## Searches for Squarks and Gluinos with ATLAS

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# Outline

#### Introduction - ATLAS and SUSY

- Status of ATLAS and SUSY in 2018
- Motivation for squark and gluino searches at ATLAS

#### Squark and gluino searches at ATLAS

- The anatomy of a typical SUSY search
- Some recent searches and results
  - Long-lived particles
  - Leptonic final states

#### Summary and Outlook

• Current limits and potential sensitivity at the end of Run 2 and beyond

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# Introduction - ATLAS and SUSY

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# Introduction

#### The LHC and ATLAS - where are we now?

- SUSY searches are a key part of the ATLAS physics program
- Huge variety of ATLAS searches with 2015+2016 and  $\sqrt{s} = 13$  TeV data completed
- SPOILER ALERT no sign of SUSY, yet ;)
- Analysis ongoing for 2017 data
- 2018 data taking under way with higher than ever instantaneous luminosity



# Squarks and Gluinos

# SUPERSYMMETRY

**SUSY** particles



#### Standard particles

#### Why squarks and gluinos?

- Natural SUSY spectrum without fine tuning requires light stop and gluino masses
- Should be accessible to the LHC
- Squarks and gluinos  $\rightarrow$  higher cross sections

#### Why SUSY?

- Cancel quadratic divergences to Higgs mass calculation
- Unification of gauge interactions
- LSP excellent candidate for dark matter



https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/ SUSY/FeynmanGraphs/CrossSections/CrossSections/ Ecm\_13TeV/xsec\_plot.png

# SUSY Models

Gauge eigenstates			Mass eigenstates		
bino, wino, higgsinos wino, higgsinos	10, wino, higgsinos $\tilde{B}, \tilde{W}^0, \tilde{H}^0_u, \tilde{H}^0_d$ wino, higgsinos $\tilde{W}^+, \tilde{H}^+$		$ ilde{\chi}^{0}_{1},  ilde{\chi}^{0}_{2},  ilde{\chi}^{0}_{3},  ilde{\chi}^{0}_{4}$ $ ilde{\chi}^{+}_{1},  ilde{\chi}^{+}_{2}$	neutralinos charginos	
	$\tilde{W}^-, \tilde{H}_d^-$	$\frac{1}{2}$	$\tilde{\chi}_1^+, \tilde{\chi}_2^-$	5	
gluinos	ĝ	$\frac{1}{2}$	no mixing		
selectron, smuon, stau sneutrinos	$\begin{array}{l} \tilde{e}_{L,R}, \tilde{\mu}_{L,R}, \tilde{\tau}_{L,R} \\ \tilde{\nu}_e, \tilde{\nu}_\mu, \tilde{\nu}_\tau \end{array}$	0 0	$\begin{array}{c} \tilde{e}_{L,R}, \tilde{\mu}_{L,R}, \tilde{\tau}_1, \tilde{\tau}_2 \\ \tilde{\nu}_e, \tilde{\nu}_\mu, \tilde{\nu}_\tau \end{array}$	sleptons sleptons or sneutrinos	
sup, scharm, stop sdown, sstrange, sbottom	$ \begin{array}{l} \tilde{u}_{L,R}, \tilde{c}_{L,R}, \tilde{t}_{L,R} \\ \tilde{d}_{L,R}, \tilde{s}_{L,R}, \tilde{b}_{L,R} \end{array} $	0 0	$ \begin{array}{l} \tilde{u}_{L,R}, \tilde{c}_{L,R}, \tilde{t}_1, \tilde{t}_2 \\ \tilde{d}_{L,R}, \tilde{s}_{L,R}, \tilde{b}_1, \tilde{b}_2 \end{array} $	squarks	

arXiv:1609.01686v1 [hep-ex]

#### Rich phenomenology

- MSSM has  $> 100 \text{ parameters} \rightarrow \text{cannot}$  scan all of parameter space
- Simplified models sparticle masses not involved in signature of interest set to high values

#### Typical signal topologies

- R-parity conserving (RPC)
  - Pair produced
  - Decaying to LSP
  - Large  $E_{\rm T}^{\rm miss}$
- R-parity violating (RPV)
  - Decay to SM
  - Lifetime depends on  $\lambda$  coupling



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# Squark and gluino searches at ATLAS

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# Typical search strategy

#### Design kinematic selections for background rejection and estimation

- Signal regions rich in signal events
- Control regions rich in specific background process
- Validation regions lie somewhere between SRs and CRs

#### **Background estimation**

- Reducible backgrounds determine from data
- Irreducible backgrounds
  - Dominant normalise in CR
  - Subdominant MC estimation





#### Simultaneous fit

 Combined fit of all regions including experimental and theoretical uncertainties

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observable 1

SR2

SR3

arXiv:1410.1280v1 [hep-ex] 6 Oct 2014

# Long-lived searches

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# Disappearing track

SUSY scenario with  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_1^0$ 

- Almost pure wino  $\rightarrow$  nearly mass degenerate
- $\tilde{\chi}_1^{\pm}$  long-lived  $\mathcal{O}(\mathsf{ns})$
- Soft  $\pi^+ \rightarrow \textit{disappearing track}$
- Run 2 improvement reconstruct shorter tracklets



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# Disappearing track

#### **Exclusion limits**

- Limits depend on the lifetime of the chargino
- Disappearing track analysis sensitive when the chargino decays in the inner detector



# **Displaced** vertex

#### Split SUSY models

- Large squark masses  $\rightarrow$  gluino long-lived
- Hadronise into R-hadron leading to displaced vertex (DV)
- Reconstruct displaced vertex using large radius tracking
- Hadronic interactions in material rich detector regions - map-based veto



PAPERS/SUSY-2016-08/fig\_07a.png





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Gluino mass limit vs. lifetime

- Reinterpretation of existing ATLAS analyses in the context of R-hadrons from split SUSY
- RPC 0L + jets strong limits until decay of R-hadron reaches calorimeters
- Displaced vertex most powerful analysis for intermediate lifetimes
- Pixel dE/dx sensitivity when R-hadron track can be reconstructed before decay



https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/CombinedSummaryPlots/ SUSY/ATLAS\_SUSY\_LLP/ATLAS\_SUSY\_LLP.png

# Reinterpret existing ATLAS searches in the context of RPV SUSY

• SUSY model - light gluinos and LSP

$$\mathcal{W}_{\mathsf{RPV}} = rac{\lambda_{ijk}}{2} L_i L_j ar{E}_k + \lambda_{ijk}' L_i Q_j ar{D}_k + rac{\lambda_{ijk}''}{2} ar{U}_i ar{D}_j ar{D}_k + \kappa_i L_i H_u$$

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• Non-zero  $\lambda_{323}^{\prime\prime},$  all other RPV couplings set to zero



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# Leptonic final states

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# Stop pair production with taus

#### Gauge-mediated SUSY breaking

- G is the LSP
- Ditau final state semileptonic and fully hadronic SRs
- lep-had region single lepton trigger
- had-had region E<sup>miss</sup>trigger
- Dominant background misidentified taus
  - lep-had Fake factor
  - had-had Control Region



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https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/ PAPERS/SUSY-2016-19/fig\_05d.png

1200 1300

m(t̃,) [GeV]

# Same sign or three leptons

Same sign / three leptons present in many SUSY scenarios

- Low SM background
- Reducible background: charge-flip electrons - reject with BDT
- Extract charge-flip probability from  $Z \rightarrow ee$  electrons



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### Same sign or three leptons

#### **Exclusion limits**

Improved sensitivity for compressed scenarios



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https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/ PAPERS/SUSY-2016-14/ATLAS\_SUSY\_Strong\_2step.png

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### Same sign or three leptons



• Sensitivity to direct sbottom pair production via  $\tilde{\chi}_1^{\pm}$ 



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https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/ PAPERS/SUSY-2016-14/fig\_04d.png

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#### ATLAS SUSY Searches\* - 95% CL Lower Limits

December 2017

	Model	$e, \mu, \tau, \gamma$	Jets	E <sup>miss</sup> <sub>T</sub>	∫£ dt[fb	<sup>1</sup> ] Mass limit	$\sqrt{s} = 7, 8$	TeV $\sqrt{s} = 13 \text{ TeV}$
Inclusive Searches	$\begin{array}{l} \overline{q}\partial_{i}, \overline{q} \rightarrow q \xi_{i}^{0} \\ \overline{q}\partial_{i}, \overline{q}\partial_{i$	0 mono-jet 0 ee, μμ 3 e, μ 0 1-2 τ + 0-1 ℓ 2 γ γ 0	2-6 jets 1-3 jets 2-6 jets 2-6 jets 2 jets 4 jets 7-11 jets 0-2 jets 2 jets mono-jet	Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 36.1 14.7 36.1 36.1 3.2 36.1 36.1 36.1 20.3	4 710 GeV 8 710 GeV 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8	1.57 TeV 2.02 TeV 2.01 TeV 1.7 TeV 1.87 TeV 1.8 TeV 2.0 TeV 2.15 TeV 2.05 TeV	$\begin{split} n(\xi_1^0) &\sim 200  \text{GeV}, \ n([1^m] \text{gen. 4}) - m(2^m)  \xi_2^{(m)} = 0 \\ m(\xi_1^0) - m(\xi_1^0) - 5  \text{GeV} \\ m(\xi_1^0) - 200  \text{GeV}, \ m(\xi_1^0) - 0.5  (m(\xi_1^0) + m(\xi_2)) \\ m(\xi_1^0) - 200  \text{GeV}, \ m(\xi_1^0) - 0.0  \text{GeV} \\ m(\xi_1^0) - 400  \text{GeV} \\ m(\xi_1^0) - 400  \text{GeV} \\ m(\xi_1^0) - 100  \text{GeV}, \ m(\xi_1^0$
3 <sup>rd</sup> gen. <u>§</u> med.	$\bar{g}\bar{g}, \bar{g} \rightarrow b\bar{b}\bar{\chi}_{1}^{0}$ $\bar{g}\bar{g}, \bar{g} \rightarrow t\bar{t}\bar{\chi}_{1}^{0}$	0 0-1 e,µ	3 b 3 b	Yes Yes	36.1 36.1	ğ ğ	1.92 TeV 1.97 TeV	m(k̂ <sup>0</sup> <sub>1</sub> )<600 GeV m(k̂ <sup>0</sup> <sub>1</sub> )<200 GeV
3rd gen. squarks direct production	$ \begin{split} \bar{b}_1 \bar{b}_1 , \bar{b}_1 \to b \bar{k}_1^{0} \\ \bar{b}_1 \bar{b}_1 , \bar{b}_1 \to b \bar{k}_1^{-1} \\ \bar{b}_1 \bar{b}_1 , \bar{b}_1 \to b \bar{k}_1^{-1} \\ \bar{t}_1 \bar{t}_1 , \bar{t}_1 \to b \bar{k}_1^{-1} \\ \bar{t}_1 \bar{t}_1 , \bar{t}_1 \to b \bar{k}_1^{0} \\ \bar{t}_1 \bar{t}_1 , \bar{t}_1 \to c \bar{k}_1^{0} \\ \bar{t}_1 \bar{t}_1 , \bar{t}_1 \to c \bar{k}_1^{0} \\ \bar{t}_2 \bar{t}_2 , \bar{t}_2 \to \bar{t}_1 + Z \\ \bar{t}_2 \bar{t}_2 , \bar{t}_2 \to \bar{t}_1 + A \end{split} $	0 2 $e, \mu$ (SS) 0 $\cdot 2 e, \mu$ 0 $\cdot 2 e, \mu$ 0 2 $e, \mu$ (Z) 3 $e, \mu$ (Z) 1 $\cdot 2 e, \mu$	2 b 1 b 1-2 b 0-2 jets/1-2 mono-jet 1 b 1 b 4 b	Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 .7/13.3 10.3/36.1 36.1 20.3 36.1 36.1 36.1	5r.         950 GeV           117-170 GeV         275-700 GeV           4, 91196 GeV         0.185-10 TeV           7, 90-190 GeV         0.185-10 TeV           7, 90-190 GeV         0.185-10 TeV           7, 90-190 GeV         0.195-10 TeV           7, 90-190 GeV         0.195-00 GeV           7, 90-790 GeV         100-900 GeV           7, 90-790 GeV         100-900 GeV           7, 90-790 GeV         100-900 GeV		$\begin{split} m(\tilde{c}_{1}^{2}) &< 420  \text{GeV} \\ m(\tilde{c}_{1}^{2}) &< 200  \text{GeV}, m(\tilde{c}_{1}^{2}) &= m(\tilde{c}_{1}^{2}) + 100  \text{GeV} \\ m(\tilde{c}_{1}^{2}) &= 2m(\tilde{c}_{1}^{2}), m(\tilde{c}_{1}^{2}) = 55  \text{GeV} \\ m(\tilde{c}_{1}) - m(\tilde{c}_{1}^{2}) &= 56  \text{GeV} \\ m(\tilde{c}_{1}^{2}) - 150  \text{GeV} \\ m(\tilde{c}_{1}^{2}) &= 0  \text{GeV} \\ m(\tilde{c}_{1}^{2}) &= 0  \text{GeV} \end{split}$
Long-lived particles	$ \begin{array}{l} \mbox{Stable}, \mbox{stopped} \ \Bar{g} \ \R-hadron \\ \mbox{Stable} \ \Bar{g} \ \R-hadron \\ \mbox{Metastable} \ \R-hadron \\ \R-hadron \ \R-hadron \\ \R-hadron \ \R-hadron \\ \R-hadron \ \R-hadron \ \R-hadron \\ \R-hadron \ \ \R-hadron \ \ $	0 trk dE/dx trk displ. vtx 1-2 μ 2 γ displ. ee/eμ/μμ	1-5 jets - - - - - μ -	Yes - Yes - Yes -	27.9 3.2 32.8 19.1 20.3 20.3	홍 850 GeV 홍 홍 홍 홍 우 537 GeV 옥 같 440 GeV 옷	1.58 TeV 1.57 TeV 2.37 T	$\begin{split} m(\tilde{k}_{1}^{0}) &= 100 \; GeV, \; 10 \; \mu s \! < \! \tau(\tilde{g}) \! < \! 1000 \; s \\ m(\tilde{k}_{1}^{0}) &= 100 \; GeV, \; \tau \! > \! 10 \; ns \\ \hline GV_{1}^{0}(\tilde{g}) \! = \! 0.17 \; ns, \; m(\tilde{k}_{1}^{0}) \! = 100 \; GeV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \! < \! tangk \! - \! SV \\ 10 \!$
RPV	$ \begin{array}{l} LFV pp {\rightarrow} \bar{v}_r + X_r \bar{v}_r {\rightarrow} ept/ert/\mu\tau \\ Bilnear RPV CMSSM \\ \bar{X}_1^r, \bar{X}_1^r \rightarrow MS_1^{r} \tilde{\mathcal{M}}_1^{r} {\rightarrow} exer, epr, \mu\mu\nu \\ \bar{X}_1^r, \bar{X}_1^r \rightarrow MS_1^{r} \tilde{\mathcal{M}}_1^{r} {\rightarrow} erre, \\ \bar{B}_1, \bar{X}_1^r \rightarrow MS_1^{r} \tilde{\mathcal{M}}_1^{r} {\rightarrow} eqq \\ \bar{B}_2, \bar{B}_2, \bar{B}_1, \bar{B}_1, \bar{D} {\rightarrow} S_1 \\ \bar{B}_1, \bar{A}_1 {\rightarrow} DS \\ \bar{I}_1\bar{I}_1, \bar{I}_1 {\rightarrow} DS \\ \bar{I}_1\bar{I}_1, \bar{I}_1 {\rightarrow} DS \end{array} $	$e\mu,e\tau,\mu\tau$ $2 e,\mu$ (SS) $4 e,\mu$ $3 e,\mu + \tau$ $0 4-1 e,\mu$ 8- $1 e,\mu$ 8- $0 2 e,\mu$	- 0-3 b - 5 large-R jo -10 jets/0-4 -10 jets/0-4 2 jets + 2 b 2 b	Yes Yes Yes ts - b - b -	3.2 20.3 13.3 20.3 36.1 36.1 36.1 36.7 36.1	5, 4 P 41 430 GeV 8 8 8 8 7, 100-470 GeV 480-610 GeV 7, 0,4 0,4 0,4 0,4 0,4 0,4 0,4 0,4 0,4 0,	1.9 TeV 1.45 TeV 1.875 TeV 2.1 TeV 1.65 TeV 1.45 TeV	$\begin{split} \lambda_{311}^{\prime} = & 0.11, \lambda_{112}(_{131})_{233} = 0.07 \\ m(\beta) = m(\beta), \ e_{T,S} = c + 1 \\ m(\beta)^{\prime} = & 0.026, \ e_{T,S} = c + 1, 2 \\ m(\beta^{\prime}) = & 0.226, \ m(\beta^{\prime}) = $
Other	Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	2 c	Yes	20.3	č 510 GeV		m(\$\vec{k}_1^0)<200 GeV
Only a pheno simpl	a selection of the available mas omena is shown. Many of the l ified models, c.f. refs. for the a	s limits on r imits are bas ssumptions	new state sed on made.	s or	10	) <sup>-1</sup> 1		Mass scale [TeV]

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## Conclusions

Rich and varied program of searches for squarks and gluinos at ATLAS

- No sign of SUSY yet in 2015 + 2016 data
- Stop limits approaching the 1 TeV level
- Gluino limits approaching the 2 TeV level
- SUSY is running out of places to hide...

#### R&D under way for 2017 + 2018 data

- By the end of the year the total integrated luminosity will be  $> 100 \text{ fb}^{-1}$
- High luminosity LHC much much more!
- Watch this space...



PUBNOTES/ATL-PHYS-PUB-2014-010/