

# Luminosity measurement with the ATLAS experiment at the LHC ICHEP 2020

Joey Carter, on behalf of the ATLAS Collaboration

University of Toronto

July 29, 2020

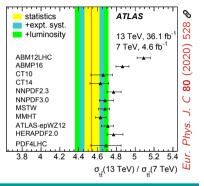






# Introduction

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- Online luminosity with accuracy of ~ 5 % was achieved in Run 2: required for operating the accelerator and the experiments (e.g. for performance optimization, levelling, trigger optimization).
- Precise offline luminosity measurement is important for all analyses, particularly for precision cross section measurements:
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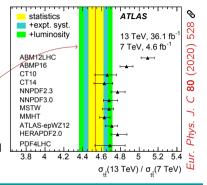
### New for ICHEP 2020: Preliminary luminosity measurements for low-pileup pp collision datasets

The Run 2 physics programme also included dedicated periods of low-pileup running at  $\sqrt{s}=5.02$  and 13 TeV:

- $\rightarrow$  13 TeV: for precision W boson measurements, where high pileup degrades the detector resolution of  $W \rightarrow \ell \nu$  decays.
- $\rightarrow$  5.02 TeV: served as pp reference data with the same nucleon-nucleon collision energy as the Run 2 PbPb dataset for the LHC heavy-ion programme (also useful for precision W/Z physics measurements).

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# The Run 2 pp Datasets

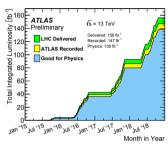
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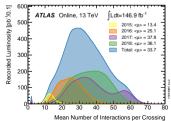
### 13 TeV high- $\mu$

Parameter	2015	2016	2017	2018
Maximum number of colliding bunch pairs $(n_b)$ Bunch spacing $(ns)$	2232 25	2208 25	2544/1909 25/8b4e	2544 25
Typical bunch population (10 <sup>11</sup> protons)	1.1	1.1	1.1/1.2	1.1
$\beta^*$ (m)	0.8	0.4	0.3	0.3-0.25
Peak luminosity $\mathcal{L}_{peak}$ ( $10^{33}$ cm <sup>-2</sup> s <sup>-1</sup> ) Peak number of inelastic interactions/crossing ( $\langle \mu \rangle_{peak}$ )	$5 \sim 16$	$^{13}$ $\sim 41$	$^{16}$ $\sim 45/60$	$^{19}$ $\sim$ 55
Luminosity-weighted mean inelastic interactions/crossing	13	25	38	36
Total delivered integrated luminosity (fb <sup>-1</sup> )	4.0	38.5	50.2	63.4

### 5.02 TeV and 13 TeV low- $\mu$

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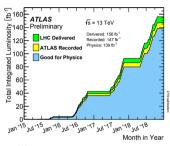
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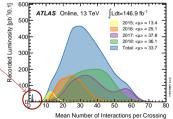
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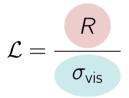
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Maximum number of colliding bunch pairs $(n_b)$	2232	2208	2544/1909	2544
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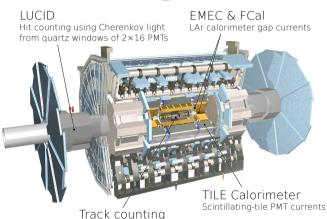
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# Visible cross section

The detector's calibration constant

# Measuring Event Rates: ATLAS Luminometers





Multiplicity of reconstructed charged particles in the silicon layers of the Inner Detector

# LUCID

- Primary ATLAS luminometer throughout Run 2
- Bunch-by-bunch measurements, integrated over "luminosity blocks" of typically 60 seconds
- Generally used a hit-counting algorithm of 2×4 PMTs for the baseline offline luminosity



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→ Per-bunch luminosity  $\mathcal{L}_b$  from bunch revolution frequency  $f_r$ , bunch populations  $n_1$  and  $n_2$ , and proton transverse-density distributions  $\hat{\rho}_{1,2}(x,y)$ .

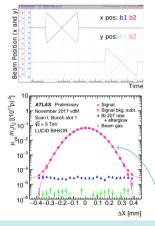
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$$\Sigma_{\scriptscriptstyle X} = rac{1}{\sqrt{2\pi}}rac{\int R(\Delta x)\,d\Delta x}{R(0)}, \quad \mathcal{L}_{\scriptscriptstyle B} = rac{f_{\scriptscriptstyle \Gamma} n_1 n_2}{2\pi\Sigma_{\scriptscriptstyle X}\Sigma_{\scriptscriptstyle Y}} = rac{\mu_{\scriptscriptstyle extsigma}f_{\scriptscriptstyle \Gamma}}{\sigma_{\scriptscriptstyle extsigma}}$$



Scan curve fitted with Gaussian  $\times$  polynomial to compute overlap integral.

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5 / 21

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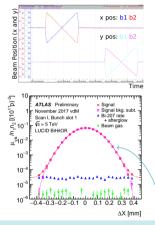
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■ Calibration constant  $\sigma_{vis}$  from visible interaction rate at peak of scan:

$$\sigma_{\mathsf{vis}} = \mu_{\mathsf{vis}}^{\mathsf{max}} \frac{2\pi \Sigma_{\mathsf{x}} \Sigma_{\mathsf{y}}}{n_1 n_2}$$



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# vdM Details and Uncertainties

- Scan-to-scan variations: expect same  $\sigma_{vis}$  for different bunch pairs and scan sets: take max difference between extreme scans as uncertainty.
  - → 1.2 % in 2017 at 13 TeV, typically half that in other years.
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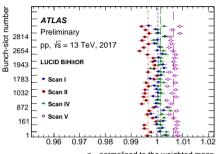
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Treatment of these corrections is under review: correction on  $\sigma_{\text{vis}}$  had been overestimated by  $\sim 1\%$  (until 2019):

- → 13 TeV datasets use **original** correction (+1.3–1.7 %, depending on scan)
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 $\sigma_{\text{vis}}$  normalised to the weighted mean

- Non-factorization:  $\hat{\rho}(x, y) \neq \rho_x(x)\rho_y(y)$ 
  - Correction applied from combined fits to the beam-separation dependence of the luminosity and of the parameters of the 3D luminous region, in both on- and off-axis vdM scans.
  - → Correction factor to  $\sigma_{vis}$  typically < 1 %, uncertainties of 0.2–0.5 %.

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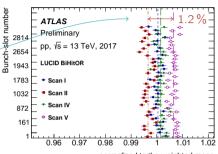
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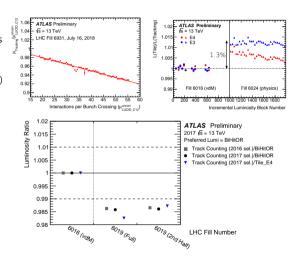
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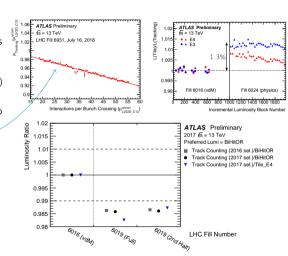
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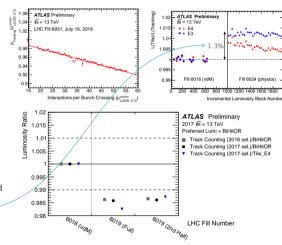
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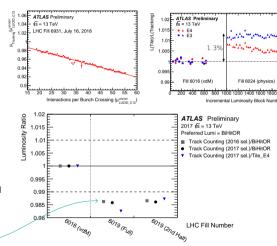
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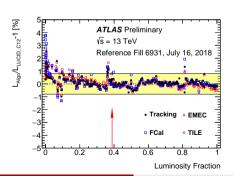
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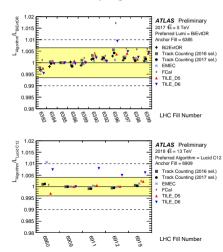
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  - → 1.3 % systematic assigned in 2017 due to non-linearity.
- At low- $\mu$ , a calibration transfer is still required, although the size of the correction is smaller (only  $\sim 1 \%$ ).

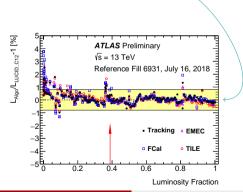


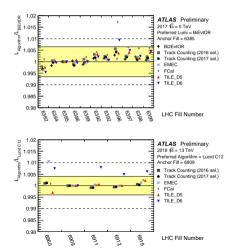
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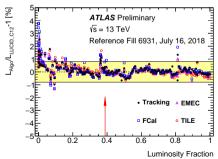


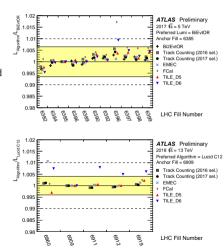
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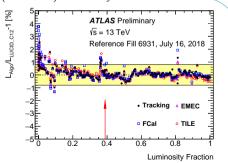


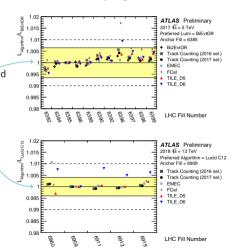
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# Final Uncertainties and Combination

Treat datasets with common CoM energy and running conditions from different years as a single combined dataset.

→ Need to combine luminosity and uncertainties, accounting for inter-year correlations

Data Sample		13 TeV	High- $\mu$		13 TeV Low- $\mu$			5.02 TeV Low- $\mu$		
	2015+16	2017	2018	Comb.	2017	2018	Comb.	2015	2017	Comb.
Integrated luminosity (fb <sup>-1</sup> )	36.2	44.3	58.5	139.0	0.1449	0.1902	0.3352	0.0251	0.2569	0.2820
Total uncertainty (fb $^{-1}$ )	0.8	1.0	1.2	2.4	0.0030	0.0029	0.0050	0.0005	0.0041	0.0043
Absolute vdM calibration	1.1	1.5	1.2	_	1.5	1.2	_	1.2	0.8	_
Calibration transfer	1.6	1.3	1.3	1.3	1.2	0.8	1.0	1.0	1.2	1.2
Afterglow and beam-halo subtraction	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1
Long-term stability	0.7	1.3	0.8	0.6	0.8	0.4	0.4	1.0	0.7	0.6
Tracking efficiency time-dependence	0.6	0.0	0.0	0.2	0.0	0.0	0.0	0.0	0.0	0.0
Total uncertainty (%)	2.1	2.4	2.0	1.7	2.1	1.5	1.5	1.9	1.6	1.5

Total luminosity and uncertainty for the ATLAS Run 2 pp datasets:

13 TeV high-
$$\mu$$

$$\mathcal{L}_{\text{tot}} = 139 \, \text{fb}^{-1}, \ \delta \mathcal{L}/\mathcal{L} = \pm 1.7 \, \%$$

13 TeV low-
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$$\mathcal{L}_{tot} = 335 \, pb^{-1}, \ \delta \mathcal{L}/\mathcal{L} = \pm 1.5 \, \%$$

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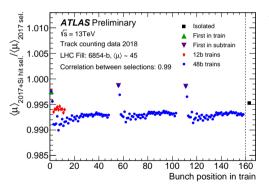
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The single largest uncertainty arises from the calibration transfer procedure. Possible avenue for improvement:

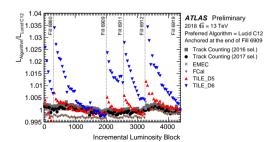


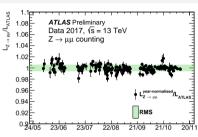
### Track counting

- The calibration transfer relies on the track-counting luminosity being linear with  $\mu$  in a broad range of LHC running conditions.
- Performance of a track-counting algorithm depends on the choice of track-selection working points.
- Currently investigating other working points (e.g. changing requirements on number of silicon hits).
  - Provides a means to monitor the stability and internal consistency of the track-counting measurements.

### Z boson counting

- **Z** boson production rate at the LHC is sufficiently high that counting  $Z \to \ell\ell$  ( $\ell = e, \mu$ ) events can act as a "luminometer."
- Serves as an additional check on the stability of the primary ATLAS luminosity measurement from LUCID.



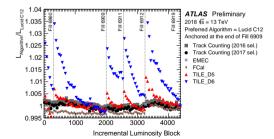


### Calorimeter Activation

- Date in 2017
- Running with head-on collisions activates the calorimeter materials, leading to an artificially high luminosity measurement that gradually decays away (order of a few hours).
- Important to understand these activation effects for both the calibration transfer and the long-term stability.
- Exploring both data-driven and simulation-based approaches to model and correct for these effects

### Z boson counting

- **Z** boson production rate at the LHC is sufficiently high that counting  $Z \to \ell\ell$  ( $\ell = e, \mu$ ) events can act as a "luminometer."
- Serves as an additional check on the stability of the primary ATLAS luminosity measurement from LUCID.



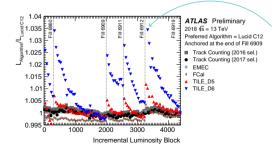
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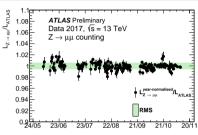
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# Summary

# ATLAS Run 2 luminosity results to date:1

Year	Dataset	Uncert.	Status
2015-18	$pp \sqrt{s} = 13  \text{TeV high-} \mu$	1.7 %	Prelim. ATLAS-CONF-2019-021 &
2015+17 2015+18	$pp \sqrt{s} = 5 \text{ TeV low-}\mu$ $pp \sqrt{s} = 13 \text{ TeV low-}\mu$	1.5 % 1.5 %	Prelim. ATLAS-CONF-2020-023 €
2015	Pb+Pb	1.5 %	Final
2016	<i>p</i> +Pb	2.4 %	Final
2018	Pb+Pb	4.1%	Prelim., update in progress

### ATLAS Luminosity Posters at ICHEP

- P. Moder: Measuring luminosity with track counting in the ATLAS experiment *Q*
- S. G. Fernandez: Measurements of Luminosity in ATLAS with Tile Calorimeter *&*
- M. O'Keefe: Luminosity Determination using  $Z \to \ell \ell$  Counting for Run-2 ATLAS Data  $\varrho$

<sup>&</sup>lt;sup>1</sup>Analysis of high- $\beta^*$  datasets is ongoing.

# Backup

# ATLAS Luminosity Public Results

Public Results TWiki: twiki.cern.ch/twiki/bin/view/AtlasPublic/LuminosityPublicResultsRun2 &

- Luminosity determination for low-pileup datasets at  $\sqrt{s}=5$  and 13 TeV using the ATLAS detector at the LHC  $\theta$ 
  - → Preliminary (Summer 2020) Run 2 luminosity results for low-pileup datasets [ATLAS-CONF-2020-023]
- Luminosity determination in pp collisions at  $\sqrt{s}=13\,\text{TeV}$  using the ATLAS detector at the LHC  $\pmb{\vartheta}$ 
  - → Preliminary (Spring 2019) Run 2 luminosity results [ATLAS-CONF-2019-021]
- Luminosity determination in pp collisions at  $\sqrt{s}=8\,\text{TeV}$  using the ATLAS detector at the LHC  $\pmb{\mathscr{O}}$ 
  - → Final 2012 luminosity results [*Eur. Phys. J. C* **76** (2016) 653]
- Improved luminosity determination in pp collisions at  $\sqrt{s}=7\,\text{TeV}$  using the ATLAS detector at the LHC  $\ref{LHC}$ 
  - → Final 2010 and 2011 luminosity results [Eur. Phys. J. C 73 (2013) 2518]

LHC Lumi Days Workshop, June 4–5 2019: indico.cern.ch/event/813285/ ATLAS Presentations:

- R. Hawkings: Overview of ATLAS luminosity determination methodology in Run 2 &
- W. Kozanecki: Impact of orbit drifts & magnetic-hysteresis effects on vdM & length-scale calibrations &
- M. Dyndal: Non factorization in ATLAS & ALICE vdM scans
- V. Lang: Long-term monitoring of delivered luminosity & calibration stability in ATLAS &
- R. Hawkings: Combination of luminosity uncertainties for the full Run-2 dataset: the ATLAS example &

# 5 TeV and 13 TeV low- $\mu$ Datasets

Dataset	Date	LHC fill(s)	Bunch config.	β* (m)	Peak $\mathcal{L}$ (cm $^{-2}$ s $^{-1}$ )	Typical $\langle \mu  angle$
$\sqrt{s} = 5.02  \text{TeV}$ - vdM scan - calibration transfer - physics	2017 11 Nov 13 Nov 12–20 Nov	6380 6385 6381–6399	22, isolated 1828, 8b4e 512–1828, 8b4e	3.1 3.1 3.1	$4 \times 10^{30}$ $1.2 \times 10^{33}$ $0.5-1.3 \times 10^{33}$	0.75-1.0 1-4 0.5-4
$\sqrt{s} = 13  \text{TeV}$ - vdM scan - calibration transfer - physics	2017 28 Jul 29 Jul 21–26 Nov	6016 6019 6404–6417	32, isolated 591 25 ns trains 644–1866, 8b4e	19 0.4 0.4	$ 2.7 \times 10^{30} \\ 8 \times 10^{31} \\ 5 \times 10^{32} $	0.6 1 (level) 2, 1 (level)
$\sqrt{s} = 13  \text{TeV}$ - vdM scan - calibration transfer - physics	2018 30 Jun 28 Jun 28 Jun + 9–13 Jul	6868 6860 6860 + 6909-6919	124, isolated 2448, 25 ns trains 2448–2544, 25 ns trains	19 0.3 0.3	$ 8 \times 10^{30} \\ 7 \times 10^{32} \\ 7 \times 10^{32} $	0.5 2 (level) 2 (level)

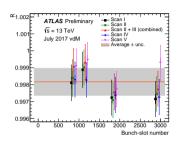
# vdM Details I

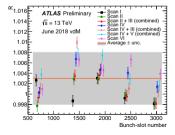
- Determining bunch populations  $n_1$  and  $n_2$ 
  - → **DC C**urrent **T**ransformers to obtain the total current in each beam.
  - → Fast Beam-Current Transformers to resolve the total current into the fractions in individual bunches.
- Orbit drifts during scans
  - → Measured using LHC arc and triplet (DOROS) Beam Position Monitors.
  - $\rightarrow$  Have observed drifts up to  $\mathcal{O}(10\,\mu\text{m})$  in the position of one or both beams during the course of a single vdM scan.
- Beam position jitter
  - → Beam movement within one scan step.
  - → BPMs constrain possible movement within a scan step, input to simulated vdM scans.
- Emittance growth
  - $\rightarrow$  Accounts for possible variations of the horizontal and vertical beam emittances, and therefore the convolved beam sizes  $\Sigma_x$  and  $\Sigma_y$ , during the course of an x-y scan pair.
  - → Creates bias only if horizontal and vertical emittances grow at different rates (which they do).
  - → Uncertainty carried over from Run 1 analysis.

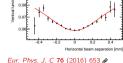
# vdM Details II

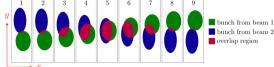
### Non-factorization

Clear signature of non-factorization: strong horizontal-separation dependence of the vertical luminous size (from reconstructed vertices).









(1d) Offset vdM scan in v direction performed in 9 steps

Data (Centred x-scan IV July 2012)

LMC Ellipage Simulated profile of each beam

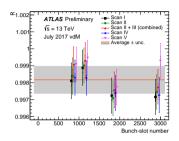
3-D double Gaussian

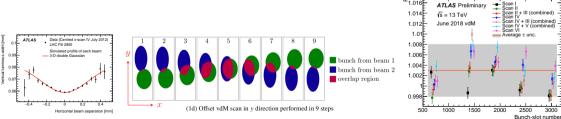
# vdM Details II

### Non-factorization

- → Clear signature of non-factorization: strong horizontal-separation dependence of the vertical luminous size (from reconstructed vertices).
- → To estimate the correction to  $\sigma_{\text{vis}}$  due to non-factorization effects, compute the quantity R, such that  $\sigma_{\text{vis}}^{\text{corr}} = \sigma_{\text{vis}}/R$ :

$$R = \frac{\text{True Luminosity}}{\text{Factorized Luminosity}} = \frac{\int \hat{\rho}_1(x,y) \hat{\rho}_2(x,y) \, dx dy}{\int \rho_1(x) \rho_2(x) \, dx \int \rho_1(y) \rho_2(y) \, dy},$$





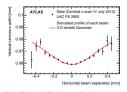
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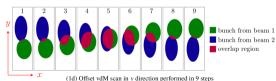
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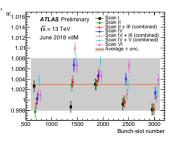
→ Do combined fits to the beam-separation dependence of the luminosity and of the parameters of the 3D luminous region, in both on- and off-axis vdM scans, to extract non-factorisable single-beam luminosity profiles, which are fed into simulated vdM scans to compute the true beam overlap integral.





ATLAS Preliminary

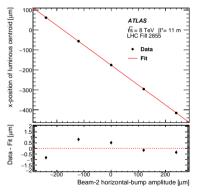
| Scan | Sc



# vdM Details III

### ■ Length scale calibration

- → Relation between *nominal* (i.e. requested) and *actual* beam displacement at IP.
- → Displace both beams in same direction.
- → Reconstruct luminous centroid position using vertices reconstructed in ATLAS inner detector.
  - Perform a mini-scan in beam-2 *x*-pos around fixed beam-1 *x*-pos to find peak position.
- → Fit linear relation between bump amplitude and luminous centroid to find calibration
- $\rightarrow$  Repeat for each of beam-1 x, y and beam-2 x, y.
  - Since Nov 2017, use same directions of movement as in vdM scan, to get same hysteresis effect.
- Additional systematics from ID alignment
- $\rightarrow$  Assessed by considering "realistic" misalignment scenarios, giving a  $\sim 0.1 \,\%$  uncertainty.



# Combination Methodology and Inter-Year Correlations

- Final uncertainty calculated using simple error propagation:
  - → Total integrated luminosity is sum of all years:  $\mathcal{L}_{tot} = \Sigma_i \mathcal{L}_i$
  - $\rightarrow$  Variance of the total depends on covariance matrix  $V_L$  encoding the errors on individual years:

$$(\sigma_{\mathcal{L}_{tot}})^2 = \mathbf{GV}_L \widetilde{\mathbf{G}}, \quad \text{where } \mathbf{G} = \left( \frac{\partial \mathcal{L}_{tot}}{\partial \mathcal{L}_1}, \frac{\partial \mathcal{L}_{tot}}{\partial \mathcal{L}_2}, \frac{\partial \mathcal{L}_{tot}}{\partial \mathcal{L}_3}, \cdots \right) = (1, 1, 1, \ldots)$$

- **E**valuation of the covariance matrix  $V_i$ :
  - $\rightarrow$  Sum of individual sources with uncertainties  $\sigma_i$  in each year (many separate uncorrelated and correlated sources):

$$V_{L} = \underbrace{\begin{pmatrix} \sigma_{1}^{2} & 0 & 0 \\ 0 & \sigma_{2}^{2} & 0 \\ 0 & 0 & \sigma_{3}^{2} \end{pmatrix}}_{\text{uncorrelated}} + \underbrace{\begin{pmatrix} \sigma_{1}^{2} & \sigma_{1}\sigma_{2} & \sigma_{1}\sigma_{3} \\ \sigma_{2}\sigma_{1} & \sigma_{2}^{2} & \sigma_{2}\sigma_{3} \\ \sigma_{3}\sigma_{1} & \sigma_{3}\sigma_{2} & \sigma_{3}^{2} \end{pmatrix}}_{\text{correlated}} + \cdots$$

- → Sources with both correlated and uncorrelated parts are broken into two separate contributions to V<sub>L</sub>.
- $\rightarrow$  "Random" uncertainties taken as uncorrelated, e.g. scan-to-scan  $\sigma_{\text{vis}}$  consistency, long-term stability.
- → "Systematic" uncertainties (and those obtained using the same methodology) taken as correlated—always have the same bias, e.g. non-factorization, beam-beam effects, calibration transfer.

# Individual and Combined Dataset Uncertainties

		13 TeV	High-μ			13 TeV Low-	ц	5.02 TeV Low-μ		
Data Sample	2015+16	2017	2018	Comb.	2017	2018	Comb.	2015	2017	Comb.
Integrated luminosity (fb <sup>-1</sup> )	36.2	44.3	58.5	139.0	0.1449	0.1902	0.3352	0.0251	0.2569	0.2820
Total uncertainty (fb <sup>-1</sup> )	8.0	1.0	1.2	2.4	0.0030	0.0029	0.0050	0.0005	0.0041	0.0043
Uncertainty contributions (%):										
DCCT calibration <sup>†</sup>	0.2	0.2	0.2	0.1	0.2	0.2	0.1	0.4	0.2	0.2
FBCT bunch-by-bunch fractions	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.0
Ghost-charge correction*	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.2	0.1	0.1
Satellite correction <sup>†**</sup>	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.1	0.0	0.0
Scan curve fit model <sup>†**</sup>	0.5	0.4	0.5	0.4	0.4	0.5	0.5	0.6	0.2	0.2
Background subtraction	0.2	0.2	0.2	0.1	0.2	0.2	0.1	0.1	0.1	0.0
Orbit-drift correction	0.1	0.2	0.1	0.1	0.2	0.1	0.1	0.2	0.2	0.2
Beam position jitter†()*	0.3	0.3	0.2	0.2	0.3	0.2	0.2	0.3	0.1	0.1
Beam-beam effects*	0.3	0.3	0.2	0.3	0.3	0.2	0.2	0.2	0.2	0.2
Emittance growth correction*	0.2	0.2	0.2	0.2	0.2	0.2	0.2	0.4	0.2	0.2
Non-factorization effects*	0.4	0.2	0.5	0.4	0.2	0.5	0.4	0.2	0.3	0.3
Length-scale calibration	0.3	0.3	0.4	0.2	0.3	0.4	0.3	0.5	0.4	0.4
ID length scale*	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1
Bunch-by-bunch $\sigma_{vis}$ consistency	0.2	0.2	0.4	0.2	0.2	0.4	0.2	0.4	0.2	0.2
Scan-to-scan reproducibility	0.5	1.2	0.6	0.5	1.2	0.6	0.6	0.3	0.3	0.2
Reference specific luminosity	0.2	0.2	0.4	0.2	0.2	0.4	0.3	0.1	0.4	0.3
Subtotal for absolute vdM calibration	1.1	1.5	1.2	-	1.5	1.2	-	1.2	0.8	-
Calibration transfer <sup>†**</sup>	1.6	1.3	1.3	1.3	1.2	0.8	1.0	1.0	1.2	1.2
Afterglow and beam-halo subtraction*	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1
Long-term stability	0.7	1.3	0.8	0.6	0.8	0.4	0.4	1.0	0.7	0.6
Tracking efficiency time-dependence	0.6	0.0	0.0	0.2	0.0	0.0	0.0	0.0	0.0	0.0
Total uncertainty (%)	2.1	2.4	2.0	1.7	2.1	1.5	1.5	1.9	1.6	1.5

# Low- $\mu$ 13 TeV and 5 TeV Combined Uncertainties

