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## DESIGN AND SYSTEM INTEGRATION OF THE SUPERCONDUCTING WIGGLER MAGNETS FOR THE CLIC DAMPING RINGS

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## Abstract

To achieve high luminosity at the collision point of the Compact Linear Collider (CLIC) the normalized horizontal and vertical emittances of the electron and positron beams must be reduced to 500 nm and 4 nm before the beams enter the 1.5TeV linear accelerators. An effective way to accomplish ultra-low emittances with only small effects on the electron polarization is using damping rings operating at 2.86 GeV equipped with superconducting wiggler magnets. This paper describes a technical design concept for the CLIC damping wigglers.

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### I. INTRODUCTION

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CLIC (Compact Linear Collider) is a study for an 12 electron-positron collider in the TeV range. Fundamen-13 tal physics data complementary to the Large Hadron 14 Collider (LHC) and lower-energy linear electron-positron 15 colliders may be obtained with measurements at the col-16 lision point of CLIC. Detailed information on the CLIC 17 design study can be found in [1]. 18

To achieve the required luminosity at the collision 19 point of CLIC, the normalized horizontal emittance in-20 cluding the effect of intrabeam scattering has to be less 21 than 500 nm rad before entering the linear collider. Intra-22 23 beam scattering is a small angle multiple Coulomb scattering effect [3], which causes a beam emittance growth. 24 Moreover, the normalized vertical zero-current emittance 25 has to be 4 nm rad and the normalized longitudinal zero-26 current emittance has to be 6000 eVm. For the reduc-27 tion of the emittance, damping rings are foreseen. Both 28 the electron and the positron beams enter first the pre-29 damping rings and afterwards the main damping rings. 30 The damping rings will be racetrack shaped rings with 26 31 wigglers placed in each straight section. The horizontal 32 equilibrium emittance is designed to be around one order 33 of magnitude smaller than in other planned or built rings 34 (Fig. 1), which can be only achieved with superconduct-35 ing damping wiggler magnets. 36

The baseline design foresees Nb-Ti superconducting 37 wigglers. But by using the more challenging Nb<sub>3</sub>Sn tech-38 nology the magnetic flux density amplitude  $B_{\rm w}$  can be 39 increased (see Fig. 2 for the parameters of the wiggler 40 magnets), which may open the possibility to shorten the 41 42 43 44



FIG. 1. Comparison of vertical versus horizontal emittances of different storage rings below 5 GeV with the CLIC damping rings [2].

48 decision can be taken.

To test the wiggler system with beam, two prototypes 49 <sup>50</sup> of two different concepts of superconducting wigglers are <sup>51</sup> foreseen to be installed in the ANKA storage ring [4]. <sup>52</sup> The operational parameters for Nb-Ti are:  $B_{\rm w} = 3.0 \,{\rm T}$ , <sup>53</sup> 35 periods, 56 mm period length, a beam stay clear of  $_{54}$  13 mm and a magnetic gap q of 18 mm. The design values <sup>55</sup> for a Nb<sub>3</sub>Sn wiggler magnet are (except for the magnetic  $_{56}$  flux density which is  $B_{\rm w} = 4.0\,{\rm T}$ ) identical. The test <sup>57</sup> of the Nb<sub>3</sub>Sn damping wiggler can be performed in the 58 same cryostat.

59 The schematic layout of the cryostat is shown in Fig. 3. <sup>60</sup> The essential of the design is that the vacuum chamber 61 and superconducting coils have minimum thermal con-<sup>62</sup> tact and that they are cooled by using separate cooling damping rings. Moreover, the enthalpy margin for  $Nb_3Sn_{63}$  circuits. In between the vacuum chamber and the coils wiggler magnets with the same loadline margin as Nb-Ti  $_{64}$  is a small vacuum (10<sup>-4</sup> Pa) gap. The vacuum chamwiggler magnets is larger, which will result in more sta- 65 ber is spaced to the coils by using small spacers with 45 ble operation. First promising tests show that Nb<sub>3</sub>Sn 66 a low thermal conductivity. So, if the vacuum cham-46 wiggler magnets can reach the high magnetic flux den- 67 ber receives heat load from the beam the coils are only 47 sity, but more detailed studies are required before a final 68 slightly affected. The vacuum chamber with a vacuum



FIG. 2. Parameters of the wiggler magnets.



FIG. 3. Conceptual design of cryostat.

70 71 72 73 <sup>74</sup> ditionally necessary to coat the vacuum chambers with <sup>115</sup> and the list of the parameters for the CLIC damping rings 75 specialized coatings to reduce electron cloud effects (see 116 can be found in [8, 9]. Three effects are determining the <sup>76</sup> section IV B 2). The coils are cooled by a flow of liq- 117 equilibrium emittance: uid helium at atmospheric pressure flowing through heat 77 exchangers. This cooling technique reduces significantly <sup>118</sup> 78 the amount of stored helium. Since the liquid helium <sup>119</sup> 79 flows through the center of the coils the existing coils act 80 <sup>81</sup> as a rigid mechanical barrier between the vacuum and the liquid helium. Therefore, no additional pressure barrier <sup>122</sup> 82 between coils and beam vacuum as in standard wiggler 83 magnets with bath cooling (summarized in [6]) is needed 84 to ensure that the beam pipe is not plastically deformed 85 in case of quench due to the mechanical load of the pres-86 sure increase caused by the vaporizing helium (usually 87 up to  $10^6$  Pa). Therefore, the wiggler magnet with the 88 proposed cooling system can be operated with the same 89 beam stay clear but at a smaller pole distance resulting 90 in a considerably larger magnetic field compared to bath 91 cooled magnets because of the nonlinear magnetic flux 131 92 density to gap relation given in Equation (3). 93

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FIG. 4. Horizontal (top) and vertical (bottom) racetrack wiggler magnets.

97 quickly an optimized wiggler magnet design for damp-98 ing rings. Moreover, a full analysis of the heat load in <sup>99</sup> racetrack damping rings equipped with superconducting 100 damping wiggler magnets in series in a racetrack damp-<sup>101</sup> ing ring is derived. The heat load analysis shows that 102 conduction cooling has to be applied to superconducting wiggler magnets. So far, to the knowledge of the au-103 thors, all existing superconducting damping wiggler mag-104 <sup>105</sup> nets are cooled by using bath cooling. A design concept <sup>106</sup> for conduction cooling is presented.

### Α. **Design** requirements

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In the following section we discuss the choice of wig-108 <sup>109</sup> gler parameters and technology required to achieve the  $_{69}$  level smaller than  $10^{-7}$  Pa [5] is cooled by gaseous he-  $_{110}$  ultra-low emittance. Both bending magnets and wiggler lium and can be operated at temperature levels between 111 magnets contribute to the damping time and the emit-20 K and 80 K, cooling at 80 K would reduce the power 112 tances. A detailed general mathematical analysis of the consumption by about a factor of 5 compared to cooling 113 dependence of the emittances and the damping times on at 20 K. For the positron damping ring it might be ad- 114 the various parameters can be found in [7]. The concept

- Excitation and damping occurs due to the emission of synchrotron radiation. The damping is based on the effect that emission of synchrotron radiation reduces both the longitudinal and the transverse momentum of the electron. However, the cavities only restore the longitudinal momentum leading to a reduction of the transverse momentum. The equilibrium emittance calculated by considering these two effects is called zero-current emittance.
- Intrabeam scattering causes an increase of the zerocurrent emittance (determined by the effects mentioned above) due to small angle multiple Coulomb scattering.

The physical analysis of damping and excitation, with-<sup>132</sup> out considering the effect of intrabeam scattering, allows In this paper, a conceptual design for superconducting 133 the following conclusions to be drawn: in order to make vertical and horizontal racetrack damping wiggler mag- 134 the wiggler section as short as possible, the magnetic flux  $_{96}$  nets (see Fig. 4) is presented, which allows to derive very  $_{135}$  density  $B_{
m w}$  has to be large. Also the damping time au is



FIG. 5. Vertical and horizontal normalized emittances versus beam energy [3].

<sup>136</sup> shorter the larger  $B_{\rm w}$  is. But a high wiggler magnetic flux  $_{137}$  density  $B_{\rm w}$  only produces low emittances when the prod- $_{\tt 138}$  uct of  $|B^3_{\rm w}|$  and the period length  $\lambda^2_{\rm w}$  is small. Therefore, 139 the ideal damping ring has wigglers with a large magnetic flux density  $B_{\rm w}$  at a short period length  $\lambda_{\rm w}$ . 140

However, the effective emittance in the ring will be 141 larger than the zero-current emittance because of the ef-142 fect of intrabeam scattering. The final normalized emittance of the beam  $\gamma \epsilon_u$ , with u = x, y, z is given by: 144

$$\gamma \epsilon_u = \gamma \epsilon_{u,0} + \gamma \epsilon_{u,\text{IBS}}.$$
 (1)

The normalized emittance growth from intrabeam 177 145 146 147 that is,

$$\gamma \epsilon_{u,\text{IBS}} \propto \frac{N^Q}{\epsilon_x^{b_1} \epsilon_y^{b_2} \epsilon_z^{b_3}},\tag{2}$$

where N is the number of particles per bunch and 148  $\{Q, b_1, b_2, b_3\}$  is a set of positive real numbers. The effect 149 of intrabeam scattering in existing rings is small, but in 150 the first design studies [8] of the CLIC damping rings, 151 the magnitude of the final emittance was dominated by 152 intrabeam scattering. Therefore, in [3] the damping rings 153 were re-optimized such that  $\gamma \epsilon_x / \gamma \epsilon_{x,0} \leq 1.9$ . 154

Fig. 5 shows the calculated emittance taking into ac-155 count all the previously mentioned effects for beam ener-156 gies between 2 and 4 GeV. The emittances have a broad 157 minimum between 2 and 3 GeV. From Fig. 5, an energy 158 of 2.86 GeV was chosen. A more detailed analysis can be 159 found in [3]. 160

In the next step,  $B_{\rm w}$  and  $\lambda_{\rm w}$  for the damping wigglers 161 have to be specified. In Fig. 6 the area is marked where 162 the horizontal, vertical, and longitudinal emittance re-163 quirements of the CLIC damping rings are met. 164

The maximum achievable flux density strength  $B_{\rm w}$  as 165 166 a function of the gap to period length ratio  $g/\lambda_{\rm w}$  and the pole magnetic flux density  $B_{\rm p}$  can be derived in free 167 <sup>168</sup> space from Maxwell's magneto-static equations:

$$B_{\rm w} = \frac{B_{\rm p}}{\cosh\left(\pi \frac{g}{\lambda_{\rm w}}\right)},\tag{3}$$

TABLE I. Parameters of wiggler magnet prototypes.

	Nb-Ti HR Nb <sub>3</sub> Sn V					
	Short Sample					
$B_{\rm w}$	$3.6\mathrm{T}$	$5.5\mathrm{T}$				
max stored $E$	$80  \mathrm{kJ}$	$700  \mathrm{kJ}$				
$I_{\rm c}$ at $B_{\rm w}$	$740\mathrm{A}$	$1250\mathrm{A}$				
$B_{\rm s}$ at $B_{\rm w}$	$6.3\mathrm{T}$	$9.9\mathrm{T}$				
	Operational Parameters					
$B_{\rm w}^{\rm d}$	$3.0\mathrm{T}$	$4.0\mathrm{T}$				
$T_{\rm c}^{\rm d}$	$5.1\mathrm{K}$	$11.9\mathrm{K}$				
$I^{ ext{d}}$	$630\mathrm{A}$	$855\mathrm{A}$				
$B_{\rm s}^{\rm d}$	$5.4\mathrm{T}$	$6.7\mathrm{T}$				
oper. T	$4.2\mathrm{K}$	$4.2\mathrm{K}$				
gap	$18\mathrm{mm}$	$18\mathrm{mm}$				
$\lambda_{ m w}$	$56\mathrm{mm}$	$56\mathrm{mm}$				
K	15.7	20.9				
Strand	Bochvar Institute [12]	OST RRP $[13]$				

<sup>169</sup> where  $B_{\rm p}$  is the pole field on the iron (see Fig. 2).

The gap of superconducting wiggler magnets is deter-170 mined by the required aperture and space to intercept 171 <sup>172</sup> the heat and radiation load from the beam. An aperture 173 of 13 mm is foreseen in the damping rings. To intercept <sup>174</sup> the heat and radiation load from the beam, 2.5 mm space <sup>175</sup> on each side is required. Therefore, the magnetic gap was set to 18 mm for these calculations. 176

The normalized horizontal and vertical emittances scattering is proportional to the brightness of the beam, 178 have to be pre-damped from  $100 \,\mu \text{m} \, \text{rad}$  for electrons and  $9.7 \times 10^3 \,\mu\mathrm{m}\,\mathrm{rad}$  for positrons to  $63 \,\mu\mathrm{m}\,\mathrm{rad}$  (horizon-180 tal) and  $1.5 \,\mu\mathrm{m}\,\mathrm{rad}$  (vertical) in pre-damping rings [10]. <sup>181</sup> For the emittance calculations the normalized vertical 182 zero-current emittance was set to 4 nm rad, and the nor-183 malized longitudinal zero-current emittance was set to 184 6000 eV m. The total wiggler length in one ring was set 185 to 104 m. Fig. 6 shows the results of this study. Fig. 6 186 (top) shows the normalized horizontal emittance as it is 187 the most critical and can be used to decide on the re-188 quired wiggler parameters. The red and the blue curves 189 show the maximum achievable field for superconducting <sup>190</sup> wiggler magnets with Nb<sub>3</sub>Sn and Nb-Ti wire technology, <sup>191</sup> respectively. The red and the blue points present the pro-<sup>192</sup> posed Nb<sub>3</sub>Sn and Nb-Ti working points. The parameters <sup>193</sup> of the wiggler magnets represented by these points are 194 presented in Tab. I. Fig. 6 (bottom) presents the effect <sup>195</sup> of intrabeam scattering (IBS) on the normalized emit-<sup>196</sup> tance, for example a value of 1.2 means that 20% of the <sup>197</sup> emittance is generated by the effect of IBS.

> Fig. 6 (top) shows that only wigglers with a period 198 <sup>199</sup> length  $\lambda_{\rm w}$  of less than 80 mm and with a sinusoidal wig- $_{200}$  gler field with a magnetic flux density amplitude  $B_{\rm w}$ <sup>201</sup> larger than 2.2-2.5 T fulfill the requirements of the CLIC 202 damping rings. For period lengths  $\lambda_{\rm w}$  of 50 to 80 mm and  $_{\rm 203}$  magnetic flux densities  $B_{\rm w}$  of 2.8-4.5 T the effect of in-<sup>204</sup> trabeam scattering (Fig. 6 (bottom)) can be minimized. In the prototype phase the wiggler magnets will be 205

> <sup>206</sup> tested in the ANKA storage ring and will also be used 207 as a light source. Therefore, a smaller period length



FIG. 6. Equilibrium normalized horizontal emittance  $\gamma \epsilon_x$ (top) and the effect of IBS  $(\gamma \epsilon_x / \gamma \epsilon_{x,0})$ . The red and the blue curves show the maximum achievable magnetic flux density for superconducting wiggler magnets with Nb<sub>3</sub>Sn and Nb-Ti wire technology, respectively.

 $_{209}$  CLIC. The Nb-Ti wiggler magnet will be operated with  $_{252}$  the maximum achievable pole field  $B_p^*$  for a period length <sup>210</sup> a loadline margin of 85% and the Nb<sub>3</sub>Sn wiggler mag- <sup>253</sup>  $\lambda_{\rm w}$  at the engineering current density  $J_{\rm eng}^*$ . In this first <sup>211</sup> net with a loadline margin of 68%. The loadline margin <sup>254</sup> step of the optimization, the design pole field  $B_p^d$  is set  $_{212}$  for Nb<sub>3</sub>Sn was chosen larger for additional margin be-  $_{255}$  to the maximum achievable pole field  $B_p^*$  for a given pe-<sup>213</sup> cause the manufacturing and operation of Nb<sub>3</sub>Sn wiggler <sup>256</sup> riod length  $\lambda_w$  at the engineering current density  $J_{eng}^*$ . <sup>214</sup> magnets is purely R&D. Each of the 104 superconduct-<sup>257</sup> In Fig. 7 a gap of 18 mm was assumed, which can be <sup>215</sup> ing wiggler magnets has 34 periods. 26 wigglers will be <sup>258</sup> adjusted by using Equation (4). <sup>216</sup> installed in each of the straight sections of the two damp-<sup>259</sup> After determining the required engineering current  $_{217}$  ing rings. This discussion shows that the development  $_{260}$  density  $J_{eng}^*$  for the desired pole field in Fig. 7 (top), <sup>218</sup> of short-period superconducting wiggler magnets with a  $_{261} B_{\rm p}^{\rm d} = B_{\rm p}^*$ ; the maximum field on the conductor  $B_{\rm s}^*$  can  $_{219}$  high magnetic flux density strength  $B_{\rm w}$  and small gap  $g_{262}$  be read off from the intersection of  $J_{\rm eng}^*$  and the period 220 is required.

221 reach the required large magnetic flux densities. The 266 RRP) it can be, in principle, realized. 222  $_{223}$  magnetic saturation induction of iron (B = 2.15 T) can  $_{267}$  If a strand is used that is different from the two shown <sup>224</sup> be considered as the theoretical upper limit for the pole <sup>268</sup> in Fig. 7 the engineering current density has to be con-<sup>225</sup> field of hybrid-permanent magnet wigglers.

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### II. WIGGLER MAGNET DESIGN

In this section we present a conceptual design for su-227 228 perconducting wiggler magnets. From the requirements <sup>229</sup> of the CLIC damping rings, the gap size q, the period  $_{230}$  length  $\lambda_{\rm w}$  and the minimal required magnetic flux den- $_{231}$  sity in the center of the gap  $B_{\rm w}$  are assumed to be given. <sup>232</sup> The choice of the wiggler coil design, the choice of the <sup>233</sup> strand technology (Nb-Ti or Nb<sub>3</sub>Sn) and the choice of the wire bundle dimension are discussed in the follow-235 ing. Further on, in the case of vertical wiggler magnets, <sup>236</sup> the bending radii of the end-coils have to be determined. 237 The parameters required for the optimization are sum-<sup>238</sup> marized in Tab. II and illustrated in Fig. 2, where  $B_{\rm w}$ <sup>239</sup> refers to the amplitude of the magnetic flux density,  $B_{\rm p}$  $_{240}$  to the pole field and  $B_{\rm s}$  to the surface field on the conduc-<sup>241</sup> tor. We chose not to investigate the graded coils (coils <sup>242</sup> with changing engineering current density over the wire <sup>243</sup> bundle) because they require the use of multiple power <sup>244</sup> supplies and current leads. All magnetic calculation were <sup>245</sup> performed by using the Opera software package [11].

We assume that the design parameters  $B_{\rm w}^{\rm d}$ ,  $\lambda_{\rm w}^{\rm d}$  and  $_{247}g^{\rm d}$  are given; the design pole field  $B_{\rm p}^{\rm d}$  can be directly 248 calculated by using

$$B_{\rm p}^{\rm d} = B_{\rm w}^{\rm d} \cosh \pi \frac{g^{\rm d}}{\lambda_{\rm w}^{\rm d}}.$$
 (4)

249 Fig. 7 (top) can be used to find the required engineering  $_{250}$  current density  $J_{eng}^*$  from the intersection of  $B_{\rm p}^{\rm d} = B_{\rm p}^*$  $_{208} \lambda_{\rm w}$  was chosen to satisfy the needs of both ANKA and  $_{251}$  and the design period length  $\lambda_{\rm w}^{\rm d} = \lambda_{\rm w}$ , where  $_{*}^{*}$  denotes

 $_{263}$  length  $\lambda_{\rm w}$  in Fig. 7 (bottom). If the intersection point of  $_{264}$   $J_{\rm eng}^*$  and the period length  $\dot{\lambda}_{\rm w}^*$  are below the critical sur-Hybrid-permanent wiggler magnets are not able to 265 face shown for Nb-Ti (Bochvar Strand) or Nb<sub>3</sub>Sn (OST

<sup>269</sup> verted correspondingly into the current in the strand with



FIG. 7. Top: Optimized field at the pole tip  $B_{\rm p}^*$  versus engineering current density. Bottom: Load lines for wiggler magnets with a coil width  $(0.3\lambda_w = y_c = z_c)$  and a gap 18 mm. tion would yield a maximum pole field  $B_{\rm p}^* = 5.5 \,{\rm T}$ .

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$$I = \frac{1}{\kappa} J_{\text{eng}},\tag{5}$$

273 274 275 276 277 magnets, the layer jump can be conveniently placed on 312 from the wiggler body occur and have to be supported the side averted from the beam. In horizontal racetrack  $_{313}$  by an adequate mechanical structure. 278 coils, usually the layer jump is placed in the straight part, <sup>314</sup> Wiggler magnets have to be transparent for the beam. 279 280 resulting in an around 10% smaller filling factor com- 315 That means, the first and second field integrals pared to that of vertical wiggler magnets. 281

We have to introduce the critical surface in order to 282 determine if the chosen strand is superconducting at the 283 chosen  $B_{\rm s}$  (shown for the two sample strands in black in Fig. 7 (bottom)). A superconducting filament in a 285 <sup>286</sup> magnetic field is superconducting as long as the current is smaller than the critical current. For a Nb-Ti strand, <sup>316</sup> over the wiggler magnet have to vanish. Two different 288 the critical current is [14]

$$I_{\rm c} = C B_{\rm s}^{\alpha - 1} \left( 1 - \frac{B_{\rm s}}{B_{\rm c2}(T)} \right)^{\beta}.$$
 (6)



FIG. 8. Optimized end-coil radius r for Nb-Ti and Nb<sub>3</sub>Sn wiggler magnets with zero force in the end-coil.

<sup>289</sup> For the Nb-Ti Bochvar Institute strand,  $C = 3300, \alpha =$ 290 0.72,  $\beta = 1.1$ , and  $B_{c2}(4.2) = 10.68$ . For a Nb<sub>3</sub>Sn  $_{291}$  strand, the critical current is [15, 16]

$$I_{\rm c} = \frac{C}{\sqrt{B_{\rm s}}} \left( 1 - \frac{B_{\rm s}}{B_{\rm c2}(T,\epsilon)} \right)^2. \tag{7}$$

 $_{292}$  For the OST RRP 0.8 mm strand, C 11030, $_{293} B_{c2}(4.2,0) = 24.92.$ 

In vertical racetrack wigglers, the bending radius can 294  $_{\rm 295}$  be chosen at will. It is useful to choose it in a way that For example, in the bottom plot, we find that an engineering  $_{296}$  the Lorentz force acting on the wire bundle is compres-current density  $J_{eng}^* = 975 \,\mathrm{Amm}^{-2}$  yields a maximum sur-  $_{297}$  sive. Fig. 8 shows the bending radii for Nb-Ti and Nb<sub>3</sub>Sn face field  $B_s^* = 6 \,\mathrm{T}$  for a Nb-Ti strand at a period length  $_{296}$  wiggler magnets at which the force becomes zero. For  $\lambda_{\rm w} = 50$  mm. The top plot shows that this wiggler configura- 299 larger bending radii, the Lorentz force is compressive; for 300 smaller bending radii, the Lorentz force is tensile. In hor-<sup>301</sup> izontal racetrack wiggler magnets this option obviously 302 does not exist because the bending radii are given by 303 the pole width. Because the pole width is much smaller 304 than the radii shown in Fig. 8, the Lorentz forces in the 305 end-coils of horizontal wiggler magnets are tensile. In the  $_{271}$  where we define the filling factor  $\kappa$  for the wiggler mag-  $_{306}$  straight section, the wire bundles are compressed towards nets as the average number of wires per unit area. For 307 the unsaturated iron. If no iron is used, the forces will the calculations presented here, a filling factor of 1.24 308 be tensile. A yoke on top of a vertical wiggler magnet strands/mm<sup>2</sup> was assumed. As Fig. 7 reveals,  $B_{\rm p} \propto J_{\rm eng}$ . 309 reduces the resulting force on the wire bundles on the Therefore, the part of the coil generating the field seen 310 upper straight part. One has to be careful at the first by the beam must be densely wound. In vertical wiggler <sup>311</sup> and last coils of the wiggler magnets. Forces acting away

$$I_y = \int_{s_0}^{s_1} B_y(z) \mathrm{d}z,\tag{8}$$

$$II_{y} = \int_{s_{0}}^{s_{1}} \int_{s_{0}}^{s} B_{y}(z') dz' dz, \qquad (9)$$

317 designs are commonly used to compensate the first and <sup>318</sup> second field integral, either a symmetric (odd number of <sup>319</sup> poles) or an antisymmetric (even number of poles) design. <sup>320</sup> In superconducting wiggler magnets the antisymmetric <sup>321</sup> design is preferred. The first field integral automatically becomes zero and reduces the possibility of beam trips 322 in the case of a quench [17]. Further, the integrals of 323 the higher even multipoles (sextupole, decapole, ...) are 324 canceled automatically because they are equal in size but 325 different in sign. The second field integral has to be min-326 imized by varying the number of conductors or the ge-327 ometry of the last coils by using a numerical calculation 328 method. If the magnet is operated at different current 329 levels, the first and last pole should be thick enough to 330 not saturate. 331

In the last paragraph, we saw that all even multi-332 poles are canceled automatically. Finally we have to 333 ensure that the influence of the odd numbered multi-334 poles (quadrupoles, octupoles,...) remains also accept-335 ably small. Tolerances for the dynamic field integral  $\int B_u ds$  have to be determined by performing tracking 337  $_{338}$  studies to specify the allowable range, where s denotes  $_{339}$  the trajectory of the particles. It is shown in [18] that  $_{378}$  can be used for compensation of the field integrals. 340 it is not sufficient to consider  $\int B_y dz$ , where z follows the global coordinate of the overall beam direction. The 341 good-field region is defined as the region where the rela-342 tive change in magnetic flux density is less than  $1 \times 10^{-4}$ . 343 Fig. 9 shows this region for typical wiggler designs. The 344 maximum deflection of the beam in a wiggler magnet in 345 346 all values have to be entered in SI units. For  $B_{\rm w}=3\,{\rm T}$ 347 we find around  $\pm 30 \,\mu\text{m}$ . The beam size is smaller than 348 1.2 mm. We suggest to choose a good-field region of at 349 least 4 mm. 350

The final step in the magnet design of a superconduct-351 <sup>352</sup> ing wiggler is the quench analysis. A quench is a transition from the superconducting to the normal-conducting 353 state followed by a thermal runaway. When the super-354 conductor becomes normal conducting, the current flows 355 through the stabilizing copper of the strand causing re-356 sistive heating of the magnet, which has to be switched 357 off. Wigglers can be built and protected in modules; in 358 the extreme each half period can be protected separately 359 <sup>360</sup> by means of a parallel resistor within the magnet cold-361 mass, as was done for the LHC beam diagnostics undulator [19]. If a coil of the wiggler quenches, the parallel 362 resistor acts both as energy extraction for the quenched 363 coil and by-pass for the current. This protection scheme 364 365 is feasible because the stored energy in a single coil of a wiggler magnet is relatively small ( $< 10 \, \text{kJ}$ ). A PSpice 366 simulation was performed. As input parameter the re-367 sistance increase value after the occurrence of a quench 368 was extrapolated from the CERN-KIT short model wig-369 gler (Section III). The analysis has shown that even the 370 protection of many wiggler magnets powered in series is 371 372 feasible.

373

#### Influence of field errors Α.

374 375 sate both for small errors of the first and second field 414 nets with cryocoolers.



FIG. 9. Example of rolloff of wiggler magnets ( $\lambda_{\rm w} = 56 \,\mathrm{mm}$ ).

376 integrals and for small multi-pole errors due to mechani-<sup>377</sup> cal errors in the wiggler magnets [18]. Additional steerers

379 The horizontal equilibrium emittance and the damping <sup>380</sup> time [7] without considering the effect of intrabeam scat-<sup>381</sup> tering vary less than 5% over a hypothetical, one million 382 damping rings with Gaussian distributed errors of the  $_{383}$  magnetic flux density amplitude  $B_{\rm w}$  and period length  $_{384} \lambda_{\rm w}$ . The standard deviation for this study was set to: the CLIC damping rings is  $x = 2.65 \times 10^{-3} B_{\rm w} \lambda_{\rm w}^2$ , where  $_{385} \sigma(B_{\rm w}) = 0.2 \,\mathrm{T}$  and  $\sigma(\lambda_{\rm w}) = 1 \,\mathrm{mm}$ . A standard devia- $_{386}$  tion of  $\sigma(B_{\rm w}) = 0.2 \,{\rm T}$  corresponds to an error in the pole  $_{\rm 387}$  height of  $\pm 1.5\,{\rm mm}$  if  $B_{\rm p}={\rm const}$  and  $\lambda_{\rm w}={\rm const.}$  These <sup>388</sup> tolerances can be easily achieved. Therefore, no special 389 care has to be taken to maintain small tolerances during <sup>390</sup> the manufacturing of the wiggler magnets.

Notice that if the yaw, pitch, or roll angles obtained 391 <sup>392</sup> with magnet alignment do not meet the target values, <sup>393</sup> higher-order field components are introduced which are <sup>394</sup> not represented in the field model described above.

#### в. Temperature level of SC coils

395

The operation at 1.9 K instead of 4.3 K increases the 396 <sup>397</sup> performance of a Nb-Ti wiggler by around 20% ( $\lambda_{\rm w} =$ <sup>398</sup> 56 mm). However, a Nb<sub>3</sub>Sn wiggler with currently tested  $_{\rm 399}$  strands reaches almost the same  $B_{\rm w}$  at 4.3 K and 1.9 K  $_{400}$  due to self-field instabilities. Only the latest Nb<sub>3</sub>Sn strand development may reach larger currents at 1.9 K 401 [20]. Cooling at around 4.2 K can be economically pro-402 403 vided by standard cryocoolers. Therefore, cooling at  $_{404}$  1.9 K is not pursued. The power consumption at  $1.9\,\mathrm{K}$  $_{405}$  would be above 200 kW instead of 76.9 kW and could not be provided by cryocoolers (Section IV). 406

The wire bundle shown in Fig. 2 is encased in epoxy 408 resin. Therefore, the superfluid helium cannot come in <sup>409</sup> direct contact with the superconducting wire. As a result 410 the large thermal conductivity of superfluid helium at <sup>411</sup> 1.9 K is only of limited advantage. Therefore, we choose <sup>412</sup> to operate the wiggler magnets at 4.2 K, which eases cool-Magic fingers magnetic shims can be used to compen- 413 ing considerably and allows for cooling the wiggler mag-

TABLE III. Parameters of short models.

	CERN-KIT	CERN-BINP
Period, mm	40	50
Stored Energy, kJ	1	10
Gap, mm	16	20
$B_{\rm w}$ , T at $4.2{\rm K}$	1.9	2.2
$B_{\rm w}$ , T at $1.9{\rm K}$	2.4	—
$I_{\rm c}$ at 4.2 K, A	730	700
$B_{\rm s}$ at 4.2 K, T	4.8	6.5
$I_{\rm c}$ at $1.9{\rm K}$	910	—
Cu/Sc ratio	1.8/1	1/1.5
# poles	6	8
Strand	LHC #3 [22]	Bochvar Institute [12]

#### CONCEPTUAL DESIGN VERIFICATION TTT 415

Two Nb-Ti vertical racetrack short-models were suc-416 417 cessfully manufactured and tested in two independent collaborations. In the framework of the CERN col-418 <sup>419</sup> laboration with the Karlsruhe Institute of Technology (KIT), Germany, a 40 mm period, Nb-Ti short-model was manufactured reaching up to 2.4 T at 1.9 K. The 421 <sup>422</sup> magnet reached short sample current after 13 quenches. The results are published in [21]. In the framework of  $_{463}$ 423 the CERN-Budker Institute of Nuclear Physics (BINP), 424 <sup>425</sup> Russia, collaboration, Nb-Ti short-models with a period <sup>426</sup> length of 50 mm were manufactured reaching up to 2.2 T <sup>427</sup> at 4.2 K. The magnet reached short sample current af-428 ter 20 quenches. The parameters of the two success-<sup>429</sup> fully tested short models are summarized in Tab. III. <sup>430</sup> These tests have proven that Nb-Ti wiggler magnets in <sup>431</sup> the proposed parameter space can be manufactured and 432 operated.

433

### THERMAL DESIGN IV.

This section presents a thermal design for the CLIC 434 damping wigglers. It will be the first time that 26 su-435 perconducting wiggler magnets will be operated in one 436 437 straight section. In this chapter we address the chal-438 lenges related to this operation mode. For all following 439 calculations the baseline parameters were used. The wig-440 gler baseline design is the Nb-Ti HR design presented in 477 The presented values for the image current power depo-Tab. I (left column). The baseline energy of the CLIC 441 <sup>442</sup> damping rings is 2.86 GeV.

### 443 A. Magnetically and electrically induced heat loads

444 <sup>445</sup> flux density during the ramp-up or a quench of the wig-<sup>446</sup> gler magnets. For the test device in ANKA the ramping <sup>447</sup> time is crucial, because during injection the wiggler has 448 to be shut down. Therefore, the test device will be built

TABLE IV. Comparison of different joint technologies [23, 24].

	Electrolytic Cu	Soldered	US welded	Cold welding
Nb-Ti	$20\mathrm{n}\Omega\mathrm{cm}$	$40n\Omegacm$	$8\mathrm{n}\Omega\mathrm{cm}$	$< 1 \mathrm{p\Omega}\mathrm{cm}$
$Nb_3Sn$	$20\mathrm{n}\Omega\mathrm{cm}$	$40n\Omegacm$	-	${<}10\mathrm{n}\Omega\mathrm{cm}$

449 by using laminated iron to minimize eddy currents during 450 the ramp-up.

Resistive joints cannot be avoided. The horizontal 451 <sup>452</sup> racetrack wiggler magnets will have around 75 joints per <sup>453</sup> meter wiggler length. Tab. IV summarizes the achieved <sup>454</sup> resistance values normalized to a 1-cm overlap length. Cold welded Nb-Ti filaments joints lead to the smallest 456 resistances:

$$\frac{P}{L} < \frac{nRI^2}{L} < 1 \,\frac{\mathrm{mW}}{\mathrm{m}},\tag{10}$$

457 where P is the power, L the length of the wiggler, n the 458 number of joints, R the resistance of one joint, and I the 459 current in the strand. In the case of vertical racetrack <sup>460</sup> wiggler magnets, the only joints are the interconnections 461 to the current leads. This would yield for 4 interconnec-<sup>462</sup> tions operated at 855 A approximately 15 mW heat load.

#### В. Beam induced heat loads

### 1. Image currents

46/

In [25] the anomalous skin effect was measured for met-465 466 als with a small resistance at high frequencies. In this regime Ohm's law can no longer be applied because the 468 free path of the conduction electrons is similar to the  $_{469}$  penetration depth of the electric field [26].

At cryogenic temperatures the electrical resistivity of <sup>471</sup> pure metals such as OFHC copper is up to a factor of 300 472 lower than at room-temperature. Therefore, the average <sup>473</sup> power deposition per unit length due to the wakefield of 474 the beam in the extreme anomalous skin effect regime 475 of a cold beam pipe for aluminum or copper (coating) is 476 given by [27]:

$$P/L = \frac{\Gamma(\frac{5}{6})cZ_0}{4b\pi^2} \frac{I_{\rm av}^2}{\sigma_z^{\frac{5}{3}}\eta f_{\rm RF}} B_{\rm Mat} \approx 1 \,\frac{\rm W}{\rm m}.$$
 (11)

 $_{\rm 478}$  sition were calculated with the values defined and given 479 in Tab. V.

While good conductors do enter the anomalous skin ef-480 481 fect regime at cryogenic temperatures and high frequen-482 cies, poor conductors do not. For poor conductors such 483 as uncoated stainless steel or TiZrV ternary alloy, called Eddy currents will occur due to the varying magnetic 484 Non Evaporable Getter (NEG) coating, the heat load can <sup>485</sup> be estimated by normal skin effect formulas:

$$P/L = \frac{\Gamma(\frac{3}{4})c\sqrt{Z_0}}{\sqrt{32}b\pi^2} \frac{I_{av}^2}{\sigma_z^{\frac{3}{2}}\eta f_{\rm RF}} \frac{1}{\sqrt{\sigma_c}} \gtrsim 32 \,\frac{\rm W}{\rm m}.$$
 (12)

TABLE V. Parameters for image current calculations, see Equations (11) and (12); parameter values from [9, 27].

Parameter	Value	Unit	Explanation
$\Gamma(\frac{5}{c})$	1.13		Gamma-function
BAI	$3.3 \times 10^{-7}$	$m^{\frac{2}{3}}$	Material constant, Al
$B_{\rm Cu}$	$3.9 \times 10^{-7}$	$m^{\frac{2}{3}}$	Material constant, Cu
$\sigma_c$	$2 \times 10^6$	$\mathrm{Sm}^{-1}$	St. steel conductivity (4.3 K
$Z_0$	$120\pi$	Ω	Free space impedance
c	$3 \times 10^8$	m	Speed of light
$I_{\rm av}$	0.15	Å	Average current
b	$5.5 \times 10^{-3}$	m	Beam pipe radius (reduced)
$\sigma_z$	$1.4 \times 10^{-3}$	m	Bunch length
$\eta f_{\rm RF}$	$0.22\times1\times10^9$	$_{\rm Hz}$	Fraction of the ring
	circumference	occupi	ed by a bunch train for 1 GH
	312		Number of bunches
			$(2 \text{ trains} \times 156 \text{ bunches})$
	$4.1 \times 10^{9}$		Bunch population
	420	m	Ring circumference
	$1.4 \times 10^{-6}$	$\mathbf{S}$	Orbital period
	$10^{-9}$	$\mathbf{S}$	Bunch separation

486 A material with poor conductivity is not acceptable be-<sup>487</sup> cause of its high heat load. Therefore, all the following <sup>488</sup> calculations assume a high-conductivity material for the 489 beam pipe.

### 490

#### 2.Electron clouds

The heat load estimation due to electron clouds in the 491 electron and positron damping rings were performed with 492 the ECLOUD code [28]. Fig. 10 (top) shows the heat load 493 induced by electron clouds in the electron damping ring. 494 Multipacting is a phenomenon of resonant electron multi-495 plication in which a large number of electrons is built up, 496 leading to remarkable power losses, heating of the beam 497 pipe and beam instabilities [29]. The electron beam is 498 not affected by multipacting for values of a secondary 499 electron emission yield  $\delta_{\text{max}}$  up to 2.4. 500

501 502 503 504 505 506 507 508 509 510 511 512 <sup>514</sup> the beam pipe conductivity; surface treated grooved cop-<sup>515</sup> per or amorphous carbon might be another solutions [30]. <sup>534</sup> beam pipe. Rectangular absorbers are proposed, because <sup>516</sup> The effect of e-cloud at low emittance rings is studied in <sup>535</sup> of manufacturing constraints. The power emitted from <sup>517</sup> detail at CesrTA [31].



FIG. 10. Heat load induced by electron clouds in the electron damping ring (top). Heat load induced by electron clouds in the positron damping ring (bottom).

#### Synchrotron radiation 3.

518

Most synchrotron radiation generated in the damping 519 In the positron ring, multipacting appears for a sec- 520 wigglers has to be absorbed at ambient temperature to ondary electron emission yield  $\delta_{\text{max}} > 1.3$  and causes  $_{521}$  avoid heating of the superconducting coils. Fig. 11 shows significantly stronger e-cloud effects over one train pas- 522 the principle of synchrotron radiation emission and absage for values above 1.4–1.5; see Fig. 10 (bottom). As 523 sorption. The radiating charged particle is moving on electron clouds cause not only heat load but also beam in- 524 a sinusoidal trajectory in the horizontal xz-plane. The stabilities, low secondary electron emission yield coating 525 angles of observation in horizontal and vertical direcis needed for the positron damping ring. The needs for 526 tions are  $\theta$  and  $\psi$ . The emitted light, which will not the beam-pipe coating to avoid heating from e-clouds and  $_{527}$  be absorbed by up-stream absorbers (yz-plane), shown in from image currents are contradictory, because a large 528 cones, will partly irradiate the beam pipe. As an example electron emission yield means usually a low resistivity of  $_{529}$   $\psi_3^{max}$  is shown, the maximum angle at which wiggler W3 the beam-pipe coating. A thin-film low-secondary-yield 530 can irradiate synchrotron radiation on the beam-pipe. coating may be sufficient to reduce the secondary elec- 531 The heat load on the beam pipe is calculated on nodes, tron emission, such a thin-film coating may not decrease 532 which are equidistantly spaced. In this example only the <sup>536</sup> wiggler magnets is mainly concentrated within a light-



FIG. 11. Principle of synchrotron radiation emission (left), absorber scheme (middle), and absorber layout (right).

567

TABLE VI. Heat load on last wiggler beam pipe after a horizontal absorber for different wiggler lengths L.

L, m	1	2	3	4	5	6	
P, W	7	40	120	280	560	980	

TABLE VII. Lattice for heat load calculation of synchrotron radiation.

Element	Length, m	$2s_l$ , mm	$2r_l$ , mm	Shape
Ho	rizontal abso	orber		
Wiggler	2	13	80	Elliptical
Transition and quadrupole	0.25	13	80	Elliptical
Absorber	0.5	13.5	12.3	Rect.
Transition	0.25	13	40	Elliptical
V	ertical absor	ber		
Wiggler	2	13	80	Elliptical
Transition and quadrupole	0.25	13.5	40	Elliptical
Absorber	0.5	9.5	12.5	Rect.
Transition	0.25	13.5	40	Elliptical
Beam pipe (heat load)	2	13	80	Elliptical

537 cone of small opening angle  $K/\gamma \approx 3 \,\mathrm{mrad}$ , where K is the wiggler's deflection parameter and  $\gamma$  is the relative 538 <sup>539</sup> energy. However, the distance between the first wiggler 540 and the last wiggler is around 80 m, which implies that a large part of the generated synchrotron radiation would 541 542 irradiate the beam pipes of the up-stream superconducting wigglers. A detailed mathematical description of the 543 calculation method is given in Appendix A. 544

The heat load depends on the distance between the 545 absorbers and therefore on the wiggler length (Tab. VI). 546 From Tab. VI results the choice of two meter long wiggler 547 magnets, because the heat load from synchrotron radia-548 tion increases rapidly. Shorter wiggler magnets are not 549 chosen, because each cryostat requires an approximately 550 two times 0.5 m long cold-warm transition which reduces 551 the compactness of the ring. The corresponding lattice 552 design is presented in Tab. VII. 553

Fig. 12 shows the heat load distribution on the beam 554 pipe of the two last installed wiggler magnets (num-555 ber 25 and 26) downstream of absorbers installed after 556 quadrupoles aligned with their focusing planes. The hor-557 izontal aperture of the wiggler has to be large enough 558 to reduce the heat load from synchrotron radiation 568 559 (Tab. VIII). Fig. 13 shows the contribution of each wig-560 <sup>561</sup> gler to the total heat load on the last and next-to-last <sup>562</sup> wigglers. The total heat load from synchrotron radiation



FIG. 12. Spatial distribution of synchrotron radiation on the beam pipe for the 25th wiggler with a total heat load of 40 W (top) and the 26th wiggler with a total heat load of 2 W (bottom) for the Nb-Ti baseline design (Units in Figure:  $W/mm^2$ ).

TABLE VIII. Heat load on the 25th beam pipe for different horizontal apertures (widths).

Width, mm	30	40	50	60	70	80
Heat load,W	62.8	47.9	43.1	40.9	39.6	38.9

563 on the beam pipe downstream of a vertical absorber is <sup>564</sup> less than 2 W and around 40 W downstream of a horizon-565 tal absorber. The first six wiggler magnets in the straight <sup>566</sup> section of the damping rings are subject to less heat load.

### 4. Absorption of synchrotron radiation in the beam pipe

The critical photon energy  $\epsilon_c$  of the CLIC damping <sup>569</sup> wiggler can be calculated according to [32]:

$$\{\epsilon_c\}_{\rm keV} = 0.665 \{E\}_{\rm GeV}^2 \{B_{\rm w}\}_{\rm T} \approx 16 \,{\rm keV}.$$
 (13)



FIG. 13. Heat load on the 25th and 26th beam pipe of the wiggler magnets due to synchrotron radiation from upstream wiggler magnets.

 $_{\rm 570}$  The spectral intensity decreases rapidly for photon ener-  $_{\rm 571}$  gies above  $\epsilon_c.$ 

The measured intensity I transmitted through a layer of material with thickness l is related to the initial intensity  $I_0$  according to the Beer-Lambert law [33]:

$$\frac{I(l)}{I_0} = e^{-\mu l},$$
(14)

where l denotes the penetration depth. The attenua-575 tion coefficient is  $\mu$ . If we limit the transmitted energy 576 to 1% of the initial intensity, the penetration depth can 578 be calculated by  $l_{1\%} = 4.6 \frac{1}{\mu}$ . For a photon energy of  $_{579}$   $3\epsilon_c \approx 50 \,\mathrm{keV}$ , the attenuation coefficients for common 580 beam pipe materials and the corresponding penetration 581 depths are given in Tab. IX. The maximum penetration depth  $l_{\text{max}}$  at the full radiation cone opening angle, 582 583 shown in Fig. 14, is also given in Tab. IX. It was assumed that the synchrotron radiation changes the angle 584 to the beam axis when entering the vacuum chamber by 585 Snell's law. In the following it was assumed that this an-586 gle changes by a factor of 5. The maximum penetration 587 depth is much smaller than the material thickness of the 588 beam pipe (1 mm). Therefore, all synchrotron radiation 589 <sup>590</sup> will be absorbed in the inner part of the beam pipe and <sup>591</sup> no heat will be deposited directly into the superconduct-592 ing coils.



FIG. 14. Maximum penetration depth  $l_{\text{max}}$  of synchrotron radiation depending on the maximum opening angle of the radiation cone.

TABLE IX. Attenuation coefficients, densities and penetration depths for a photon energy of  $5 \times 10^{-2}$  MeV [33].

Material	$\mu/\rho,  \mathrm{cm}^2  \mathrm{g}^{-1}$	$\rho,{\rm gcm^{-3}}$	$\mu$ , cm <sup>-1</sup>	$l,  \mathrm{cm}$	$l_{\max},$	$\mu \mathrm{m}$
Al	0.3681	2.699	0.994	4.628	454	
Cu	2.613	8.960	23.41	0.1965	19	
Fe	1.958	7.874	15.42	0.298	29	

### 5. Absorber heat load

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604

The absorber design was optimized such that the heat for load on the vertical and horizontal absorbers is approximately balanced and reaches, after some 5 to 10 absorbers, its maximum value. Fig. 11 shows the principle solution of the absorbers. They will be water-cooled. After the last wiggler magnet a dump will be installed to absorb the remaining synchrotron radiation. Fig. 15 shows the heat load on each of the 26 absorbers.

<sup>602</sup> A detailed mathematical description of the calculation <sup>603</sup> method is given in Appendix B.

### C. Cooling concept and related heat loads

Fig. 16 compares the standard cooling concept (bath cooling, left) with the cooling concept for the CLIC damping wigglers proposed in the introduction (indirect cooling, right and Fig. 3). For indirect cooling the whole wiggler magnet is in an insulation vacuum. The helium is contained in the heat exchangers. In bath cool-



FIG. 15. Absorber load distribution including the dump at the end of the straight section.



FIG. 16. Two cooling concepts. Left: Bath cooling. The magnet is immersed in liquid helium (LHe). Right: Indirect cooling. The whole wiggler magnet is in vacuum. The helium is contained in the heat exchangers.

<sup>611</sup> ing the magnet is immersed in liquid helium (LHe). To 612 extract the heat a copper liner cooled with gaseous he-<sup>613</sup> lium contained in heat exchangers is required (depicted in red). Further, a mechanically stable pipe that can with-614 stand the pressure increase during a quench (in black) is 615 needed. This pipe also insulates the beam vacuum from 616 the surrounding liquid helium. Although indirect cooling 617 requires a detailed thermal design of the superconducting 618 coils it comes with a number of advantages compared to 619 bath cooling: 620

- Mechanically less demanding beam pipe. Bath 621 cooling would require a stainless steel beam pipe 622 to sustain the pressure increase during a quench. 623 Without this requirement a 2 mm smaller gap can 624 be realized, resulting in a more than 10% higher 625 gap-field  $B_{\rm w}$ . 626
- 627 lium is contained within the heat exchangers result-628 ing in smaller helium mass, smaller valves, tubes, 629 etc. 630
- Cryostat design for exchangeable coils and vacuum 631 pipes for maintenance and repairs is less complex. 632

An insulating vacuum of  $10^{-4}$  Pa will be established to  $\tilde{A} \ge 2 \times 10^{-3}$  m<sup>2</sup>. 633 minimize the heat transfer to the superconducting coils 657 634 635 636 637 638 639 640 small emissivities to reduce heat transfer by radiation. 641

642  $_{643}$  around 20 K where the heat conductivity k of copper is  $_{666}$  beam pipe can be cooled at a higher temperature level <sup>644</sup> highest (Fig. 17) resulting in an iso-thermal beam pipe. <sup>667</sup> (20 K), which greatly reduces the consumed power for 645 Cryocoolers for this temperature level are available, and 668 cooling. Tab. X gives a summary of all heat load sources 646 radiation to the superconducting coils is kept to a mini- 669 occurring in the CLIC damping wigglers at the different 647 mum.



FIG. 17. Heat conductivity of OFHC copper versus temperature. Data from [39].

TABLE X. Summary of the average heat loads (all sources) at the different temperature levels in Watts.

	Shield	Beam pipe	SC coils
	$80\mathrm{K}$	$20\mathrm{K}$	$4.2\mathrm{K}$
Synchrotron radiation (HA)		40.0	
Synchrotron radiation (VA)		2.0	
Average SR		21.0	
Image currents		2.0	
E-Clouds		0.02	
Radiation	16.3		0
Convection	1.0		0.10
Conduction	0.9		0.64
Joints	-		0.02
Current leads	72.0		0.16
Total (Average)	90.2	23.0	0.9

The heat load is deposited in the center of the cross-648 • Less complex cryogenic structure because all he- 649 section of the 2 m long beam pipe (Fig. 12) and has to be  $_{650}$  transported over  ${\lesssim}50\,\mathrm{mm}$  to the outer side of the beam <sup>651</sup> pipe. The resulting temperature difference is smaller <sup>652</sup> than  $\Delta T < \frac{1}{2} \frac{\dot{Q}l}{kA} \approx 0.3 \,\mathrm{K}$ , with the heat load  $\dot{Q} = 50 \,\mathrm{W}$ , <sup>653</sup> the distance from the center of the beam-pipe to the <sub>654</sub> heat exchanger  $l = 50 \,\mathrm{mm}$ , the averaged heat conductiv-655 ity  $k = 2000 \text{ W}(\text{m K})^{-1}$  (RRR = 80), and cross-section

An overview of heat leak calculation methods for by convection. The superconducting coils and the beam  $_{658}$  cryostats is given in [34–38]. The coils are in an insupipe will be mounted with minimal thermal contact and 659 lation vacuum; therefore, the total heat load on the coils with materials such as glass-fiber reinforced plastics or 660 results mainly from conduction (current leads, and fixing kevlar stripes with a large strength to heat conductiv- 661 structure) and is less than 1 W. Radiation and convecity ratio. All surfaces should be wrapped with polished 662 tion can be minimized by using a 60 to 80 K shield. No well-conducting metal foils or aluminized Mylar foils with 663 heat load induced by the beam is deposited directly into <sup>664</sup> the coils because all heat load can be intercepted by a The temperature of the beam pipe will be stabilized at 665 metallic beam pipe of a thickness of about 0.5 mm. The 670 temperature levels.

TABLE XI. Power consumption of the cryogenic plants at the different temperature levels. The total power consumption is around 450 kW.

80	20	4.2
11.3	2.9	0.113
33.3	6.7	1.3
23.0	19.3	11.3
147.5	224.3	76.9
	80 11.3 33.3 23.0 147.5	80         20           11.3         2.9           33.3         6.7           23.0         19.3           147.5         224.3

#### Large scale cooling concept and power D. 671 consumption 672

A total of 104 wiggler systems is needed for the two 673 CLIC damping rings. Two cooling schemes are investi-674 gated: (1) A large scale cryogenic system connected to 675 <sup>676</sup> all wiggler magnets with cryogenic transfer lines. (2) Each wiggler magnet is independently cooled with small 677 cryocoolers. In recent years cryocoolers have shown a considerable increase in performance, reliability and a 679 dramatic decrease in costs. 680

The efficiency of large cryogenic plants can be calcu-681  $_{682}$  lated according to the Carnot efficiency [40] and with efficiency data of large-scale cryogenic plants taken from 683 [41]. These data are in good agreement with more recent CERN experience (for example [42]): 685

$$P(T_{\rm op}) \approx \frac{\frac{T_{\rm op}}{T} - 1}{\alpha C^{\beta} - \gamma} C \quad , \tag{15}$$

686 where  $T_{\rm op} = 320 \,{\rm K}, \, C = [10, \dots, 10^6] \,{\rm W}$  is the capacity 687 of the cryoplant and  $T = [4.2, \ldots, 80]$  K the operating temperature of the cryoplant. The fitting parameters 688 are:  $\alpha = 35.7$ ,  $\beta = 0.05$ , and  $\gamma = 33.9$ . 689

Tab. XI presents the expected power consumption of 690 the two rings with large scale cryogenic plants. Losses 691 in the transfer-line can be kept small, for a well-shielded 692 line 0.05 W/m [43]. In this calculation the losses of the 693 distribution system were considered by adding 20% to 694 the total required heat budget. 695

As an alternative to a large scale cryogenic plant, three 696 different cryocoolers per wiggler with a total input power 698 of around 24 kW [44], yielding a total power consumption of around 2.5 MW for the two damping rings are 699 proposed. 700

701 702 energy consumption and will consume around 5 times 759 load from synchrotron radiation on the beam pipes of the 703 704 from the others, which may increase the reliability of the 762 the absorbers is derived and presented. 705 CLIC damping rings. The CLIC damping rings should 763 706 707 708 <sup>709</sup> using only a certain percentage of the damping wiggler <sup>766</sup> ally derived. Two large scale cooling concepts are pre-710 magnets.

### V. CONCLUSION

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A conceptual design of superconducting damping wig-712 713 glers for the CLIC damping rings was presented. Nb-Ti  $_{\rm 714}$  and Nb\_3Sn superconducting CLIC damping wiggler mag-<sup>715</sup> nets operated at 4.2 K meet the specifications. Hybrid-<sup>716</sup> permanent magnet wiggler magnets cannot achieve the 717 required magnetic fields. From the requirements of the CLIC damping rings, the gap size g, the period length 718  $_{719}$   $\lambda_{\rm w}$  and the minimal required magnetic flux density in the  $_{720}$  center of the gap  $B_{\rm w}$  of the wiggler magnets are given. 721 Then, the choice of the wiggler coil design, the choice of <sup>722</sup> the strand technology (Nb-Ti or Nb<sub>3</sub>Sn) and the choice 723 of the wire bundle dimension was discussed in detail in <sup>724</sup> this paper. By using the presented methods and figures 725 the superconducting wiggler design becomes straight forward. The quench behavior is not critical, the dipole field 726 quality in the wiggler magnets required for the damp-727 728 ing rings can be met easily. Further, was discussed that 729 operation at 1.9 K is only of limited advantage and will <sup>730</sup> therefore not be pursued. After showing that the magnet 731 design of wiggler magnets for the CLIC damping rings is 732 feasible, the design was verified by building two Nb-Ti short models, which both reached short sample current. The usage of Nb<sub>3</sub>Sn superconductor wire may increase 734 <sup>735</sup> the mid plane flux density  $B_{\rm w}$  by 50%, leading to a bet-<sup>736</sup> ter beam quality. However, the manufacturing of Nb<sub>3</sub>Sn <sup>737</sup> wiggler magnets is challenging due to the required heat 738 treatment and the brittleness of the strand afterwards. <sup>739</sup> The first trial coils showed that Nb<sub>3</sub>Sn wiggler magnets 740 are within the technological reach. The results will be published elsewhere. 741

The system integration of the superconducting damping wiggler was discussed in detail. Therefore, all known 744 heat load sources from the beam and from operation are 745 qualitatively and quantitatively discussed. Eddy cur-746 rents and the heat load from resistive joints are small. 747 The major source of heat load is beam-induced. Im-748 age currents and electron clouds can be kept small, but 749 have contradictory requirements. To reduce the electron <sup>750</sup> clouds, especially in the positron ring, the beam pipe 751 should be coated with a material with small electron <sup>752</sup> emission yield, usually a material with poor conductivity. 753 On the other hand, for the reduction of image currents <sup>754</sup> the surface of the beam pipe should have high conduc-<sup>755</sup> tivity. Surface treated grooved copper requires further 756 studies. The most important beam induced heat load is <sup>757</sup> from synchrotron radiation. In this paper an optimized Large scale cryogenic plants are favorable in terms of <sup>758</sup> absorber scheme is presented with the corresponding heat less electrical energy than small cryocoolers. But cry- 760 wiggler magnets. Further, the calculation scheme for the ocoolers allow each wiggler to be operated independently 761 heat load on the beam pipes of the wiggler magnets and

The heat load study is used to discuss the cooling conbe designed with a certain amount of redundancy, that 764 cept. Standard bath cooling is not an option; therefore, is, the emittance and damping time can also be met by 765 an advanced cooling concept is presented and conceptu-<sup>767</sup> sented and compared to each other. Large scale cryogenic

TABLE XII. Nomenclature for synchrotron radiation calcualtion

Symbol	Description
a, b	1/2 of beam pipe's x and y dimension
d	Distance
$d_{\mathrm{e}},d_{\mathrm{z}}$	Element size in $xy$ and $xz$ -plane
k,l	Sequence number of wiggler and absorber
n,m	Number of elements in $xy$ and $xz$ -plane
p	Number of downstream wigglers
$r_l, s_l$	1/2 of <i>l</i> -th absorber's $x$ and $y$ dimension
$X_{i,j}, Y_{i,j}, Z_{i,j,k}$	Node's coordinates
$ heta,\psi$	Angle between ray and $xz$ and $yz$ -plane
$\theta_k^{\max},  \psi_k^{\max}$	Max angle of $k$ 's ray irradiating beam pipe
$ heta_{k,l}^{\max},\psi_{k,l}^{\max}$	Max angle of $k$ 's ray irradiating $l$

<sup>768</sup> plants are favorable in terms of energy consumption and <sup>769</sup> will consume around 5 times less electrical energy than <sup>770</sup> small cryocoolers. But cryocoolers allow independent opr71 eration of each wiggler, which may increase the reliability 772 of the CLIC damping rings. The technical concept will <sup>773</sup> be tested at the ANKA storage ring in Karlsruhe.

### Appendix A: Calculation method of synchrotron 774 radiation heat load 775

The radiation power emitted by an electron moving on 799 776 a sinusoidal trajectory is [45, 46]:

$$\frac{\mathrm{d}P}{\mathrm{d}\Omega} = \frac{\mathrm{d}^2 P}{\mathrm{d}\theta \mathrm{d}\psi} = 3\frac{\gamma^2}{\pi^2} P_{\mathrm{T}} f_{\mathrm{K}}(\gamma\theta,\gamma\psi), \qquad (A1)$$

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<sup>778</sup> where  $\theta$  and  $\psi$  are the observation angles in the horizon-<sup>779</sup> tal and vertical directions (Fig. 11), and  $P_{\rm T}$  is the total 780 power integrated over all angles and frequencies:

$$\{P_{\rm T}\}_{\rm kW} \approx 0.633 \{E\}_{\rm GeV}^2 \{B_{\rm w}\}_{\rm T}^2 \{L\}_{\rm m} \{I\}_{\rm A},$$
 (A2)

where L is the total length of the wiggler and I is the  $^{807}$ 781 average current of the beam. The total dissipated power 782 from one Nb-Ti wiggler magnet will be 14 kW. 783

Tab. XII describes the nomenclature used in the fol-784 lowing derivation. The angular dependence is given by: <sup>811</sup> the positive quadrant are given by: 785 786

$$f_{\rm K}(\gamma\theta,\gamma\psi) = \int_{-\pi}^{\pi} \sin^2 \alpha \left(\frac{1}{D^3} - \frac{4\left(\gamma\theta - K\cos\alpha\right)^2}{D^5}\right) \mathrm{d}\alpha,\tag{A3}$$

788 parameter K is given by:

$$K = 0.934 \{\lambda_{\rm w}\}_{\rm cm} \{B_{\rm w}\}_{\rm T} \approx 15.7.$$
 (A4)

Equation (A1) cannot be integrated analytically and 789 therefore a numerical method was developed to calculate 790 the heat load on the beam pipe. 791

Fig. 18 shows the cross section of the elliptic beam 792 <sup>793</sup> pipe with the axes a and b. The ellipse is divided into a



FIG. 19. Example for an absorber scheme. Left: Heat load from each wiggler magnet k (here k = 1) on each absorber l (here l = 1, 2, 3, 4). The heat load from additional wiggler magnets can be superimposed. Right: Absorber.

<sup>794</sup> mesh of 4n equidistant segments with a length of  $d_{\rm e} = \frac{C}{4n}$  $_{795}$  (C is the circumference of the ellipse). In the example <sup>796</sup> shown in Fig. 18 n is 5. In a next step it is investigated <sup>797</sup> how much radiation hits the beam pipe segment from a downstream wiggler. In these calculations it is assumed 798 that absorbers as shown in Fig. 19 are in position. A <sup>800</sup> computer program was written to calculate the radiation <sup>801</sup> generated in each longitudinal segment and also to calcu-<sup>802</sup> late if and where the radiation hits the beam pipe. The <sup>803</sup> total power deposited on the beam pipe is finally inte-<sup>804</sup> grated and allows calculating the total deposited power 805 on the vacuum chamber.

### Appendix B: Heat load on absorbers

The heat load on each quadrant of the absorber l ir-<sup>808</sup> radiated by each wiggler k can be calculated by triple <sup>809</sup> numerical integration of Equation (A1). The positive <sup>\$10</sup> boundaries  $\theta_{k,l}$  and  $\psi_{k,l}$  (Fig. 19) for the calculation of

$$\theta_{k,l} = \arctan \frac{r_l}{d_{k,l}},\tag{B1}$$

$$\psi_{k,l} = \arctan \frac{s_l}{d_{k,l}},\tag{B2}$$

where  $D = 1 + (\gamma \psi)^2 + (\gamma \theta - K \cos \alpha)^2$  and the deflection where  $d_{k,l}$  is the distance between the wiggler k and the absorber l. The parameters  $r_l$  and  $s_l$  are the horizontal  $_{\rm 814}$  and vertical apertures of the absorber l.

The maximum angles  $\theta_{k,l}^{\max}$  and  $\psi_{k,l}^{\max}$  at which syns16 chrotron radiation from wiggler k hits the absorber l is:

$$\theta_{k,l}^{\max} = \min_{i=[k,l-1]} \theta_{k,i} \quad \text{with} \quad \theta_{k,k} = \frac{\pi}{2}, \qquad (B3)$$

$$\psi_{k,l}^{\max} = \min_{i=[k,l-1]} \psi_{k,i}$$
 with  $\psi_{k,k} = \frac{\pi}{2}$ . (B4)

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