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# AIDA-2020

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## Poster

# Beam Telescopes at DESY II Test Beam Facility

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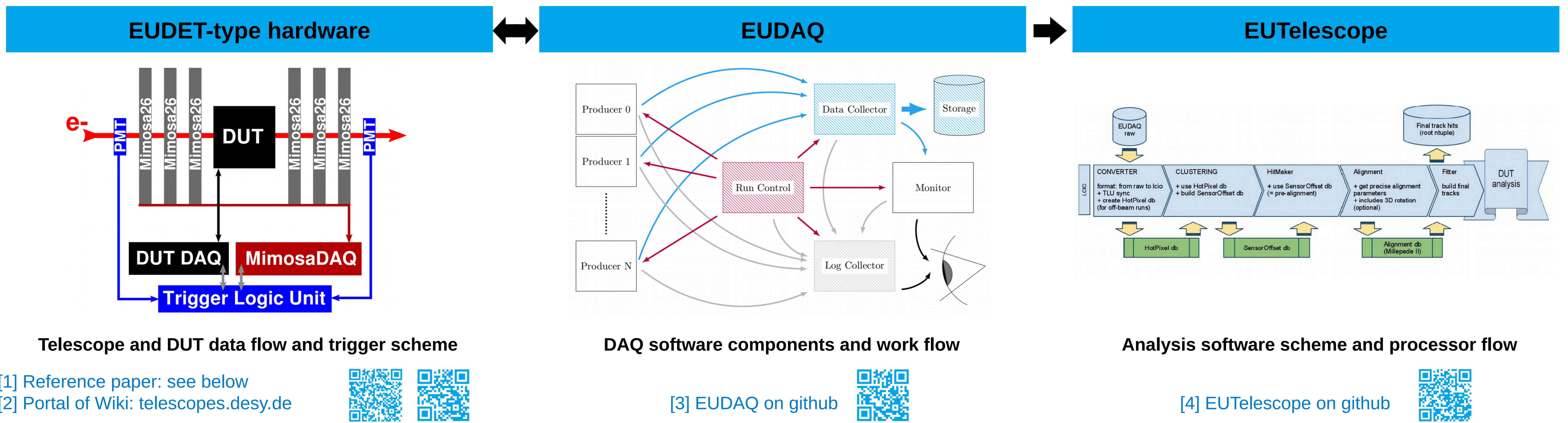
This work is part of AIDA-2020 Work Package 15: **Upgrade of beam and irradiation test infrastructure.**

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# Beam Telescopes

## A common tool: the EUDET-type beam telescope package

A high resolution ( $\sigma \approx 2 \mu\text{m}$ ) beam telescope based on monolithic active pixel sensors (Mimosa26) was developed within the EUDET collaboration [1,2]. In the last decade it has become a in-beam tool for many different high-energy physics groups, largely due to its precise spatial resolution, reliable operation and user device-under-test (DUT) integration capabilities. Besides the hardware components [1,2], two software frameworks play a central role: EUDAQ [3], a multi-platform data acquisition system that allows easy integration of the device-under-test, and EUTelescope [4], a group of processors running in ILCSoft's Marlin framework that allows the spatial reconstruction of particle tracks and the final data analysis.



## DATURA in TB21



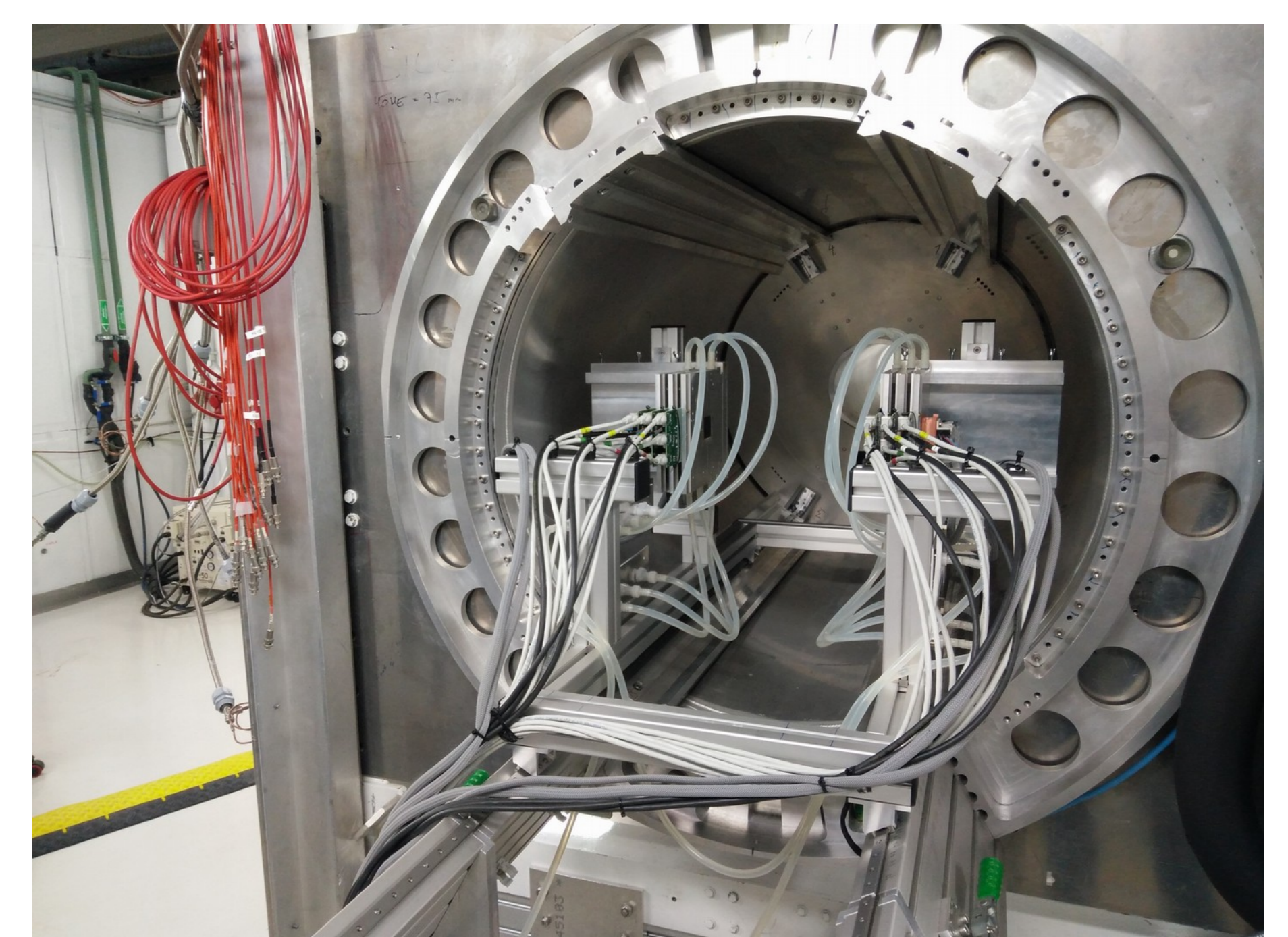
One of 7<sup>th</sup> EUDET-type telescopes worldwide is permanently installed in area TB21. The Big Red Magnet can be used to deflect particles, for example for energy measurements.

## DURANTA in TB22



A second EUDET-type telescope is permanently installed in TB22. This area is the largest area at the DESY II Test Beam facility, which is useful for large user setups.

## Options for TB24



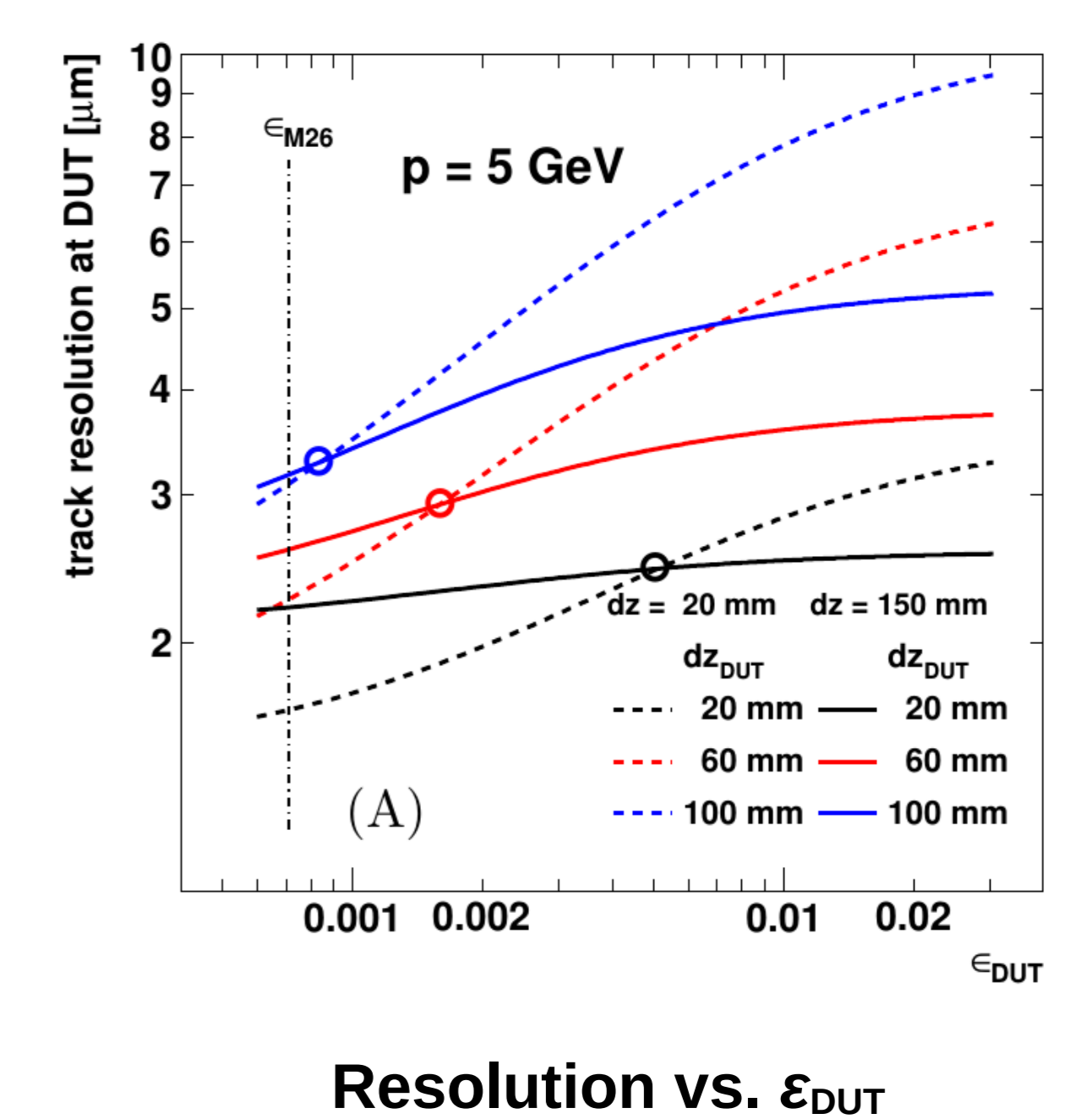
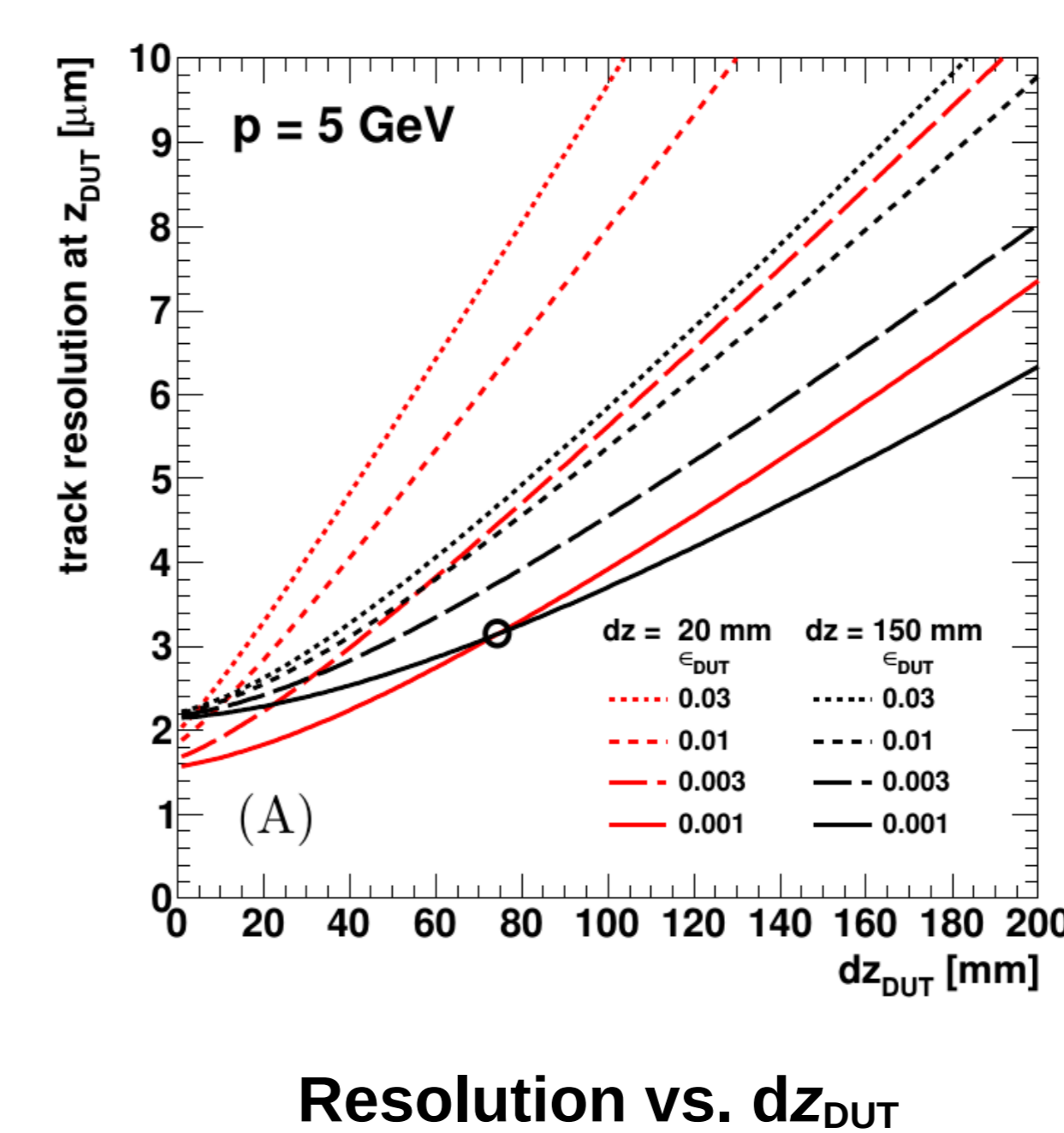
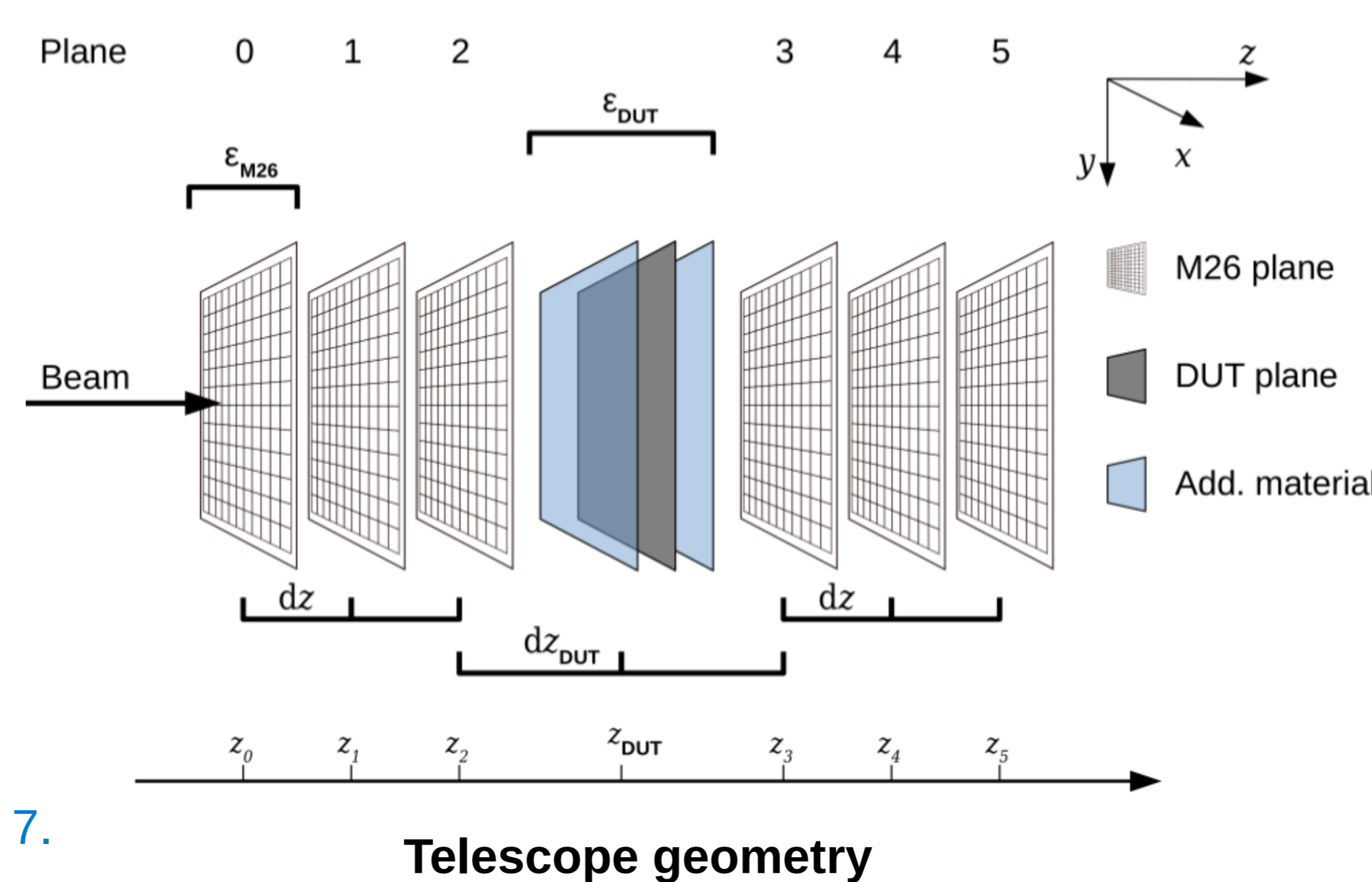
EUDET-type telescope planes based on Mimosa26 sensors can optionally be installed in the PCMAG in TB24/1. In 2019 a permanently installed strip telescope (LYCORIS) is available.

## Resolution predictions for DUT integrations

Using the measured intrinsic resolution, the material budget in the beam (telescope planes, air, DUT) and GBL fitting, predictions of the track resolution at the actual DUT position  $z_{\text{DUT}}$  are possible. The mean intrinsic resolution of a Mimosa26 sensor is measured to be  $\sigma_{\text{M26}} = (3.24 \pm 0.09) \mu\text{m}$  @ thr. 6 and the normalised material budget of one telescope plane is  $\epsilon_{\text{M26}} = 7.5 \cdot 10^{-4}$  including  $54 \mu\text{m}$  Silicon and  $2 \times 25 \mu\text{m}$  Kapton foil.

Using an EUDET-type telescope at the DESY II test beam facility (1-6 GeV), the optimum telescope geometry can be predicted by using:

1. a small  $dz_{\text{DUT}}$  (distance between inner telescope plane and DUT)
2. the narrow ( $dz = 20 \text{ mm}$ ) or the wide ( $dz = 150 \text{ mm}$ ) setup depending on the total DUT's budget  $\epsilon_{\text{DUT}}$



[1] Jansen, H. et al. EPJ Techn Instrum (2016) 3: 7.