those passing a $Z \rightarrow ee$ selection, a more generic event selection is applied: for each region, events are considered if at least one cell of that region has a recorded transverse momentum of $p_T > 10$ GeV.

Four variables of interest were considered. The cell occupancy N^{cells} and the cells' total energy E^{cells} are respectively defined as the number and total energy of cells in a given region that are part of a topo-cluster: they quantify whether single cells are lost due to the miscalibrated channel. The topo-cluster occupancy N^{clus} and topo-cluster total energy E^{clus} are the number and total energy of reconstructed topo-clusters that share at least one cell with a given region: they provide a way to evaluate whether entire topo-clusters are lost or their energy is significantly impacted by the time cut. In order to better quantify the effect of the time cut, the fraction $\mathcal{F}_{\text{unchanged}}$ of events for which a given variable in a given region is unchanged by the cut is computed. For each event, N^{cells} (N^{clus}) is considered unchanged if the number of cells in the region (topo-clusters overlapping with the region) remains the same when applying the time cut. The total energies E^{cells} and E^{clus} are considered unchanged if the difference between the total energies computed with and without the time cut ($E_{\text{cut}}^{\text{cells(clus)}}$ and $E_{\text{no cut}}^{\text{cells(clus)}}$ respectively) satisfies:

$$|E_{\text{cut}}^{\text{cells(clus)}} - E_{\text{no cut}}^{\text{cells(clus)}}| < \sigma(E_{\text{no cut}}^{\text{cells(clus)}}),$$

where $\sigma(E_{\text{no cut}}^{\text{cells(clus)}})$ is the expected calorimeter resolution computed for either the cells' total energy (E^{cells}) or total cluster energy (E^{clus}) without any cut. The resolution is taken from Ref. [1] to be:

$$\frac{\sigma(E_{\text{no cut}}^{\text{cells(clus)}})}{E_{\text{no cut}}^{\text{cells(clus)}}} = \frac{10\%}{\sqrt{|E_{\text{no cut}}^{\text{cells(clus)}}|}} \oplus 0.7\%.$$

Figures 17 and 18 show $\mathcal{F}_{\text{unchanged}}$ for the four variables of interest. Figure 17 is indicative of the time cut's impact on the cell content for clusters in the Test Region compared to that for clusters in the Control Regions. The fraction of events with unchanged N^{cells} is shown in Figures 17(a) and 17(b): for the Seed Extended cut, $\mathcal{F}_{\text{unchanged}}$ in the Test Region is smaller than the average of the control values by slightly more than 1σ . Even though the phenomenon has a small significance due to the limited sample size, the effect of the time cut can be seen to disappear when the same events are processed with the Seed Extended plus $X_{\text{UL}} = 20$ cut, confirming that cell losses are cured by the Upper Limit. The same behaviour is visible for E^{cells} (Figures 17(c) and 17(d)), indicating that the loss of a cell, when it occurs, causes a significant energy variation, while no decrease in E^{cells} can be seen in the case of the Seed Extended plus $X_{\text{UL}} = 20$ cut. Figure 18 shows the fraction of events in which the topo-cluster occupancy and total energy are unchanged, thus investigating the effect of the time cut on topo-clusters as a whole, when they overlap the Test or Control Regions. The topo-cluster occupancy (Figures 18(a) and 18(b)) remains consistent with the control values even for the Seed Extended cut, confirming that the signal topo-cluster is not entirely lost. The topo-cluster total energy E^{clus} is impacted by the Seed Extended cut (Figure 18(c)), reflecting the loss of a cell in a way similar to that causing losses in the cells' total energy, while also being restored to its original value by applying the Upper Limit (Figure 18(d)).

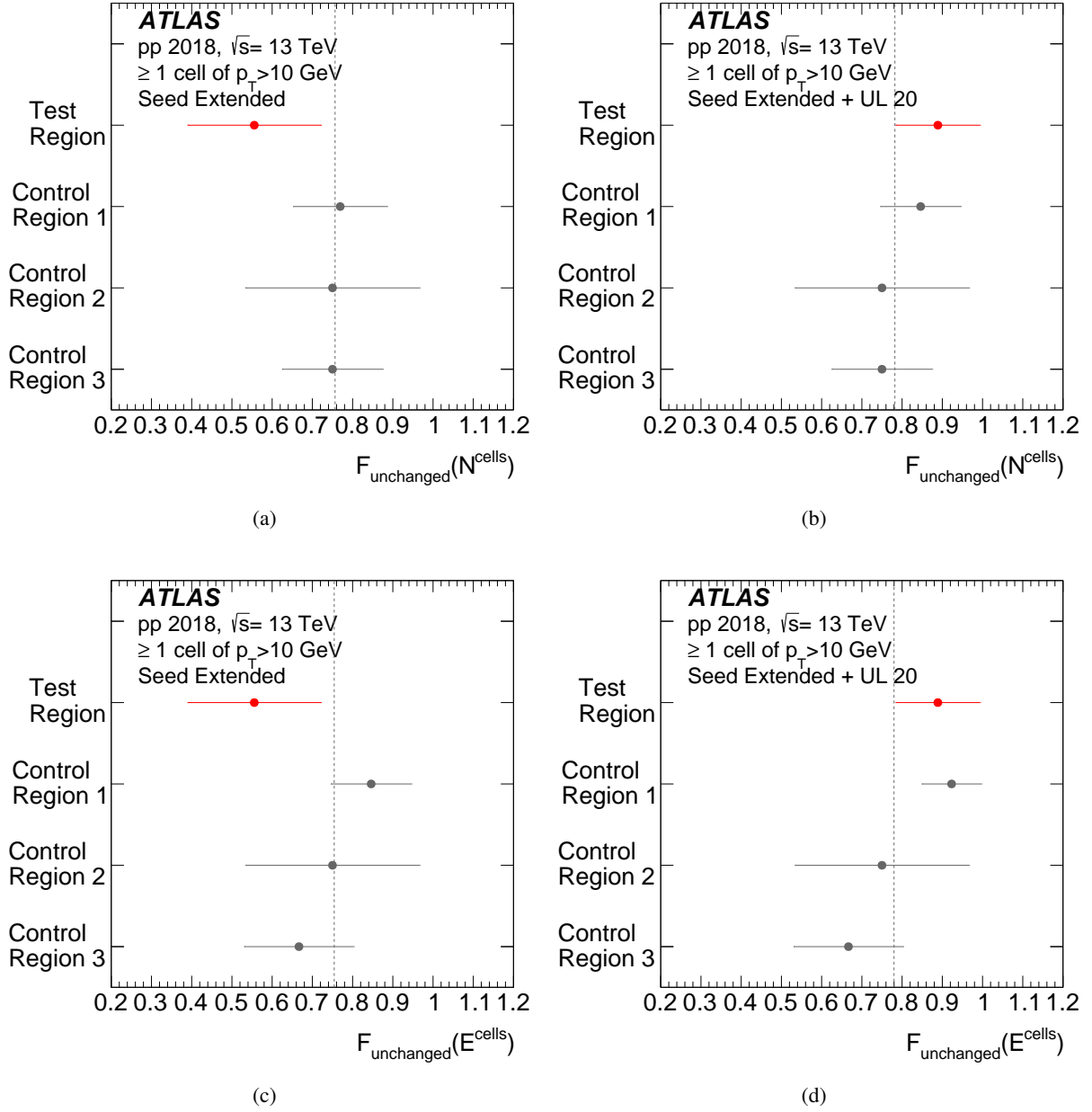


Figure 17: Comparison between the Test Region and three Control Regions for (a) the fraction of events in which the cell occupancy N^{cells} is left unchanged by the Seed Extended cut, (b) the fraction of events in which N^{cells} is left unchanged by the Seed Extended plus $X_{\text{UL}} = 20$ cut, (c) the fraction of events in which the cells' total energy E^{cells} is left unchanged by the Seed Extended cut, and (d) the fraction of events in which E^{cells} is left unchanged by the Seed Extended plus $X_{\text{UL}} = 20$ cut. The cell occupancy (cells' total energy) in a given region is defined as the number (energy sum) of cells in the region that are included in a topo-cluster. Uncertainties on the event fractions are computed as the standard deviation of a binomial distribution. The dashed line represents the average of the three Control Regions.

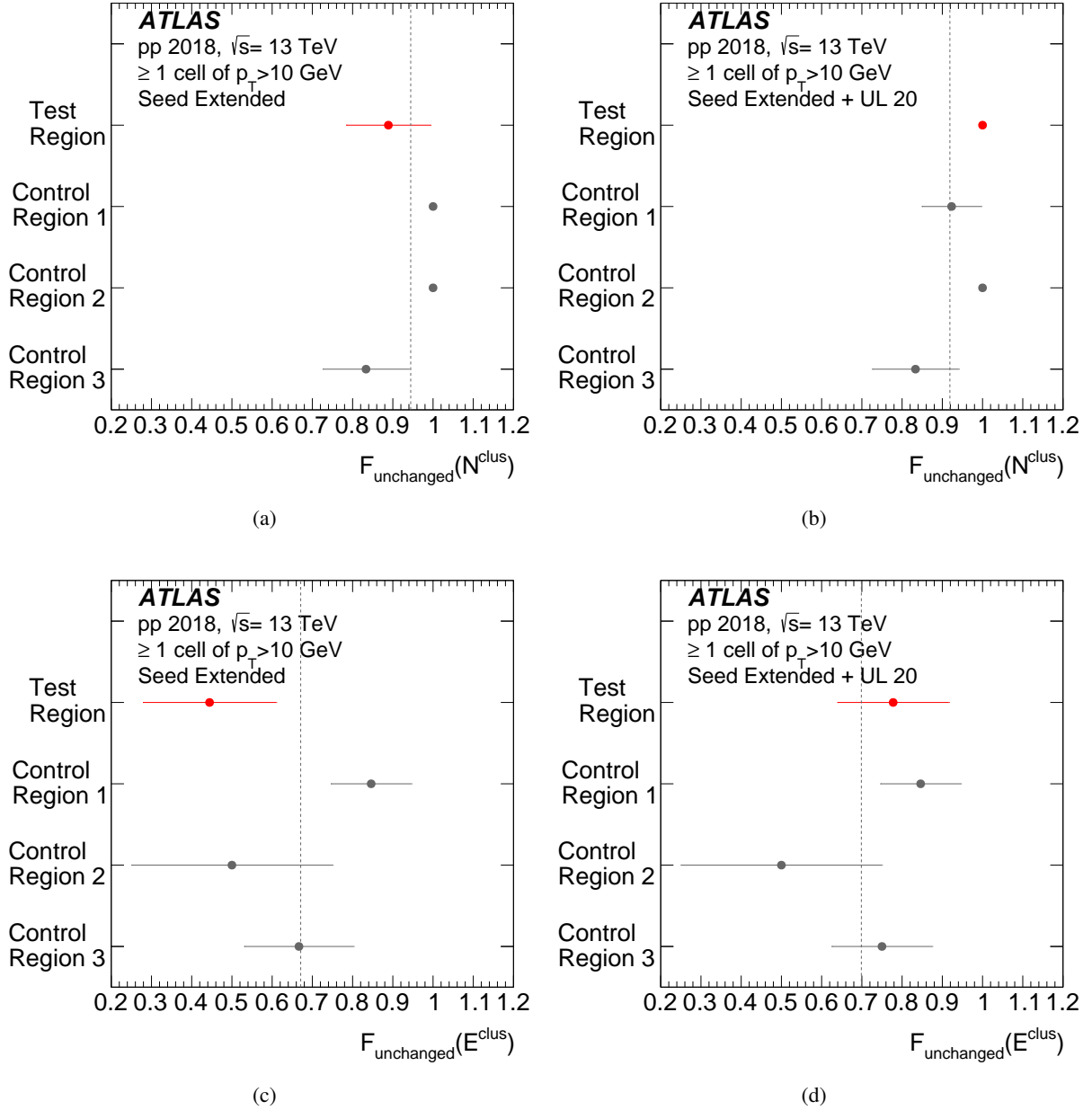


Figure 18: Comparison between the Test Region and three Control Regions for (a) the fraction of events in which the topo-cluster occupancy N^{clus} is left unchanged by the Seed Extended cut, (b) the fraction of events in which N^{clus} is left unchanged by the Seed Extended plus $X_{\text{UL}} = 20$ cut, (c) the fraction of events in which the topo-cluster total energy E^{clus} is left unchanged by the Seed Extended cut, and (d) the fraction of events in which E^{clus} is left unchanged by the Seed Extended plus $X_{\text{UL}} = 20$ cut. The topo-cluster occupancy (total energy) in a given region is defined as the number (energy sum) of topo-clusters that contain at least one cell belonging to the region. Uncertainties on the event fractions are computed as the standard deviation of a binomial distribution. The dashed line represents the average of the three Control Regions.

7.2 Impact on the ATLAS event size

Calorimeter topo-clusters are used as input to the reconstruction of particles other than jets, most notably electrons, photons and τ -leptons. The time cut will then have pile-up-suppressing effects on these particles as well. A complete discussion of the effects of the time cut on electrons, photons and τ -leptons is beyond the scope of this paper. However, the widespread usage of topo-clusters in the ATLAS event reconstruction implies that the time cut can have a large beneficial effect on the overall consumption of computing resources by ATLAS data and MC samples.

The change in the ATLAS event size due to the time cut was evaluated for the introduction of the Seed Extended plus $X_{UL} = 20$ time cut into the default ATLAS reconstruction at the beginning of Run 3, before the start of 2023 data-taking. The test is performed by reconstructing one data sample and one MC sample ($t\bar{t}$ production in the fully hadronic decay) before and after the introduction of the time cut and later comparing the average event sizes on disk. In both samples the detector conditions are those of Run 3. Results are shown in Tables 3 and 4. The principal particle collections are shown, together with their relative size change. As expected, the largest effects are observed for the particle-flow object and topo-cluster collections, which are reduced in size by $\sim 16\%$ to $\sim 18\%$, while smaller differences are present for electrons/photons and τ -leptons⁷. The overall event size is reduced by about 6% in data and 7% in MC simulation.

Table 3: Collection size changes for data events on disk with the introduction of the time cut. Changes of $\mathcal{O}(100\text{ b})$ are not significant for most categories since they can arise from changes in the alignment of the (otherwise unchanged) objects and the subsequent compression when writing to disk. Larger size changes are reported relative to the original size on disk. The ‘Cells in topo-clusters’ line is limited to clusters matched to tracks, electrons and muons.

| Category | Fraction (before cut) | Size on disk before cut | Size on disk after cut | Size change |
|--|--------------------------|----------------------------|---------------------------|----------------|
| Total | 1.000 | 288.7 kb | 271.1 kb | -6.1% |
| Trigger | 0.302 | 87.11 kb | 87.01 kb | - |
| Tracking | 0.218 | 62.95 kb | 62.84 kb | - |
| Topo-clusters | 0.150 | 43.21 kb | 35.95 kb | -17% |
| Particle-flow [42, 43] objects | 0.103 | 29.75 kb | 24.50 kb | -18% |
| τ -leptons | 0.078 | 22.42 kb | 20.58 kb | -8.2% |
| Electrons / photons | 0.071 | 20.37 kb | 18.79 kb | -7.8% |
| Cells in topo-clusters | 0.057 | 16.60 kb | 15.13 kb | -8.8% |
| Muons | 0.018 | 5.293 kb | 5.295 kb | - |
| ATLAS forward proton detector [51, 52] | 0.003 | 0.810 kb | 0.810 kb | - |
| Metadata | 0.001 | 0.154 kb | 0.154 kb | - |

⁷ A very small difference in the ‘Tracking’ category is to be expected since specific information related to electron tracks is grouped in this category.

Table 4: Collection size changes for MC events on disk with the introduction of the time cut. Changes of $\mathcal{O}(100\text{ b})$ are not significant for most categories since they can arise from changes in the alignment of the (otherwise unchanged) objects and the subsequent compression when writing to disk. The ‘Cells in topo-clusters’ line is limited to clusters matched to tracks, electrons and muons.

| Category | Fraction (before cut) | Disk Size before cut | Disk Size after cut | Size change |
|--|--------------------------|-------------------------|------------------------|----------------|
| Total | 1.000 | 428.8 kb | 398.3 kb | -7.1% |
| Tracking | 0.222 | 94.98 kb | 94.51 kb | -0.5% |
| Topo-clusters | 0.160 | 68.79 kb | 57.34 kb | -17% |
| Particle-flow [42, 43] objects | 0.114 | 49.00 kb | 39.94 kb | -18% |
| τ -leptons | 0.108 | 46.38 kb | 43.24 kb | -6.8% |
| Trigger | 0.106 | 45.26 kb | 45.25 kb | - |
| Generator-level event | 0.103 | 44.28 kb | 44.29 kb | - |
| Electrons / photons | 0.086 | 36.94 kb | 33.74 kb | -8.6% |
| Cells in topo-clusters | 0.082 | 34.98 kb | 31.80 kb | -9.1% |
| Muons | 0.016 | 6.927 kb | 6.931 kb | - |
| Metadata | 0.003 | 1.201 kb | 1.202 kb | - |
| ATLAS forward proton detector [51, 52] | < 0.001 | 0.041 kb | 0.041 kb | - |

8 Conclusion

This paper presents the design and evaluation of a cell-level time criterion used in the ATLAS topo-cluster reconstruction algorithm. This criterion (the ‘time cut’) removes cells compatible with out-of-time pile-up signals. A second requirement, the Upper Limit, is imposed to prevent cells with a large amount of energy from being removed, including those expected from long-lived particles. The impact of the time cut on the reconstruction of hadronic signals was studied at both cluster level and jet level, using a combination of MC samples and ATLAS data recorded in 2017 and 2018.

The time cut was found to significantly reduce the contribution from out-of-time pile-up while not hindering signal reconstruction. Studies of MC events indicate that the multiplicity of out-of-time pile-up jets is reduced by $\sim 50\%$ at $p_T \sim 20\text{ GeV}$ and by $\sim 80\%$ above $p_T \sim 50\text{ GeV}$, across the rapidity region $|y| \lesssim 3.5$. Studies of data events found that the time cut has negligible effect on events accepted by a single-jet trigger, requiring one jet with $p_T > 45\text{ GeV}$ and seeded by a single jet at L1. On the other hand, the time cut does affect events enriched in out-of-time pile-up, typically those accepted by a random L1 trigger. In this case, the effect of the time cut is largest for $p_T \lesssim 30\text{ GeV}$ and it is found to have a stronger effect on data events than on MC events, typical differences being of $\mathcal{O}(10\%)$. The effect is believed to be caused by the absence of pile-up-only events in the MC simulation. Additional tests for signal inefficiency in data were carried out and show that the time cut would not cause a loss of clusters in the case of a calorimeter channel miscalibration. The addition of the Upper Limit is also found to prevent the loss of individual cells within clusters in such a case.

The time cut was also found to impact the jet calibration, a repetition of the pile-up correction step being necessary. The time cut provides percent-level improvement in the jet resolution at $p_T \lesssim 30\text{ GeV}$. More extensive investigation is needed to evaluate in detail the effect of the time cut on the jet energy resolution

and jet energy scale uncertainty, e.g. the pile-up-related JES uncertainty components. The full calibration chain will be re-evaluated after the introduction of the time cut into the ATLAS reconstruction software.

The Seed Extended plus $X_{UL} = 20$ cut was found to be the best option among those studied. This cut was adopted as the default for offline topo-cluster reconstruction in ATLAS Run 3 data. The introduction of the Seed Extended plus $X_{UL} = 20$ cut into the ATLAS offline reconstruction is also found to reduce the ATLAS event size on disk by 6% in data and by 7% in MC simulation. Further studies are necessary before the cut can be introduced at trigger level.

Acknowledgements

We thank CERN for the very successful operation of the LHC, as well as the support staff from our institutions without whom ATLAS could not be operated efficiently.

We acknowledge the support of ANPCyT, Argentina; YerPhI, Armenia; ARC, Australia; BMWFW and FWF, Austria; ANAS, Azerbaijan; CNPq and FAPESP, Brazil; NSERC, NRC and CFI, Canada; CERN; ANID, Chile; CAS, MOST and NSFC, China; Minciencias, Colombia; MEYS CR, Czech Republic; DNRf and DNSRC, Denmark; IN2P3-CNRS and CEA-DRF/IRFU, France; SRNSFG, Georgia; BMBF, HGF and MPG, Germany; GSRI, Greece; RGC and Hong Kong SAR, China; ISF and Benozzi Center, Israel; INFN, Italy; MEXT and JSPS, Japan; CNRST, Morocco; NWO, Netherlands; RCN, Norway; MEiN, Poland; FCT, Portugal; MNE/IFA, Romania; MESTD, Serbia; MSSR, Slovakia; ARRS and MIZŠ, Slovenia; DSI/NRF, South Africa; MICINN, Spain; SRC and Wallenberg Foundation, Sweden; SERI, SNSF and Cantons of Bern and Geneva, Switzerland; MOST, Taipei; TENMAK, Türkiye; STFC, United Kingdom; DOE and NSF, United States of America. In addition, individual groups and members have received support from BCKDF, CANARIE, CRC and DRAC, Canada; PRIMUS 21/SCI/017 and UNCE SCI/013, Czech Republic; COST, ERC, ERDF, Horizon 2020, ICSC-NextGenerationEU and Marie Skłodowska-Curie Actions, European Union; Investissements d'Avenir Labex, Investissements d'Avenir IDEX and ANR, France; DFG and AvH Foundation, Germany; Herakleitos, Thales and Aristeia programmes co-financed by EU-ESF and the Greek NSRF, Greece; BSF-NSF and MINERVA, Israel; Norwegian Financial Mechanism 2014-2021, Norway; NCN and NAWA, Poland; La Caixa Banking Foundation, CERCA Programme Generalitat de Catalunya and PROMETEO and GenT Programmes Generalitat Valenciana, Spain; Göran Gustafssons Stiftelse, Sweden; The Royal Society and Leverhulme Trust, United Kingdom.

The crucial computing support from all WLCG partners is acknowledged gratefully, in particular from CERN, the ATLAS Tier-1 facilities at TRIUMF/SFU (Canada), NDGF (Denmark, Norway, Sweden), CC-IN2P3 (France), KIT/GridKA (Germany), INFN-CNAF (Italy), NL-T1 (Netherlands), PIC (Spain), RAL (UK) and BNL (USA), the Tier-2 facilities worldwide and large non-WLCG resource providers. Major contributors of computing resources are listed in Ref. [53].

References

- [1] ATLAS Collaboration, *The ATLAS Experiment at the CERN Large Hadron Collider*, [JINST 3 \(2008\) S08003](#).
- [2] L. Evans and P. Bryant, *LHC Machine*, [JINST 3 \(2008\) S08001](#).

- [3] ATLAS Collaboration, *Topological cell clustering in the ATLAS calorimeters and its performance in LHC Run 1*, *Eur. Phys. J. C* **77** (2017) 490, arXiv: [1603.02934 \[hep-ex\]](#).
- [4] ATLAS Collaboration, *Luminosity determination in pp collisions at $\sqrt{s} = 13$ TeV using the ATLAS detector at the LHC*, *Eur. Phys. J. C* **83** (2023) 982, arXiv: [2212.09379 \[hep-ex\]](#).
- [5] D. Krohn, J. Thaler and L.-T. Wang, *Jet Trimming*, *JHEP* **02** (2010) 084, arXiv: [0912.1342 \[hep-ph\]](#).
- [6] A. J. Larkoski, S. Marzani, G. Soyez and J. Thaler, *Soft Drop*, *JHEP* **05** (2014) 146, arXiv: [1402.2657 \[hep-ph\]](#).
- [7] M. Dasgupta, A. Fregoso, S. Marzani and G. P. Salam, *Towards an understanding of jet substructure*, *JHEP* **09** (2013) 029, arXiv: [1307.0007 \[hep-ph\]](#).
- [8] P. Berta, M. Spusta, D. W. Miller and R. Leitner, *Particle-level pileup subtraction for jets and jet shapes*, *JHEP* **06** (2014) 092, arXiv: [1403.3108 \[hep-ex\]](#).
- [9] M. Cacciari, G. P. Salam and G. Soyez, *SoftKiller, a particle-level pileup removal method*, *Eur. Phys. J. C* **75** (2015) 59, arXiv: [1407.0408 \[hep-ph\]](#).
- [10] ATLAS Collaboration, *Optimisation of large-radius jet reconstruction for the ATLAS detector in 13 TeV proton–proton collisions*, *Eur. Phys. J. C* **81** (2021) 334, arXiv: [2009.04986 \[hep-ex\]](#).
- [11] ATLAS Collaboration, *ATLAS Insertable B-Layer: Technical Design Report*, ATLAS-TDR-19; CERN-LHCC-2010-013, 2010, URL: <https://cds.cern.ch/record/1291633>, Addendum: ATLAS-TDR-19-ADD-1; CERN-LHCC-2012-009, 2012, URL: <https://cds.cern.ch/record/1451888>.
- [12] B. Abbott et al., *Production and integration of the ATLAS Insertable B-Layer*, *JINST* **13** (2018) T05008, arXiv: [1803.00844 \[physics.ins-det\]](#).
- [13] ATLAS Collaboration, *Performance of the ATLAS trigger system in 2015*, *Eur. Phys. J. C* **77** (2017) 317, arXiv: [1611.09661 \[hep-ex\]](#).
- [14] ATLAS Collaboration, *The ATLAS Collaboration Software and Firmware*, ATLAS-SOFT-PUB-2021-001, 2021, URL: <https://cds.cern.ch/record/2767187>.
- [15] N. J. Buchanan et al., *Design and implementation of the Front End Board for the readout of the ATLAS liquid argon calorimeters*, *JINST* **3** (2008) P03004.
- [16] W. E. Cleland and E. G. Stern, *Signal processing considerations for liquid ionization calorimeters in a high rate environment*, *Nucl. Instrum. Meth. A* **338** (1994) 467.
- [17] ATLAS Collaboration, *Readiness of the ATLAS liquid argon calorimeter for LHC collisions*, *Eur. Phys. J. C* **70** (2010) 723, arXiv: [0912.2642 \[hep-ex\]](#).
- [18] ATLAS Collaboration, *LAr Time Resolution Plots for 2015 Data*, URL: <https://twiki.cern.ch/twiki/bin/view/AtlasPublic/LArCaloPublicResults2015>.
- [19] ATLAS Collaboration, *Monitoring and data quality assessment of the ATLAS liquid argon calorimeter*, *JINST* **9** (2014) P07024, arXiv: [1405.3768 \[hep-ex\]](#).

- [20] ATLAS Collaboration, *Readiness of the ATLAS Tile Calorimeter for LHC collisions*, *Eur. Phys. J. C* **70** (2010) 1193, arXiv: 1007.5423 [hep-ex].
- [21] M. N. Agaras et al., *Laser calibration of the ATLAS Tile Calorimeter during LHC Run 2*, *JINST* **18** (2023) P06023, arXiv: 2303.00121 [physics.ins-det].
- [22] ATLAS Collaboration, *Timing calibration and performance public plots*, URL: <https://twiki.cern.ch/twiki/bin/view/AtlasPublic/TileCaloPublicResultsTiming>.
- [23] ATLAS Collaboration, *ATLAS data quality operations and performance for 2015–2018 data-taking*, *JINST* **15** (2020) P04003, arXiv: 1911.04632 [physics.ins-det].
- [24] G. Avoni et al., *The new LUCID-2 detector for luminosity measurement and monitoring in ATLAS*, *JINST* **13** (2018) P07017.
- [25] T. Sjöstrand et al., *An introduction to PYTHIA 8.2*, *Comput. Phys. Commun.* **191** (2015) 159, arXiv: 1410.3012 [hep-ph].
- [26] NNPDF Collaboration, R. D. Ball et al., *Parton distributions with LHC data*, *Nucl. Phys. B* **867** (2013) 244, arXiv: 1207.1303 [hep-ph].
- [27] ATLAS Collaboration, *ATLAS Pythia 8 tunes to 7 TeV data*, ATL-PHYS-PUB-2014-021, 2014, URL: <https://cds.cern.ch/record/1966419>.
- [28] B. Andersson, G. Gustafson, G. Ingelman and T. Sjöstrand, *Parton fragmentation and string dynamics*, *Phys. Rept.* **97** (1983) 31.
- [29] T. Sjöstrand, *Jet fragmentation of multiparton configurations in a string framework*, *Nucl. Phys. B* **248** (1984) 469.
- [30] T. Sjöstrand, S. Mrenna and P. Skands, *A brief introduction to PYTHIA 8.1*, *Comput. Phys. Commun.* **178** (2008) 852, arXiv: 0710.3820 [hep-ph].
- [31] ATLAS Collaboration, *The Pythia 8 A3 tune description of ATLAS minimum bias and inelastic measurements incorporating the Donnachie–Landshoff diffractive model*, ATL-PHYS-PUB-2016-017, 2016, URL: <https://cds.cern.ch/record/2206965>.
- [32] S. Frixione, G. Ridolfi and P. Nason, *A positive-weight next-to-leading-order Monte Carlo for heavy flavour hadroproduction*, *JHEP* **09** (2007) 126, arXiv: 0707.3088 [hep-ph].
- [33] P. Nason, *A new method for combining NLO QCD with shower Monte Carlo algorithms*, *JHEP* **11** (2004) 040, arXiv: [hep-ph/0409146](https://arxiv.org/abs/hep-ph/0409146).
- [34] S. Frixione, P. Nason and C. Oleari, *Matching NLO QCD computations with parton shower simulations: the POWHEG method*, *JHEP* **11** (2007) 070, arXiv: 0709.2092 [hep-ph].
- [35] S. Alioli, P. Nason, C. Oleari and E. Re, *A general framework for implementing NLO calculations in shower Monte Carlo programs: the POWHEG BOX*, *JHEP* **06** (2010) 043, arXiv: 1002.2581 [hep-ph].
- [36] NNPDF Collaboration, R. D. Ball et al., *Parton distributions for the LHC run II*, *JHEP* **04** (2015) 040, arXiv: 1410.8849 [hep-ph].
- [37] ATLAS Collaboration, *Studies on top-quark Monte Carlo modelling for Top2016*, ATL-PHYS-PUB-2016-020, 2016, URL: <https://cds.cern.ch/record/2216168>.

- [38] D. J. Lange, *The EvtGen particle decay simulation package*, *Nucl. Instrum. Meth. A* **462** (2001) 152.
- [39] J. Alimena et al., *Searching for long-lived particles beyond the Standard Model at the Large Hadron Collider*, *J. Phys. G* **47** (2020) 090501, arXiv: [1903.04497 \[hep-ex\]](#).
- [40] L. Lee, C. Ohm, A. Soffer and T.-T. Yu, *Collider Searches for Long-Lived Particles Beyond the Standard Model*, *Prog. Part. Nucl. Phys.* **106** (2019) 210, arXiv: [1810.12602 \[hep-ph\]](#), Erratum: *Prog. Part. Nucl. Phys.* **122** (2022) 103912, ISSN: 0146-6410.
- [41] M. Cacciari, G. P. Salam and G. Soyez, *The anti- k_r jet clustering algorithm*, *JHEP* **04** (2008) 063, arXiv: [0802.1189 \[hep-ph\]](#).
- [42] ATLAS Collaboration, *Jet reconstruction and performance using particle flow with the ATLAS Detector*, *Eur. Phys. J. C* **77** (2017) 466, arXiv: [1703.10485 \[hep-ex\]](#).
- [43] ATLAS Collaboration, *Jet energy scale and resolution measured in proton–proton collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector*, *Eur. Phys. J. C* **81** (2021) 689, arXiv: [2007.02645 \[hep-ex\]](#).
- [44] S. D. Ellis and D. E. Soper, *Successive combination jet algorithm for hadron collisions*, *Phys. Rev. D* **48** (7 1993) 3160.
- [45] S. Catani, Y. Dokshitzer, M. Seymour and B. Webber, *Longitudinally-invariant k_{\perp} -clustering algorithms for hadron-hadron collisions*, *Nucl. Phys. B* **406** (1993) 187, ISSN: 0550-3213.
- [46] ATLAS Collaboration, *Tagging and suppression of pileup jets with the ATLAS detector*, ATLAS-CONF-2014-018, 2014, URL: <https://cds.cern.ch/record/1700870>.
- [47] ATLAS Collaboration, *Emulating the impact of additional proton–proton interactions in the ATLAS simulation by pre-sampling sets of inelastic Monte Carlo events*, *Comput. Softw. Big Sci.* **6** (2022) 3, arXiv: [2102.09495 \[hep-ex\]](#).
- [48] ATLAS Collaboration, *Search for displaced photons produced in exotic decays of the Higgs boson using 13 TeV pp collisions with the ATLAS detector*, *Phys. Rev. D* **108** (2023) 032016, arXiv: [2209.01029 \[hep-ex\]](#).
- [49] ATLAS Collaboration, *Search in diphoton and dielectron final states for displaced production of Higgs or Z bosons with the ATLAS detector in $\sqrt{s} = 13$ TeV pp collisions*, *Phys. Rev. D* **108** (2023) 012012, arXiv: [2304.12885 \[hep-ex\]](#).
- [50] ATLAS Collaboration, *Electron and photon performance measurements with the ATLAS detector using the 2015–2017 LHC proton–proton collision data*, *JINST* **14** (2019) P12006, arXiv: [1908.00005 \[hep-ex\]](#).
- [51] ATLAS Collaboration, *ATLAS Forward Proton Phase-I Upgrade: Technical Design Report*, ATLAS-TDR-024; CERN-LHCC-2015-009, 2015, URL: <https://cds.cern.ch/record/2017378>.
- [52] ATLAS Collaboration, *Proton tagging with the one arm AFP detector*, ATL-PHYS-PUB-2017-012, 2017, URL: <https://cds.cern.ch/record/2273274>.

- [53] ATLAS Collaboration, *ATLAS Computing Acknowledgements*, ATL-SOFT-PUB-2023-001, 2023,
URL: <https://cds.cern.ch/record/2869272>.

The ATLAS Collaboration

G. Aad ¹⁰², B. Abbott ¹²⁰, K. Abeling ⁵⁵, N.J. Abicht ⁴⁹, S.H. Abidi ²⁹, A. Aboulhorma ^{35e}, H. Abramowicz ¹⁵¹, H. Abreu ¹⁵⁰, Y. Abulaiti ¹¹⁷, B.S. Acharya ^{69a,69b,m}, C. Adam Bourdarios ⁴, L. Adamczyk ^{86a}, S.V. Addepalli ²⁶, M.J. Addison ¹⁰¹, J. Adelman ¹¹⁵, A. Adiguzel ^{21c}, T. Adye ¹³⁴, A.A. Affolder ¹³⁶, Y. Afik ³⁶, M.N. Agaras ¹³, J. Agarwala ^{73a,73b}, A. Aggarwal ¹⁰⁰, C. Agheorghiesei ^{27c}, A. Ahmad ³⁶, F. Ahmadov ^{38,y}, W.S. Ahmed ¹⁰⁴, S. Ahuja ⁹⁵, X. Ai ^{62a}, G. Aielli ^{76a,76b}, A. Aikot ¹⁶³, M. Ait Tamlihat ^{35e}, B. Aitbenchikh ^{35a}, I. Aizenberg ¹⁶⁹, M. Akbiyik ¹⁰⁰, T.P.A. Åkesson ⁹⁸, A.V. Akimov ³⁷, D. Akiyama ¹⁶⁸, N.N. Akolkar ²⁴, S. Aktas ^{21a}, K. Al Houry ⁴¹, G.L. Alberghi ^{23b}, J. Albert ¹⁶⁵, P. Albicocco ⁵³, G.L. Albouy ⁶⁰, S. Alderweireldt ⁵², M. Aleksa ³⁶, I.N. Aleksandrov ³⁸, C. Alexa ^{27b}, T. Alexopoulos ¹⁰, F. Alfonsi ^{23b}, M. Algren ⁵⁶, M. Alhroob ¹²⁰, B. Ali ¹³², H.M.J. Ali ⁹¹, S. Ali ¹⁴⁸, S.W. Alibocus ⁹², M. Aliev ¹⁴⁵, G. Alimonti ^{71a}, W. Alkakh ⁵⁵, C. Allaire ⁶⁶, B.M.M. Allbrooke ¹⁴⁶, J.F. Allen ⁵², C.A. Allendes Flores ^{137f}, P.P. Allport ²⁰, A. Aloisio ^{72a,72b}, F. Alonso ⁹⁰, C. Alpigiani ¹³⁸, M. Alvarez Estevez ⁹⁹, A. Alvarez Fernandez ¹⁰⁰, M. Alves Cardoso ⁵⁶, M.G. Alviggi ^{72a,72b}, M. Aly ¹⁰¹, Y. Amaral Coutinho ^{83b}, A. Ambler ¹⁰⁴, C. Amelung ³⁶, M. Amerl ¹⁰¹, C.G. Ames ¹⁰⁹, D. Amidei ¹⁰⁶, S.P. Amor Dos Santos ^{130a}, K.R. Amos ¹⁶³, V. Ananiev ¹²⁵, C. Anastopoulos ¹³⁹, T. Andeen ¹¹, J.K. Anders ³⁶, S.Y. Andrean ^{47a,47b}, A. Andreazza ^{71a,71b}, S. Angelidakis ⁹, A. Angerami ^{41,ab}, A.V. Anisenkov ³⁷, A. Annovi ^{74a}, C. Antel ⁵⁶, M.T. Anthony ¹³⁹, E. Antipov ¹⁴⁵, M. Antonelli ⁵³, F. Anulli ^{75a}, M. Aoki ⁸⁴, T. Aoki ¹⁵³, J.A. Aparisi Pozo ¹⁶³, M.A. Aparo ¹⁴⁶, L. Aperio Bella ⁴⁸, C. Appelt ¹⁸, A. Apyan ²⁶, N. Aranzabal ³⁶, S.J. Arbiol Val ⁸⁷, C. Arcangeletti ⁵³, A.T.H. Arce ⁵¹, E. Arena ⁹², J-F. Arguin ¹⁰⁸, S. Argyropoulos ⁵⁴, J.-H. Arling ⁴⁸, O. Arnaez ⁴, H. Arnold ¹¹⁴, G. Artoni ^{75a,75b}, H. Asada ¹¹¹, K. Asai ¹¹⁸, S. Asai ¹⁵³, N.A. Asbah ⁶¹, K. Assamagan ²⁹, R. Astalos ^{28a}, S. Atashi ¹⁶⁰, R.J. Atkin ^{33a}, M. Atkinson ¹⁶², H. Atmani ^{35f}, P.A. Atlasiddha ¹²⁸, K. Augsten ¹³², S. Auricchio ^{72a,72b}, A.D. Auriol ²⁰, V.A. Austrup ¹⁰¹, G. Avolio ³⁶, K. Axiotis ⁵⁶, G. Azuelos ^{108,af}, D. Babal ^{28b}, H. Bachacou ¹³⁵, K. Bachas ^{152,p}, A. Bachi ³⁴, F. Backman ^{47a,47b}, A. Badea ⁶¹, T.M. Baer ¹⁰⁶, P. Bagnaia ^{75a,75b}, M. Bahmani ¹⁸, A.J. Bailey ¹⁶³, V.R. Bailey ¹⁶², J.T. Baines ¹³⁴, L. Baines ⁹⁴, O.K. Baker ¹⁷², E. Bakos ¹⁵, D. Bakshi Gupta ⁸, V. Balakrishnan ¹²⁰, R. Balasubramanian ¹¹⁴, E.M. Baldin ³⁷, P. Balek ^{86a}, E. Ballabene ^{23b,23a}, F. Balli ¹³⁵, L.M. Baltes ^{63a}, W.K. Balunas ³², J. Balz ¹⁰⁰, E. Banas ⁸⁷, M. Bandieramonte ¹²⁹, A. Bandyopadhyay ²⁴, S. Bansal ²⁴, L. Barak ¹⁵¹, M. Barakat ⁴⁸, E.L. Barberio ¹⁰⁵, D. Barberis ^{57b,57a}, M. Barbero ¹⁰², M.Z. Barel ¹¹⁴, K.N. Barends ^{33a}, T. Barillari ¹¹⁰, M-S. Barisits ³⁶, T. Barklow ¹⁴³, P. Baron ¹²², D.A. Baron Moreno ¹⁰¹, A. Baroncelli ^{62a}, G. Barone ²⁹, A.J. Barr ¹²⁶, J.D. Barr ⁹⁶, L. Barranco Navarro ^{47a,47b}, F. Barreiro ⁹⁹, J. Barreiro Guimarães da Costa ^{14a}, U. Barron ¹⁵¹, M.G. Barros Teixeira ^{130a}, S. Barsov ³⁷, F. Bartels ^{63a}, R. Bartoldus ¹⁴³, A.E. Barton ⁹¹, P. Bartos ^{28a}, A. Basan ¹⁰⁰, M. Baselga ⁴⁹, A. Bassalat ^{66,b}, M.J. Basso ^{156a}, C.R. Basson ¹⁰¹, R.L. Bates ⁵⁹, S. Batlamous ^{35e}, J.R. Batley ³², B. Batool ¹⁴¹, M. Battaglia ¹³⁶, D. Battulga ¹⁸, M. Bauge ^{75a,75b}, M. Bauer ³⁶, P. Bauer ²⁴, L.T. Bazzano Hurrell ³⁰, J.B. Beacham ⁵¹, T. Beau ¹²⁷, J.Y. Beauchamp ⁹⁰, P.H. Beauchemin ¹⁵⁸, F. Becherer ⁵⁴, P. Bechtel ²⁴, H.P. Beck ^{19,o}, K. Becker ¹⁶⁷, A.J. Beddall ⁸², V.A. Bednyakov ³⁸, C.P. Bee ¹⁴⁵, L.J. Beemster ¹⁵, T.A. Beermann ³⁶, M. Begalli ^{83d}, M. Begel ²⁹, A. Behera ¹⁴⁵, J.K. Behr ⁴⁸, J.F. Beirer ³⁶, F. Beisiegel ²⁴, M. Belfkir ¹⁵⁹, G. Bella ¹⁵¹, L. Bellagamba ^{23b}, A. Bellerive ³⁴, P. Bellos ²⁰, K. Beloborodov ³⁷, D. Benchechroun ^{35a}, F. Bendebba ^{35a}, Y. Benhammou ¹⁵¹, M. Benoit ²⁹, J.R. Bensinger ²⁶, S. Bentvelsen ¹¹⁴, L. Beresford ⁴⁸, M. Beretta ⁵³,

E. Bergeaas Kuutmann ¹⁶¹, N. Berger ⁴, B. Bergmann ¹³², J. Beringer ^{17a}, G. Bernardi ⁵,
 C. Bernius ¹⁴³, F.U. Bernlochner ²⁴, F. Bernon ^{36,102}, A. Berrocal Guardia ¹³, T. Berry ⁹⁵,
 P. Berta ¹³³, A. Berthold ⁵⁰, I.A. Bertram ⁹¹, S. Bethke ¹¹⁰, A. Betti ^{75a,75b}, A.J. Bevan ⁹⁴,
 N.K. Bhalla ⁵⁴, M. Bhamjee ^{33c}, S. Bhatta ¹⁴⁵, D.S. Bhattacharya ¹⁶⁶, P. Bhattacharai ¹⁴³,
 V.S. Bhopatkar ¹²¹, R. Bi ^{29,ai}, R.M. Bianchi ¹²⁹, G. Bianco ^{23b,23a}, O. Biebel ¹⁰⁹, R. Bielski ¹²³,
 M. Biglietti ^{77a}, M. Bindi ⁵⁵, A. Bingul ^{21b}, C. Bini ^{75a,75b}, A. Biondini ⁹², C.J. Birch-sykes ¹⁰¹,
 G.A. Bird ^{20,134}, M. Birman ¹⁶⁹, M. Biros ¹³³, S. Biryukov ¹⁴⁶, T. Bisanz ⁴⁹, E. Bisceglie ^{43b,43a},
 J.P. Biswal ¹³⁴, D. Biswas ¹⁴¹, A. Bitadze ¹⁰¹, K. Bjørke ¹²⁵, I. Bloch ⁴⁸, A. Blue ⁵⁹,
 U. Blumenschein ⁹⁴, J. Blumenthal ¹⁰⁰, G.J. Bobbink ¹¹⁴, V.S. Bobrovnikov ³⁷, M. Boehler ⁵⁴,
 B. Boehm ¹⁶⁶, D. Bogavac ³⁶, A.G. Bogdanchikov ³⁷, C. Bohm ^{47a}, V. Boisvert ⁹⁵, P. Bokan ⁴⁸,
 T. Bold ^{86a}, M. Bomben ⁵, M. Bona ⁹⁴, M. Boonekamp ¹³⁵, C.D. Booth ⁹⁵, A.G. Borbély ⁵⁹,
 I.S. Bordulev ³⁷, H.M. Borecka-Bielska ¹⁰⁸, G. Borissov ⁹¹, D. Bortoletto ¹²⁶, D. Boscherini ^{23b},
 M. Bosman ¹³, J.D. Bossio Sola ³⁶, K. Bouaouda ^{35a}, N. Bouchhar ¹⁶³, J. Boudreau ¹²⁹,
 E.V. Bouhova-Thacker ⁹¹, D. Boumediene ⁴⁰, R. Bouquet ¹⁶⁵, A. Boveia ¹¹⁹, J. Boyd ³⁶,
 D. Boye ²⁹, I.R. Boyko ³⁸, J. Bracinek ²⁰, N. Brahimi ^{62d}, G. Brandt ¹⁷¹, O. Brandt ³²,
 F. Braren ⁴⁸, B. Brau ¹⁰³, J.E. Brau ¹²³, R. Brenner ¹⁶⁹, L. Brenner ¹¹⁴, R. Brenner ¹⁶¹,
 S. Bressler ¹⁶⁹, D. Britton ⁵⁹, D. Britzger ¹¹⁰, I. Brock ²⁴, G. Brooijmans ⁴¹, W.K. Brooks ^{137f},
 E. Brost ²⁹, L.M. Brown ¹⁶⁵, L.E. Bruce ⁶¹, T.L. Bruckler ¹²⁶, P.A. Bruckman de Renstrom ⁸⁷,
 B. Brüers ⁴⁸, A. Bruni ^{23b}, G. Bruni ^{23b}, M. Bruschi ^{23b}, N. Bruscino ^{75a,75b}, T. Buanes ¹⁶,
 Q. Buat ¹³⁸, D. Buchin ¹¹⁰, A.G. Buckley ⁵⁹, O. Bulekov ³⁷, B.A. Bullard ¹⁴³, S. Burdin ⁹²,
 C.D. Burgard ⁴⁹, A.M. Burger ⁴⁰, B. Burghgrave ⁸, O. Burlayenko ⁵⁴, J.T.P. Burr ³²,
 C.D. Burton ¹¹, J.C. Burzynski ¹⁴², E.L. Busch ⁴¹, V. Büscher ¹⁰⁰, P.J. Bussey ⁵⁹,
 J.M. Butler ²⁵, C.M. Buttar ⁵⁹, J.M. Butterworth ⁹⁶, W. Buttinger ¹³⁴, C.J. Buxo Vazquez ¹⁰⁷,
 A.R. Buzykaev ³⁷, S. Cabrera Urbán ¹⁶³, L. Cadamuro ⁶⁶, D. Caforio ⁵⁸, H. Cai ¹²⁹,
 Y. Cai ^{14a,14e}, Y. Cai ^{14c}, V.M.M. Cairo ³⁶, O. Cakir ^{3a}, N. Calace ³⁶, P. Calafiura ^{17a},
 G. Calderini ¹²⁷, P. Calfayan ⁶⁸, G. Callea ⁵⁹, L.P. Caloba ^{83b}, D. Calvet ⁴⁰, S. Calvet ⁴⁰,
 T.P. Calvet ¹⁰², M. Calvetti ^{74a,74b}, R. Camacho Toro ¹²⁷, S. Camarda ³⁶, D. Camarero Munoz ²⁶,
 P. Camarri ^{76a,76b}, M.T. Camerlingo ^{72a,72b}, D. Cameron ³⁶, C. Camincher ¹⁶⁵, M. Campanelli ⁹⁶,
 A. Camplani ⁴², V. Canale ^{72a,72b}, A. Canesse ¹⁰⁴, J. Cantero ¹⁶³, Y. Cao ¹⁶², F. Capocasa ²⁶,
 M. Capua ^{43b,43a}, A. Carbone ^{71a,71b}, R. Cardarelli ^{76a}, J.C.J. Cardenas ⁸, F. Cardillo ¹⁶³,
 G. Carducci ^{43b,43a}, T. Carli ³⁶, G. Carlino ^{72a}, J.I. Carlotto ¹³, B.T. Carlson ^{129,q},
 E.M. Carlson ^{165,156a}, L. Carminati ^{71a,71b}, A. Carnelli ¹³⁵, M. Carnesale ^{75a,75b}, S. Caron ¹¹³,
 E. Carquin ^{137f}, S. Carrá ^{71a}, G. Carratta ^{23b,23a}, F. Carrio Argos ^{33g}, J.W.S. Carter ¹⁵⁵,
 T.M. Carter ⁵², M.P. Casado ^{13,i}, M. Caspar ⁴⁸, F.L. Castillo ⁴, L. Castillo Garcia ¹³,
 V. Castillo Gimenez ¹⁶³, N.F. Castro ^{130a,130e}, A. Catinaccio ³⁶, J.R. Catmore ¹²⁵, V. Cavaliere ²⁹,
 N. Cavalli ^{23b,23a}, V. Cavalinni ^{74a,74b}, Y.C. Cekmecelioglu ⁴⁸, E. Celebi ^{21a}, F. Celli ¹²⁶,
 M.S. Centonze ^{70a,70b}, V. Cepaitis ⁵⁶, K. Cerny ¹²², A.S. Cerqueira ^{83a}, A. Cerri ¹⁴⁶,
 L. Cerrito ^{76a,76b}, F. Cerutti ^{17a}, B. Cervato ¹⁴¹, A. Cervelli ^{23b}, G. Cesarini ⁵³, S.A. Cetin ⁸²,
 D. Chakraborty ¹¹⁵, J. Chan ¹⁷⁰, W.Y. Chan ¹⁵³, J.D. Chapman ³², E. Chapon ¹³⁵,
 B. Chargeishvili ^{149b}, D.G. Charlton ²⁰, T.P. Charman ⁹⁴, M. Chatterjee ¹⁹, C. Chauhan ¹³³,
 S. Chekanov ⁶, S.V. Chekulaev ^{156a}, G.A. Chelkov ^{38,a}, A. Chen ¹⁰⁶, B. Chen ¹⁵¹, B. Chen ¹⁶⁵,
 H. Chen ^{14c}, H. Chen ²⁹, J. Chen ^{62c}, J. Chen ¹⁴², M. Chen ¹²⁶, S. Chen ¹⁵³, S.J. Chen ^{14c},
 X. Chen ^{62c,135}, X. Chen ^{14b,ae}, Y. Chen ^{62a}, C.L. Cheng ¹⁷⁰, H.C. Cheng ^{64a}, S. Cheong ¹⁴³,
 A. Cheplakov ³⁸, E. Cheremushkina ⁴⁸, E. Cherepanova ¹¹⁴, R. Cherkaoui El Moursli ^{35e},
 E. Cheu ⁷, K. Cheung ⁶⁵, L. Chevalier ¹³⁵, V. Chiarella ⁵³, G. Chiarelli ^{74a}, N. Chiedde ¹⁰²,
 G. Chiodini ^{70a}, A.S. Chisholm ²⁰, A. Chitan ^{27b}, M. Chitishvili ¹⁶³, M.V. Chizhov ³⁸,
 K. Choi ¹¹, A.R. Chomont ^{75a,75b}, Y. Chou ¹⁰³, E.Y.S. Chow ¹¹³, T. Chowdhury ^{33g},

K.L. Chu ¹⁶⁹, M.C. Chu ^{64a}, X. Chu ^{14a,14e}, J. Chudoba ¹³¹, J.J. Chwastowski ⁸⁷, D. Cieri ¹¹⁰,
 K.M. Ciesla ^{86a}, V. Cindro ⁹³, A. Ciocio ^{17a}, F. Ciroto ^{72a,72b}, Z.H. Citron ^{169,k}, M. Citterio ^{71a},
 D.A. Ciubotaru ^{27b}, A. Clark ⁵⁶, P.J. Clark ⁵², C. Clarry ¹⁵⁵, J.M. Clavijo Columbie ⁴⁸,
 S.E. Clawson ⁴⁸, C. Clement ^{47a,47b}, J. Clercx ⁴⁸, Y. Coadou ¹⁰², M. Cobal ^{69a,69c},
 A. Coccaro ^{57b}, R.F. Coelho Barrue ^{130a}, R. Coelho Lopes De Sa ¹⁰³, S. Coelli ^{71a},
 A.E.C. Coimbra ^{71a,71b}, B. Cole ⁴¹, J. Collot ⁶⁰, P. Conde Muiño ^{130a,130g}, M.P. Connell ^{33c},
 S.H. Connell ^{33c}, I.A. Connelly ⁵⁹, E.I. Conroy ¹²⁶, F. Conventi ^{72a,ag}, H.G. Cooke ²⁰,
 A.M. Cooper-Sarkar ¹²⁶, A. Cordeiro Oudot Choi ¹²⁷, L.D. Corpe ⁴⁰, M. Corradi ^{75a,75b},
 F. Corriveau ^{104,w}, A. Cortes-Gonzalez ¹⁸, M.J. Costa ¹⁶³, F. Costanza ⁴, D. Costanzo ¹³⁹,
 B.M. Cote ¹¹⁹, G. Cowan ⁹⁵, K. Cranmer ¹⁷⁰, D. Cremonini ^{23b,23a}, S. Crépe-Renaudin ⁶⁰,
 F. Crescioli ¹²⁷, M. Cristinziani ¹⁴¹, M. Cristoforetti ^{78a,78b}, V. Croft ¹¹⁴, J.E. Crosby ¹²¹,
 G. Crosetti ^{43b,43a}, A. Cueto ⁹⁹, T. Cuhadar Donszelmann ¹⁶⁰, H. Cui ^{14a,14e}, Z. Cui ⁷,
 W.R. Cunningham ⁵⁹, F. Curcio ^{43b,43a}, P. Czodrowski ³⁶, M.M. Czurylo ^{63b},
 M.J. Da Cunha Sargedas De Sousa ^{57b,57a}, J.V. Da Fonseca Pinto ^{83b}, C. Da Via ¹⁰¹,
 W. Dabrowski ^{86a}, T. Dado ⁴⁹, S. Dahbi ^{33g}, T. Dai ¹⁰⁶, D. Dal Santo ¹⁹, C. Dallapiccola ¹⁰³,
 M. Dam ⁴², G. D'amen ²⁹, V. D'Amico ¹⁰⁹, J. Damp ¹⁰⁰, J.R. Dandoy ³⁴, M.F. Daneri ³⁰,
 M. Danninger ¹⁴², V. Dao ³⁶, G. Darbo ^{57b}, S. Darmora ⁶, S.J. Das ^{29,ai}, S. D'Auria ^{71a,71b},
 C. David ^{156b}, T. Davidek ¹³³, B. Davis-Purcell ³⁴, I. Dawson ⁹⁴, H.A. Day-hall ¹³², K. De ⁸,
 R. De Asmundis ^{72a}, N. De Biase ⁴⁸, S. De Castro ^{23b,23a}, N. De Groot ¹¹³, P. de Jong ¹¹⁴,
 H. De la Torre ¹¹⁵, A. De Maria ^{14c}, A. De Salvo ^{75a}, U. De Sanctis ^{76a,76b}, A. De Santo ¹⁴⁶,
 J.B. De Vivie De Regie ⁶⁰, D.V. Dedovich ³⁸, J. Degens ¹¹⁴, A.M. Deiana ⁴⁴, F. Del Corso ^{23b,23a},
 J. Del Peso ⁹⁹, F. Del Rio ^{63a}, F. Deliot ¹³⁵, C.M. Delitzsch ⁴⁹, M. Della Pietra ^{72a,72b},
 D. Della Volpe ⁵⁶, A. Dell'Acqua ³⁶, L. Dell'Asta ^{71a,71b}, M. Delmastro ⁴, P.A. Delsart ⁶⁰,
 S. Demers ¹⁷², M. Demichev ³⁸, S.P. Denisov ³⁷, L. D'Eramo ⁴⁰, D. Derendarz ⁸⁷, F. Derue ¹²⁷,
 P. Dervan ⁹², K. Desch ²⁴, C. Deutsch ²⁴, F.A. Di Bello ^{57b,57a}, A. Di Ciaccio ^{76a,76b},
 L. Di Ciaccio ⁴, A. Di Domenico ^{75a,75b}, C. Di Donato ^{72a,72b}, A. Di Girolamo ³⁶,
 G. Di Gregorio ³⁶, A. Di Luca ^{78a,78b}, B. Di Micco ^{77a,77b}, R. Di Nardo ^{77a,77b}, C. Diaconu ¹⁰²,
 M. Diamantopoulou ³⁴, F.A. Dias ¹¹⁴, T. Dias Do Vale ¹⁴², M.A. Diaz ^{137a,137b},
 F.G. Diaz Capriles ²⁴, M. Didenko ¹⁶³, E.B. Diehl ¹⁰⁶, L. Diehl ⁵⁴, S. Díez Cornell ⁴⁸,
 C. Diez Pardos ¹⁴¹, C. Dimitriadi ^{161,24}, A. Dimitrievska ^{17a}, J. Dingfelder ²⁴, I-M. Dinu ^{27b},
 S.J. Dittmeier ^{63b}, F. Dittus ³⁶, F. Djama ¹⁰², T. Djobava ^{149b}, J.I. Djuvsland ¹⁶,
 C. Doglioni ^{101,98}, A. Dohnalova ^{28a}, J. Dolejsi ¹³³, Z. Dolezal ¹³³, K.M. Dona ³⁹,
 M. Donadelli ^{83c}, B. Dong ¹⁰⁷, J. Donini ⁴⁰, A. D'Onofrio ^{72a,72b}, M. D'Onofrio ⁹²,
 J. Dopke ¹³⁴, A. Doria ^{72a}, N. Dos Santos Fernandes ^{130a}, P. Dougan ¹⁰¹, M.T. Dova ⁹⁰,
 A.T. Doyle ⁵⁹, M.A. Draguet ¹²⁶, E. Dreyer ¹⁶⁹, I. Drivas-koulouris ¹⁰, M. Drnevich ¹¹⁷,
 A.S. Drobac ¹⁵⁸, M. Drozdova ⁵⁶, D. Du ^{62a}, T.A. du Pree ¹¹⁴, F. Dubinin ³⁷, M. Dubovsky ^{28a},
 E. Duchovni ¹⁶⁹, G. Duckeck ¹⁰⁹, O.A. Ducu ^{27b}, D. Duda ⁵², A. Dudarev ³⁶, E.R. Duden ²⁶,
 M. D'uffizi ¹⁰¹, L. Duflot ⁶⁶, M. Dührssen ³⁶, C. Dülsen ¹⁷¹, A.E. Dumitriu ^{27b}, M. Dunford ^{63a},
 S. Dungs ⁴⁹, K. Dunne ^{47a,47b}, A. Duperrin ¹⁰², H. Duran Yildiz ^{3a}, M. Düren ⁵⁸,
 A. Durglishvili ^{149b}, B.L. Dwyer ¹¹⁵, G.I. Dyckes ^{17a}, M. Dyndal ^{86a}, B.S. Dziedzic ⁸⁷,
 Z.O. Earnshaw ¹⁴⁶, G.H. Eberwein ¹²⁶, B. Eckerova ^{28a}, S. Eggebrecht ⁵⁵,
 E. Egidio Purcino De Souza ¹²⁷, L.F. Ehrke ⁵⁶, G. Eigen ¹⁶, K. Einsweiler ^{17a}, T. Ekelof ¹⁶¹,
 P.A. Ekman ⁹⁸, S. El Farkh ^{35b}, Y. El Ghazali ^{35b}, H. El Jarrari ³⁶, A. El Moussaouy ¹⁰⁸,
 V. Ellajosyula ¹⁶¹, M. Ellert ¹⁶¹, F. Ellinghaus ¹⁷¹, N. Ellis ³⁶, J. Elmsheuser ²⁹, M. Elsing ³⁶,
 D. Emelianov ¹³⁴, Y. Enari ¹⁵³, I. Ene ^{17a}, S. Epari ¹³, J. Erdmann ⁴⁹, P.A. Erland ⁸⁷,
 M. Errenst ¹⁷¹, M. Escalier ⁶⁶, C. Escobar ¹⁶³, E. Etzion ¹⁵¹, G. Evans ^{130a}, H. Evans ⁶⁸,
 L.S. Evans ⁹⁵, M.O. Evans ¹⁴⁶, A. Ezhilov ³⁷, S. Ezzarqtouni ^{35a}, F. Fabbri ⁵⁹, L. Fabbri ^{23b,23a},

G. Facini ⁹⁶, V. Fadeyev ¹³⁶, R.M. Fakhruddinov ³⁷, S. Falciano ^{75a}, L.F. Falda Ulhoa Coelho ³⁶, P.J. Falke ²⁴, J. Faltova ¹³³, C. Fan ¹⁶², Y. Fan ^{14a}, Y. Fang ^{14a,14e}, M. Fanti ^{71a,71b}, M. Faraj ^{69a,69b}, Z. Farazpay ⁹⁷, A. Farbin ⁸, A. Farilla ^{77a}, T. Farooque ¹⁰⁷, S.M. Farrington ⁵², F. Fassi ^{35e}, D. Fassouliotis ⁹, M. Faucci Giannelli ^{76a,76b}, W.J. Fawcett ³², L. Fayard ⁶⁶, P. Federic ¹³³, P. Federicova ¹³¹, O.L. Fedin ^{37,a}, G. Fedotov ³⁷, M. Feickert ¹⁷⁰, L. Feligioni ¹⁰², D.E. Fellers ¹²³, C. Feng ^{62b}, M. Feng ^{14b}, Z. Feng ¹¹⁴, M.J. Fenton ¹⁶⁰, A.B. Fenyuk ³⁷, L. Ferencz ⁴⁸, R.A.M. Ferguson ⁹¹, S.I. Fernandez Luengo ^{137f}, P. Fernandez Martinez ¹³, M.J.V. Fernoux ¹⁰², J. Ferrando ⁴⁸, A. Ferrari ¹⁶¹, P. Ferrari ^{114,113}, R. Ferrari ^{73a}, D. Ferrere ⁵⁶, C. Ferretti ¹⁰⁶, F. Fiedler ¹⁰⁰, P. Fiedler ¹³², A. Filipčič ⁹³, E.K. Filmer ¹, F. Filthaut ¹¹³, M.C.N. Fiolhais ^{130a,130c,c}, L. Fiorini ¹⁶³, W.C. Fisher ¹⁰⁷, T. Fitschen ¹⁰¹, P.M. Fitzhugh ¹³⁵, I. Fleck ¹⁴¹, P. Fleischmann ¹⁰⁶, T. Flick ¹⁷¹, M. Flores ^{33d,ac}, L.R. Flores Castillo ^{64a}, L. Flores Sanz De Acedo ³⁶, F.M. Follega ^{78a,78b}, N. Fomin ¹⁶, J.H. Foo ¹⁵⁵, B.C. Forland ⁶⁸, A. Formica ¹³⁵, A.C. Forti ¹⁰¹, E. Fortin ³⁶, A.W. Fortman ⁶¹, M.G. Foti ^{17a}, L. Fountas ^{9,j}, D. Fournier ⁶⁶, H. Fox ⁹¹, P. Francavilla ^{74a,74b}, S. Francescato ⁶¹, S. Franchellucci ⁵⁶, M. Franchini ^{23b,23a}, S. Franchino ^{63a}, D. Francis ³⁶, L. Franco ¹¹³, V. Franco Lima ³⁶, L. Franconi ⁴⁸, M. Franklin ⁶¹, G. Frattari ²⁶, A.C. Freegard ⁹⁴, W.S. Freund ^{83b}, Y.Y. Frid ¹⁵¹, J. Friend ⁵⁹, N. Fritzsche ⁵⁰, A. Froch ⁵⁴, D. Froidevaux ³⁶, J.A. Frost ¹²⁶, Y. Fu ^{62a}, S. Fuenzalida Garrido ^{137f}, M. Fujimoto ¹⁰², K.Y. Fung ^{64a}, E. Furtado De Simas Filho ^{83b}, M. Furukawa ¹⁵³, J. Fuster ¹⁶³, A. Gabrielli ^{23b,23a}, A. Gabrielli ¹⁵⁵, P. Gadow ³⁶, G. Gagliardi ^{57b,57a}, L.G. Gagnon ^{17a}, E.J. Gallas ¹²⁶, B.J. Gallop ¹³⁴, K.K. Gan ¹¹⁹, S. Ganguly ¹⁵³, Y. Gao ⁵², F.M. Garay Walls ^{137a,137b}, B. Garcia ²⁹, C. García ¹⁶³, A. Garcia Alonso ¹¹⁴, A.G. Garcia Caffaro ¹⁷², J.E. García Navarro ¹⁶³, M. Garcia-Sciveres ^{17a}, G.L. Gardner ¹²⁸, R.W. Gardner ³⁹, N. Garelli ¹⁵⁸, D. Garg ⁸⁰, R.B. Garg ^{143,n}, J.M. Gargan ⁵², C.A. Garner ¹⁵⁵, C.M. Garvey ^{33a}, P. Gaspar ^{83b}, V.K. Gassmann ¹⁵⁸, G. Gaudio ^{73a}, V. Gautam ¹³, P. Gauzzi ^{75a,75b}, I.L. Gavrilenko ³⁷, A. Gavrilyuk ³⁷, C. Gay ¹⁶⁴, G. Gaycken ⁴⁸, E.N. Gazis ¹⁰, A.A. Geanta ^{27b}, C.M. Gee ¹³⁶, A. Gekow ¹¹⁹, C. Gemme ^{57b}, M.H. Genest ⁶⁰, S. Gentile ^{75a,75b}, A.D. Gentry ¹¹², S. George ⁹⁵, W.F. George ²⁰, T. Geralis ⁴⁶, P. Gessinger-Befurt ³⁶, M.E. Geyik ¹⁷¹, M. Ghani ¹⁶⁷, M. Ghneimat ¹⁴¹, K. Ghorbanian ⁹⁴, A. Ghosal ¹⁴¹, A. Ghosh ¹⁶⁰, A. Ghosh ⁷, B. Giacobbe ^{23b}, S. Giagu ^{75a,75b}, T. Giani ¹¹⁴, P. Giannetti ^{74a}, A. Giannini ^{62a}, S.M. Gibson ⁹⁵, M. Gignac ¹³⁶, D.T. Gil ^{86b}, A.K. Gilbert ^{86a}, B.J. Gilbert ⁴¹, D. Gillberg ³⁴, G. Gilles ¹¹⁴, N.E.K. Gillwald ⁴⁸, L. Ginabat ¹²⁷, D.M. Gingrich ^{2,af}, M.P. Giordani ^{69a,69c}, P.F. Giraud ¹³⁵, G. Giugliarelli ^{69a,69c}, D. Giugni ^{71a}, F. Giuli ³⁶, I. Gkialas ^{9,j}, L.K. Gladilin ³⁷, C. Glasman ⁹⁹, G.R. Gledhill ¹²³, G. Glemža ⁴⁸, M. Glisic ¹²³, I. Gnesi ^{43b,f}, Y. Go ^{29,ai}, M. Goblirsch-Kolb ³⁶, B. Gocke ⁴⁹, D. Godin ¹⁰⁸, B. Gokturk ^{21a}, S. Goldfarb ¹⁰⁵, T. Golling ⁵⁶, M.G.D. Gololo ^{33g}, D. Golubkov ³⁷, J.P. Gombas ¹⁰⁷, A. Gomes ^{130a,130b}, G. Gomes Da Silva ¹⁴¹, A.J. Gomez Delegido ¹⁶³, R. Gonçalves ^{130a,130c}, G. Gonella ¹²³, L. Gonella ²⁰, A. Gongadze ^{149c}, F. Gonnella ²⁰, J.L. Gonski ⁴¹, R.Y. González Andana ⁵², S. González de la Hoz ¹⁶³, S. Gonzalez Fernandez ¹³, R. Gonzalez Lopez ⁹², C. Gonzalez Renteria ^{17a}, M.V. Gonzalez Rodrigues ⁴⁸, R. Gonzalez Suarez ¹⁶¹, S. Gonzalez-Sevilla ⁵⁶, G.R. Gonzalvo Rodriguez ¹⁶³, L. Goossens ³⁶, B. Gorini ³⁶, E. Gorini ^{70a,70b}, A. Gorišek ⁹³, T.C. Gosart ¹²⁸, A.T. Goshaw ⁵¹, M.I. Gostkin ³⁸, S. Goswami ¹²¹, C.A. Gottardo ³⁶, S.A. Gotz ¹⁰⁹, M. Goughri ^{35b}, V. Goumarre ⁴⁸, A.G. Goussiou ¹³⁸, N. Govender ^{33c}, I. Grabowska-Bold ^{86a}, K. Graham ³⁴, E. Gramstad ¹²⁵, S. Grancagnolo ^{70a,70b}, M. Grandi ¹⁴⁶, C.M. Grant ^{1,135}, P.M. Gravila ^{27f}, F.G. Gravili ^{70a,70b}, H.M. Gray ^{17a}, M. Greco ^{70a,70b}, C. Grefe ²⁴, I.M. Gregor ⁴⁸, P. Grenier ¹⁴³, S.G. Grewe ¹¹⁰, C. Grieco ¹³, A.A. Grillo ¹³⁶, K. Grimm ³¹, S. Grinstein ^{13,s}, J.-F. Grivaz ⁶⁶, E. Gross ¹⁶⁹, J. Grosse-Knetter ⁵⁵, C. Grud ¹⁰⁶, J.C. Grundy ¹²⁶, L. Guan ¹⁰⁶, W. Guan ²⁹, C. Gubbels ¹⁶⁴,

J.G.R. Guerrero Rojas ¹⁶³, G. Guerrieri ^{69a,69c}, F. Guescini ¹¹⁰, R. Gugel ¹⁰⁰, J.A.M. Guhit ¹⁰⁶, A. Guida ¹⁸, E. Guilloton ^{167,134}, S. Guindon ³⁶, F. Guo ^{14a,14e}, J. Guo ^{62c}, L. Guo ⁴⁸, Y. Guo ¹⁰⁶, R. Gupta ⁴⁸, R. Gupta ¹²⁹, S. Gurbuz ²⁴, S.S. Gurdasani ⁵⁴, G. Gustavino ³⁶, M. Guth ⁵⁶, P. Gutierrez ¹²⁰, L.F. Gutierrez Zagazeta ¹²⁸, M. Gutsche ⁵⁰, C. Gutschow ⁹⁶, C. Gwenlan ¹²⁶, C.B. Gwilliam ⁹², E.S. Haaland ¹²⁵, A. Haas ¹¹⁷, M. Habedank ⁴⁸, C. Haber ^{17a}, H.K. Hadavand ⁸, A. Hadeef ⁵⁰, S. Hadzic ¹¹⁰, A.I. Hagan ⁹¹, J.J. Hahn ¹⁴¹, E.H. Haines ⁹⁶, M. Haleem ¹⁶⁶, J. Haley ¹²¹, J.J. Hall ¹³⁹, G.D. Hallewell ¹⁰², L. Halser ¹⁹, K. Hamano ¹⁶⁵, M. Hamer ²⁴, G.N. Hamity ⁵², E.J. Hampshire ⁹⁵, J. Han ^{62b}, K. Han ^{62a}, L. Han ^{14c}, L. Han ^{62a}, S. Han ^{17a}, Y.F. Han ¹⁵⁵, K. Hanagaki ⁸⁴, M. Hance ¹³⁶, D.A. Hangal ^{41,ab}, H. Hanif ¹⁴², M.D. Hank ¹²⁸, R. Hankache ¹⁰¹, J.B. Hansen ⁴², J.D. Hansen ⁴², P.H. Hansen ⁴², K. Hara ¹⁵⁷, D. Harada ⁵⁶, T. Harenberg ¹⁷¹, S. Harkusha ³⁷, M.L. Harris ¹⁰³, Y.T. Harris ¹²⁶, J. Harrison ¹³, N.M. Harrison ¹¹⁹, P.F. Harrison ¹⁶⁷, N.M. Hartman ¹¹⁰, N.M. Hartmann ¹⁰⁹, Y. Hasegawa ¹⁴⁰, R. Hauser ¹⁰⁷, C.M. Hawkes ²⁰, R.J. Hawkings ³⁶, Y. Hayashi ¹⁵³, S. Hayashida ¹¹¹, D. Hayden ¹⁰⁷, C. Hayes ¹⁰⁶, R.L. Hayes ¹¹⁴, C.P. Hays ¹²⁶, J.M. Hays ⁹⁴, H.S. Hayward ⁹², F. He ^{62a}, M. He ^{14a,14e}, Y. He ¹⁵⁴, Y. He ⁴⁸, N.B. Heatley ⁹⁴, V. Hedberg ⁹⁸, A.L. Heggelund ¹²⁵, N.D. Hehir ^{94,*}, C. Heidegger ⁵⁴, K.K. Heidegger ⁵⁴, W.D. Heidorn ⁸¹, J. Heilman ³⁴, S. Heim ⁴⁸, T. Heim ^{17a}, J.G. Heinlein ¹²⁸, J.J. Heinrich ¹²³, L. Heinrich ^{110,ad}, J. Hejbal ¹³¹, L. Helary ⁴⁸, A. Held ¹⁷⁰, S. Hellesund ¹⁶, C.M. Helling ¹⁶⁴, S. Hellman ^{47a,47b}, R.C.W. Henderson ⁹¹, L. Henkelmann ³², A.M. Henriques Correia ³⁶, H. Herde ⁹⁸, Y. Hernández Jiménez ¹⁴⁵, L.M. Herrmann ²⁴, T. Herrmann ⁵⁰, G. Herten ⁵⁴, R. Hertenberger ¹⁰⁹, L. Hervas ³⁶, M.E. Hespings ¹⁰⁰, N.P. Hessey ^{156a}, H. Hibi ⁸⁵, E. Hill ¹⁵⁵, S.J. Hillier ²⁰, J.R. Hinds ¹⁰⁷, F. Hinterkeuser ²⁴, M. Hirose ¹²⁴, S. Hirose ¹⁵⁷, D. Hirschbuehl ¹⁷¹, T.G. Hitchings ¹⁰¹, B. Hiti ⁹³, J. Hobbs ¹⁴⁵, R. Hobincu ^{27e}, N. Hod ¹⁶⁹, M.C. Hodgkinson ¹³⁹, B.H. Hodgkinson ³², A. Hoecker ³⁶, D.D. Hofer ¹⁰⁶, J. Hofer ⁴⁸, T. Holm ²⁴, M. Holzbock ¹¹⁰, L.B.A.H. Hommels ³², B.P. Honan ¹⁰¹, J. Hong ^{62c}, T.M. Hong ¹²⁹, B.H. Hooberman ¹⁶², W.H. Hopkins ⁶, Y. Horii ¹¹¹, S. Hou ¹⁴⁸, A.S. Howard ⁹³, J. Howarth ⁵⁹, J. Hoya ⁶, M. Hrabovsky ¹²², A. Hrynevich ⁴⁸, T. Hryn'ova ⁴, P.J. Hsu ⁶⁵, S.-C. Hsu ¹³⁸, Q. Hu ^{62a}, Y.F. Hu ^{14a,14e}, S. Huang ^{64b}, X. Huang ^{14c}, X. Huang ^{14a,14e}, Y. Huang ¹³⁹, Y. Huang ^{14a}, Z. Huang ¹⁰¹, Z. Hubacek ¹³², M. Huebner ²⁴, F. Huegging ²⁴, T.B. Huffman ¹²⁶, C.A. Hugli ⁴⁸, M. Huhtinen ³⁶, S.K. Huiberts ¹⁶, R. Hulsken ¹⁰⁴, N. Huseynov ¹², J. Huston ¹⁰⁷, J. Huth ⁶¹, R. Hyneman ¹⁴³, G. Iacobucci ⁵⁶, G. Iakovidis ²⁹, I. Ibragimov ¹⁴¹, L. Iconomidou-Fayard ⁶⁶, P. Iengo ^{72a,72b}, R. Iguchi ¹⁵³, T. Iizawa ¹²⁶, Y. Ikegami ⁸⁴, N. Ilic ¹⁵⁵, H. Imam ^{35a}, M. Ince Lezki ⁵⁶, T. Ingebretsen Carlson ^{47a,47b}, G. Introzzi ^{73a,73b}, M. Iodice ^{77a}, V. Ippolito ^{75a,75b}, R.K. Irwin ⁹², M. Ishino ¹⁵³, W. Islam ¹⁷⁰, C. Issever ^{18,48}, S. Istin ^{21a,ak}, H. Ito ¹⁶⁸, J.M. Iturbe Ponce ^{64a}, R. Iuppa ^{78a,78b}, A. Ivina ¹⁶⁹, J.M. Izen ⁴⁵, V. Izzo ^{72a}, P. Jacka ^{131,132}, P. Jackson ¹, R.M. Jacobs ⁴⁸, B.P. Jaeger ¹⁴², C.S. Jagfeld ¹⁰⁹, G. Jain ^{156a}, P. Jain ⁵⁴, K. Jakobs ⁵⁴, T. Jakoubek ¹⁶⁹, J. Jamieson ⁵⁹, K.W. Janas ^{86a}, M. Javurkova ¹⁰³, F. Jeanneau ¹³⁵, L. Jeanty ¹²³, J. Jejelava ^{149a,z}, P. Jenni ^{54,g}, C.E. Jessiman ³⁴, S. Jézéquel ⁴, C. Jia ^{62b}, J. Jia ¹⁴⁵, X. Jia ⁶¹, X. Jia ^{14a,14e}, Z. Jia ^{14c}, S. Jiggins ⁴⁸, J. Jimenez Pena ¹³, S. Jin ^{14c}, A. Jinaru ^{27b}, O. Jinnouchi ¹⁵⁴, P. Johansson ¹³⁹, K.A. Johns ⁷, J.W. Johnson ¹³⁶, D.M. Jones ³², E. Jones ⁴⁸, P. Jones ³², R.W.L. Jones ⁹¹, T.J. Jones ⁹², H.L. Joos ^{55,36}, R. Joshi ¹¹⁹, J. Jovicevic ¹⁵, X. Ju ^{17a}, J.J. Junggeburch ¹⁰³, T. Junkermann ^{63a}, A. Juste Rozas ^{13,s}, M.K. Juzek ⁸⁷, S. Kabana ^{137e}, A. Kaczmarzka ⁸⁷, M. Kado ¹¹⁰, H. Kagan ¹¹⁹, M. Kagan ¹⁴³, A. Kahn ⁴¹, A. Kahn ¹²⁸, C. Kahra ¹⁰⁰, T. Kaji ¹⁵³, E. Kajomovitz ¹⁵⁰, N. Kakati ¹⁶⁹, I. Kalaitzidou ⁵⁴, C.W. Kalderon ²⁹, A. Kamenshchikov ¹⁵⁵, N.J. Kang ¹³⁶, D. Kar ^{33g}, K. Karava ¹²⁶, M.J. Kareem ^{156b}, E. Karentzos ⁵⁴, I. Karkanias ¹⁵², O. Karkout ¹¹⁴, S.N. Karpov ³⁸,

Z.M. Karpova ³⁸, V. Kartvelishvili ⁹¹, A.N. Karyukhin ³⁷, E. Kasimi ¹⁵², J. Katzy ⁴⁸, S. Kaur ³⁴, K. Kawade ¹⁴⁰, M.P. Kawale ¹²⁰, C. Kawamoto ⁸⁸, T. Kawamoto ^{62a}, E.F. Kay ³⁶, F.I. Kaya ¹⁵⁸, S. Kazakos ¹⁰⁷, V.F. Kazanin ³⁷, Y. Ke ¹⁴⁵, J.M. Keaveney ^{33a}, R. Keeler ¹⁶⁵, G.V. Kehris ⁶¹, J.S. Keller ³⁴, A.S. Kelly ⁹⁶, J.J. Kempster ¹⁴⁶, K.E. Kennedy ⁴¹, P.D. Kennedy ¹⁰⁰, O. Kepka ¹³¹, B.P. Kerridge ¹⁶⁷, S. Kersten ¹⁷¹, B.P. Kerševan ⁹³, S. Keshri ⁶⁶, L. Keszeghova ^{28a}, S. Ketabchi Haghghat ¹⁵⁵, R.A. Khan ¹²⁹, M. Khandoga ¹²⁷, A. Khanov ¹²¹, A.G. Kharlamov ³⁷, T. Kharlamova ³⁷, E.E. Khoda ¹³⁸, M. Kholodenko ³⁷, T.J. Khoo ¹⁸, G. Khorauli ¹⁶⁶, J. Khubua ^{149b}, Y.A.R. Khwaira ⁶⁶, A. Kilgallon ¹²³, D.W. Kim ^{47a,47b}, Y.K. Kim ³⁹, N. Kimura ⁹⁶, M.K. Kingston ⁵⁵, A. Kirchhoff ⁵⁵, C. Kirfel ²⁴, F. Kirfel ²⁴, J. Kirk ¹³⁴, A.E. Kiryunin ¹¹⁰, C. Kitsaki ¹⁰, O. Kivernyk ²⁴, M. Klassen ^{63a}, C. Klein ³⁴, L. Klein ¹⁶⁶, M.H. Klein ¹⁰⁶, M. Klein ⁹², S.B. Klein ⁵⁶, U. Klein ⁹², P. Klimek ³⁶, A. Klimentov ²⁹, T. Klioutchnikova ³⁶, P. Kluit ¹¹⁴, S. Kluth ¹¹⁰, E. Kneringer ⁷⁹, T.M. Knight ¹⁵⁵, A. Knue ⁴⁹, R. Kobayashi ⁸⁸, D. Kobylanski ¹⁶⁹, S.F. Koch ¹²⁶, M. Kocian ¹⁴³, P. Kodyš ¹³³, D.M. Koeck ¹²³, P.T. Koenig ²⁴, T. Koffas ³⁴, O. Kolay ⁵⁰, I. Koletsou ⁴, T. Komarek ¹²², K. Köneke ⁵⁴, A.X.Y. Kong ¹, T. Kono ¹¹⁸, N. Konstantinidis ⁹⁶, P. Kontaxakis ⁵⁶, B. Konya ⁹⁸, R. Kopeliansky ⁶⁸, S. Koperny ^{86a}, K. Korcyl ⁸⁷, K. Kordas ^{152,e}, G. Koren ¹⁵¹, A. Korn ⁹⁶, S. Korn ⁵⁵, I. Korolkov ¹³, N. Korotkova ³⁷, B. Kortman ¹¹⁴, O. Kortner ¹¹⁰, S. Kortner ¹¹⁰, W.H. Kostecka ¹¹⁵, V.V. Kostyukhin ¹⁴¹, A. Kotskechagia ¹³⁵, A. Kotwal ⁵¹, A. Koulouris ³⁶, A. Kourkoumeli-Charalampidi ^{73a,73b}, C. Kourkoumelis ⁹, E. Kourlitis ^{110,ad}, O. Kovanda ¹⁴⁶, R. Kowalewski ¹⁶⁵, W. Kozanecki ¹³⁵, A.S. Kozhin ³⁷, V.A. Kramarenko ³⁷, G. Kramberger ⁹³, P. Kramer ¹⁰⁰, M.W. Krasny ¹²⁷, A. Krasznahorkay ³⁶, J.W. Kraus ¹⁷¹, J.A. Kremer ⁴⁸, T. Kresse ⁵⁰, J. Kretschmar ⁹², K. Kreul ¹⁸, P. Krieger ¹⁵⁵, S. Krishnamurthy ¹⁰³, M. Krivos ¹³³, K. Krizka ²⁰, K. Kroeninger ⁴⁹, H. Kroha ¹¹⁰, J. Kroll ¹³¹, J. Kroll ¹²⁸, K.S. Krowpman ¹⁰⁷, U. Kruchonak ³⁸, H. Krüger ²⁴, N. Krumnack ⁸¹, M.C. Kruse ⁵¹, O. Kuchinskaia ³⁷, S. Kuday ^{3a}, S. Kuehn ³⁶, R. Kuesters ⁵⁴, T. Kuhl ⁴⁸, V. Kukhtin ³⁸, Y. Kulchitsky ^{37,a}, S. Kuleshov ^{137d,137b}, M. Kumar ^{33g}, N. Kumari ⁴⁸, A. Kupco ¹³¹, T. Kupfer ⁴⁹, A. Kupich ³⁷, O. Kuprash ⁵⁴, H. Kurashige ⁸⁵, L.L. Kurchaninov ^{156a}, O. Kurdysh ⁶⁶, Y.A. Kurochkin ³⁷, A. Kurova ³⁷, M. Kuze ¹⁵⁴, A.K. Kvam ¹⁰³, J. Kvita ¹²², T. Kwan ¹⁰⁴, N.G. Kyriacou ¹⁰⁶, L.A.O. Laatu ¹⁰², C. Lacasta ¹⁶³, F. Lacava ^{75a,75b}, H. Lacker ¹⁸, D. Lacour ¹²⁷, N.N. Lad ⁹⁶, E. Ladygin ³⁸, B. Laforge ¹²⁷, T. Lagouri ^{137e}, F.Z. Lahbabi ^{35a}, S. Lai ⁵⁵, I.K. Lakomic ^{86a}, N. Lalloue ⁶⁰, J.E. Lambert ¹⁶⁵, S. Lammers ⁶⁸, W. Lampl ⁷, C. Lampoudis ^{152,e}, A.N. Lancaster ¹¹⁵, E. Lançon ²⁹, U. Landgraf ⁵⁴, M.P.J. Landon ⁹⁴, V.S. Lang ⁵⁴, R.J. Langenberg ¹⁰³, O.K.B. Langrekken ¹²⁵, A.J. Lankford ¹⁶⁰, F. Lanni ³⁶, K. Lantzsck ²⁴, A. Lanza ^{73a}, A. Lapertosa ^{57b,57a}, J.F. Laporte ¹³⁵, T. Lari ^{71a}, F. Lasagni Manghi ^{23b}, M. Lassnig ³⁶, V. Latonova ¹³¹, A. Laudrain ¹⁰⁰, A. Laurier ¹⁵⁰, S.D. Lawlor ¹³⁹, Z. Lawrence ¹⁰¹, R. Lazaridou ¹⁶⁷, M. Lazzaroni ^{71a,71b}, B. Le ¹⁰¹, E.M. Le Boulicaut ⁵¹, B. Leban ⁹³, A. Lebedev ⁸¹, M. LeBlanc ¹⁰¹, F. Ledroit-Guillon ⁶⁰, A.C.A. Lee ⁹⁶, S.C. Lee ¹⁴⁸, S. Lee ^{47a,47b}, T.F. Lee ⁹², L.L. Leeuw ^{33c}, H.P. Lefebvre ⁹⁵, M. Lefebvre ¹⁶⁵, C. Leggett ^{17a}, G. Lehmann Miotto ³⁶, M. Leigh ⁵⁶, W.A. Leight ¹⁰³, W. Leinonen ¹¹³, A. Leisos ^{152,r}, M.A.L. Leite ^{83c}, C.E. Leitgeb ⁴⁸, R. Leitner ¹³³, K.J.C. Leney ⁴⁴, T. Lenz ²⁴, S. Leone ^{74a}, C. Leonidopoulos ⁵², A. Leopold ¹⁴⁴, C. Leroy ¹⁰⁸, R. Les ¹⁰⁷, C.G. Lester ³², M. Levchenko ³⁷, J. Levêque ⁴, D. Levin ¹⁰⁶, L.J. Levinson ¹⁶⁹, M.P. Lewicki ⁸⁷, D.J. Lewis ⁴, A. Li ⁵, B. Li ^{62b}, C. Li ^{62a}, C-Q. Li ¹¹⁰, H. Li ^{62a}, H. Li ^{62b}, H. Li ^{14c}, H. Li ^{14b}, H. Li ^{62b}, J. Li ^{62c}, K. Li ¹³⁸, L. Li ^{62c}, M. Li ^{14a,14e}, Q.Y. Li ^{62a}, S. Li ^{14a,14e}, S. Li ^{62d,62c,d}, T. Li ⁵, X. Li ¹⁰⁴, Z. Li ¹²⁶, Z. Li ¹⁰⁴, Z. Li ⁹², Z. Li ^{14a,14e}, S. Liang ^{14a,14e}, Z. Liang ^{14a}, M. Liberatore ¹³⁵, B. Liberti ^{76a}, K. Lie ^{64c}, J. Lieber Marin ^{83b}, H. Lien ⁶⁸, K. Lin ¹⁰⁷, R.E. Lindley ⁷, J.H. Lindon ², E. Lipeles ¹²⁸, A. Lipniacka ¹⁶,

A. Lister ¹⁶⁴, J.D. Little ⁴, B. Liu ^{14a}, B.X. Liu ¹⁴², D. Liu ^{62d,62c}, J.B. Liu ^{62a}, J.K.K. Liu ³²,
 K. Liu ^{62d,62c}, M. Liu ^{62a}, M.Y. Liu ^{62a}, P. Liu ^{14a}, Q. Liu ^{62d,138,62c}, X. Liu ^{62a}, Y. Liu ^{14d,14e},
 Y.L. Liu ^{62b}, Y.W. Liu ^{62a}, J. Llorente Merino ¹⁴², S.L. Lloyd ⁹⁴, E.M. Lobodzinska ⁴⁸,
 P. Loch ⁷, T. Lohse ¹⁸, K. Lohwasser ¹³⁹, E. Loiacono ⁴⁸, M. Lokajicek ^{131,*}, J.D. Lomas ²⁰,
 J.D. Long ¹⁶², I. Longarini ¹⁶⁰, L. Longo ^{70a,70b}, R. Longo ¹⁶², I. Lopez Paz ⁶⁷,
 A. Lopez Solis ⁴⁸, N. Lorenzo Martinez ⁴, A.M. Lory ¹⁰⁹, G. Löschke Centeno ¹⁴⁶, O. Loseva ³⁷,
 X. Lou ^{47a,47b}, X. Lou ^{14a,14e}, A. Lounis ⁶⁶, J. Love ⁶, P.A. Love ⁹¹, G. Lu ^{14a,14e}, M. Lu ⁸⁰,
 S. Lu ¹²⁸, Y.J. Lu ⁶⁵, H.J. Lubatti ¹³⁸, C. Luci ^{75a,75b}, F.L. Lucio Alves ^{14c}, A. Lucotte ⁶⁰,
 F. Luehring ⁶⁸, I. Luise ¹⁴⁵, O. Lukianchuk ⁶⁶, O. Lundberg ¹⁴⁴, B. Lund-Jensen ¹⁴⁴,
 N.A. Luongo ⁶, M.S. Lutz ¹⁵¹, A.B. Lux ²⁵, D. Lynn ²⁹, H. Lyons ⁹², R. Lysak ¹³¹, E. Lytken ⁹⁸,
 V. Lyubushkin ³⁸, T. Lyubushkina ³⁸, M.M. Lyukova ¹⁴⁵, H. Ma ²⁹, K. Ma ^{62a}, L.L. Ma ^{62b},
 W. Ma ^{62a}, Y. Ma ¹²¹, D.M. Mac Donell ¹⁶⁵, G. Maccarrone ⁵³, J.C. MacDonald ¹⁰⁰,
 P.C. Machado De Abreu Farias ^{83b}, R. Madar ⁴⁰, W.F. Mader ⁵⁰, T. Madula ⁹⁶, J. Maeda ⁸⁵,
 T. Maeno ²⁹, H. Maguire ¹³⁹, V. Maiboroda ¹³⁵, A. Maio ^{130a,130b,130d}, K. Maj ^{86a},
 O. Majersky ⁴⁸, S. Majewski ¹²³, N. Makovec ⁶⁶, V. Maksimovic ¹⁵, B. Malaescu ¹²⁷,
 Pa. Malecki ⁸⁷, V.P. Maleev ³⁷, F. Malek ⁶⁰, M. Mali ⁹³, D. Malito ⁹⁵, U. Mallik ⁸⁰,
 S. Maltezos ¹⁰, S. Malyukov ³⁸, J. Mamuzic ¹³, G. Mancini ⁵³, G. Manco ^{73a,73b}, J.P. Mandalia ⁹⁴,
 I. Mandić ⁹³, L. Manhaes de Andrade Filho ^{83a}, I.M. Maniatis ¹⁶⁹, J. Manjarres Ramos ^{102,aa},
 D.C. Mankad ¹⁶⁹, A. Mann ¹⁰⁹, B. Mansoulie ¹³⁵, S. Manzoni ³⁶, L. Mao ^{62c}, X. Mapekula ^{33c},
 A. Marantis ^{152,r}, G. Marchiori ⁵, M. Marcisovsky ¹³¹, C. Marcon ^{71a}, M. Marinescu ²⁰,
 S. Marium ⁴⁸, M. Marjanovic ¹²⁰, E.J. Marshall ⁹¹, Z. Marshall ^{17a}, S. Marti-Garcia ¹⁶³,
 T.A. Martin ¹⁶⁷, V.J. Martin ⁵², B. Martin dit Latour ¹⁶, L. Martinelli ^{75a,75b}, M. Martinez ^{13,s},
 P. Martinez Agullo ¹⁶³, V.I. Martinez Outschoorn ¹⁰³, P. Martinez Suarez ¹³, S. Martin-Haugh ¹³⁴,
 V.S. Martoiu ^{27b}, A.C. Martyniuk ⁹⁶, A. Marzin ³⁶, D. Mascione ^{78a,78b}, L. Masetti ¹⁰⁰,
 T. Mashimo ¹⁵³, J. Masik ¹⁰¹, A.L. Maslennikov ³⁷, L. Massa ^{23b}, P. Massarotti ^{72a,72b},
 P. Mastrandrea ^{74a,74b}, A. Mastroberardino ^{43b,43a}, T. Masubuchi ¹⁵³, T. Mathisen ¹⁶¹,
 J. Matousek ¹³³, N. Matsuzawa ¹⁵³, J. Maurer ^{27b}, B. Maček ⁹³, D.A. Maximov ³⁷, R. Mazini ¹⁴⁸,
 I. Maznas ¹⁵², M. Mazza ¹⁰⁷, S.M. Mazza ¹³⁶, E. Mazzeo ^{71a,71b}, C. Mc Ginn ²⁹,
 J.P. Mc Gowan ¹⁰⁴, S.P. Mc Kee ¹⁰⁶, E.F. McDonald ¹⁰⁵, A.E. McDougall ¹¹⁴, J.A. Mcfayden ¹⁴⁶,
 R.P. McGovern ¹²⁸, G. Mchedlidze ^{149b}, R.P. Mckenzie ^{33g}, T.C. Mclachlan ⁴⁸,
 D.J. McLaughlin ⁹⁶, S.J. McMahon ¹³⁴, C.M. Mcpartland ⁹², R.A. McPherson ^{165,w},
 S. Mehlhase ¹⁰⁹, A. Mehta ⁹², D. Melini ¹⁵⁰, B.R. Mellado Garcia ^{33g}, A.H. Melo ⁵⁵,
 F. Meloni ⁴⁸, A.M. Mendes Jacques Da Costa ¹⁰¹, H.Y. Meng ¹⁵⁵, L. Meng ⁹¹, S. Menke ¹¹⁰,
 M. Mentink ³⁶, E. Meoni ^{43b,43a}, G. Mercado ¹¹⁵, C. Merlassino ^{69a,69c}, L. Merola ^{72a,72b},
 C. Meroni ^{71a,71b}, G. Merz ¹⁰⁶, O. Meshkov ³⁷, J. Metcalfe ⁶, A.S. Mete ⁶, C. Meyer ⁶⁸,
 J-P. Meyer ¹³⁵, R.P. Middleton ¹³⁴, L. Mijović ⁵², G. Mikenberg ¹⁶⁹, M. Mikestikova ¹³¹,
 M. Mikuž ⁹³, H. Mildner ¹⁰⁰, A. Milic ³⁶, C.D. Milke ⁴⁴, D.W. Miller ³⁹, L.S. Miller ³⁴,
 A. Milov ¹⁶⁹, D.A. Milstead ^{47a,47b}, T. Min ^{14c}, A.A. Minaenko ³⁷, I.A. Minashvili ^{149b}, L. Mince ⁵⁹,
 A.I. Mincer ¹¹⁷, B. Mindur ^{86a}, M. Mineev ³⁸, Y. Mino ⁸⁸, L.M. Mir ¹³, M. Miralles Lopez ¹⁶³,
 M. Mironova ^{17a}, A. Mishima ¹⁵³, M.C. Missio ¹¹³, A. Mitra ¹⁶⁷, V.A. Mitsou ¹⁶³,
 Y. Mitsumori ¹¹¹, O. Miu ¹⁵⁵, P.S. Miyagawa ⁹⁴, T. Mkrtchyan ^{63a}, M. Mlinarevic ⁹⁶,
 T. Mlinarevic ⁹⁶, M. Mlynarikova ³⁶, S. Mobius ¹⁹, P. Moder ⁴⁸, P. Mogg ¹⁰⁹,
 A.F. Mohammed ^{14a,14e}, S. Mohapatra ⁴¹, G. Mokgatitwane ^{33g}, L. Moleri ¹⁶⁹, B. Mondal ¹⁴¹,
 S. Mondal ¹³², K. Mönig ⁴⁸, E. Monnier ¹⁰², L. Monsonis Romero ¹⁶³, J. Montejo Berlingen ¹³,
 M. Montella ¹¹⁹, F. Montekali ^{77a,77b}, F. Monticelli ⁹⁰, S. Monzani ^{69a,69c}, N. Morange ⁶⁶,
 A.L. Moreira De Carvalho ^{130a}, M. Moreno Llácer ¹⁶³, C. Moreno Martinez ⁵⁶, P. Morettini ^{57b},
 S. Morgenstern ³⁶, M. Morii ⁶¹, M. Morinaga ¹⁵³, A.K. Morley ³⁶, F. Morodei ^{75a,75b},

L. Morvaj ³⁶, P. Moschovakos ³⁶, B. Moser ³⁶, M. Mosidze ^{149b}, T. Moskalets ⁵⁴,
 P. Moskvitina ¹¹³, J. Moss ^{31,1}, E.J.W. Moyses ¹⁰³, O. Mtintsilana ^{33g}, S. Muanza ¹⁰²,
 J. Mueller ¹²⁹, D. Muenstermann ⁹¹, R. Müller ¹⁹, G.A. Mullier ¹⁶¹, A.J. Mullin ³², J.J. Mullin ¹²⁸,
 D.P. Mungo ¹⁵⁵, D. Munoz Perez ¹⁶³, F.J. Munoz Sanchez ¹⁰¹, M. Murin ¹⁰¹, W.J. Murray ^{167,134},
 A. Murrone ^{71a,71b}, M. Muškinja ^{17a}, C. Mwewa ²⁹, A.G. Myagkov ^{37,a}, A.J. Myers ⁸,
 G. Myers ⁶⁸, M. Myska ¹³², B.P. Nachman ^{17a}, O. Nackenhorst ⁴⁹, A. Nag ⁵⁰, K. Nagai ¹²⁶,
 K. Nagano ⁸⁴, J.L. Nagle ^{29,ai}, E. Nagy ¹⁰², A.M. Nairz ³⁶, Y. Nakahama ⁸⁴, K. Nakamura ⁸⁴,
 K. Nakkalil ⁵, H. Nanjo ¹²⁴, R. Narayan ⁴⁴, E.A. Narayanan ¹¹², I. Naryshkin ³⁷, M. Naseri ³⁴,
 S. Nasri ¹⁵⁹, C. Nass ²⁴, G. Navarro ^{22a}, J. Navarro-Gonzalez ¹⁶³, R. Nayak ¹⁵¹, A. Nayaz ¹⁸,
 P.Y. Nechaeva ³⁷, F. Nechansky ⁴⁸, L. Nedic ¹²⁶, T.J. Neep ²⁰, A. Negri ^{73a,73b}, M. Negrini ^{23b},
 C. Nellist ¹¹⁴, C. Nelson ¹⁰⁴, K. Nelson ¹⁰⁶, S. Nemecek ¹³¹, M. Nessi ^{36,h}, M.S. Neubauer ¹⁶²,
 F. Neuhaus ¹⁰⁰, J. Neundorf ⁴⁸, R. Newhouse ¹⁶⁴, P.R. Newman ²⁰, C.W. Ng ¹²⁹, Y.W.Y. Ng ⁴⁸,
 B. Ngair ^{35e}, H.D.N. Nguyen ¹⁰⁸, R.B. Nickerson ¹²⁶, R. Nicolaidou ¹³⁵, J. Nielsen ¹³⁶,
 M. Niemeyer ⁵⁵, J. Niermann ^{55,36}, N. Nikiforou ³⁶, V. Nikolaenko ^{37,a}, I. Nikolic-Audit ¹²⁷,
 K. Nikolopoulos ²⁰, P. Nilsson ²⁹, I. Ninca ⁴⁸, H.R. Nindhito ⁵⁶, G. Ninio ¹⁵¹, A. Nisati ^{75a},
 N. Nishu ², R. Nisius ¹¹⁰, J-E. Nitschke ⁵⁰, E.K. Nkadimeng ^{33g}, T. Nobe ¹⁵³, D.L. Noel ³²,
 T. Nommensen ¹⁴⁷, M.B. Norfolk ¹³⁹, R.R.B. Norisam ⁹⁶, B.J. Norman ³⁴, M. Noury ^{35a},
 J. Novak ⁹³, T. Novak ⁴⁸, L. Novotny ¹³², R. Novotny ¹¹², L. Nozka ¹²², K. Ntekas ¹⁶⁰,
 N.M.J. Nunes De Moura Junior ^{83b}, E. Nurse ⁹⁶, J. Ocariz ¹²⁷, A. Ochi ⁸⁵, I. Ochoa ^{130a},
 S. Oerdek ⁴⁸, J.T. Offermann ³⁹, A. Ogrodnik ¹³³, A. Oh ¹⁰¹, C.C. Ohm ¹⁴⁴, H. Oide ⁸⁴,
 R. Oishi ¹⁵³, M.L. Ojeda ⁴⁸, M.W. O'Keefe ⁹², Y. Okumura ¹⁵³, L.F. Oleiro Seabra ^{130a},
 S.A. Olivares Pino ^{137d}, D. Oliveira Damazio ²⁹, D. Oliveira Goncalves ^{83a}, J.L. Oliver ¹⁶⁰,
 Ö.O. Öncel ⁵⁴, A.P. O'Neill ¹⁹, A. Onofre ^{130a,130e}, P.U.E. Onyisi ¹¹, M.J. Oreglia ³⁹,
 G.E. Orellana ⁹⁰, D. Orestano ^{77a,77b}, N. Orlando ¹³, R.S. Orr ¹⁵⁵, V. O'Shea ⁵⁹,
 L.M. Osojnak ¹²⁸, R. Ospanov ^{62a}, G. Otero y Garzon ³⁰, H. Otono ⁸⁹, P.S. Ott ^{63a},
 G.J. Ottino ^{17a}, M. Ouchrif ^{35d}, J. Ouellette ²⁹, F. Ould-Saada ¹²⁵, M. Owen ⁵⁹, R.E. Owen ¹³⁴,
 K.Y. Oyulmaz ^{21a}, V.E. Ozcan ^{21a}, F. Ozturk ⁸⁷, N. Ozturk ⁸, S. Ozturk ⁸², H.A. Pacey ¹²⁶,
 A. Pacheco Pages ¹³, C. Padilla Aranda ¹³, G. Padovano ^{75a,75b}, S. Pagan Griso ^{17a},
 G. Palacino ⁶⁸, A. Palazzo ^{70a,70b}, S. Palestini ³⁶, J. Pan ¹⁷², T. Pan ^{64a}, D.K. Panchal ¹¹,
 C.E. Pandini ¹¹⁴, J.G. Panduro Vazquez ⁹⁵, H.D. Pandya ¹, H. Pang ^{14b}, P. Pani ⁴⁸,
 G. Panizzo ^{69a,69c}, L. Paolozzi ⁵⁶, C. Papadatos ¹⁰⁸, S. Parajuli ⁴⁴, A. Paramonov ⁶,
 C. Paraskevopoulos ¹⁰, D. Paredes Hernandez ^{64b}, K.R. Park ⁴¹, T.H. Park ¹⁵⁵, M.A. Parker ³²,
 F. Parodi ^{57b,57a}, E.W. Parrish ¹¹⁵, V.A. Parrish ⁵², J.A. Parsons ⁴¹, U. Parzefall ⁵⁴,
 B. Pascual Dias ¹⁰⁸, L. Pascual Dominguez ¹⁵¹, E. Pasqualucci ^{75a}, S. Passaggio ^{57b}, F. Pastore ⁹⁵,
 P. Pasuwan ^{47a,47b}, P. Patel ⁸⁷, U.M. Patel ⁵¹, J.R. Pater ¹⁰¹, T. Pauly ³⁶, J. Pearkes ¹⁴³,
 M. Pedersen ¹²⁵, R. Pedro ^{130a}, S.V. Peleganchuk ³⁷, O. Penc ³⁶, E.A. Pender ⁵²,
 K.E. Pensi ¹⁰⁹, M. Penzin ³⁷, B.S. Peralva ^{83d}, A.P. Pereira Peixoto ⁶⁰, L. Pereira Sanchez ^{47a,47b},
 D.V. Perepelitsa ^{29,ai}, E. Perez Codina ^{156a}, M. Perganti ¹⁰, L. Perini ^{71a,71b,*}, H. Pernegger ³⁶,
 O. Perrin ⁴⁰, K. Peters ⁴⁸, R.F.Y. Peters ¹⁰¹, B.A. Petersen ³⁶, T.C. Petersen ⁴², E. Petit ¹⁰²,
 V. Petousis ¹³², C. Petridou ^{152,e}, A. Petrukhin ¹⁴¹, M. Pettee ^{17a}, N.E. Pettersson ³⁶,
 A. Petukhov ³⁷, K. Petukhova ¹³³, R. Pezoa ^{137f}, L. Pezzotti ³⁶, G. Pezzullo ¹⁷², T.M. Pham ¹⁷⁰,
 T. Pham ¹⁰⁵, P.W. Phillips ¹³⁴, G. Piacquadio ¹⁴⁵, E. Pianori ^{17a}, F. Piazza ¹²³, R. Piegai ³⁰,
 D. Pietreanu ^{27b}, A.D. Pilkington ¹⁰¹, M. Pinamonti ^{69a,69c}, J.L. Pinfeld ²,
 B.C. Pinheiro Pereira ^{130a}, A.E. Pinto Pinoargote ^{100,135}, L. Pintucci ^{69a,69c}, K.M. Piper ¹⁴⁶,
 A. Pirttikoski ⁵⁶, D.A. Pizzi ³⁴, L. Pizzimento ^{64b}, A. Pizzini ¹¹⁴, M.-A. Pleier ²⁹, V. Plesanovs ⁵⁴,
 V. Pleskot ¹³³, E. Plotnikova ³⁸, G. Poddar ⁴, R. Poettgen ⁹⁸, L. Poggioli ¹²⁷, I. Pokharel ⁵⁵,
 S. Polacek ¹³³, G. Polesello ^{73a}, A. Poley ^{142,156a}, R. Polifka ¹³², A. Polini ^{23b}, C.S. Pollard ¹⁶⁷,

Z.B. Pollock ¹¹⁹, V. Polychronakos ²⁹, E. Pompa Pacchi ^{75a,75b}, D. Ponomarenko ¹¹³,
L. Pontecorvo ³⁶, S. Popa ^{27a}, G.A. Popeneciu ^{27d}, A. Poreba ³⁶, D.M. Portillo Quintero ^{156a},
S. Pospisil ¹³², M.A. Postill ¹³⁹, P. Postolache ^{27c}, K. Potamianos ¹⁶⁷, P.A. Potepa ^{86a},
I.N. Potrap ³⁸, C.J. Potter ³², H. Potti ¹, T. Poulsen ⁴⁸, J. Poveda ¹⁶³, M.E. Pozo Astigarraga ³⁶,
A. Prades Ibanez ¹⁶³, J. Pretel ⁵⁴, D. Price ¹⁰¹, M. Primavera ^{70a}, M.A. Principe Martin ⁹⁹,
R. Privara ¹²², T. Procter ⁵⁹, M.L. Proffitt ¹³⁸, N. Proklova ¹²⁸, K. Prokofiev ^{64c}, G. Proto ¹¹⁰,
S. Protopopescu ²⁹, J. Proudfoot ⁶, M. Przybycien ^{86a}, W.W. Przygoda ^{86b}, J.E. Puddefoot ¹³⁹,
D. Pudzha ³⁷, D. Pyatiizbyantseva ³⁷, J. Qian ¹⁰⁶, D. Qichen ¹⁰¹, Y. Qin ¹⁰¹, T. Qiu ⁵²,
A. Quadt ⁵⁵, M. Queitsch-Maitland ¹⁰¹, G. Quetant ⁵⁶, R.P. Quinn ¹⁶⁴, G. Rabanal Bolanos ⁶¹,
D. Rafanoharana ⁵⁴, F. Ragusa ^{71a,71b}, J.L. Rainbolt ³⁹, J.A. Raine ⁵⁶, S. Rajagopalan ²⁹,
E. Ramakoti ³⁷, I.A. Ramirez-Berend ³⁴, K. Ran ^{48,14e}, N.P. Rapheeha ^{33g}, H. Rasheed ^{27b},
V. Raskina ¹²⁷, D.F. Rassloff ^{63a}, S. Rave ¹⁰⁰, B. Ravina ⁵⁵, I. Ravinovich ¹⁶⁹, M. Raymond ³⁶,
A.L. Read ¹²⁵, N.P. Readioff ¹³⁹, D.M. Rebutzi ^{73a,73b}, G. Redlinger ²⁹, A.S. Reed ¹¹⁰,
K. Reeves ²⁶, J.A. Reidelsturz ¹⁷¹, D. Reikher ¹⁵¹, A. Rej ⁴⁹, C. Rembser ³⁶, A. Renardi ⁴⁸,
M. Renda ^{27b}, M.B. Rendel ¹¹⁰, F. Renner ⁴⁸, A.G. Rennie ¹⁶⁰, A.L. Rescia ⁴⁸, S. Resconi ^{71a},
M. Ressegotti ^{57b,57a}, S. Rettie ³⁶, J.G. Reyes Rivera ¹⁰⁷, E. Reynolds ^{17a}, O.L. Rezanova ³⁷,
P. Reznicek ¹³³, N. Ribaric ⁹¹, E. Ricci ^{78a,78b}, R. Richter ¹¹⁰, S. Richter ^{47a,47b},
E. Richter-Was ^{86b}, M. Ridel ¹²⁷, S. Ridouani ^{35d}, P. Rieck ¹¹⁷, P. Riedler ³⁶, E.M. Riefel ^{47a,47b},
J.O. Rieger ¹¹⁴, M. Rijssenbeek ¹⁴⁵, A. Rimoldi ^{73a,73b}, M. Rimoldi ³⁶, L. Rinaldi ^{23b,23a},
T.T. Rinn ²⁹, M.P. Rinnagel ¹⁰⁹, G. Ripellino ¹⁶¹, I. Riu ¹³, P. Rivadeneira ⁴⁸,
J.C. Rivera Vergara ¹⁶⁵, F. Rizatdinova ¹²¹, E. Rizvi ⁹⁴, B.A. Roberts ¹⁶⁷, B.R. Roberts ^{17a},
S.H. Robertson ^{104,w}, D. Robinson ³², C.M. Robles Gajardo ^{137f}, M. Robles Manzano ¹⁰⁰,
A. Robson ⁵⁹, A. Rocchi ^{76a,76b}, C. Roda ^{74a,74b}, S. Rodriguez Bosca ^{63a}, Y. Rodriguez Garcia ^{22a},
A. Rodriguez Rodriguez ⁵⁴, A.M. Rodríguez Vera ^{156b}, S. Roe ³⁶, J.T. Roemer ¹⁶⁰,
A.R. Roepe-Gier ¹³⁶, J. Roggel ¹⁷¹, O. Røhne ¹²⁵, R.A. Rojas ¹⁰³, C.P.A. Roland ¹²⁷,
J. Roloff ²⁹, A. Romaniouk ³⁷, E. Romano ^{73a,73b}, M. Romano ^{23b}, A.C. Romero Hernandez ¹⁶²,
N. Rompotis ⁹², L. Roos ¹²⁷, S. Rosati ^{75a}, B.J. Rosser ³⁹, E. Rossi ¹²⁶, E. Rossi ^{72a,72b},
L.P. Rossi ^{57b}, L. Rossini ⁵⁴, R. Rosten ¹¹⁹, M. Rotaru ^{27b}, B. Rottler ⁵⁴, C. Rougier ^{102,aa},
D. Rousseau ⁶⁶, D. Rousso ³², A. Roy ¹⁶², S. Roy-Garand ¹⁵⁵, A. Rozanov ¹⁰²,
Z.M.A. Rozario ⁵⁹, Y. Rozen ¹⁵⁰, X. Ruan ^{33g}, A. Rubio Jimenez ¹⁶³, A.J. Ruby ⁹²,
V.H. Ruelas Rivera ¹⁸, T.A. Ruggeri ¹, A. Ruggiero ¹²⁶, A. Ruiz-Martinez ¹⁶³, A. Rummler ³⁶,
Z. Rurikova ⁵⁴, N.A. Rusakovich ³⁸, H.L. Russell ¹⁶⁵, G. Russo ^{75a,75b}, J.P. Rutherford ⁷,
S. Rutherford Colmenares ³², K. Rybacki ⁹¹, M. Rybar ¹³³, E.B. Rye ¹²⁵, A. Ryzhov ⁴⁴,
J.A. Sabater Iglesias ⁵⁶, P. Sabatini ¹⁶³, H.F.W. Sadrozinski ¹³⁶, F. Safai Tehrani ^{75a},
B. Safarzadeh Samani ¹³⁴, M. Safdari ¹⁴³, S. Saha ¹⁶⁵, M. Sahinsoy ¹¹⁰, A. Saibel ¹⁶³,
M. Saimpert ¹³⁵, M. Saito ¹⁵³, T. Saito ¹⁵³, D. Salamani ³⁶, A. Salnikov ¹⁴³, J. Salt ¹⁶³,
A. Salvador Salas ¹⁵¹, D. Salvatore ^{43b,43a}, F. Salvatore ¹⁴⁶, A. Salzburger ³⁶, D. Sammel ⁵⁴,
D. Sampsonidis ^{152,e}, D. Sampsonidou ¹²³, J. Sánchez ¹⁶³, A. Sanchez Pineda ⁴,
V. Sanchez Sebastian ¹⁶³, H. Sandaker ¹²⁵, C.O. Sander ⁴⁸, J.A. Sandesara ¹⁰³, M. Sandhoff ¹⁷¹,
C. Sandoval ^{22b}, D.P.C. Sankey ¹³⁴, T. Sano ⁸⁸, A. Sansoni ⁵³, L. Santi ^{75a,75b}, C. Santoni ⁴⁰,
H. Santos ^{130a,130b}, S.N. Santpur ^{17a}, A. Santra ¹⁶⁹, K.A. Saoucha ^{116b}, J.G. Saraiva ^{130a,130d},
J. Sardain ⁷, O. Sasaki ⁸⁴, K. Sato ¹⁵⁷, C. Sauer ^{63b}, F. Sauerburger ⁵⁴, E. Sauvan ⁴,
P. Savard ^{155,af}, R. Sawada ¹⁵³, C. Sawyer ¹³⁴, L. Sawyer ⁹⁷, I. Sayago Galvan ¹⁶³, C. Sbarra ^{23b},
A. Sbrizzi ^{23b,23a}, T. Scanlon ⁹⁶, J. Schaarschmidt ¹³⁸, P. Schacht ¹¹⁰, U. Schäfer ¹⁰⁰,
A.C. Schaffer ^{66,44}, D. Schaile ¹⁰⁹, R.D. Schamberger ¹⁴⁵, C. Scharf ¹⁸, M.M. Schefer ¹⁹,
V.A. Schegelsky ³⁷, D. Scheirich ¹³³, F. Schenck ¹⁸, M. Schernau ¹⁶⁰, C. Scheulen ⁵⁵,
C. Schiavi ^{57b,57a}, E.J. Schioppa ^{70a,70b}, M. Schioppa ^{43b,43a}, B. Schlag ^{143,n}, K.E. Schleicher ⁵⁴,

S. Schlenker ³⁶, J. Schmeing ¹⁷¹, M.A. Schmidt ¹⁷¹, K. Schmieden ¹⁰⁰, C. Schmitt ¹⁰⁰,
 N. Schmitt ¹⁰⁰, S. Schmitt ⁴⁸, L. Schoeffel ¹³⁵, A. Schoening ^{63b}, P.G. Scholer ⁵⁴, E. Schopf ¹²⁶,
 M. Schott ¹⁰⁰, J. Schovancova ³⁶, S. Schramm ⁵⁶, F. Schroeder ¹⁷¹, T. Schroer ⁵⁶,
 H-C. Schultz-Coulon ^{63a}, M. Schumacher ⁵⁴, B.A. Schumm ¹³⁶, Ph. Schune ¹³⁵, A.J. Schuy ¹³⁸,
 H.R. Schwartz ¹³⁶, A. Schwartzman ¹⁴³, T.A. Schwarz ¹⁰⁶, Ph. Schwemling ¹³⁵,
 R. Schwienhorst ¹⁰⁷, A. Sciandra ¹³⁶, G. Sciolla ²⁶, F. Scuri ^{74a}, C.D. Sebastiani ⁹²,
 K. Sedlaczek ¹¹⁵, P. Seema ¹⁸, S.C. Seidel ¹¹², A. Seiden ¹³⁶, B.D. Seidlitz ⁴¹, C. Seitz ⁴⁸,
 J.M. Seixas ^{83b}, G. Sekhniaidze ^{72a}, S.J. Sekula ⁴⁴, L. Selem ⁶⁰, N. Semprini-Cesari ^{23b,23a},
 D. Sengupta ⁵⁶, V. Senthilkumar ¹⁶³, L. Serin ⁶⁶, L. Serkin ^{69a,69b}, M. Sessa ^{76a,76b},
 H. Severini ¹²⁰, F. Sforza ^{57b,57a}, A. Sfyra ⁵⁶, E. Shabalina ⁵⁵, R. Shaheen ¹⁴⁴,
 J.D. Shahinian ¹²⁸, D. Shaked Renous ¹⁶⁹, L.Y. Shan ^{14a}, M. Shapiro ^{17a}, A. Sharma ³⁶,
 A.S. Sharma ¹⁶⁴, P. Sharma ⁸⁰, S. Sharma ⁴⁸, P.B. Shatalov ³⁷, K. Shaw ¹⁴⁶, S.M. Shaw ¹⁰¹,
 A. Shcherbakova ³⁷, Q. Shen ^{62c,5}, D.J. Sheppard ¹⁴², P. Sherwood ⁹⁶, L. Shi ⁹⁶, X. Shi ^{14a},
 C.O. Shimmin ¹⁷², J.D. Shinner ⁹⁵, I.P.J. Shipsey ¹²⁶, S. Shirabe ^{56,h}, M. Shiyakova ^{38,u},
 J. Shlomi ¹⁶⁹, M.J. Shochet ³⁹, J. Shojaii ¹⁰⁵, D.R. Shope ¹²⁵, B. Shrestha ¹²⁰, S. Shrestha ^{119,aj},
 E.M. Shrif ^{33g}, M.J. Shroff ¹⁶⁵, P. Sicho ¹³¹, A.M. Sickles ¹⁶², E. Sideras Haddad ^{33g},
 A. Sidoti ^{23b}, F. Siegert ⁵⁰, Dj. Sijacki ¹⁵, F. Sili ⁹⁰, J.M. Silva ²⁰, M.V. Silva Oliveira ²⁹,
 S.B. Silverstein ^{47a}, S. Simion ⁶⁶, R. Simoniello ³⁶, E.L. Simpson ⁵⁹, H. Simpson ¹⁴⁶,
 L.R. Simpson ¹⁰⁶, N.D. Simpson ⁹⁸, S. Simsek ⁸², S. Sindhu ⁵⁵, P. Sinervo ¹⁵⁵, S. Singh ¹⁵⁵,
 S. Sinha ⁴⁸, S. Sinha ¹⁰¹, M. Sioli ^{23b,23a}, I. Siral ³⁶, E. Sitnikova ⁴⁸, S.Yu. Sivoklov ^{37,*},
 J. Sjölin ^{47a,47b}, A. Skaf ⁵⁵, E. Skorda ²⁰, P. Skubic ¹²⁰, M. Slawinska ⁸⁷, V. Smakhtin ¹⁶⁹,
 B.H. Smart ¹³⁴, J. Smiesko ³⁶, S.Yu. Smirnov ³⁷, Y. Smirnov ³⁷, L.N. Smirnova ^{37,a},
 O. Smirnova ⁹⁸, A.C. Smith ⁴¹, E.A. Smith ³⁹, H.A. Smith ¹²⁶, J.L. Smith ⁹², R. Smith ¹⁴³,
 M. Smizanska ⁹¹, K. Smolek ¹³², A.A. Snesev ³⁷, S.R. Snider ¹⁵⁵, H.L. Snoek ¹¹⁴,
 S. Snyder ²⁹, R. Sobie ^{165,w}, A. Soffer ¹⁵¹, C.A. Solans Sanchez ³⁶, E.Yu. Soldatov ³⁷,
 U. Soldevila ¹⁶³, A.A. Solodkov ³⁷, S. Solomon ²⁶, A. Soloshenko ³⁸, K. Solovieva ⁵⁴,
 O.V. Solovyanov ⁴⁰, V. Solovyev ³⁷, P. Sommer ³⁶, A. Sonay ¹³, W.Y. Song ^{156b},
 J.M. Sonneveld ¹¹⁴, A. Sopczak ¹³², A.L. Sopio ⁹⁶, F. Sopkova ^{28b}, I.R. Sotarriva Alvarez ¹⁵⁴,
 V. Sothilingam ^{63a}, O.J. Soto Sandoval ^{137c,137b}, S. Sottocornola ⁶⁸, R. Soualah ^{116b},
 Z. Soumami ^{35e}, D. South ⁴⁸, N. Soybelman ¹⁶⁹, S. Spagnolo ^{70a,70b}, M. Spalla ¹¹⁰,
 D. Sperlich ⁵⁴, G. Spigo ³⁶, S. Spinali ⁹¹, D.P. Spiteri ⁵⁹, M. Spousta ¹³³, E.J. Staats ³⁴,
 A. Stabile ^{71a,71b}, R. Stamen ^{63a}, A. Stampekis ²⁰, M. Standke ²⁴, E. Stanecka ⁸⁷,
 M.V. Stange ⁵⁰, B. Stanislaus ^{17a}, M.M. Stanitzki ⁴⁸, B. Stapf ⁴⁸, E.A. Starchenko ³⁷,
 G.H. Stark ¹³⁶, J. Stark ^{102,aa}, D.M. Starke ^{156b}, P. Staroba ¹³¹, P. Starovoitov ^{63a}, S. Stärz ¹⁰⁴,
 R. Staszewski ⁸⁷, G. Stavropoulos ⁴⁶, J. Steentoft ¹⁶¹, P. Steinberg ²⁹, B. Stelzer ^{142,156a},
 H.J. Stelzer ¹²⁹, O. Stelzer-Chilton ^{156a}, H. Stenzel ⁵⁸, T.J. Stevenson ¹⁴⁶, G.A. Stewart ³⁶,
 J.R. Stewart ¹²¹, M.C. Stockton ³⁶, G. Stoica ^{27b}, M. Stolarski ^{130a}, S. Stonjek ¹¹⁰,
 A. Straessner ⁵⁰, J. Strandberg ¹⁴⁴, S. Strandberg ^{47a,47b}, M. Stratmann ¹⁷¹, M. Strauss ¹²⁰,
 T. Streblner ¹⁰², P. Strizenec ^{28b}, R. Ströhmer ¹⁶⁶, D.M. Strom ¹²³, R. Stroynowski ⁴⁴,
 A. Strubig ^{47a,47b}, S.A. Stucci ²⁹, B. Stugu ¹⁶, J. Stupak ¹²⁰, N.A. Styles ⁴⁸, D. Su ¹⁴³,
 S. Su ^{62a}, W. Su ^{62d}, X. Su ^{62a,66}, K. Sugizaki ¹⁵³, V.V. Sulim ³⁷, M.J. Sullivan ⁹²,
 D.M.S. Sultan ^{78a,78b}, L. Sultanaliyeva ³⁷, S. Sultansoy ^{3b}, T. Sumida ⁸⁸, S. Sun ¹⁰⁶, S. Sun ¹⁷⁰,
 O. Sunneborn Gudnadottir ¹⁶¹, N. Sur ¹⁰², M.R. Sutton ¹⁴⁶, H. Suzuki ¹⁵⁷, M. Svatos ¹³¹,
 M. Swiatlowski ^{156a}, T. Swirski ¹⁶⁶, I. Sykora ^{28a}, M. Sykora ¹³³, T. Sykora ¹³³, D. Ta ¹⁰⁰,
 K. Tackmann ^{48,t}, A. Taffard ¹⁶⁰, R. Tafirout ^{156a}, J.S. Tafuya Vargas ⁶⁶, E.P. Takeva ⁵²,
 Y. Takubo ⁸⁴, M. Talby ¹⁰², A.A. Talyshev ³⁷, K.C. Tam ^{64b}, N.M. Tamir ¹⁵¹, A. Tanaka ¹⁵³,
 J. Tanaka ¹⁵³, R. Tanaka ⁶⁶, M. Tanasini ^{57b,57a}, Z. Tao ¹⁶⁴, S. Tapia Araya ^{137f},

S. Tapprogge ¹⁰⁰, A. Tarek Abouelfadl Mohamed ¹⁰⁷, S. Tarem ¹⁵⁰, K. Tariq ^{14a}, G. Tarna ^{102,27b},
 G.F. Tartarelli ^{71a}, P. Tas ¹³³, M. Tasevsky ¹³¹, E. Tassi ^{43b,43a}, A.C. Tate ¹⁶², G. Tateno ¹⁵³,
 Y. Tayalati ^{35e,v}, G.N. Taylor ¹⁰⁵, W. Taylor ^{156b}, A.S. Tee ¹⁷⁰, R. Teixeira De Lima ¹⁴³,
 P. Teixeira-Dias ⁹⁵, J.J. Teoh ¹⁵⁵, K. Terashi ¹⁵³, J. Terron ⁹⁹, S. Terzo ¹³, M. Testa ⁵³,
 R.J. Teuscher ^{155,w}, A. Thaler ⁷⁹, O. Theiner ⁵⁶, N. Themistokleous ⁵², T. Theveneaux-Pelzer ¹⁰²,
 O. Thielmann ¹⁷¹, D.W. Thomas ⁹⁵, J.P. Thomas ²⁰, E.A. Thompson ^{17a}, P.D. Thompson ²⁰,
 E. Thomson ¹²⁸, Y. Tian ⁵⁵, V. Tikhomirov ^{37,a}, Yu.A. Tikhonov ³⁷, S. Timoshenko ³⁷,
 D. Timoshyn ¹³³, E.X.L. Ting ¹, P. Tipton ¹⁷², S.H. Tlou ^{33g}, A. Tnourji ⁴⁰, K. Todome ¹⁵⁴,
 S. Todorova-Nova ¹³³, S. Todt ⁵⁰, M. Togawa ⁸⁴, J. Tojo ⁸⁹, S. Tokár ^{28a}, K. Tokushuku ⁸⁴,
 O. Toldaiev ⁶⁸, R. Tombs ³², M. Tomoto ^{84,111}, L. Tompkins ^{143,n}, K.W. Topolnicki ^{86b},
 E. Torrence ¹²³, H. Torres ^{102,aa}, E. Torró Pastor ¹⁶³, M. Toscani ³⁰, C. Tosciri ³⁹, M. Tost ¹¹,
 D.R. Tovey ¹³⁹, A. Traeet ¹⁶, I.S. Trandafir ^{27b}, T. Trefzger ¹⁶⁶, A. Tricoli ²⁹, I.M. Trigger ^{156a},
 S. Trincaz-Duvoid ¹²⁷, D.A. Trischuk ²⁶, B. Trocmé ⁶⁰, C. Troncon ^{71a}, L. Truong ^{33c},
 M. Trzebinski ⁸⁷, A. Trzupiek ⁸⁷, F. Tsai ¹⁴⁵, M. Tsai ¹⁰⁶, A. Tsiamis ^{152,e}, P.V. Tsiareshka ³⁷,
 S. Tsigaridas ^{156a}, A. Tsirigotis ^{152,r}, V. Tsiskaridze ¹⁵⁵, E.G. Tskhadadze ^{149a},
 M. Tsopoulou ^{152,e}, Y. Tsujikawa ⁸⁸, I.I. Tsukerman ³⁷, V. Tsulaia ^{17a}, S. Tsuno ⁸⁴, K. Tsurii ¹¹⁸,
 D. Tsybychev ¹⁴⁵, Y. Tu ^{64b}, A. Tudorache ^{27b}, V. Tudorache ^{27b}, A.N. Tuna ⁶¹,
 S. Turchikhin ^{57b,57a}, I. Turk Cakir ^{3a}, R. Turra ^{71a}, T. Turtuvshin ^{38,x}, P.M. Tuts ⁴¹,
 S. Tzamarias ^{152,e}, P. Tzanis ¹⁰, E. Tzovara ¹⁰⁰, F. Ukegawa ¹⁵⁷, P.A. Ulloa Poblete ^{137c,137b},
 E.N. Umaka ²⁹, G. Unal ³⁶, M. Unal ¹¹, A. Undrus ²⁹, G. Unel ¹⁶⁰, J. Urban ^{28b},
 P. Urquijo ¹⁰⁵, P. Urrejola ^{137a}, G. Usai ⁸, R. Ushioda ¹⁵⁴, M. Usman ¹⁰⁸, Z. Uysal ^{21b},
 V. Vacek ¹³², B. Vachon ¹⁰⁴, K.O.H. Vadla ¹²⁵, T. Vafeiadis ³⁶, A. Vaitkus ⁹⁶, C. Valderanis ¹⁰⁹,
 E. Valdes Santurio ^{47a,47b}, M. Valente ^{156a}, S. Valentinetti ^{23b,23a}, A. Valero ¹⁶³,
 E. Valiente Moreno ¹⁶³, A. Vallier ^{102,aa}, J.A. Valls Ferrer ¹⁶³, D.R. Van Arneman ¹¹⁴,
 T.R. Van Daalen ¹³⁸, A. Van Der Graaf ⁴⁹, P. Van Gemmeren ⁶, M. Van Rijnbach ^{125,36},
 S. Van Stroud ⁹⁶, I. Van Vulpen ¹¹⁴, M. Vanadia ^{76a,76b}, W. Vandelli ³⁶, M. Vandenbroucke ¹³⁵,
 E.R. Vandewall ¹²¹, D. Vannicola ¹⁵¹, L. Vannoli ^{57b,57a}, R. Vari ^{75a}, E.W. Varnes ⁷,
 C. Varni ^{17b}, T. Varol ¹⁴⁸, D. Varouchas ⁶⁶, L. Varriale ¹⁶³, K.E. Varvell ¹⁴⁷, M.E. Vasile ^{27b},
 L. Vaslin ⁸⁴, G.A. Vasquez ¹⁶⁵, A. Vasyukov ³⁸, F. Vazeille ⁴⁰, T. Vazquez Schroeder ³⁶,
 J. Veatch ³¹, V. Vecchio ¹⁰¹, M.J. Veen ¹⁰³, I. Veliscek ¹²⁶, L.M. Veloce ¹⁵⁵, F. Veloso ^{130a,130c},
 S. Veneziano ^{75a}, A. Ventura ^{70a,70b}, S. Ventura Gonzalez ¹³⁵, A. Verbytskyi ¹¹⁰,
 M. Verducci ^{74a,74b}, C. Vergis ²⁴, M. Verissimo De Araujo ^{83b}, W. Verkerke ¹¹⁴,
 J.C. Vermeulen ¹¹⁴, C. Vernieri ¹⁴³, M. Vessella ¹⁰³, M.C. Vetterli ^{142,af}, A. Vgenopoulos ^{152,e},
 N. Viaux Maira ^{137f}, T. Vickey ¹³⁹, O.E. Vickey Boeriu ¹³⁹, G.H.A. Viehhauser ¹²⁶, L. Vignani ^{63b},
 M. Villa ^{23b,23a}, M. Villaplana Perez ¹⁶³, E.M. Villhauer ⁵², E. Vilucchi ⁵³, M.G. Vincter ³⁴,
 G.S. Virdee ²⁰, A. Vishwakarma ⁵², A. Visibile ¹¹⁴, C. Vittori ³⁶, I. Vivarelli ¹⁴⁶,
 E. Voevodina ¹¹⁰, F. Vogel ¹⁰⁹, J.C. Voigt ⁵⁰, P. Vokac ¹³², Yu. Volkotrub ^{86a}, J. Von Ahnen ⁴⁸,
 E. Von Toerne ²⁴, B. Vormwald ³⁶, V. Vorobel ¹³³, K. Vorobev ³⁷, M. Vos ¹⁶³, K. Voss ¹⁴¹,
 J.H. Vossebeld ⁹², M. Vozak ¹¹⁴, L. Vozdecky ⁹⁴, N. Vranjes ¹⁵, M. Vranjes Milosavljevic ¹⁵,
 M. Vreeswijk ¹¹⁴, R. Vuillermet ³⁶, O. Vujanovic ¹⁰⁰, I. Vukotic ³⁹, S. Wada ¹⁵⁷, C. Wagner ¹⁰³,
 J.M. Wagner ^{17a}, W. Wagner ¹⁷¹, S. Wahdan ¹⁷¹, H. Wahlberg ⁹⁰, M. Wakida ¹¹¹, J. Walder ¹³⁴,
 R. Walker ¹⁰⁹, W. Walkowiak ¹⁴¹, A. Wall ¹²⁸, T. Wamorkar ⁶, A.Z. Wang ¹³⁶, C. Wang ¹⁰⁰,
 C. Wang ^{62c}, H. Wang ^{17a}, J. Wang ^{64a}, R.-J. Wang ¹⁰⁰, R. Wang ⁶¹, R. Wang ⁶,
 S.M. Wang ¹⁴⁸, S. Wang ^{62b}, T. Wang ^{62a}, W.T. Wang ⁸⁰, W. Wang ^{14a}, X. Wang ^{14c},
 X. Wang ¹⁶², X. Wang ^{62c}, Y. Wang ^{62d}, Y. Wang ^{14c}, Z. Wang ¹⁰⁶, Z. Wang ^{62d,51,62c},
 Z. Wang ¹⁰⁶, A. Warburton ¹⁰⁴, R.J. Ward ²⁰, N. Warrack ⁵⁹, A.T. Watson ²⁰, H. Watson ⁵⁹,
 M.F. Watson ²⁰, E. Watton ^{59,134}, G. Watts ¹³⁸, B.M. Waugh ⁹⁶, C. Weber ²⁹, H.A. Weber ¹⁸,

M.S. Weber ¹⁹, S.M. Weber ^{63a}, C. Wei ^{62a}, Y. Wei ¹²⁶, A.R. Weidberg ¹²⁶, E.J. Weik ¹¹⁷, J. Weingarten ⁴⁹, M. Weirich ¹⁰⁰, C. Weiser ⁵⁴, C.J. Wells ⁴⁸, T. Wenaus ²⁹, B. Wendland ⁴⁹, T. Wengler ³⁶, N.S. Wenke ¹¹⁰, N. Wermes ²⁴, M. Wessels ^{63a}, A.M. Wharton ⁹¹, A.S. White ⁶¹, A. White ⁸, M.J. White ¹, D. Whiteson ¹⁶⁰, L. Wickremasinghe ¹²⁴, W. Wiedenmann ¹⁷⁰, C. Wiel ⁵⁰, M. Wielers ¹³⁴, C. Wiglesworth ⁴², D.J. Wilbern ¹²⁰, H.G. Wilkens ³⁶, D.M. Williams ⁴¹, H.H. Williams ¹²⁸, S. Williams ³², S. Willocq ¹⁰³, B.J. Wilson ¹⁰¹, P.J. Windischhofer ³⁹, F.I. Winkel ³⁰, F. Winklmeier ¹²³, B.T. Winter ⁵⁴, J.K. Winter ¹⁰¹, M. Wittgen ¹⁴³, M. Wobisch ⁹⁷, Z. Wolffs ¹¹⁴, J. Wollrath ¹⁶⁰, M.W. Wolter ⁸⁷, H. Wolters ^{130a,130c}, A.F. Wongel ⁴⁸, E.L. Woodward ⁴¹, S.D. Worm ⁴⁸, B.K. Wosiek ⁸⁷, K.W. Woźniak ⁸⁷, S. Wozniowski ⁵⁵, K. Wraight ⁵⁹, C. Wu ²⁰, J. Wu ^{14a,14e}, M. Wu ^{64a}, M. Wu ¹¹³, S.L. Wu ¹⁷⁰, X. Wu ⁵⁶, Y. Wu ^{62a}, Z. Wu ¹³⁵, J. Wuerzinger ^{110,ad}, T.R. Wyatt ¹⁰¹, B.M. Wynne ⁵², S. Xella ⁴², L. Xia ^{14c}, M. Xia ^{14b}, J. Xiang ^{64c}, M. Xie ^{62a}, X. Xie ^{62a}, S. Xin ^{14a,14e}, A. Xiong ¹²³, J. Xiong ^{17a}, D. Xu ^{14a}, H. Xu ^{62a}, L. Xu ^{62a}, R. Xu ¹²⁸, T. Xu ¹⁰⁶, Y. Xu ^{14b}, Z. Xu ⁵², Z. Xu ^{14c}, B. Yabsley ¹⁴⁷, S. Yacoob ^{33a}, Y. Yamaguchi ¹⁵⁴, E. Yamashita ¹⁵³, H. Yamauchi ¹⁵⁷, T. Yamazaki ^{17a}, Y. Yamazaki ⁸⁵, J. Yan ^{62c}, S. Yan ¹²⁶, Z. Yan ²⁵, H.J. Yang ^{62c,62d}, H.T. Yang ^{62a}, S. Yang ^{62a}, T. Yang ^{64c}, X. Yang ³⁶, X. Yang ^{14a}, Y. Yang ⁴⁴, Y. Yang ^{62a}, Z. Yang ^{62a}, W-M. Yao ^{17a}, Y.C. Yap ⁴⁸, H. Ye ^{14c}, H. Ye ⁵⁵, J. Ye ^{14a}, S. Ye ²⁹, X. Ye ^{62a}, Y. Yeh ⁹⁶, I. Yeletsikh ³⁸, B.K. Yeo ^{17b}, M.R. Yexley ⁹⁶, P. Yin ⁴¹, K. Yorita ¹⁶⁸, S. Younas ^{27b}, C.J.S. Young ³⁶, C. Young ¹⁴³, C. Yu ^{14a,14e,ah}, Y. Yu ^{62a}, M. Yuan ¹⁰⁶, R. Yuan ^{62b}, L. Yue ⁹⁶, M. Zaazoua ^{62a}, B. Zabinski ⁸⁷, E. Zaid ⁵², Z.K. Zak ⁸⁷, T. Zakareishvili ^{149b}, N. Zakharchuk ³⁴, S. Zambito ⁵⁶, J.A. Zamora Saa ^{137d,137b}, J. Zang ¹⁵³, D. Zanzi ⁵⁴, O. Zaplatilek ¹³², C. Zeitnitz ¹⁷¹, H. Zeng ^{14a}, J.C. Zeng ¹⁶², D.T. Zenger Jr ²⁶, O. Zenin ³⁷, T. Ženiš ^{28a}, S. Zenz ⁹⁴, S. Zerradi ^{35a}, D. Zerwas ⁶⁶, M. Zhai ^{14a,14e}, B. Zhang ^{14c}, D.F. Zhang ¹³⁹, J. Zhang ^{62b}, J. Zhang ⁶, K. Zhang ^{14a,14e}, L. Zhang ^{14c}, P. Zhang ^{14a,14e}, R. Zhang ¹⁷⁰, S. Zhang ¹⁰⁶, S. Zhang ⁴⁴, T. Zhang ¹⁵³, X. Zhang ^{62c}, X. Zhang ^{62b}, Y. Zhang ^{62c,5}, Y. Zhang ⁹⁶, Y. Zhang ^{14c}, Z. Zhang ^{17a}, Z. Zhang ⁶⁶, H. Zhao ¹³⁸, T. Zhao ^{62b}, Y. Zhao ¹³⁶, Z. Zhao ^{62a}, A. Zhemchugov ³⁸, J. Zheng ^{14c}, K. Zheng ¹⁶², X. Zheng ^{62a}, Z. Zheng ¹⁴³, D. Zhong ¹⁶², B. Zhou ¹⁰⁶, H. Zhou ⁷, N. Zhou ^{62c}, Y. Zhou ⁷, C.G. Zhu ^{62b}, J. Zhu ¹⁰⁶, Y. Zhu ^{62c}, Y. Zhu ^{62a}, X. Zhuang ^{14a}, K. Zhukov ³⁷, V. Zhulanov ³⁷, N.I. Zimine ³⁸, J. Zinsser ^{63b}, M. Ziolkowski ¹⁴¹, L. Živković ¹⁵, A. Zoccoli ^{23b,23a}, K. Zoch ⁶¹, T.G. Zorbas ¹³⁹, O. Zormpa ⁴⁶, W. Zou ⁴¹, L. Zwalinski ³⁶.

¹Department of Physics, University of Adelaide, Adelaide; Australia.

²Department of Physics, University of Alberta, Edmonton AB; Canada.

^{3(a)}Department of Physics, Ankara University, Ankara; ^(b)Division of Physics, TOBB University of Economics and Technology, Ankara; Türkiye.

⁴LAPP, Université Savoie Mont Blanc, CNRS/IN2P3, Annecy; France.

⁵APC, Université Paris Cité, CNRS/IN2P3, Paris; France.

⁶High Energy Physics Division, Argonne National Laboratory, Argonne IL; United States of America.

⁷Department of Physics, University of Arizona, Tucson AZ; United States of America.

⁸Department of Physics, University of Texas at Arlington, Arlington TX; United States of America.

⁹Physics Department, National and Kapodistrian University of Athens, Athens; Greece.

¹⁰Physics Department, National Technical University of Athens, Zografou; Greece.

¹¹Department of Physics, University of Texas at Austin, Austin TX; United States of America.

¹²Institute of Physics, Azerbaijan Academy of Sciences, Baku; Azerbaijan.

¹³Institut de Física d'Altes Energies (IFAE), Barcelona Institute of Science and Technology, Barcelona; Spain.

- ^{14(a)}Institute of High Energy Physics, Chinese Academy of Sciences, Beijing; ^(b)Physics Department, Tsinghua University, Beijing; ^(c)Department of Physics, Nanjing University, Nanjing; ^(d)School of Science, Shenzhen Campus of Sun Yat-sen University; ^(e)University of Chinese Academy of Science (UCAS), Beijing; China.
- ¹⁵Institute of Physics, University of Belgrade, Belgrade; Serbia.
- ¹⁶Department for Physics and Technology, University of Bergen, Bergen; Norway.
- ^{17(a)}Physics Division, Lawrence Berkeley National Laboratory, Berkeley CA; ^(b)University of California, Berkeley CA; United States of America.
- ¹⁸Institut für Physik, Humboldt Universität zu Berlin, Berlin; Germany.
- ¹⁹Albert Einstein Center for Fundamental Physics and Laboratory for High Energy Physics, University of Bern, Bern; Switzerland.
- ²⁰School of Physics and Astronomy, University of Birmingham, Birmingham; United Kingdom.
- ^{21(a)}Department of Physics, Bogazici University, Istanbul; ^(b)Department of Physics Engineering, Gaziantep University, Gaziantep; ^(c)Department of Physics, Istanbul University, Istanbul; Türkiye.
- ^{22(a)}Facultad de Ciencias y Centro de Investigaciones, Universidad Antonio Nariño, Bogotá; ^(b)Departamento de Física, Universidad Nacional de Colombia, Bogotá; Colombia.
- ^{23(a)}Dipartimento di Fisica e Astronomia A. Righi, Università di Bologna, Bologna; ^(b)INFN Sezione di Bologna; Italy.
- ²⁴Physikalisches Institut, Universität Bonn, Bonn; Germany.
- ²⁵Department of Physics, Boston University, Boston MA; United States of America.
- ²⁶Department of Physics, Brandeis University, Waltham MA; United States of America.
- ^{27(a)}Transilvania University of Brasov, Brasov; ^(b)Horia Hulubei National Institute of Physics and Nuclear Engineering, Bucharest; ^(c)Department of Physics, Alexandru Ioan Cuza University of Iasi, Iasi; ^(d)National Institute for Research and Development of Isotopic and Molecular Technologies, Physics Department, Cluj-Napoca; ^(e)National University of Science and Technology Politehnica, Bucharest; ^(f)West University in Timisoara, Timisoara; ^(g)Faculty of Physics, University of Bucharest, Bucharest; Romania.
- ^{28(a)}Faculty of Mathematics, Physics and Informatics, Comenius University, Bratislava; ^(b)Department of Subnuclear Physics, Institute of Experimental Physics of the Slovak Academy of Sciences, Kosice; Slovak Republic.
- ²⁹Physics Department, Brookhaven National Laboratory, Upton NY; United States of America.
- ³⁰Universidad de Buenos Aires, Facultad de Ciencias Exactas y Naturales, Departamento de Física, y CONICET, Instituto de Física de Buenos Aires (IFIBA), Buenos Aires; Argentina.
- ³¹California State University, CA; United States of America.
- ³²Cavendish Laboratory, University of Cambridge, Cambridge; United Kingdom.
- ^{33(a)}Department of Physics, University of Cape Town, Cape Town; ^(b)iThemba Labs, Western Cape; ^(c)Department of Mechanical Engineering Science, University of Johannesburg, Johannesburg; ^(d)National Institute of Physics, University of the Philippines Diliman (Philippines); ^(e)University of South Africa, Department of Physics, Pretoria; ^(f)University of Zululand, KwaDlangezwa; ^(g)School of Physics, University of the Witwatersrand, Johannesburg; South Africa.
- ³⁴Department of Physics, Carleton University, Ottawa ON; Canada.
- ^{35(a)}Faculté des Sciences Ain Chock, Réseau Universitaire de Physique des Hautes Energies - Université Hassan II, Casablanca; ^(b)Faculté des Sciences, Université Ibn-Tofail, Kénitra; ^(c)Faculté des Sciences Semlalia, Université Cadi Ayyad, LPHEA-Marrakech; ^(d)LPMR, Faculté des Sciences, Université Mohamed Premier, Oujda; ^(e)Faculté des sciences, Université Mohammed V, Rabat; ^(f)Institute of Applied Physics, Mohammed VI Polytechnic University, Ben Guerir; Morocco.
- ³⁶CERN, Geneva; Switzerland.
- ³⁷Affiliated with an institute covered by a cooperation agreement with CERN.

- ³⁸Affiliated with an international laboratory covered by a cooperation agreement with CERN.
- ³⁹Enrico Fermi Institute, University of Chicago, Chicago IL; United States of America.
- ⁴⁰LPC, Université Clermont Auvergne, CNRS/IN2P3, Clermont-Ferrand; France.
- ⁴¹Nevis Laboratory, Columbia University, Irvington NY; United States of America.
- ⁴²Niels Bohr Institute, University of Copenhagen, Copenhagen; Denmark.
- ⁴³(^a)Dipartimento di Fisica, Università della Calabria, Rende; (^b)INFN Gruppo Collegato di Cosenza, Laboratori Nazionali di Frascati; Italy.
- ⁴⁴Physics Department, Southern Methodist University, Dallas TX; United States of America.
- ⁴⁵Physics Department, University of Texas at Dallas, Richardson TX; United States of America.
- ⁴⁶National Centre for Scientific Research "Demokritos", Agia Paraskevi; Greece.
- ⁴⁷(^a)Department of Physics, Stockholm University; (^b)Oskar Klein Centre, Stockholm; Sweden.
- ⁴⁸Deutsches Elektronen-Synchrotron DESY, Hamburg and Zeuthen; Germany.
- ⁴⁹Fakultät Physik, Technische Universität Dortmund, Dortmund; Germany.
- ⁵⁰Institut für Kern- und Teilchenphysik, Technische Universität Dresden, Dresden; Germany.
- ⁵¹Department of Physics, Duke University, Durham NC; United States of America.
- ⁵²SUPA - School of Physics and Astronomy, University of Edinburgh, Edinburgh; United Kingdom.
- ⁵³INFN e Laboratori Nazionali di Frascati, Frascati; Italy.
- ⁵⁴Physikalisches Institut, Albert-Ludwigs-Universität Freiburg, Freiburg; Germany.
- ⁵⁵II. Physikalisches Institut, Georg-August-Universität Göttingen, Göttingen; Germany.
- ⁵⁶Département de Physique Nucléaire et Corpusculaire, Université de Genève, Genève; Switzerland.
- ⁵⁷(^a)Dipartimento di Fisica, Università di Genova, Genova; (^b)INFN Sezione di Genova; Italy.
- ⁵⁸II. Physikalisches Institut, Justus-Liebig-Universität Giessen, Giessen; Germany.
- ⁵⁹SUPA - School of Physics and Astronomy, University of Glasgow, Glasgow; United Kingdom.
- ⁶⁰LPSC, Université Grenoble Alpes, CNRS/IN2P3, Grenoble INP, Grenoble; France.
- ⁶¹Laboratory for Particle Physics and Cosmology, Harvard University, Cambridge MA; United States of America.
- ⁶²(^a)Department of Modern Physics and State Key Laboratory of Particle Detection and Electronics, University of Science and Technology of China, Hefei; (^b)Institute of Frontier and Interdisciplinary Science and Key Laboratory of Particle Physics and Particle Irradiation (MOE), Shandong University, Qingdao; (^c)School of Physics and Astronomy, Shanghai Jiao Tong University, Key Laboratory for Particle Astrophysics and Cosmology (MOE), SKLPPC, Shanghai; (^d)Tsung-Dao Lee Institute, Shanghai; (^e)School of Physics and Microelectronics, Zhengzhou University; China.
- ⁶³(^a)Kirchhoff-Institut für Physik, Ruprecht-Karls-Universität Heidelberg, Heidelberg; (^b)Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg; Germany.
- ⁶⁴(^a)Department of Physics, Chinese University of Hong Kong, Shatin, N.T., Hong Kong; (^b)Department of Physics, University of Hong Kong, Hong Kong; (^c)Department of Physics and Institute for Advanced Study, Hong Kong University of Science and Technology, Clear Water Bay, Kowloon, Hong Kong; China.
- ⁶⁵Department of Physics, National Tsing Hua University, Hsinchu; Taiwan.
- ⁶⁶IJCLab, Université Paris-Saclay, CNRS/IN2P3, 91405, Orsay; France.
- ⁶⁷Centro Nacional de Microelectrónica (IMB-CNM-CSIC), Barcelona; Spain.
- ⁶⁸Department of Physics, Indiana University, Bloomington IN; United States of America.
- ⁶⁹(^a)INFN Gruppo Collegato di Udine, Sezione di Trieste, Udine; (^b)ICTP, Trieste; (^c)Dipartimento Politecnico di Ingegneria e Architettura, Università di Udine, Udine; Italy.
- ⁷⁰(^a)INFN Sezione di Lecce; (^b)Dipartimento di Matematica e Fisica, Università del Salento, Lecce; Italy.
- ⁷¹(^a)INFN Sezione di Milano; (^b)Dipartimento di Fisica, Università di Milano, Milano; Italy.
- ⁷²(^a)INFN Sezione di Napoli; (^b)Dipartimento di Fisica, Università di Napoli, Napoli; Italy.
- ⁷³(^a)INFN Sezione di Pavia; (^b)Dipartimento di Fisica, Università di Pavia, Pavia; Italy.

- ^{74(a)}INFN Sezione di Pisa; ^(b)Dipartimento di Fisica E. Fermi, Università di Pisa, Pisa; Italy.
- ^{75(a)}INFN Sezione di Roma; ^(b)Dipartimento di Fisica, Sapienza Università di Roma, Roma; Italy.
- ^{76(a)}INFN Sezione di Roma Tor Vergata; ^(b)Dipartimento di Fisica, Università di Roma Tor Vergata, Roma; Italy.
- ^{77(a)}INFN Sezione di Roma Tre; ^(b)Dipartimento di Matematica e Fisica, Università Roma Tre, Roma; Italy.
- ^{78(a)}INFN-TIFPA; ^(b)Università degli Studi di Trento, Trento; Italy.
- ⁷⁹Universität Innsbruck, Department of Astro and Particle Physics, Innsbruck; Austria.
- ⁸⁰University of Iowa, Iowa City IA; United States of America.
- ⁸¹Department of Physics and Astronomy, Iowa State University, Ames IA; United States of America.
- ⁸²Istinye University, Sariyer, Istanbul; Türkiye.
- ^{83(a)}Departamento de Engenharia Elétrica, Universidade Federal de Juiz de Fora (UFJF), Juiz de Fora; ^(b)Universidade Federal do Rio De Janeiro COPPE/EE/IF, Rio de Janeiro; ^(c)Instituto de Física, Universidade de São Paulo, São Paulo; ^(d)Rio de Janeiro State University, Rio de Janeiro; Brazil.
- ⁸⁴KEK, High Energy Accelerator Research Organization, Tsukuba; Japan.
- ⁸⁵Graduate School of Science, Kobe University, Kobe; Japan.
- ^{86(a)}AGH University of Krakow, Faculty of Physics and Applied Computer Science, Krakow; ^(b)Marian Smoluchowski Institute of Physics, Jagiellonian University, Krakow; Poland.
- ⁸⁷Institute of Nuclear Physics Polish Academy of Sciences, Krakow; Poland.
- ⁸⁸Faculty of Science, Kyoto University, Kyoto; Japan.
- ⁸⁹Research Center for Advanced Particle Physics and Department of Physics, Kyushu University, Fukuoka ; Japan.
- ⁹⁰Instituto de Física La Plata, Universidad Nacional de La Plata and CONICET, La Plata; Argentina.
- ⁹¹Physics Department, Lancaster University, Lancaster; United Kingdom.
- ⁹²Oliver Lodge Laboratory, University of Liverpool, Liverpool; United Kingdom.
- ⁹³Department of Experimental Particle Physics, Jožef Stefan Institute and Department of Physics, University of Ljubljana, Ljubljana; Slovenia.
- ⁹⁴School of Physics and Astronomy, Queen Mary University of London, London; United Kingdom.
- ⁹⁵Department of Physics, Royal Holloway University of London, Egham; United Kingdom.
- ⁹⁶Department of Physics and Astronomy, University College London, London; United Kingdom.
- ⁹⁷Louisiana Tech University, Ruston LA; United States of America.
- ⁹⁸Fysiska institutionen, Lunds universitet, Lund; Sweden.
- ⁹⁹Departamento de Física Teórica C-15 and CIAFF, Universidad Autónoma de Madrid, Madrid; Spain.
- ¹⁰⁰Institut für Physik, Universität Mainz, Mainz; Germany.
- ¹⁰¹School of Physics and Astronomy, University of Manchester, Manchester; United Kingdom.
- ¹⁰²CPPM, Aix-Marseille Université, CNRS/IN2P3, Marseille; France.
- ¹⁰³Department of Physics, University of Massachusetts, Amherst MA; United States of America.
- ¹⁰⁴Department of Physics, McGill University, Montreal QC; Canada.
- ¹⁰⁵School of Physics, University of Melbourne, Victoria; Australia.
- ¹⁰⁶Department of Physics, University of Michigan, Ann Arbor MI; United States of America.
- ¹⁰⁷Department of Physics and Astronomy, Michigan State University, East Lansing MI; United States of America.
- ¹⁰⁸Group of Particle Physics, University of Montreal, Montreal QC; Canada.
- ¹⁰⁹Fakultät für Physik, Ludwig-Maximilians-Universität München, München; Germany.
- ¹¹⁰Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), München; Germany.
- ¹¹¹Graduate School of Science and Kobayashi-Maskawa Institute, Nagoya University, Nagoya; Japan.
- ¹¹²Department of Physics and Astronomy, University of New Mexico, Albuquerque NM; United States of America.

America.

¹¹³Institute for Mathematics, Astrophysics and Particle Physics, Radboud University/Nikhef, Nijmegen; Netherlands.

¹¹⁴Nikhef National Institute for Subatomic Physics and University of Amsterdam, Amsterdam; Netherlands.

¹¹⁵Department of Physics, Northern Illinois University, DeKalb IL; United States of America.

¹¹⁶(^a)New York University Abu Dhabi, Abu Dhabi; (^b)University of Sharjah, Sharjah; United Arab Emirates.

¹¹⁷Department of Physics, New York University, New York NY; United States of America.

¹¹⁸Ochanomizu University, Otsuka, Bunkyo-ku, Tokyo; Japan.

¹¹⁹Ohio State University, Columbus OH; United States of America.

¹²⁰Homer L. Dodge Department of Physics and Astronomy, University of Oklahoma, Norman OK; United States of America.

¹²¹Department of Physics, Oklahoma State University, Stillwater OK; United States of America.

¹²²Palacký University, Joint Laboratory of Optics, Olomouc; Czech Republic.

¹²³Institute for Fundamental Science, University of Oregon, Eugene, OR; United States of America.

¹²⁴Graduate School of Science, Osaka University, Osaka; Japan.

¹²⁵Department of Physics, University of Oslo, Oslo; Norway.

¹²⁶Department of Physics, Oxford University, Oxford; United Kingdom.

¹²⁷LPNHE, Sorbonne Université, Université Paris Cité, CNRS/IN2P3, Paris; France.

¹²⁸Department of Physics, University of Pennsylvania, Philadelphia PA; United States of America.

¹²⁹Department of Physics and Astronomy, University of Pittsburgh, Pittsburgh PA; United States of America.

¹³⁰(^a)Laboratório de Instrumentação e Física Experimental de Partículas - LIP, Lisboa; (^b)Departamento de Física, Faculdade de Ciências, Universidade de Lisboa, Lisboa; (^c)Departamento de Física, Universidade de Coimbra, Coimbra; (^d)Centro de Física Nuclear da Universidade de Lisboa, Lisboa; (^e)Departamento de Física, Universidade do Minho, Braga; (^f)Departamento de Física Teórica y del Cosmos, Universidad de Granada, Granada (Spain); (^g)Departamento de Física, Instituto Superior Técnico, Universidade de Lisboa, Lisboa; Portugal.

¹³¹Institute of Physics of the Czech Academy of Sciences, Prague; Czech Republic.

¹³²Czech Technical University in Prague, Prague; Czech Republic.

¹³³Charles University, Faculty of Mathematics and Physics, Prague; Czech Republic.

¹³⁴Particle Physics Department, Rutherford Appleton Laboratory, Didcot; United Kingdom.

¹³⁵IRFU, CEA, Université Paris-Saclay, Gif-sur-Yvette; France.

¹³⁶Santa Cruz Institute for Particle Physics, University of California Santa Cruz, Santa Cruz CA; United States of America.

¹³⁷(^a)Departamento de Física, Pontificia Universidad Católica de Chile, Santiago; (^b)Millennium Institute for Subatomic physics at high energy frontier (SAPHIR), Santiago; (^c)Instituto de Investigación Multidisciplinario en Ciencia y Tecnología, y Departamento de Física, Universidad de La Serena; (^d)Universidad Andres Bello, Department of Physics, Santiago; (^e)Instituto de Alta Investigación, Universidad de Tarapacá, Arica; (^f)Departamento de Física, Universidad Técnica Federico Santa María, Valparaíso; Chile.

¹³⁸Department of Physics, University of Washington, Seattle WA; United States of America.

¹³⁹Department of Physics and Astronomy, University of Sheffield, Sheffield; United Kingdom.

¹⁴⁰Department of Physics, Shinshu University, Nagano; Japan.

¹⁴¹Department Physik, Universität Siegen, Siegen; Germany.

¹⁴²Department of Physics, Simon Fraser University, Burnaby BC; Canada.

- ¹⁴³SLAC National Accelerator Laboratory, Stanford CA; United States of America.
- ¹⁴⁴Department of Physics, Royal Institute of Technology, Stockholm; Sweden.
- ¹⁴⁵Departments of Physics and Astronomy, Stony Brook University, Stony Brook NY; United States of America.
- ¹⁴⁶Department of Physics and Astronomy, University of Sussex, Brighton; United Kingdom.
- ¹⁴⁷School of Physics, University of Sydney, Sydney; Australia.
- ¹⁴⁸Institute of Physics, Academia Sinica, Taipei; Taiwan.
- ¹⁴⁹^(a)E. Andronikashvili Institute of Physics, Iv. Javakhishvili Tbilisi State University, Tbilisi; ^(b)High Energy Physics Institute, Tbilisi State University, Tbilisi; ^(c)University of Georgia, Tbilisi; Georgia.
- ¹⁵⁰Department of Physics, Technion, Israel Institute of Technology, Haifa; Israel.
- ¹⁵¹Raymond and Beverly Sackler School of Physics and Astronomy, Tel Aviv University, Tel Aviv; Israel.
- ¹⁵²Department of Physics, Aristotle University of Thessaloniki, Thessaloniki; Greece.
- ¹⁵³International Center for Elementary Particle Physics and Department of Physics, University of Tokyo, Tokyo; Japan.
- ¹⁵⁴Department of Physics, Tokyo Institute of Technology, Tokyo; Japan.
- ¹⁵⁵Department of Physics, University of Toronto, Toronto ON; Canada.
- ¹⁵⁶^(a)TRIUMF, Vancouver BC; ^(b)Department of Physics and Astronomy, York University, Toronto ON; Canada.
- ¹⁵⁷Division of Physics and Tomonaga Center for the History of the Universe, Faculty of Pure and Applied Sciences, University of Tsukuba, Tsukuba; Japan.
- ¹⁵⁸Department of Physics and Astronomy, Tufts University, Medford MA; United States of America.
- ¹⁵⁹United Arab Emirates University, Al Ain; United Arab Emirates.
- ¹⁶⁰Department of Physics and Astronomy, University of California Irvine, Irvine CA; United States of America.
- ¹⁶¹Department of Physics and Astronomy, University of Uppsala, Uppsala; Sweden.
- ¹⁶²Department of Physics, University of Illinois, Urbana IL; United States of America.
- ¹⁶³Instituto de Física Corpuscular (IFIC), Centro Mixto Universidad de Valencia - CSIC, Valencia; Spain.
- ¹⁶⁴Department of Physics, University of British Columbia, Vancouver BC; Canada.
- ¹⁶⁵Department of Physics and Astronomy, University of Victoria, Victoria BC; Canada.
- ¹⁶⁶Fakultät für Physik und Astronomie, Julius-Maximilians-Universität Würzburg, Würzburg; Germany.
- ¹⁶⁷Department of Physics, University of Warwick, Coventry; United Kingdom.
- ¹⁶⁸Waseda University, Tokyo; Japan.
- ¹⁶⁹Department of Particle Physics and Astrophysics, Weizmann Institute of Science, Rehovot; Israel.
- ¹⁷⁰Department of Physics, University of Wisconsin, Madison WI; United States of America.
- ¹⁷¹Fakultät für Mathematik und Naturwissenschaften, Fachgruppe Physik, Bergische Universität Wuppertal, Wuppertal; Germany.
- ¹⁷²Department of Physics, Yale University, New Haven CT; United States of America.
- ^a Also Affiliated with an institute covered by a cooperation agreement with CERN.
- ^b Also at An-Najah National University, Nablus; Palestine.
- ^c Also at Borough of Manhattan Community College, City University of New York, New York NY; United States of America.
- ^d Also at Center for High Energy Physics, Peking University; China.
- ^e Also at Center for Interdisciplinary Research and Innovation (CIRI-AUTH), Thessaloniki; Greece.
- ^f Also at Centro Studi e Ricerche Enrico Fermi; Italy.
- ^g Also at CERN, Geneva; Switzerland.
- ^h Also at Département de Physique Nucléaire et Corpusculaire, Université de Genève, Genève; Switzerland.

- ⁱ Also at Departament de Fisica de la Universitat Autònoma de Barcelona, Barcelona; Spain.
- ^j Also at Department of Financial and Management Engineering, University of the Aegean, Chios; Greece.
- ^k Also at Department of Physics, Ben Gurion University of the Negev, Beer Sheva; Israel.
- ^l Also at Department of Physics, California State University, Sacramento; United States of America.
- ^m Also at Department of Physics, King's College London, London; United Kingdom.
- ⁿ Also at Department of Physics, Stanford University, Stanford CA; United States of America.
- ^o Also at Department of Physics, University of Fribourg, Fribourg; Switzerland.
- ^p Also at Department of Physics, University of Thessaly; Greece.
- ^q Also at Department of Physics, Westmont College, Santa Barbara; United States of America.
- ^r Also at Hellenic Open University, Patras; Greece.
- ^s Also at Institutio Catalana de Recerca i Estudis Avancats, ICREA, Barcelona; Spain.
- ^t Also at Institut für Experimentalphysik, Universität Hamburg, Hamburg; Germany.
- ^u Also at Institute for Nuclear Research and Nuclear Energy (INRNE) of the Bulgarian Academy of Sciences, Sofia; Bulgaria.
- ^v Also at Institute of Applied Physics, Mohammed VI Polytechnic University, Ben Guerir; Morocco.
- ^w Also at Institute of Particle Physics (IPP); Canada.
- ^x Also at Institute of Physics and Technology, Mongolian Academy of Sciences, Ulaanbaatar; Mongolia.
- ^y Also at Institute of Physics, Azerbaijan Academy of Sciences, Baku; Azerbaijan.
- ^z Also at Institute of Theoretical Physics, Ilia State University, Tbilisi; Georgia.
- ^{aa} Also at L2IT, Université de Toulouse, CNRS/IN2P3, UPS, Toulouse; France.
- ^{ab} Also at Lawrence Livermore National Laboratory, Livermore; United States of America.
- ^{ac} Also at National Institute of Physics, University of the Philippines Diliman (Philippines); Philippines.
- ^{ad} Also at Technical University of Munich, Munich; Germany.
- ^{ae} Also at The Collaborative Innovation Center of Quantum Matter (CICQM), Beijing; China.
- ^{af} Also at TRIUMF, Vancouver BC; Canada.
- ^{ag} Also at Università di Napoli Parthenope, Napoli; Italy.
- ^{ah} Also at University of Chinese Academy of Sciences (UCAS), Beijing; China.
- ^{ai} Also at University of Colorado Boulder, Department of Physics, Colorado; United States of America.
- ^{aj} Also at Washington College, Chestertown, MD; United States of America.
- ^{ak} Also at Yeditepe University, Physics Department, Istanbul; Türkiye.
- * Deceased