

Strange particle production in proton–proton collisions at $\sqrt{s} = 0.9$ TeV with ALICE at the LHC

The ALICE Collaboration

K. Aamodt¹, A. Abrahantes Quintana², D. Adamová³, A.M. Adare⁴, M.M. Aggarwal⁵, G. Aglieri Rinella⁶, A.G. Agocs⁷, S. Aguilar Salazar⁸, Z. Ahammed⁹, N. Ahmad¹⁰, A. Ahmad Masoodi¹⁰, S.U. Ahn^{11,b}, A. Akindinov¹², D. Aleksandrov¹³, B. Alessandro¹⁴, R. Alfaro Molina⁸, A. Alici^{15,16,c}, A. Alkin¹⁷, E. Almaráz Aviña⁸, T. Alt¹⁸, V. Altini^{19,20}, S. Altinpinar²¹, I. Altsybeev²², C. Andrei²³, A. Andronic²¹, V. Anguelov^{24,d,e}, C. Anson²⁵, T. Antičić²⁶, F. Antinori^{27,28}, P. Antonioli¹⁶, L. Aphecetche²⁹, H. Appelshäuser³⁰, N. Arbor^{31,32}, S. Arceci^{15,16}, A. Arend³⁰, N. Armesto³³, R. Arnaldi¹⁴, T. Aronsson⁴, I.C. Arsene²¹, A. Asryan²², A. Augustinus⁶, R. Averbek²¹, T.C. Awes³⁴, J. Äystö^{35,36}, M.D. Azmi¹⁰, M. Bach¹⁸, A. Badalà³⁷, Y.W. Baek^{11,b}, S. Bagnasco¹⁴, R. Bailhache³⁰, R. Bala^{38,14,f}, R. Baldini Ferrolli³⁹, A. Baldisseri⁴⁰, A. Baldit^{41,42}, J. Bán⁴³, R. Barbera^{44,37}, F. Barile^{19,20}, G.G. Barnaföldi⁷, L.S. Barnby⁴⁵, V. Barret^{41,42}, J. Bartke⁴⁶, M. Basile^{15,16}, N. Bastid^{41,42}, B. Bathen⁴⁷, G. Batigne²⁹, B. Batyunya⁴⁸, C. Baumann³⁰, I.G. Bearden⁴⁹, H. Beck³⁰, I. Belikov⁵⁰, F. Bellini^{15,16}, R. Bellwied^{51,g}, E. Belmont-Moreno⁸, S. Beole^{38,14}, I. Berceanu²³, A. Bercuci²³, E. Berdermann²¹, Y. Berdnikov⁵², L. Betev⁶, A. Bhasin⁵³, A.K. Bhati⁵, L. Bianchi^{38,14}, N. Bianchi⁵⁴, C. Bianchin^{27,28}, J. Bielčik⁵⁵, J. Bielčiková³, A. Bilandžić⁵⁶, E. Biolcati^{38,14}, A. Blanc^{41,42}, F. Blanco⁵⁷, F. Blanco⁵⁸, D. Blau¹³, C. Blume³⁰, M. Boccioni⁶, N. Bock²⁵, A. Bogdanov⁵⁹, H. Bøggild⁴⁹, M. Bogolyubsky⁶⁰, L. Boldizsár⁷, M. Bombara⁶¹, C. Bombonati^{27,28}, J. Book³⁰, H. Borel⁴⁰, C. Bortolin^{27,h}, S. Bose⁶², F. Bossú^{38,14}, M. Botje⁵⁶, S. Böttger²⁴, B. Boyer⁶³, P. Braun-Munzinger²¹, L. Bravina⁶⁴, M. Bregant^{65,66,i}, T. Breitner²⁴, M. Broz⁶⁷, R. Brun⁶, E. Bruna⁴, G.E. Bruno^{19,20}, D. Budnikov⁶⁸, H. Buesching³⁰, O. Busch⁶⁹, Z. Buthelezi⁷⁰, D. Caffarri^{27,28}, X. Cai⁷¹, H. Caines⁴, E. Calvo Villar⁷², P. Camerini^{65,66}, V. Canoa Roman^{6,j,k}, G. Cara Romeo¹⁶, F. Carena⁶, W. Carena⁶, F. Carminati⁶, A. Casanova Díaz⁵⁴, M. Caselle⁶, J. Castillo Castellanos⁴⁰, V. Catanescu²³, C. Cavicchioli⁶, P. Cerello¹⁴, B. Chang^{35,36}, S. Chapeland⁶, J.L. Charvet⁴⁰, S. Chattopadhyay⁶², S. Chattopadhyay⁹, M. Cherney⁷³, C. Cheshkov^{74,75}, B. Cheynis^{74,75}, E. Chiavassa¹⁴, V. Chibante Barroso⁶, D.D. Chinellato⁷⁶, P. Chochula⁶, M. Chojnacki^{77,78}, P. Christakoglou^{77,78}, C.H. Christensen⁴⁹, P. Christiansen⁷⁹, T. Chujo⁸⁰, C. Cicalo⁸¹, L. Cifarelli^{15,16}, F. Cindolo¹⁶, J. Cleymans⁷⁰, F. Coccetti³⁹, J.-P. Coffin⁵⁰, S. Coli¹⁴, G. Conesa Balbastre⁵⁴, Z. Conesa del Valle²⁹, P. Constantin⁶⁹, G. Contin^{65,66}, J.G. Contreras⁸², T.M. Cormier⁵¹, Y. Corrales Morales^{38,14}, I. Cortés Maldonado⁸³, P. Cortese^{84,85}, M.R. Cosentino⁷⁶, F. Costa⁶, M.E. Cotallo⁵⁷, E. Crescio⁸², P. Crochet^{41,42}, E. Cuautle⁸⁶, L. Cunqueiro⁵⁴, G.D. Erasmo^{19,20}, A. Dainese^{87,n}, H.H. Dalsgaard⁴⁹, A. Danu⁸⁸, D. Das⁶², I. Das⁶², A. Dash⁸⁹, S. Dash¹⁴, S. De⁹, A. De Azevedo Moregula⁵⁴, G.O.V. de Barros⁹⁰, A. De Caro^{91,92}, G. de Cataldo²⁰, J. de Cuveland¹⁸, A. De Falco^{93,81}, D. De Gruttola^{91,92}, N. De Marco¹⁴, S. De Pasquale^{91,92}, R. De Remigis¹⁴, R. de Rooij^{77,78}, H. Delagrange²⁹, Y. Delgado Mercado⁷², G. Dellacasa^{84,85,o}, A. Deloff⁹⁴, V. Demanov⁶⁸, E. Dénes⁷, A. Deppman⁹⁰, D. Di Bari^{19,20}, C. Di Giglio^{19,20}, S. Di Liberto⁹⁵, A. Di Mauro⁶, P. Di Nezza⁵⁴, T. Dietel⁴⁷, R. Divià⁶, Ø. Djuvsland¹, A. Dobrin^{51,p}, T. Dobrowolski⁹⁴, I. Domínguez⁸⁶, B. Dönigus²¹, O. Dordic⁶⁴, O. Driga²⁹, A.K. Dubey⁹, J. Dubuisson⁶, L. Ducroux^{74,75}, P. Dupieux^{41,42}, A.K. Dutta Majumdar⁶², M.R. Dutta Majumdar⁹, D. Elia²⁰, D. Emschermann⁴⁷, H. Engel²⁴, H.A. Erdal⁹⁶, B. Espagnon⁶³, M. Estienne²⁹, S. Esumi⁸⁰, D. Evans⁴⁵, S. Evrard⁶, G. Eyyubova⁶⁴, C.W. Fabjan^{6,q}, D. Fabris²⁷, J. Faivre^{31,32}, D. Falchieri^{15,16}, A. Fantoni⁵⁴, M. Fasel²¹, R. Fearick⁷⁰, A. Fedunov⁴⁸, D. Fehlker¹, V. Fekete⁶⁷, D. Felea⁸⁸, G. Feofilov²², A. Fernández Téllez⁸³, A. Ferretti^{38,14}, R. Ferretti^{84,85,r}, M.A.S. Figueredo⁹⁰, S. Filchagin⁶⁸, R. Fini²⁰, D. Finogeev⁹⁷, F.M. Fionda^{19,20}, E.M. Fiore^{19,20}, M. Floris⁶, S. Foertsch⁷⁰, P. Foka²¹, S. Fokin¹³, E. Fragiaco⁶⁶, M. Fragkiadakis⁹⁸, U. Frankenfeld²¹, U. Fuchs⁶, F. Furano⁶, C. Furget^{31,32}, M. Fusco Girard^{91,92}, J.J. Gaardhøje⁴⁹, S. Gadrat^{31,32}, M. Gagliardi^{38,14}, A. Gago⁷², M. Gallio^{38,14}, P. Ganoti^{98,s}, C. Garabatos²¹, R. Gemme^{84,85}, J. Gerhard¹⁸, M. Germain²⁹, C. Geuna⁴⁰, A. Gheata⁶, M. Gheata⁶, B. Ghidini^{19,20}, P. Ghosh⁹, M.R. Girard⁹⁹, G. Giraud¹⁴, P. Giubellino^{38,14,t}, E. Gladysz-Dziadus⁴⁶, P. Glässel⁶⁹, R. Gomez¹⁰⁰, L.H. González-Trueba⁸, P. González-Zamora⁵⁷, H. González Santos⁸³, S. Gorbunov¹⁸, S. Gotovac¹⁰¹, V. Grabski⁸, R. Grajcarek⁶⁹, A. Grelli^{77,78}, A. Grigoras⁶, C. Grigoras⁶, V. Grigoriev⁵⁹, A. Grigoryan¹⁰², S. Grigoryan⁴⁸,

B. Grinyov¹⁷, N. Grion⁶⁶, P. Gros⁷⁹, J.F. Grosse-Oetringhaus⁶, J.-Y. Grossiord^{74,75}, R. Grosso²⁷, F. Guber⁹⁷, R. Guernane^{31,32}, C. Guerra Gutierrez⁷², B. Guerzoni^{15,16}, K. Gulbrandsen⁴⁹, T. Gunji¹⁰³, A. Gupta⁵³, R. Gupta⁵³, H. Gutbrod²¹, Ø. Haaland¹, C. Hadjidakis⁶³, M. Haiduc⁸⁸, H. Hamagaki¹⁰³, G. Hamar⁷, J.W. Harris⁴, M. Hartig³⁰, D. Hasch⁵⁴, D. Hasegan⁸⁸, D. Hatzifotiadou¹⁶, A. Hayrapetyan^{102,r}, M. Heide⁴⁷, M. Heinz⁴, H. Helstrup⁹⁶, A. Herghelegiu²³, C. Hernández²¹, G. Herrera Corral⁸², N. Herrmann⁶⁹, K.F. Hetland⁹⁶, B. Hicks⁴, P.T. Hille⁴, B. Hippolyte⁵⁰, T. Horaguchi⁸⁰, Y. Hori¹⁰³, P. Hristov⁶, I. Hřivnáčová⁶³, M. Huang¹, S. Huber²¹, T.J. Humanic²⁵, D.S. Hwang¹⁰⁴, R. Ichou²⁹, R. Ilkaev⁶⁸, I. Ilkiv⁹⁴, M. Inaba⁸⁰, E. Incani^{93,81}, G.M. Innocenti^{38,14}, P.G. Innocenti⁶, M. Ippolitov¹³, M. Irfan¹⁰, C. Ivan²¹, A. Ivanov²², M. Ivanov²¹, V. Ivanov⁵², A. Jacholkowski⁶, P.M. Jacobs¹⁰⁵, L. Jancurová⁴⁸, S. Jangal⁵⁰, R. Janik⁶⁷, S.P. Jayarathna^{58,u}, S. Jena¹⁰⁶, L. Jirden⁶, G.T. Jones⁴⁵, P.G. Jones⁴⁵, P. Jovanović⁴⁵, H. Jung¹¹, W. Jung¹¹, A. Jusko⁴⁵, S. Kalcher¹⁸, P. Kaliňák⁴³, M. Kalisky⁴⁷, T. Kalliokoski^{35,36}, A. Kalweit¹⁰⁷, R. Kamermans^{77,78,o}, K. Kanaki¹, E. Kang¹¹, J.H. Kang¹⁰⁸, V. Kaplin⁵⁹, O. Karavichev⁹⁷, T. Karavicheva⁹⁷, E. Karpechev⁹⁷, A. Kazantsev¹³, U. Kbschull²⁴, R. Keidel¹⁰⁹, M.M. Khan¹⁰, A. Khanzadeev⁵², Y. Kharlov⁶⁰, B. Kileng⁹⁶, D.J. Kim^{35,36}, D.S. Kim¹¹, D.W. Kim¹¹, H.N. Kim¹¹, J.H. Kim¹⁰⁴, J.S. Kim¹¹, M. Kim¹¹, M. Kim¹⁰⁸, S. Kim¹⁰⁴, S.H. Kim¹¹, S. Kirsch^{6,v}, I. Kisel^{24,e}, S. Kiselev¹², A. Kisiel⁶, J.L. Klay¹¹⁰, J. Klein⁶⁹, C. Klein-Bösing⁴⁷, M. Kliemant³⁰, A. Klovning¹, A. Kluge⁶, M.L. Knichel²¹, K. Koch⁶⁹, M.K. Köhler²¹, R. Kolevatov⁶⁴, A. Kolojvari²², V. Kondratiev²², N. Kondratyeva⁵⁹, A. Konevskih⁹⁷, E. Kornas⁴⁶, C. Kottachchi Kankanamge Don⁵¹, R. Kour⁴⁵, M. Kowalski⁴⁶, S. Kox^{31,32}, K. Kozlov¹³, J. Kral^{35,36}, I. Králik⁴³, F. Kramer³⁰, I. Kraus^{107,w}, T. Krawutschke^{69,x}, M. Kretz¹⁸, M. Krivda^{45,y}, D. Krumbhorn⁶⁹, M. Krus⁵⁵, E. Kryshen⁵², M. Krzewicki⁵⁶, Y. Kucheriaev¹³, C. Kuhn⁵⁰, P.G. Kuijer^{56,a}, P. Kurashvili⁹⁴, A. Kurepin⁹⁷, A.B. Kurepin⁹⁷, A. Kuryakin⁶⁸, S. Kushpil³, V. Kushpil³, M.J. Kweon⁶⁹, Y. Kwon¹⁰⁸, P. La Rocca^{44,37}, P. Ladrón de Guevara^{57,z}, V. Lafage⁶³, C. Lara²⁴, D.T. Larsen¹, C. Lazzeroni⁴⁵, Y. Le Bornec⁶³, R. Lea^{65,66}, K.S. Lee¹¹, S.C. Lee¹¹, F. Lefèvre²⁹, J. Lehnert³⁰, L. Leistam⁶, M. Lenhardt²⁹, V. Lenti²⁰, I. León Monzón¹⁰⁰, H. León Vargas³⁰, P. Lévai⁷, X. Li¹¹¹, R. Lietava⁴⁵, S. Lindal⁶⁴, V. Lindenstruth^{24,e}, C. Lippmann^{6,w}, M.A. Lisa²⁵, L. Liu¹, V.R. Loggins⁵¹, V. Loginov⁵⁹, S. Lohn⁶, D. Lohner⁶⁹, X. Lopez^{41,42}, M. López Noriega⁶³, E. López Torres², G. Løvholden⁶⁴, X.-G. Lu⁶⁹, P. Luettig³⁰, M. Lunardon^{27,28}, G. Luparello^{38,14}, L. Luquin²⁹, C. Luzzi⁶, K. Ma⁷¹, R. Ma⁴, D.M. Madagodahettige-Don⁵⁸, A. Maevskaya⁹⁷, M. Mager⁶, D.P. Mahapatra⁸⁹, A. Maire⁵⁰, M. Malaev⁵², I. Maldonado Cervantes⁸⁶, D. Mal'Kevich¹², P. Malzacher²¹, A. Mamonov⁶⁸, L. Manceau^{41,42}, L. Mangotra⁵³, V. Manko¹³, F. Manso^{41,42}, V. Manzari²⁰, Y. Mao^{71,aa}, J. Mareš¹¹², G.V. Margagliotti^{65,66}, A. Margotti¹⁶, A. Marín²¹, I. Martashvili¹¹³, P. Martinengo⁶, M.I. Martínez⁸³, A. Martínez Davalos⁸, G. Martínez García²⁹, Y. Martynov¹⁷, A. Mas²⁹, S. Masciocchi²¹, M. Maserà^{38,14}, A. Masoni⁸¹, L. Massacrier^{74,75}, M. Mastro marco²⁰, A. Mastroserio⁶, Z.L. Matthews⁴⁵, A. Matyja^{46,i}, D. Mayani⁸⁶, G. Mazza¹⁴, M.A. Mazzoni⁹⁵, F. Meddi^{114,95}, A. Menchaca-Rocha⁸, P. Mendez Lorenzo⁶, J. Mercado Pérez⁶⁹, P. Mereu¹⁴, Y. Miake⁸⁰, J. Midori¹¹⁵, L. Milano³⁸, J. Milosevic^{64,ab}, A. Mischke^{77,78}, D. Miśkowiec^{21,t}, C. Mitu⁸⁸, J. Mlynarz⁵¹, B. Mohanty⁹, L. Molnar⁶, L. Montaño Zetina⁸², M. Monteno¹⁴, E. Montes⁵⁷, M. Morando^{27,28}, D.A. Moreira De Godoy⁹⁰, S. Moretto^{27,28}, A. Morsch⁶, V. Muccifora⁵⁴, E. Mudnic¹⁰¹, H. Müller⁶, S. Muhuri⁹, M.G. Munhoz⁹⁰, J. Munoz⁸³, L. Musa⁶, A. Musso¹⁴, B.K. Nandi¹⁰⁶, R. Nania¹⁶, E. Nappi²⁰, C. Natrass¹¹³, F. Navach^{19,20}, S. Navin⁴⁵, T.K. Nayak⁹, S. Nazarenko⁶⁸, G. Nazarov⁶⁸, A. Nedosekin¹², F. Nendaz^{74,75}, J. Newby¹¹⁶, M. Nicassio^{19,20}, B.S. Nielsen⁴⁹, S. Nikolaev¹³, V. Nikolic²⁶, S. Nikulin¹³, V. Nikulin⁵², B.S. Nilsen⁷³, M.S. Nilsson⁶⁴, F. Noferini¹⁶, G. Nooren^{77,78}, N. Novitzky^{35,36}, A. Nyanin¹³, A. Nyatha¹⁰⁶, C. Nygaard⁴⁹, J. Nystrand¹, H. Obayashi¹¹⁵, A. Ochirov²², H. Oeschler¹⁰⁷, S.K. Oh¹¹, J. Oleniacz⁹⁹, C. Oppedisano¹⁴, A. Ortiz Velasquez⁸⁶, G. Ortona^{38,14}, A. Oskarsson⁷⁹, P. Ostrowski⁹⁹, I. Otterlund⁷⁹, J. Otwinowski²¹, G. Øvrebek¹, K. Oyama⁶⁹, K. Ozawa¹⁰³, Y. Pachmayer⁶⁹, M. Pachr⁵⁵, F. Padilla^{38,14}, P. Pagano^{91,92}, G. Paic⁸⁶, F. Painke¹⁸, C. Pajares³³, S. Pal⁴⁰, S.K. Pal⁹, A. Palaha⁴⁵, A. Palmeri³⁷, G.S. Pappalardo³⁷, W.J. Park²¹, V. Paticchio²⁰, A. Pavlinov⁵¹, T. Pawlak⁹⁹, T. Peitzmann^{77,78}, D. Peresunko¹³, C.E. Pérez Lara⁵⁶, D. Perini⁶, D. Perrino^{19,20}, W. Peryt⁹⁹, A. Pesci¹⁶, V. Peskov⁶, Y. Pestov¹¹⁷, A.J. Peters⁶, V. Petráček⁵⁵, M. Petris²³, P. Petrov⁴⁵, M. Petrovici²³, C. Petta^{44,37}, S. Piano⁶⁶, A. Piccotti¹⁴, M. Pikna⁶⁷, P. Pillot²⁹, O. Pinazza⁶, L. Pinsky⁵⁸, N. Pitz³⁰, F. Piuze⁶, D.B. Piyarathna^{51,ac}, R. Platt⁴⁵, M. Płoskoń¹⁰⁵, J. Pluta⁹⁹, T. Pocheptsov^{48,ad}, S. Pochybova⁷, P.L.M. Podesta-Lerma¹⁰⁰, M.G. Poghosyan^{38,14}, K. Polák¹¹², B. Polichtchouk⁶⁰, A. Pop²³, V. Pospíšil⁵⁵, B. Potukuchi⁵³, S.K. Prasad^{51,ae}, R. Preghenella³⁹, F. Prino¹⁴, C.A. Pruneau⁵¹, I. Pshenichnov⁹⁷, G. Puddu^{93,81}, A. Pulvirenti^{44,37}, V. Punin⁶⁸, M. Putiš⁶¹, J. Putschke⁴, E. Quercigh⁶, H. Qvigstad⁶⁴, A. Rachevski⁶⁶, A. Rademakers⁶, O. Rademakers⁶, S. Radomski⁶⁹, T.S. Rähä^{35,36}, J. Rak^{35,36}, A. Rakotozafindrabe⁴⁰, L. Ramello^{84,85}, A. Ramírez Reyes⁸², M. Rammler⁴⁷, R. Raniwala¹¹⁸, S. Raniwala¹¹⁸, S.S. Räsänen^{35,36}, K.F. Read¹¹³, J.S. Real^{31,32}, K. Redlich⁹⁴, R. Renfordt³⁰, A.R. Reolon⁵⁴, A. Reshetin⁹⁷, F. Rettig¹⁸, J.-P. Revol⁶,

K. Reygers⁶⁹, H. Ricaud¹⁰⁷, L. Riccati¹⁴, R.A. Ricci⁸⁷, M. Richter^{1,ae}, P. Riedler⁶, W. Riegler⁶, F. Riggi^{44,37}, A. Rivetti¹⁴, M. Rodríguez Cahuantzi⁸³, D. Rohr¹⁸, D. Röhrich¹, R. Romita²¹, F. Ronchetti⁵⁴, P. Rosinsky⁶, P. Rosnet^{41,42}, S. Rossegger⁶, A. Rossi^{27,28}, F. Roukoutakis⁹⁸, S. Rousseau⁶³, C. Roy^{29,m}, P. Roy⁶², A.J. Rubio Montero⁵⁷, R. Rui^{65,66}, I. Rusanov⁶, E. Ryabinkin¹³, A. Rybicki⁴⁶, S. Sadovsky⁶⁰, K. Šafařík⁶, R. Sahoo^{27,28}, P.K. Sahu⁸⁹, P. Saiz⁶, S. Sakai¹⁰⁵, D. Sakata⁸⁰, C.A. Salgado³³, T. Samanta⁹, S. Sambyal⁵³, V. Samsonov⁵², L. Šándor⁴³, A. Sandoval⁸, M. Sano⁸⁰, S. Sano¹⁰³, R. Santo⁴⁷, R. Santoro²⁰, J. Sarkamo^{35,36}, P. Saturnini^{41,42}, E. Scapparone¹⁶, F. Scarlassara^{27,28}, R.P. Scharenberg¹¹⁹, C. Schiaua²³, R. Schicker⁶⁹, C. Schmidt²¹, H.R. Schmidt²¹, S. Schreiner⁶, S. Schuchmann³⁰, J. Schukraft⁶, Y. Schutz^{29,r}, K. Schwarz²¹, K. Schweda⁶⁹, G. Scioli^{15,16}, E. Scomparin¹⁴, P.A. Scott⁴⁵, R. Scott¹¹³, G. Segato^{27,28}, S. Senyukov^{84,85}, J. Seo¹¹, S. Serici^{93,81}, E. Serradilla⁵⁷, A. Sevcenco⁸⁸, G. Shabratova⁴⁸, R. Shahoyan⁶, N. Sharma⁵, S. Sharma⁵³, K. Shigaki¹¹⁵, M. Shimomura⁸⁰, K. Shtejer², Y. Sibirjak¹³, M. Siciliano^{38,14}, E. Sicking⁶, T. Siemiarzczuk⁹⁴, A. Silenzi^{15,16}, D. Silvermyr³⁴, G. Simonetti^{6,af}, R. Singaraju⁹, R. Singh⁵³, B.C. Sinha⁹, T. Sinha⁶², B. Sitar⁶⁷, M. Sitta^{84,85}, T.B. Skaali⁶⁴, K. Skjerdal¹, R. Smakal⁵⁵, N. Smirnov⁴, R. Snellings^{56,ag}, C. Sogaard⁴⁹, A. Soloviev⁶⁰, R. Soltz¹¹⁶, H. Son¹⁰⁴, M. Song¹⁰⁸, C. Soos⁶, F. Soramel^{27,28}, M. Spyropoulou-Stassinaki⁹⁸, B.K. Srivastava¹¹⁹, J. Stachel⁶⁹, I. Stan⁸⁸, G. Stefanek⁹⁴, G. Stefanini⁶, T. Steinbeck^{24,e}, E. Stenlund⁷⁹, G. Steyn⁷⁰, D. Stocco²⁹, R. Stock³⁰, M. Stolpovskiy⁶⁰, P. Strmen⁶⁷, A.A.P. Suaide⁹⁰, M.A. Subieta Vásquez^{38,14}, T. Sugitate¹¹⁵, C. Suire⁶³, M. Šumbera³, T. Susa²⁶, D. Swoboda⁶, T.J.M. Symons¹⁰⁵, A. Szanto de Toledo⁹⁰, I. Szarka⁶⁷, A. Szostak¹, C. Tagridis⁹⁸, J. Takahashi⁷⁶, J.D. Tapia Takaki⁶³, A. Tauro⁶, M. Tavlet⁶, G. Tejada Muñoz⁸³, A. Telesca⁶, C. Terrevoli^{19,20}, J. Thäder²¹, D. Thomas^{77,78}, J.H. Thomas²¹, R. Tieulent^{74,75}, A.R. Timmins^{51,g}, D. Tlustý⁵⁵, A. Toia⁶, H. Torii¹¹⁵, L. Toscano⁶, F. Tosello¹⁴, T. Traczyk⁹⁹, D. Truesdale²⁵, W.H. Trzaska^{35,36}, A. Tumkin⁶⁸, R. Turrisi²⁷, A.J. Turvey⁷³, T.S. Tveter⁶⁴, J. Ulery³⁰, K. Ullaland¹, A. Uras^{93,81}, J. Urbán⁶¹, G.M. Urciuoli⁹⁵, G.L. Usai^{93,81}, A. Vacchi⁶⁶, M. Vala^{48,y}, L. Valencia Palomo⁶³, S. Vallero⁶⁹, N. van der Kolk⁵⁶, M. van Leeuwen^{77,78}, P. Vande Vyvre⁶, L. Vannucci⁸⁷, A. Vargas⁸³, R. Varma¹⁰⁶, M. Vasileiou⁹⁸, A. Vasiliev¹³, V. Vechernin²², M. Venaruzzo^{65,66}, E. Vercellin^{38,14}, S. Vergara⁸³, R. Vernet¹²⁰, M. Verweij^{77,78}, L. Vickovic¹⁰¹, G. Viesti^{27,28}, O. Vikhlyantsev⁶⁸, Z. Vilakazi⁷⁰, O. Villalobos Baillie⁴⁵, A. Vinogradov¹³, L. Vinogradov²², Y. Vinogradov⁶⁸, T. Virgili^{91,92}, Y.P. Viyogi⁹, A. Vodopyanov⁴⁸, K. Voloshin¹², S. Voloshin⁵¹, G. Volpe^{19,20}, B. von Haller⁶, D. Vranic²¹, J. Vrláková⁶¹, B. Vulpescu^{41,42}, B. Wagner¹, V. Wagner⁵⁵, R. Wan^{50,ah}, D. Wang⁷¹, Y. Wang⁶⁹, Y. Wang⁷¹, K. Watanabe⁸⁰, J.P. Wessels⁴⁷, U. Westerhoff⁴⁷, J. Wiechula⁶⁹, J. Wikne⁶⁴, M. Wilde⁴⁷, A. Wilk⁴⁷, G. Wilk⁹⁴, M.C.S. Williams¹⁶, B. Windelband⁶⁹, H. Yang⁴⁰, S. Yasnopolskiy¹³, J. Yi¹²¹, Z. Yin⁷¹, H. Yokoyama⁸⁰, I.-K. Yoo¹²¹, X. Yuan⁷¹, I. Yushmanov¹³, E. Zabrodin⁶⁴, C. Zampolli⁶, S. Zaporozhets⁴⁸, A. Zarochentsev²², P. Závada¹¹², H. Zbroszczyk⁹⁹, P. Zelnicek²⁴, A. Zenin⁶⁰, I. Zgura⁸⁸, M. Zhalov⁵², X. Zhang^{71,b}, D. Zhou⁷¹, A. Zichichi^{15,16,aj}, G. Zinoviev¹⁷, Y. Zoccarato^{74,75}, M. Zynovyev¹⁷

This publication is dedicated to the memory of our colleague René Kamermans

¹Department of Physics and Technology, University of Bergen, Bergen, Norway

²Centro de Aplicaciones Tecnológicas y Desarrollo Nuclear (CEADEN), Havana, Cuba

³Nuclear Physics Institute, Academy of Sciences of the Czech Republic, Řež u Prahy, Czech Republic

⁴Yale University, New Haven, CT, USA

⁵Physics Department, Panjab University, Chandigarh, India

⁶European Organization for Nuclear Research (CERN), Geneva, Switzerland

⁷KFKI Research Institute for Particle and Nuclear Physics, Hungarian Academy of Sciences, Budapest, Hungary

⁸Instituto de Física, Universidad Nacional Autónoma de México, Mexico City, Mexico

⁹Variable Energy Cyclotron Centre, Kolkata, India

¹⁰Department of Physics, Aligarh Muslim University, Aligarh, India

¹¹Gangneung-Wonju National University, Gangneung, South Korea

¹²Institute for Theoretical and Experimental Physics, Moscow, Russia

¹³Russian Research Centre Kurchatov Institute, Moscow, Russia

¹⁴Sezione INFN, Turin, Italy

¹⁵Dipartimento di Fisica dell'Università, Bologna, Italy

¹⁶Sezione INFN, Bologna, Italy

¹⁷Bogolyubov Institute for Theoretical Physics, Kiev, Ukraine

¹⁸Frankfurt Institute for Advanced Studies, Johann Wolfgang Goethe-Universität Frankfurt, Frankfurt, Germany

¹⁹Dipartimento Interateneo di Fisica 'M. Merlin', Bari, Italy

²⁰Sezione INFN, Bari, Italy

²¹Research Division and ExtreMe Matter Institute EMMI, GSI Helmholtzzentrum für Schwerionenforschung, Darmstadt, Germany

- ²²V. Fock Institute for Physics, St. Petersburg State University, St. Petersburg, Russia
- ²³National Institute for Physics and Nuclear Engineering, Bucharest, Romania
- ²⁴Kirchhoff-Institut für Physik, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany
- ²⁵Department of Physics, Ohio State University, Columbus, OH, USA
- ²⁶Rudjer Bošković Institute, Zagreb, Croatia
- ²⁷Dipartimento di Fisica dell'Università, Padova, Italy
- ²⁸Sezione INFN, Padova, Italy
- ²⁹SUBATECH, Ecole des Mines de Nantes, Université de Nantes, CNRS-IN2P3, Nantes, France
- ³⁰Institut für Kernphysik, Johann Wolfgang Goethe-Universität Frankfurt, Frankfurt, Germany
- ³¹Laboratoire de Physique Subatomique et de Cosmologie (LPSC), Université Joseph Fourier, Grenoble, France
- ³²Université Joseph Fourier, CNRS-IN2P3, Institut Polytechnique de Grenoble, Grenoble, France
- ³³Departamento de Física de Partículas and IGFAE, Universidad de Santiago de Compostela, Santiago de Compostela, Spain
- ³⁴Oak Ridge National Laboratory, Oak Ridge, TN, USA
- ³⁵Helsinki Institute of Physics (HIP), Jyväskylä, Finland
- ³⁶University of Jyväskylä, Jyväskylä, Finland
- ³⁷Sezione INFN, Catania, Italy
- ³⁸Dipartimento di Fisica, Sperimentale dell'Università, Turin, Italy
- ³⁹Centro Fermi—Centro Studi e Ricerche e Museo Storico della Fisica “Enrico Fermi”, Rome, Italy
- ⁴⁰Commissariat à l’Energie Atomique, IRFU, Saclay, France
- ⁴¹Laboratoire de Physique Corpusculaire (LPC), Clermont Université, Clermont-Ferrand, France
- ⁴²CNRS-IN2P3, Université Blaise Pascal, Clermont-Ferrand, France
- ⁴³Institute of Experimental Physics, Slovak Academy of Sciences, Košice, Slovakia
- ⁴⁴Dipartimento di Fisica e Astronomia dell'Università, Catania, Italy
- ⁴⁵School of Physics and Astronomy, University of Birmingham, Birmingham, UK
- ⁴⁶The Henryk Niewodniczanski Institute of Nuclear Physics, Polish Academy of Sciences, Cracow, Poland
- ⁴⁷Institut für Kernphysik, Westfälische Wilhelms-Universität Münster, Münster, Germany
- ⁴⁸Joint Institute for Nuclear Research (JINR), Dubna, Russia
- ⁴⁹Niels Bohr Institute, University of Copenhagen, Copenhagen, Denmark
- ⁵⁰Institut Pluridisciplinaire Hubert Curien (IPHC), Université de Strasbourg, CNRS-IN2P3, Strasbourg, France
- ⁵¹Wayne State University, Detroit, MI, USA
- ⁵²Petersburg Nuclear Physics Institute, Gatchina, Russia
- ⁵³Physics Department, University of Jammu, Jammu, India
- ⁵⁴Laboratori Nazionali di Frascati, INFN, Frascati, Italy
- ⁵⁵Faculty of Nuclear Sciences and Physical Engineering, Czech Technical University in Prague, Prague, Czech Republic
- ⁵⁶Nikhef, National Institute for Subatomic Physics, Amsterdam, The Netherlands
- ⁵⁷Centro de Investigaciones Energéticas Medioambientales y Tecnológicas (CIEMAT), Madrid, Spain
- ⁵⁸University of Houston, Houston, TX, USA
- ⁵⁹Moscow Engineering Physics Institute, Moscow, Russia
- ⁶⁰Institute for High Energy Physics, Protvino, Russia
- ⁶¹Faculty of Science, P.J. Šafárik University, Košice, Slovakia
- ⁶²Saha Institute of Nuclear Physics, Kolkata, India
- ⁶³Institut de Physique Nucléaire d’Orsay (IPNO), Université Paris-Sud, CNRS-IN2P3, Orsay, France
- ⁶⁴Department of Physics, University of Oslo, Oslo, Norway
- ⁶⁵Dipartimento di Fisica dell'Università, Trieste, Italy
- ⁶⁶Sezione INFN, Trieste, Italy
- ⁶⁷Faculty of Mathematics, Physics and Informatics, Comenius University, Bratislava, Slovakia
- ⁶⁸Russian Federal Nuclear Center (VNIIEF), Sarov, Russia
- ⁶⁹Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany
- ⁷⁰Physics Department, iThemba Laboratories, University of Cape Town, Cape Town, South Africa
- ⁷¹Hua-Zhong Normal University, Wuhan, China
- ⁷²Sección Física, Departamento de Ciencias, Pontificia Universidad Católica del Perú, Lima, Peru
- ⁷³Physics Department, Creighton University, Omaha, NE, USA
- ⁷⁴Université de Lyon, Université Lyon 1, Villeurbanne, France
- ⁷⁵CNRS/IN2P3, IPN-Lyon, Villeurbanne, France
- ⁷⁶Universidade Estadual de Campinas (UNICAMP), Campinas, Brazil
- ⁷⁷Nikhef, National Institute for Subatomic Physics, Utrecht, The Netherlands
- ⁷⁸Institute for Subatomic Physics of Utrecht University, Utrecht, The Netherlands
- ⁷⁹Division of Experimental High Energy Physics, University of Lund, Lund, Sweden
- ⁸⁰University of Tsukuba, Tsukuba, Japan
- ⁸¹Sezione INFN, Cagliari, Italy
- ⁸²Centro de Investigación y de Estudios Avanzados (CINVESTAV), Mexico City and Merida, Mexico
- ⁸³Benemérita Universidad Autónoma de Puebla, Puebla, Mexico
- ⁸⁴Dipartimento di Scienze e Technologie Avanzate dell'Università del Piemonte Orientale, Alessandria, Italy
- ⁸⁵Gruppo Collegato INFN, Alessandria, Italy
- ⁸⁶Instituto de Ciencias Nucleares, Universidad Nacional Autónoma de México, Mexico City, Mexico

- ⁸⁷Laboratori Nazionali di Legnaro, INFN, Legnaro, Italy
⁸⁸Institute of Space Sciences (ISS), Bucharest, Romania
⁸⁹Institute of Physics, Bhubaneswar, India
⁹⁰Universidade de São Paulo (USP), São Paulo, Brazil
⁹¹Dipartimento di Fisica ‘E.R. Caianiello’ dell’Università, Salerno, Italy
⁹²Gruppo Collegato INFN, Salerno, Italy
⁹³Dipartimento di Fisica dell’Università, Cagliari, Italy
⁹⁴Soltan Institute for Nuclear Studies, Warsaw, Poland
⁹⁵Sezione INFN, Rome, Italy
⁹⁶Faculty of Engineering, Bergen University College, Bergen, Norway
⁹⁷Institute for Nuclear Research, Academy of Sciences, Moscow, Russia
⁹⁸Physics Department, University of Athens, Athens, Greece
⁹⁹Warsaw University of Technology, Warsaw, Poland
¹⁰⁰Universidad Autónoma de Sinaloa, Culiacan, Mexico
¹⁰¹Technical University of Split FESB, Split, Croatia
¹⁰²Yerevan Physics Institute, Yerevan, Armenia
¹⁰³University of Tokyo, Tokyo, Japan
¹⁰⁴Department of Physics, Sejong University, Seoul, South Korea
¹⁰⁵Lawrence Berkeley National Laboratory, Berkeley, CA, USA
¹⁰⁶Indian Institute of Technology, Mumbai, India
¹⁰⁷Institut für Kernphysik, Technische Universität Darmstadt, Darmstadt, Germany
¹⁰⁸Yonsei University, Seoul, South Korea
¹⁰⁹Zentrum für Technologietransfer und Telekommunikation (ZTT), Fachhochschule Worms, Worms, Germany
¹¹⁰California Polytechnic State University, San Luis Obispo, CA, USA
¹¹¹China Institute of Atomic Energy, Beijing, China
¹¹²Institute of Physics, Academy of Sciences of the Czech Republic, Prague, Czech Republic
¹¹³University of Tennessee, Knoxville, TN, USA
¹¹⁴Dipartimento di Fisica dell’Università ‘La Sapienza’, Rome, Italy
¹¹⁵Hiroshima University, Hiroshima, Japan
¹¹⁶Lawrence Livermore National Laboratory, Livermore, CA, USA
¹¹⁷Budker Institute for Nuclear Physics, Novosibirsk, Russia
¹¹⁸Physics Department, University of Rajasthan, Jaipur, India
¹¹⁹Purdue University, West Lafayette, IN, USA
¹²⁰Centre de Calcul de l’IN2P3, Villeurbanne, France
¹²¹Pusan National University, Pusan, South Korea

Received: 15 December 2010 / Revised: 9 February 2011 / Published online: 30 March 2011

© CERN for the benefit of the ALICE collaboration 2011. This article is published with open access at Springerlink.com

^ae-mail: Paul.Kuijjer@nikhef.nl

^bAlso at Laboratoire de Physique Corpusculaire (LPC), Clermont Université, Université Blaise Pascal, CNRS-IN2P3, Clermont-Ferrand, France

^cNow at Centro Fermi—Centro Studi e Ricerche e Museo Storico della Fisica “Enrico Fermi”, Rome, Italy

^dNow at Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany

^eNow at Frankfurt Institute for Advanced Studies, Johann Wolfgang Goethe-Universität Frankfurt, Frankfurt, Germany

^fNow at Sezione INFN, Turin, Italy

^gNow at University of Houston, Houston, Texas, United States

^hAlso at Dipartimento di Fisica dell’Università, Udine, Italy

ⁱNow at SUBATECH, Ecole des Mines de Nantes, Université de Nantes, CNRS-IN2P3, Nantes, France

^jNow at Centro de Investigación y de Estudios Avanzados (CINVESTAV), Mexico City and Mérida, Mexico

^kNow at Benemérita Universidad Autónoma de Puebla, Puebla, Mexico

^lNow at Laboratoire de Physique Subatomique et de Cosmologie (LPSC), Université Joseph Fourier, CNRS-IN2P3, Institut Polytechnique de Grenoble, Grenoble, France

^mNow at Institut Pluridisciplinaire Hubert Curien (IPHC), Université de Strasbourg, CNRS-IN2P3, Strasbourg, France

ⁿNow at Sezione INFN, Padova, Italy

^oDeceased

^pAlso at Division of Experimental High Energy Physics, University of Lund, Lund, Sweden

^qAlso at University of Technology and Austrian Academy of Sciences, Vienna, Austria

^rAlso at European Organization for Nuclear Research (CERN), Geneva, Switzerland

^sNow at Oak Ridge National Laboratory, Oak Ridge, TN, USA

^tNow at European Organization for Nuclear Research (CERN), Geneva, Switzerland

^uAlso at Wayne State University, Detroit, MI, USA

^vAlso at Frankfurt Institute for Advanced Studies, Johann Wolfgang Goethe-Universität Frankfurt, Frankfurt, Germany

^wNow at Research Division and ExtreMe Matter Institute EMMI, GSI Helmholtzzentrum für Schwerionenforschung, Darmstadt, Germany

^xAlso at Fachhochschule Köln, Köln, Germany

^yAlso at Institute of Experimental Physics, Slovak Academy of Sciences, Košice, Slovakia

Abstract The production of mesons containing strange quarks (K_S^0 , ϕ) and both singly and doubly strange baryons (Λ , $\bar{\Lambda}$, and $\Xi^- + \bar{\Xi}^+$) are measured at mid-rapidity in pp collisions at $\sqrt{s} = 0.9$ TeV with the ALICE experiment at the LHC. The results are obtained from the analysis of about 250 k minimum bias events recorded in 2009. Measurements of yields (dN/dy) and transverse momentum spectra at mid-rapidity for inelastic pp collisions are presented. For mesons, we report yields ($\langle dN/dy \rangle$) of $0.184 \pm 0.002(\text{stat.}) \pm 0.006(\text{syst.})$ for K_S^0 and $0.021 \pm 0.004(\text{stat.}) \pm 0.003(\text{syst.})$ for ϕ . For baryons, we find $\langle dN/dy \rangle = 0.048 \pm 0.001(\text{stat.}) \pm 0.004(\text{syst.})$ for Λ , $0.047 \pm 0.002(\text{stat.}) \pm 0.005(\text{syst.})$ for $\bar{\Lambda}$ and $0.0101 \pm 0.0020(\text{stat.}) \pm 0.0009(\text{syst.})$ for $\Xi^- + \bar{\Xi}^+$. The results are also compared with predictions for identified particle spectra from QCD-inspired models and provide a baseline for comparisons with both future pp measurements at higher energies and heavy-ion collisions.

1 Introduction

The production of hadrons at high transverse momenta in high energy proton–proton collisions is reasonably well described by perturbative Quantum Chromodynamics (pQCD) in terms of hard parton–parton scattering (large momentum transfers) followed by fragmentation [1, 2]. However, the low-momentum region, where most particles are produced and which therefore contributes most to the underlying event, is dominated by soft interactions. In the soft regime, it has been found that particle production can be described effectively by models based on emission from an equilibrated system at a specific temperature and baryochemical potential, with additional accounting of conserved quantities [3–5]. It can also be treated in the framework of

QCD inspired phenomenological models, that include multi-parton processes, extrapolated to very low-momentum transfers [6]. The contribution and evolution of multi-parton processes as a function of \sqrt{s} is difficult to establish. Measurements of identified particles at the beam injection energy of the LHC and in the low transverse momentum (p_T) region, along with their comparison with QCD-inspired models, constitute a baseline for comparisons with higher centre-of-mass energies. The low p_T cutoff achievable through the low material budget, low central barrel magnetic field (0.5 T) and excellent particle identification (PID) of the ALICE detectors, allows an accurate measurement of the low momentum region at mid-rapidity.

The differential transverse momentum yields (p_T spectra) and integrated yields at mid-rapidity of K_S^0 , ϕ , Λ , $\bar{\Lambda}$ and $\Xi^- + \bar{\Xi}^+$ have been measured by the ALICE experiment during the commissioning phase of the LHC (December 2009) [7] with the very first proton–proton collisions [8] and are reported in this article. A sample of 250 k minimum bias pp collisions at $\sqrt{s} = 0.9$ TeV has been selected with triggers combining several fast detectors [9]. Measurements are performed using the tracking devices and the main PID detectors of ALICE in the rapidity region $|y| < 0.8$. A comparison of the transverse momentum shapes (mass dependence and mean transverse momentum) with PYTHIA [1] and PHOJET [2] is provided.

This article is organized as follows. Section 2 presents the experimental conditions, the minimum bias event selection as well as a brief description of the main detectors and the associated event reconstruction tools used for the analysis. Section 3 is dedicated to the data analysis, including track and topological selections, signal extraction methods and the corresponding efficiency corrections. The determination of the systematic uncertainties are also described in this section. In Sect. 4, the p_T spectra and the integrated yields of the studied particle species are given and compared with previous measurements and model predictions. Conclusions are given in Sect. 5.

2 Experimental set-up and data collection

A detailed description of the ALICE experimental setup and its detector subsystems can be found in [10].

2.1 Main detectors and reconstruction techniques used for the analyses

The central barrel of ALICE covers polar angles from 45° to 135° over the full azimuth. It is embedded in the large L3 solenoidal magnet, providing a nominal magnetic field B of 0.5 T. Within the barrel, the two tracking detectors used in these present analyses consist of an Inner Tracking System (ITS), composed of 6 cylindrical layers of high-resolution

^zNow at Instituto de Ciencias Nucleares, Universidad Nacional Autónoma de México, Mexico City, Mexico

^{aa}Also at Laboratoire de Physique Subatomique et de Cosmologie (LPSC), Université Joseph Fourier, CNRS-IN2P3, Institut Polytechnique de Grenoble, Grenoble, France

^{ab}Also at “Vinča” Institute of Nuclear Sciences, Belgrade, Serbia

^{ac}Also at University of Houston, Houston, TX, USA

^{ad}Also at Department of Physics, University of Oslo, Oslo, Norway

^{ae}Also at Variable Energy Cyclotron Centre, Kolkata, India

^{af}Now at Department of Physics, University of Oslo, Oslo, Norway

^{ag}Also at Dipartimento Interateneo di Fisica ‘M. Merlin’ and Sezione INFN, Bari, Italy

^{ah}Now at Nikhef, National Institute for Subatomic Physics and Institute for Subatomic Physics of Utrecht University, Utrecht, The Netherlands

^{ai}Also at Hua-Zhong Normal University, Wuhan, China

^{aj}Also at Centro Fermi—Centro Studi e Ricerche e Museo Storico della Fisica “Enrico Fermi”, Rome, Italy

Silicon detectors and a cylindrical Time Projection Chamber (TPC). PID is performed using secondary (displaced) vertex reconstruction, invariant mass analysis and single track PID methods, which include the measurement of specific ionization in the ITS and the TPC, and the information from the Time-Of-Flight detector (TOF).

2.1.1 The inner tracking system

The Silicon Pixel Detector (SPD) corresponds to the two innermost ITS layers. These two layers have a very high granularity, with a total of about 9.8 million pixels, each with a size of $50 \times 425 \mu\text{m}^2$. They are located at radii of 3.9 and 7.6 cm and the pseudo-rapidity coverages are $|\eta| < 2.0$ and $|\eta| < 1.4$ respectively. The detector provides a position resolution of $12 \mu\text{m}$ in the $r\phi$ direction and about $100 \mu\text{m}$ in the direction along the beam axis. It can also deliver a signal for the first level of trigger (L0) in less than 850 ns. The two layers of the Silicon Drift Detector (SDD), located at radii of 15.0 and 23.9 cm, are composed of 260 sensors, including 133 000 collection anodes with a pitch of $294 \mu\text{m}$. They provide a charge deposit measurement and a position measurement with a resolution of about $35 \mu\text{m}$ in the $r\phi$ direction and about $25 \mu\text{m}$ in the beam direction [11]. The Silicon Strip Detector (SSD) consists of 1698 double-sided sensors (with a strip pitch of $95 \mu\text{m}$ and a stereo angle of 35 mrad) arranged in 2 layers located at radii of 38 and 43 cm. It provides a measurement of the charge deposited in each of the 2.6 million strips, as well as a position measurement with a resolution of $20 \mu\text{m}$ in the $r\phi$ direction and about $800 \mu\text{m}$ in the beam direction.

The ITS sensor modules were aligned using survey information and tracks from cosmic-ray muons and pp collisions. The corresponding methods are described in [11].

The percentage of operational channels in the ITS during the 2009 run is 82% for the SPD, 91% for the SDD and 90% for the SSD. This information is used in the Monte Carlo simulations, in the reconstruction of both Monte Carlo and real data, and taken into account in the efficiency corrections.

2.1.2 The time projection chamber

The ALICE TPC is a cylindrical drift detector with a pseudo-rapidity coverage of $|\eta| \leq 0.9$ [12]. It has a field cage filled with 90 m^3 of $\text{Ne}/\text{CO}_2/\text{N}_2$ (85.7/9.5/4.8%). The inner and outer radii of the active volume are of 85 cm and 247 cm respectively and the length along the beam direction is 500 cm. Inside the field cage, ionization electrons produced when charged particles traverse the active volume on either side of the central electrode (a high voltage membrane at -100 kV) migrate to the end plates in less than $94 \mu\text{s}$. A total of 72 multi-wire proportional chambers, with cathode pad readout, instrument the two end plates of the TPC

which are segmented in 18 sectors and amount to a total of 557 568 readout pads. The ALICE TPC ReadOut (ALTRO) chip, employing a 10 bit ADC at 10 MHz sampling rate and digital filtering circuits, allows for precise position and linear energy loss measurements with a gas gain of the order of 10^4 .

The position resolution in the $r\phi$ direction varies from $1100 \mu\text{m}$ to $800 \mu\text{m}$ when going from the inner to the outer radius whereas the resolution along the beam axis ranges between $1250 \mu\text{m}$ and $1100 \mu\text{m}$.

2.1.3 The time-of-flight detector

The ALICE Time-Of-Flight detector [13] is a cylindrical assembly of Multi-gap Resistive Plate Chambers (MRPC) with an inner radius of 370 cm and an outer radius of 399 cm, a pseudo-rapidity range $|\eta| < 0.9$ and full azimuth angle, except for the region $260^\circ < \phi < 320^\circ$ at η near zero, where no TOF modules were installed to reduce the material in front of the Photon Spectrometer. The basic unit of the TOF system is a 10-gap double-stack MRPC strip 122 cm long and 13 cm wide, with an active area of $120 \times 7.4 \text{ cm}^2$ subdivided into two rows of 48 pads of $3.5 \times 2.5 \text{ cm}^2$. Five modules of three different types are needed to cover the full cylinder along the z direction. All modules have the same structure and width (128 cm) but differ in length. The overall TOF barrel length is 741 cm (active region). It has 152 928 readout channels and an average thickness of 25–30% of a radiation length, depending on the detector zone. For pp collisions, such a segmentation leads to an occupancy smaller than 0.02%. Its front-end electronics is designed to comply with the basic characteristics of a MRPC detector, i.e. very fast differential signals from the anode and cathode readout. Test beam results demonstrated a time resolution below 50 ps, dominated by the jitter in the electronic readout.

2.1.4 The VZERO counters

The VZERO counters are two scintillator hodoscopes located along the beam direction at -0.9 m and 3.3 m from the geometrical centre of the experiment. They correspond to a coverage of $-3.7 < \eta < -1.7$ and $2.8 < \eta < 5.1$ respectively and have a time resolution close to 0.5 ns. They are used as trigger detectors and help to remove beam-gas interaction background.

2.1.5 Track reconstruction and particle identification

The global tracking system in the ALICE central barrel (combining the ITS and the TPC) covers the pseudo-rapidity window $|\eta| < 0.9$.

The reconstruction in the tracking detectors begins with charge cluster finding. The two coordinates of the crossing

points (space points) between tracks and detector sensitive elements (pad rows in the TPC, and silicon sensors in the ITS) are calculated as the centres of gravity of the clusters. The errors on the space point positions are parametrized as functions of the cluster size and of the deposited charge. In the TPC, these errors are further corrected during the tracking, using the crossing angles between tracks and the pad rows.

The space points reconstructed at the two innermost ITS layers (pixel detector, SPD) are then used for the reconstruction of the primary vertex. One space point from the first SPD layer and one from the second layer are combined into pairs called “tracklets”. The primary vertex is consequently reconstructed in 3D as the location that minimizes the sum of the squared distances to all the tracklet extrapolations. If this fails, the algorithm instead reconstructs the z coordinate of the vertex by correlating the z coordinates of the SPD space points, while for x and y the average position of the beam in the transverse plane (measured by a dedicated calibration procedure on a run-by-run basis) is assumed.

Track reconstruction in ALICE is based on the Kalman filter approach and is discussed in detail in [14]. The initial approximations for the track parameters (the “seeds”) are constructed using pairs of space points taken at two outer TPC pad rows separated by a few pad rows and the primary vertex. The a priori uncertainty of the primary vertex position used in the seeding procedure is considered to be equal to the radius of the beam pipe (3 cm) to limit the bias for tracks coming from particles decaying inside this volume. The seeds for the secondary tracks are created without using the primary vertex, since such a constraint would unnecessarily reduce the V0 finding efficiency. The additional space points used as a seed are then searched along the straight line segment connecting the pairs of points taken at those two outer TPC pad rows.

Once the track seeds are created, they are sorted according to the estimate of their transverse momentum (p_T). Then they are extended from one pad row to another in the TPC and from one layer to another in the ITS towards the primary vertex. Every time a space point is found within a prolongation path defined by the current estimate of the covariance matrix, the track parameters and the covariance matrix are updated using the Kalman filter. For each tracking step, the estimates of the track parameters and the covariance matrix are also corrected for the mean energy loss and Coulomb multiple scattering in the traversed material. The decision on the particle mass to be used for these corrections is based on the dE/dx information given by the TPC when available. If the information is missing or not conclusive, a pion mass is assumed. Only five particle hypotheses are considered: electrons, muons, pions, kaons and protons.

All the tracks are then propagated outwards, through the ITS and the TPC. When possible, they are matched with the

hits reconstructed in the TOF detector. During this tracking phase, the track length and five time-of-flight hypotheses per track (corresponding to the electron, muon, pion, kaon and proton masses) are calculated. This information is later used for the TOF PID procedure. The track parameters are then re-estimated at the distance of closest approach (DCA) to the primary vertex applying the Kalman filter to the space points already attached. Finally, the primary vertex is fitted once again, now using reconstructed tracks and the information about the average position and spread of the beam-beam interaction region estimated for this run.

In pp collisions, the track reconstruction efficiency in the acceptance of TPC saturates at about 90% because of the effect of the dead zones between its sectors. It goes down to about 75% around $p_T = 1$ GeV/ c and drops to 45% at 0.15 GeV/ c . It is limited by particle decays (for kaons), track bending at low p_T and absorption in the detector material. The amount of material traversed by particles near $\eta = 0$ is about 11% of a radiation length including the beam pipe, the ITS and the TPC (with services and support).

The overall p_T resolution is at least as good as the TPC-standalone resolution, which is typically 1% for momenta of 1 GeV/ c and 7% for momenta of 10 GeV/ c , and follows the parameterization:

$$(\sigma(p_T)/p_T)^2 = (0.01)^2 + (0.007p_T)^2$$

where p_T is expressed in GeV/ c (see [15] for the details).

The resolution of the track transverse impact parameter (the minimal distance between a track and the primary vertex in the transverse plane) depends on the precision of track and primary vertex reconstruction. These in turn depend on the momentum, and, in the case of the vertex, on the number of contributing tracks. As it was estimated from the data, the transverse impact parameter resolution in a typical pp event could be parameterized as $\sigma(p_T) = 50 + 60/(p_T)^{0.9}$ (σ is in μm , and p_T is in GeV/ c), which was defined by the level of the ITS alignment achieved in 2009.

The dE/dx resolution of the TPC is estimated to be about 5% for tracks with 159 clusters [12], which is better than the design value [14]. When averaged over all reconstructed tracks, this resolution is about 6.5%.

During the run, the preliminary calibration of the TOF detector corresponds to a resolution of 180 ps, which includes 140 ps due to the jitter in the absolute time of the collisions. This contribution is reduced to about 85 ps for those events with at least 3 tracks reaching the TOF, in which case an independent time zero determination is possible. The matching efficiency with TPC tracks (which includes geometry, decays and interaction with material) is on average 60% for protons and pions and reaches 65% above $p_T = 1$ GeV/ c . For kaons it remains slightly lower [16]. Above $p_T = 0.5$ GeV/ c , the TOF PID has an efficiency larger than 60% with a very small contamination.

2.2 LHC running conditions and triggers

For the first collisions provided by the Large Hadron Collider, four low intensity proton bunches (10^9 protons per bunch, giving the luminosity of the order of $10^{26} \text{ cm}^{-2} \text{ s}^{-1}$) per beam were circulated, and two pairs of them crossed at the ALICE interaction point. Under such conditions, the rate for multiple events in a given bunch-crossing (“pile-up”) was negligible. The energy in the centre of mass corresponded to twice the beam injection energy, that is $\sqrt{s} = 0.9 \text{ TeV}$. The data acquisition of ALICE was triggered by requiring two coincidence conditions: (i) the LHC bunch-crossing signal together with the two beam pick-up monitors (BPTX); (ii) ALICE minimum bias (MB) trigger requiring a combination of signals from the SPD and from the VZERO counters. For these analyses, the MB_{OR} was used, which is fulfilled when at least one of the VZEROs or the SPD trigger is fired [9]. The corresponding data rate was $\sim 10 \text{ Hz}$.

3 Data analysis

3.1 Event and track selection

The primary vertex is reconstructed using either SPD tracklets [8] (5% of the events) or global tracks (95% of the events). Events are selected by requiring that the distance between the position of primary vertex and the geometrical centre of the apparatus along the beam axis be less than 10 cm ($\bar{z} = -0.40 \text{ cm}$ and $\text{rms}_z = 4.24 \text{ cm}$, where \bar{z} is the average position of the primary vertex along the beam axis). Events with less centred primary vertices ($|z| > 10 \text{ cm}$) are

discarded in order to minimize acceptance and efficiency biases for tracks at the edge of the TPC detection volume. The average position and dispersion for both horizontal and vertical directions are found to be $\bar{x} = -0.35 \text{ mm}$ ($\bar{y} = +1.63 \text{ mm}$) and $\text{rms}_x = 0.23 \text{ mm}$ ($\text{rms}_y = 0.27 \text{ mm}$). No conditions were applied on the x and y position of the vertex. The total number of events used for obtaining the particle spectra and yields is about 250 k events. Figure 1 shows the primary vertex distribution along the beam axis (left panel) and for the x and y directions (right panel). The dashed lines indicate the limits of the selected vertex region.

The normalization to the number of inelastic events (INEL) is obtained in the same way as other ALICE analyses [8, 16]. It leads to a correction for the normalization of $\sim 5\%$ with an uncertainty of 2%. This uncertainty is added to the ones described in Sect. 3.5 and mainly related to the modeling of the fraction of diffractive events with several Monte Carlo event generators.

Several quality criteria are defined for track selection. Each track is required to have been reconstructed in the TPC in the initial outward-in step of tracking and then successfully refitted in the final back-propagation to the primary vertex as described in Sect. 2.1.5. It is also required that each track has at least 80 TPC clusters out of a maximum of 159. At the reconstruction level, split tracks are rejected as well as those which may correspond to daughters of kaons decaying in the TPC.

As the ϕ particle is a strongly decaying resonance, its daughters are indistinguishable from primary particles at the reconstruction level and therefore primary track selections are used. As a first step, each track is propagated to the reconstructed primary vertex. If this operation is successful,

Fig. 1 Primary vertex distributions for the analysed events. The *left panel* shows the distributions along the beam axis. Selected events (*full symbols*) are required to have a reconstructed primary vertex with $|z| < 10 \text{ cm}$. The *right panel* corresponds to the directions perpendicular to the beam axis: *horizontally* (i.e. x -direction, *squares* and *full line*) and *vertically* (i.e. y -direction, *triangles* and *dashed line*)

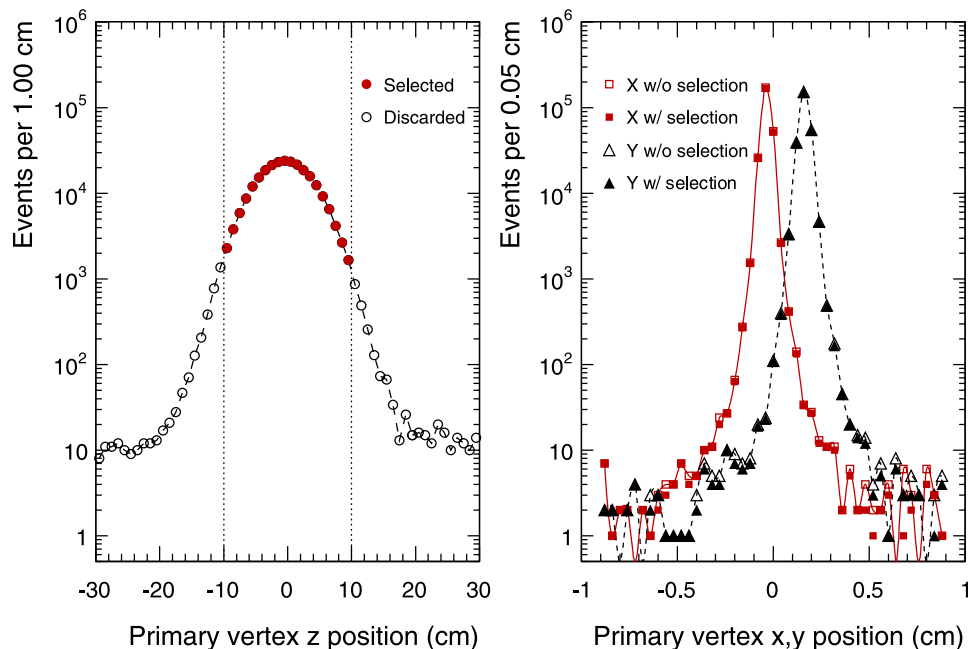


Table 1 Track selection criteria

Common selections	
Detectors required for track rec./fit	ITS, TPC
Number of TPC clusters ^a	>80
$N(\sigma)$ dE/dx (TPC PID)	3 to 5
Primary track selections	
χ^2 per cluster	<4
DCA to primary vertex (r, z)	<(0.5, 3.0) cm
Number of SPD clusters ^b	≥ 1
Secondary track selections	
Transverse momentum ^c	>160 MeV/c

^aMaximum number for the TPC is 159;

^bMaximum number for the SPD is 2;

^cIn the cases of K_S^0 , Λ and $\bar{\Lambda}$

the track is kept if it has a DCA smaller than 5 mm (3 cm) in the transverse (longitudinal) direction with the additional constraints of having at least one SPD cluster and a χ^2 of less than 4 per cluster assignment (for each cluster, the χ^2 has two degrees of freedom).

Depending on its lifetime, a particle may cross several layers of the ITS before weakly decaying. The probability that the daughter tracks of K_S^0 , Λ , $\bar{\Lambda}$ and $\Xi^- + \bar{\Xi}^+$ have a hit in this detector decreases accordingly. Therefore, no specific condition on the number of ITS hits is required for the daughter tracks of the reconstructed secondary vertices. However, other quality criteria are applied for selecting the daughter tracks of weakly decaying particles which are not considered as primaries. These selection criteria are based on the ones defined in [14] and optimized for efficiency rather than purity, owing to the modest statistics. They are summarized in Table 1.

The measurement of differential yields in rapidity and p_T bins cannot be performed simultaneously for the particles considered due to the small available statistics. Therefore the rapidity ranges are chosen such that (i) the efficiency does not vary strongly for each species and (ii) the rapidity distribution is sufficiently flat for it to be possible to rely on the Monte Carlo to obtain the corrections.

3.2 Particle reconstruction and identification methods

3.2.1 Topological reconstruction of K_S^0 , Λ , $\bar{\Lambda}$ and $\Xi^- + \bar{\Xi}^+$

The K_S^0 , Λ , $\bar{\Lambda}$ and $\Xi^- + \bar{\Xi}^+$ are identified by applying selections on the characteristics of their daughter tracks (see Table 2) and using their weak decay topologies in the channels listed in Table 3.

The measurement of K_S^0 , Λ and $\bar{\Lambda}$ is based on the reconstruction of the secondary vertex (V0) associated to their

Table 2 Secondary vertex selection criteria

Common selections	
Minimum transverse decay radius	= 0.2 cm
Maximum transverse decay radius	= 100 cm
V0 vertex selections (K_S^0 , Λ and $\bar{\Lambda}$)	
DCA of V0 daughter track to primary vertex	>0.05 cm
DCA between V0 daughter tracks	<0.50 cm
Cosine of V0 pointing angle (Λ and $\bar{\Lambda}$)	>0.99
Cascade vertex selections	
DCA of cascade daughter track to primary vertex ^a	>0.01 cm
DCA between V0 daughter tracks	<3.0 cm
Cosine of V0 pointing angle	>0.97
DCA of V0 to primary vertex	>0.001 cm
V0 invariant mass	>1110 MeV/c ²
V0 invariant mass	<1122 MeV/c ²
DCA between V0 and bachelor track	<3.0 cm
Cosine of cascade pointing angle	>0.85

^aFor bachelor and each V0 daughter

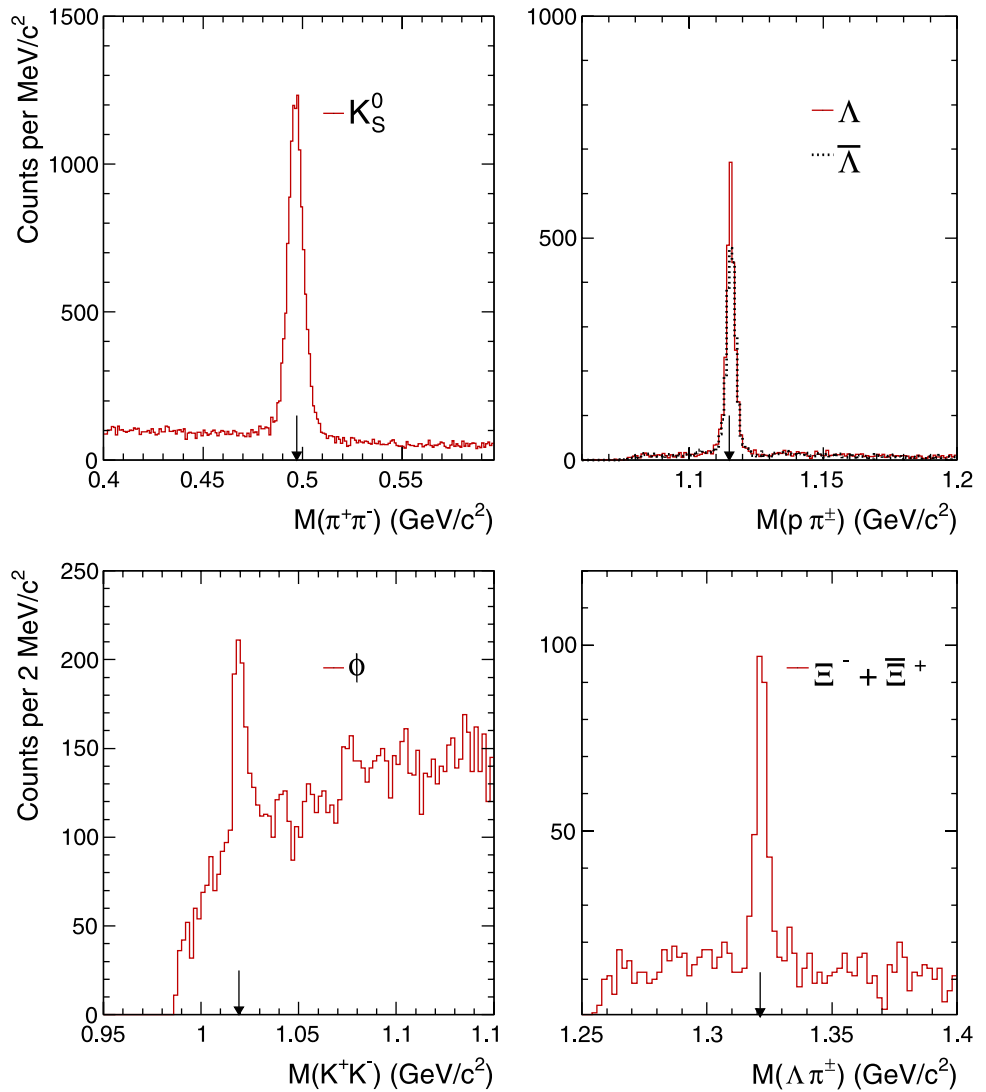
weak decay. The V0 finding procedure starts with the selection of secondary tracks, i.e. tracks having a sufficiently large impact parameter with respect to the primary vertex. All possible combinations between two secondary tracks of opposite charge are then examined. They are accepted as V0 candidates only if the DCA between them is smaller than 0.5 cm. The minimization of the distance between the tracks is performed numerically using helix parametrizations in 3D. The V0 vertex position is a point on the line connecting the points of closest approach between the two tracks. Its distance from each daughter track is taken to be proportional to the precision of the track parameter estimations. Once their position is determined, only the V0 candidates located inside a given fiducial volume are kept. The inner boundary of this fiducial volume is at a radius of 0.2 cm from the primary vertex, while the outer limit is set at 100 cm. Finally, for Λ and $\bar{\Lambda}$ reconstruction, the V0 finding procedure checks whether the particle momentum (\mathbf{p}) associated with the V0 candidate (calculated as the sum of the track momenta extrapolated to the position of the DCA) points back to the primary vertex. This is achieved by applying a cut on the cosine of the angle (pointing angle θ_p) between \mathbf{p} and a vector connecting the primary vertex and the V0 position ($\cos\theta_p > 0.99$). The invariant mass of each candidate can then be calculated either under the K_S^0 or the Λ hypothesis.

The TPC PID helps substantially to remove the combinatorial background for the Λ and $\bar{\Lambda}$ (mainly for the baryon daughter identification, while it is not needed for the K_S^0 decaying into pions). TPC PID is described in Sect. 3.2.3.

Table 3 Main characteristics of the reconstructed particles: valence quark content, mass, $c\tau$ and charged decay branching ratio (B.R.) [18]

	Particles	Mass (MeV/c ²)	$c\tau$	Charged decay	B.R. (%)
Mesons	K_S^0	497.61	2.68 cm	$K_S^0 \rightarrow \pi^+ + \pi^-$	69.2
	$\phi (s\bar{s})$	1019.46	45 fm	$\phi \rightarrow K^+ + K^-$	49.2
Baryons	$\Lambda (uds)$ and $\bar{\Lambda} (\bar{u}\bar{d}\bar{s})$	1115.68	7.89 cm	$\Lambda \rightarrow p + \pi^-$ and $\bar{\Lambda} \rightarrow \bar{p} + \pi^+$	63.9
	$\Xi^- (dss)$ and $\bar{\Xi}^+ (\bar{d}s\bar{s})$	1321.71	4.91 cm	$\Xi^- \rightarrow \Lambda + \pi^-$ and $\bar{\Xi}^+ \rightarrow \bar{\Lambda} + \pi^+$	99.9

Fig. 2 Invariant mass distributions of K_S^0 , Λ and $\bar{\Lambda}$, ϕ and the sum $\Xi^- + \bar{\Xi}^+$. The vertical arrows indicate the nominal mass values from PDG



The selections here concern the proton daughter only and have been chosen to be looser for the daughter track with momentum below 0.7 GeV/c ($\pm 5\sigma$) and tighter for higher momentum ($\pm 3\sigma$).

The $\Xi^- + \bar{\Xi}^+$ particles are identified via their “cascade” decay topology. The cascade finding procedure starts from the V0 finding procedure for the Λ daughter but with less stringent selection criteria (see Table 2 and Cascade vertex selections). This is done to increase the efficiency and to

allow for the fact that the daughter Λ 's do not have to point back to the primary vertex.

The V0 candidates found within the Λ mass window (1116 ± 6 MeV/c²) are combined with all possible secondary tracks (bachelor candidates) with the exception of both V0 daughter tracks. A cut on the impact parameter of the bachelor track is applied to reject the primary particles which increase the combinatorial background.

A V0-bachelor association is performed if the distance of closest approach between the bachelor track and the V0 tra-

jectory (DCA between V0 and bachelor track) is small (less than 3 cm). Finally, this cascade candidate is selected if its reconstructed momentum points back to the primary vertex (cosine of cascade pointing angle). The cascade finding is limited to the fiducial region used for V0 reconstruction (see Table 2).

In addition to topological selections, the reconstruction of cascades also makes use of the single-track PID information delivered by the TPC. This is considered for each of the three daughters (both pions and the proton). For each track, a loose selection is required ($\pm 4\sigma$ over the whole momentum range) to reject the combinatorial background in part. The resulting invariant mass distributions are presented in Fig. 2.

3.2.2 Additional quality checks for K_S^0 , Λ , $\bar{\Lambda}$

A significant fraction of the reconstructed V0 come from γ conversion in the detector material. This can be clearly seen in the Armenteros-Podolanski distribution [17] shown in Fig. 3 where p_L^+ and p_L^- are the longitudinal components of the total momentum for the positive and negative daughters respectively, relative to the direction of the V0 momentum vector. The K_S^0 , Λ and $\bar{\Lambda}$ signal regions are symmetric and clearly distinguishable.

The lifetime (τ) distributions for K_S^0 , Λ and $\bar{\Lambda}$ are also checked. All V0 candidates within a $\pm 3\sigma$ effective mass region around the nominal value are used in the distribution without further residual background subtraction. The corresponding distributions of $c\tau = L \frac{m}{p}$ are obtained, where L is defined as the distance between primary and V0 vertices, and m and p are the particle mass and momentum. Because of the acceptance, the single track efficiency and the topological selections applied at reconstruction level, the

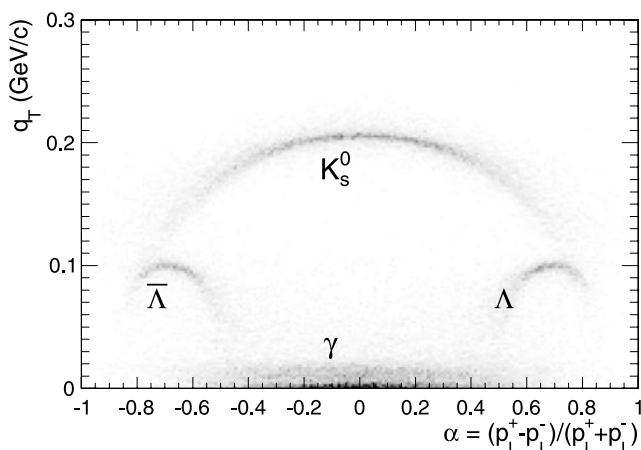


Fig. 3 Armenteros–Podolanski distribution for V0 candidates showing a clear separation between K_S^0 , Λ and $\bar{\Lambda}$. The γ converting to e^+e^- with the detector material are located in the low q_T region, where q_T is the momentum component perpendicular to the parent momentum vector

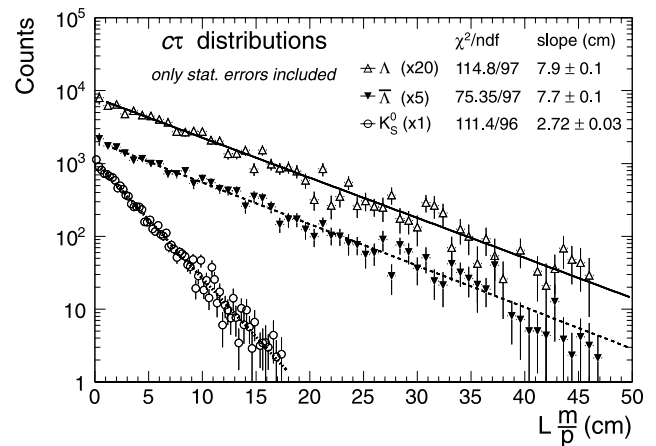


Fig. 4 K_S^0 , Λ and $\bar{\Lambda}$ lifetime distributions obtained for the candidates selected by the invariant mass within a $\pm 3\sigma$ region around the nominal mass and corrected for detection efficiency. The distributions are scaled for visibility and fitted to an exponential distribution (straight lines). Only statistical uncertainties are shown

reconstruction efficiency as a function of the decay length is not constant. The corresponding corrections are extracted from the reconstruction of full Monte Carlo simulations (see Sect. 3.4). The corrected $c\tau$ distributions are fitted using exponential functions. The results are shown with the statistical uncertainties in Fig. 4. The extracted decay lengths of 7.9 ± 0.1 cm, 7.7 ± 0.1 cm and 2.72 ± 0.03 cm for Λ , $\bar{\Lambda}$ and K_S^0 , respectively, are compatible with the PDG values given in Table 3.

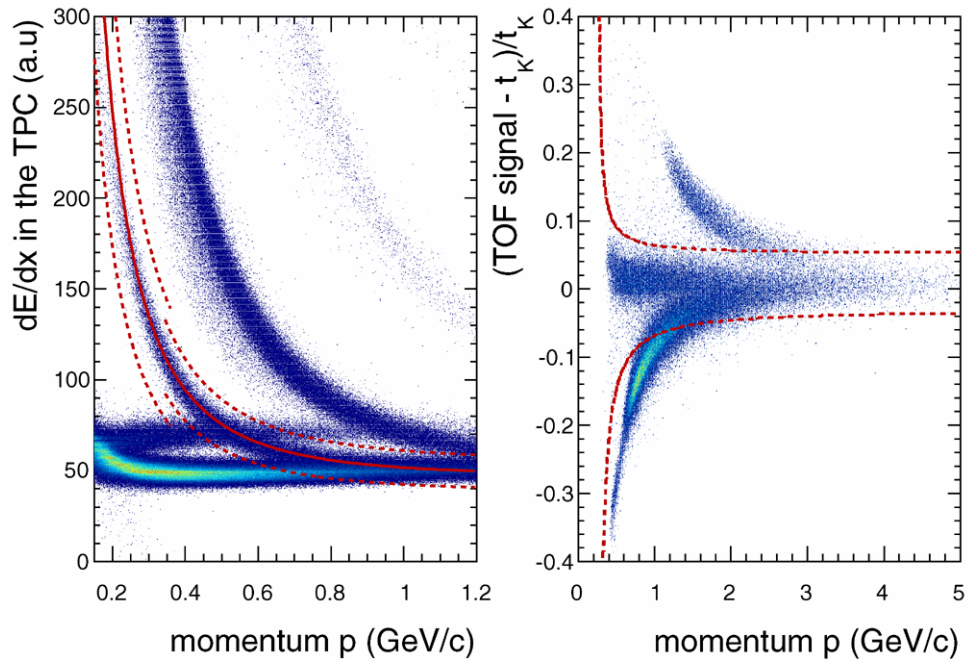
3.2.3 ϕ reconstruction

The ϕ resonance is reconstructed through its principal decay channel $\phi \rightarrow K^+K^-$ (see Table 3). With a $c\tau$ of 45 fm, its decay vertex is indistinguishable from the primary collision vertex. Therefore the selection criteria adopted for the candidate daughter tracks are the ones used for primaries, as specified in Table 1.

A crucial issue for the ϕ reconstruction, as for any strongly decaying resonance, is the combinatorial background determination. In the present analysis PID is used to select kaons, rejecting most of the background while leading to a very small loss in efficiency. For this purpose, tracks are selected if the PID information from the TPC is compatible with a kaon signal and using the TOF signal when available.

For each track, the expected energy loss is calculated using a parametrised response based on the Bethe–Bloch formula [19] computed with a kaon mass hypothesis. It is compared with the TPC specific ionization dE/dx measured via truncated mean (the reconstructed momentum being evaluated at the inner radius of the TPC). With the current TPC calibration for this data set, the assumed dE/dx resolution is 6%. For momenta smaller than 350 MeV/c, the species are well separated so the window is set to $\pm 5\sigma$ with little or no

Fig. 5 *Left panel* shows the distribution of the measured energy-deposit of charged particles versus their momentum in the TPC. *The dashed lines* delimit successively a $\pm 5\sigma$ then a $\pm 3\sigma$ selection of the kaon tracks using the ALEPH parameterization [19] of the Bethe–Bloch curve (*solid line*). *Right panel* shows the relative difference between TOF measured times and that corresponding to a kaon mass hypothesis. *The dashed lines* delimit a coarse fiducial region compatible with this kaon hypothesis



contamination; above 350 MeV/c, it is set instead to $\pm 3\sigma$ as shown in the left panel of Fig. 5.

The accepted band for TOF kaon identification is defined with two hyperbolas as shown in the right panel of Fig. 5.

3.3 Background evaluation and signal extraction

For minimum bias pp collisions, the signals for all particles are clearly distinguishable from the combinatorial background as shown in Fig. 2. Two different methods are used to extract the invariant mass signal from the background. For the single strange particles (K_S^0 , Λ and $\bar{\Lambda}$), the signal is first approximated by a Gaussian on a second order polynomial background. This gives an estimate of the signal mean and width although the invariant mass signal is not strictly Gaussian. Then the background is sampled on each side of the signal by using both sampled regions that are more than 6σ away from the Gaussian mean. The assumption that no reconstructed signal is included in these regions is checked using Monte Carlo data. The width of the background regions can vary depending on the p_T interval considered in the invariant mass distributions. The sum of signal and background ($S + B$) is sampled in the region defined by the mean $\pm 4\sigma$.

The sampling method is illustrated in Fig. 6 for the K_S^0 . Two methods are used to evaluate the background and give consistent results. The background areas are either (i) fitted simultaneously with polynomial functions (from first to third order) or (ii) averaged by simply counting the number of entries (“bin-counting”). The background B under the signal S is estimated using the normalized area sampled on both sides of the signal region (Gaussian mean $\pm 4\sigma$). The

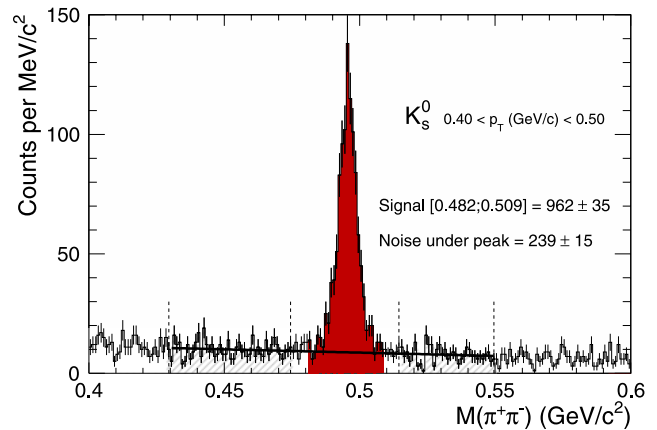


Fig. 6 Plot illustrating the “bin-counting” method used to extract the raw yields. It corresponds to the invariant mass distribution of K_S^0 for the p_T bin 0.4–0.5 GeV/c. *The hashed regions* delimited with *dashed lines* show where the background is sampled; they are chosen to be 6σ away from the signal approximated with a Gaussian distribution. The averaged or fitted background is subtracted from the signal region of $\pm 4\sigma$

signal yield $S = (B + S) - B$ is thus evaluated without any assumption as to its shape. Systematic effects such as signal asymmetry are taken into account by varying the size of the signal and background intervals up to 1σ . The difference between the two methods (fit and bin-counting) contributes to the evaluation of the systematic uncertainties associated to the signal extraction.

In the case of the $\Xi^- + \bar{\Xi}^+$, statistical uncertainties are significant so that, in parallel to the bin-counting method, the background level is simply estimated by a straight line fit.

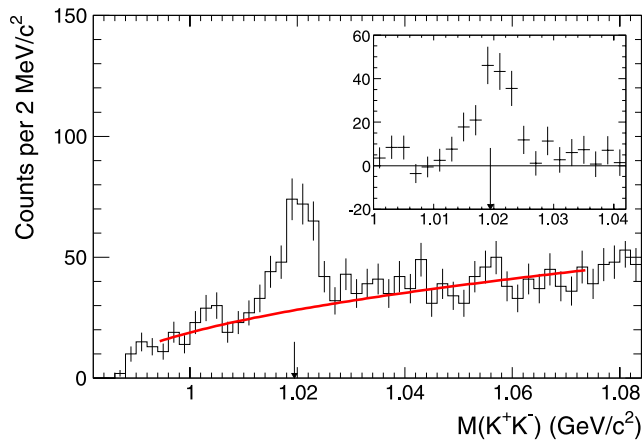


Fig. 7 Background evaluation for the ϕ corresponding to the p_T bin [1.0–1.5] GeV/c. The inset shows the ϕ signal after background subtraction. The vertical arrows indicate the nominal mass value from PDG

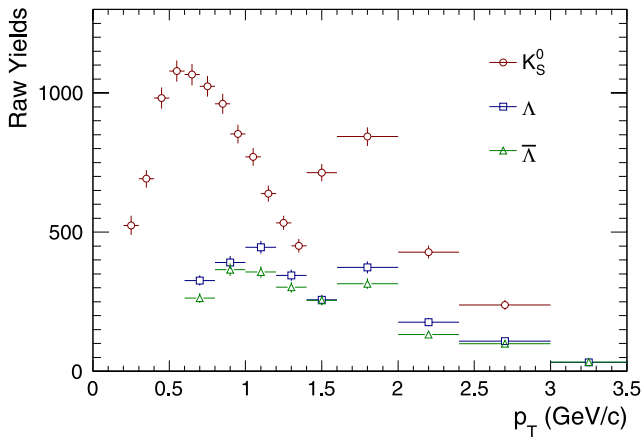


Fig. 8 Reconstructed (raw) yields of K_S^0 (open circles), Λ (open squares) and $\bar{\Lambda}$ (open triangles) as a function of p_T . The change of bin size results in successive offsets of the raw counts at $p_T = [1.4, 1.6, 2.4]$ GeV/c for K_S^0 and $p_T = [1.6, 2.4, 3.0]$ GeV/c for Λ and $\bar{\Lambda}$. Uncertainties correspond to the statistics and the systematics from the signal extraction. They are represented by the vertical error bars. The horizontal error bars give the bin width

The ϕ invariant mass distribution has a larger combinatorial background and a function reproducing both the background and the signal is preferred. It is found that the background can be well reproduced by a function $f(M) = a\sqrt{M-b}$, while the peak has the shape of a Gaussian. The peak range is defined as $\pm 4\sigma$ around the PDG mass of the ϕ , where $\sigma = \Gamma/2.35$ and Γ is the nominal value of the resonance full width at half maximum ($4.5 \text{ MeV}/c^2$) [18]. For each analyzed p_T bin, several fit ranges are investigated. It is found that the fitted width matches that extracted from a full Monte Carlo simulation (as defined in Sect. 3.4) within 5%, except for the last p_T bin where it is broader ($\sim 10\%$). While fluctuations of the fit values as a function of the fit range are taken into account for the systematic error (see Sect. 3.5.1),

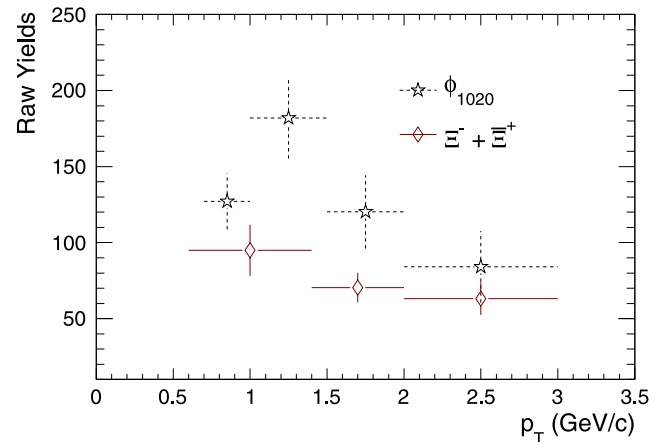


Fig. 9 Reconstructed (raw) yields of ϕ (stars) and $\Xi^- + \Xi^+$ (diamonds) as a function of p_T . With the current statistics, 4 p_T bins are used for the ϕ ([0.7–1.0], [0.7–1.5], [1.5–2.0] and [2.0–3.0] GeV/c) and 3 p_T bins for the $\Xi^- + \Xi^+$ ([0.6–1.4], [1.4–2.0] and [2.0–3.0] GeV/c). Uncertainties correspond to the statistics (i.e. the number of reconstructed particles) and the systematics from the signal extraction. They are represented by the vertical error bars (the horizontal ones give the bin width)

the fit values used for all subsequent steps in the analysis are those that minimize the difference $|\chi^2/\text{NDF} - 1|$. Figure 7 illustrates the method for the [1.0–1.5] GeV/c p_T bin. Every unlike-sign track pair passing all selection criteria and falling within the ϕ invariant mass peak range is counted. The total number of ϕ is estimated by subtracting the integral of the background function alone, computed in the same invariant mass range.

The signal counts (raw yields) for each of the p_T bins are histogrammed as a function of p_T for K_S^0 , Λ , $\bar{\Lambda}$ in Fig. 8 and for ϕ and $\Xi^- + \Xi^+$ in Fig. 9. The uncertainties correspond to both the statistical errors related to the number of counts and the systematics from the bin-counting and fit methods used to extract the signal from the background.

3.4 Efficiency corrections

The efficiency corrections are obtained by analysing Monte Carlo (MC) events in exactly the same way as for the real events. Little dependence is found on the several MC generators which are used. Therefore the corrections presented here are obtained using the event generator PYTHIA 6.4 (tune D6T) [1, 20] and GEANT3 [21, 22] for particle transport through the ALICE detectors.

The MC information is propagated through the whole reconstruction and identification procedure to generate the differential p_T efficiencies as shown in Fig. 10 for K_S^0 , Λ and $\bar{\Lambda}$ and in Fig. 11 for ϕ and $\Xi^- + \Xi^+$. The uncertainties correspond to the statistics of Monte Carlo samples used to compute the corrections. For all particles, the global efficiency is limited at low p_T because of the acceptance of at least two

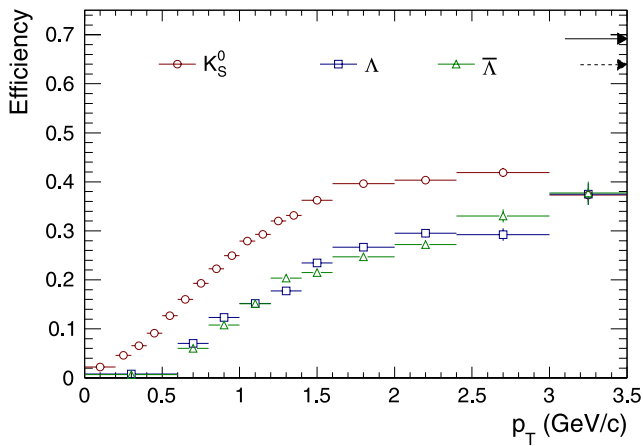


Fig. 10 Efficiency of K_S^0 (open circles), Λ (open squares) and $\bar{\Lambda}$ (open triangles) as a function of p_T . The uncertainties correspond to the statistics in Monte Carlo samples used to compute the corrections. The efficiency is limited by the branching ratio represented by a solid arrow for K_S^0 (0.692) and by a dashed arrow for Λ and $\bar{\Lambda}$ (0.639)

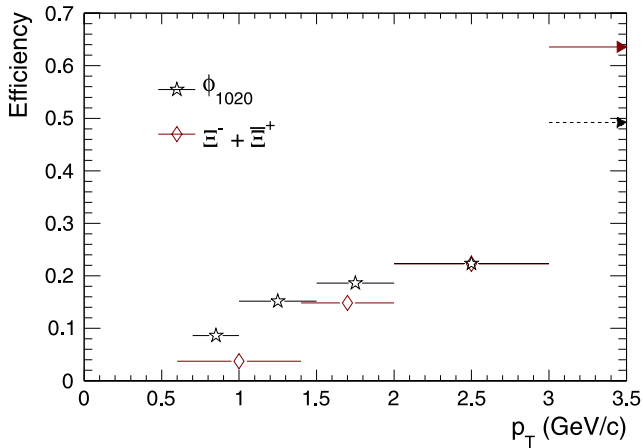


Fig. 11 Efficiency of ϕ (stars) and $\Xi^- + \Xi^+$ (diamonds) as a function of p_T . The uncertainties correspond to the statistics in the Monte Carlo sample used to compute the corrections. The efficiency is limited by the branching ratio represented by a solid arrow for $\Xi^- + \Xi^+$ (0.636) and by a dashed arrow for ϕ (0.492)

charged daughter tracks in the detection volume of the TPC (three tracks in the case of $\Xi^- + \Xi^+$). It rapidly increases with p_T but cannot exceed the asymptotic limits given by the charged particle decay branching ratios presented in Table 3. The difference between the Λ and $\bar{\Lambda}$ reflects the absorption of the anti-proton daughter of the $\bar{\Lambda}$. For all the variables used to select the particles and improve the signal over noise ratio (see Tables 1 and 2), it is verified that data and MC distributions match, thus possible efficiency biases can be properly managed. Examples of such distributions are presented in Fig. 12.

3.5 Estimation of the systematic uncertainties

Systematic uncertainties are discussed in the following sections, where details are given on the contributions due to topological selections and signal extraction methods, as well as those due to material budget and feed-down. As for efficiency corrections, MC data are generated with PYTHIA 6.4 (tune D6T) [1, 20] and transported with GEANT3 [21, 22]. At low p_T , the anti-proton absorption cross section in GEANT3 is known to be too large [23–26]. GEANT4 (with the absorption cross-sections of [27]) was then used to correct the anti-proton tracking efficiency. The information is summarized in Table 4. In addition to these point-to-point systematic uncertainties, there is also a 2% systematic error on the global normalization coming from the evaluation of the total number of inelastic events.

3.5.1 Systematic uncertainties due to track or topological selections and signal extraction

Systematic uncertainties due to tracking and topological identification are determined by varying the track and topological (for secondary vertices) selections, as well as the definition of the regions sampled for signal extraction. To assess the different systematic uncertainties, only the deviations that are statistically significant are taken into account (more than 2 standard deviations away from the central value on the corrected spectrum).

The systematic variation of track and topological selections results in a variation of the amount of signal extracted from invariant mass distribution in both data and the Monte Carlo simulation mentioned above. The difference between these amounts of signal corresponds indirectly to the accuracy with which the MC simulation reproduces the characteristics of real events, from the simulation of the detector response to the background shape and composition considered for the extracted signal. It is estimated that the point-to-point uncertainties in the p_T spectra are at most 4.6%, 3.3%, 4.7%, 6% and 13.9% for the K_S^0 , Λ , $\bar{\Lambda}$, ϕ and $\Xi^- + \Xi^+$, respectively. The systematic uncertainties of the signal extraction for K_S^0 , Λ , $\bar{\Lambda}$ and $\Xi^- + \Xi^+$ are p_T -dependent and estimated by varying the invariant mass regions where the signal and the background are sampled using the bin-counting method described in Sect. 3.3. For the ϕ signal, the systematic uncertainties from background subtraction are estimated using three different criteria. First of all, the function reproducing the background is replaced by a second or third order polynomial. Moreover, the fit is repeated fixing the width parameter to $\pm 10\%$ of the value obtained in the default procedure (described in Sect. 3.3), and also to the value obtained when fitting the Monte Carlo sample and to $\pm 10\%$ of this. Finally, the fit range is also varied. All of these computations result in a variation of the raw counts with respect to those

Fig. 12 (Color online) Comparison between data (*red circles*) and Monte Carlo (*black open triangles*) for several topological variables used to select secondary vertices. The *top panels* correspond to the K_S^0 candidates selected in a $\pm 20 \text{ MeV}/c^2$ invariant mass window around the nominal mass. The distribution of the DCA between the positive daughter track and the primary vertex and the DCA distribution between the two daughters tracks are displayed in the *left* and the *right top panels* respectively. On the *bottom panels*, the same distributions are shown for the Λ candidates selected in a $\pm 8 \text{ MeV}/c^2$ invariant mass window around the nominal mass

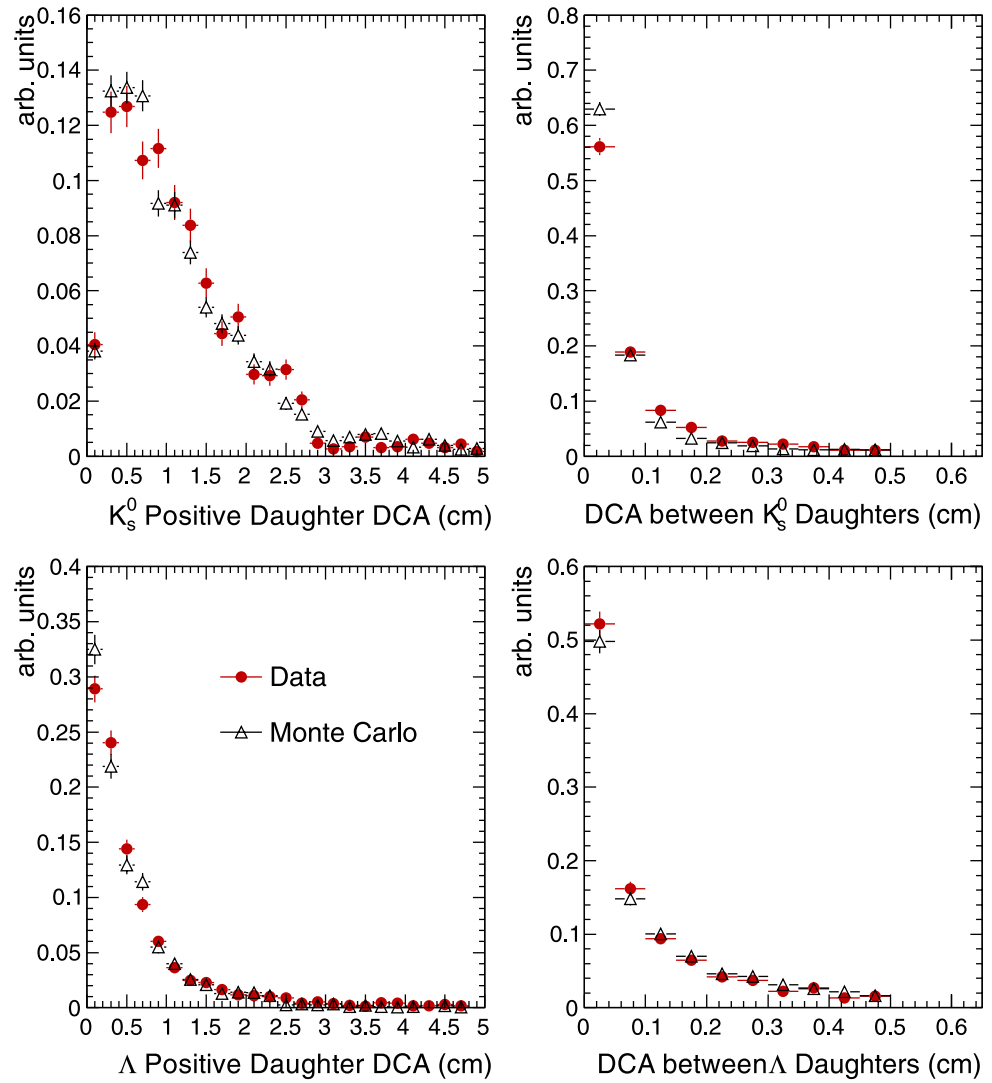


Table 4 Point-to-point systematic uncertainties expressed in percentage for p_T spectra of different particles. For each particle, the reported values correspond to the effect on the lowest p_T bin, the average and the highest p_T bin, except for the feed-down contributions where values are estimated as being constant versus p_T or where the effect is found to be negligible (less than 2 standard deviations from the default value on the corrected spectrum)

Systematic effects (%)	K_S^0	Λ	$\bar{\Lambda}$	ϕ	$\Xi^- + \bar{\Xi}^+$
Selections					
tracks	[4.6–1.1–2.1]	[2.6–2.0–2.5]	[3.0–2.0–4.1]	[0.9–3.1–6.0]	[negl.–5.4–negl.]
topological	[3.8–1.4–1.3]	[3.3–3.3–1.5]	[4.7–4.7–3.8]	–	[6.8–11.6–13.9]
Signal extraction	[4.5–1.5–1.5]	[3.0–2.0–5.0]	[3.0–2.0–5.0]	[3.2–4.3–7.0]	[5.6–negl.–2.5]
TPC dE/dx	–	[5–negl.]	[5–negl.]	[1.8–2.9–3.6]	[negl.]
Efficiency					
material budget	[1.5–1.5–1.1]	[3.4–1.0–1.6]	[3.7–2.0–4.5]	[4.7–4.0–2.3]	[2.7–1.5–3.6]
\bar{p} cross-section	–	< 1	< 2	–	< 2
Feed-down	–	1.7	1	–	–

shown in Fig. 9. Although a quite large compatibility region is requested for PID (at least 3σ) the effects of varying the dE/dx selections are taken into account for the corresponding efficiency calculation. For both ϕ and $\Xi^- + \bar{\Xi}^+$, statistical errors dominate after signal extraction (see Sect. 3.3) and consequently, some systematic effects due to PID are extrapolated from single track and V0 measurements. The TOF PID selection is applied only to reject the ϕ background. No systematic effects are observed on the ϕ signal (i) for the Monte Carlo data sample, when the selection is applied to the ϕ daughters in addition to all other cuts; (ii) for real events, when comparing the ϕ statistics before and after applying the selection.

3.5.2 Systematic uncertainties due to material budget and absorption cross-section

A dedicated study involved the variation of the detector material thickness crossed by particles. The material budget uncertainty, based on γ conversion measurements, is estimated to be 7% in terms of radiation length [23]. The efficiency variation due to this material budget uncertainty depends on the momentum of each of the decay daughters. Although such a variation is also correlated with the momentum of the parent particle, the corresponding systematic uncertainties are reported as point-to-point errors in Table 4 for the lowest, the average and the highest p_T bin and eventually added in quadrature to the total systematic errors.

Specific uncertainties are related to the (anti-)proton absorption and scattering cross-sections used for propagating these particles through the geometry of the detectors with both GEANT3 [21, 22] (and its default absorption cross-sections) and GEANT4 (using the absorption cross-sections of [27]). More details about the modifications can be found in [23–26] and references therein. The corresponding corrections are taken into account in the efficiency versus p_T assuming that absorption cross-sections are identical for the (anti-) hyperon and its (anti-) proton daughter. The uncertainties associated with these corrections are derived from the (anti-)proton cross-section uncertainties and the values are estimated as constant and lower than 1% (2%) for Λ ($\bar{\Lambda}$) and 2% for $\Xi^- + \bar{\Xi}^+$.

3.5.3 Systematic uncertainties for Λ and $\bar{\Lambda}$ due to feed-down

Some of the reconstructed Λ and $\bar{\Lambda}$ particles come from decays of Ξ -hyperons. The proportion of reconstructed secondary Λ and $\bar{\Lambda}$ depends on the selection criteria used. For the parameters listed in Table 2 (V0 vertex part), the impact of the Ξ feed-down on the final spectra is evaluated to be 13% for Λ and 12% for $\bar{\Lambda}$. No p_T dependence is found within uncertainties.

This assessment results in a global correction of the spectra, applied as an additional factor in the overall normalization. Provided that both primary and secondary Λ have similar spectral shapes, such integrated correction is applicable. This is tested directly using Monte Carlo data, but also with real data, changing the fraction of the secondary Λ by varying the DCA of reconstructed candidates. Within the available statistics and p_T reach, no significant change in spectral shape is observed.

Using Monte Carlo, the ratio $r_{\text{feed-down}}$ of the reconstructed Ξ^- ($\bar{\Xi}^+$) candidates to the number of reconstructed Λ ($\bar{\Lambda}$) candidates from Ξ decays is:

$$r_{\text{feed-down}} = \frac{(N_{\Xi^-})_{\text{MC}}}{(N_{\Lambda \leftarrow \Xi^-})_{\text{MC}}}$$

Assuming that this ratio is the same in both Monte Carlo and data, the whole feed-down contribution to the spectra is estimated by dividing the number of reconstructed Ξ^- ($\bar{\Xi}^+$) in data by the ratio extracted from Monte Carlo:

$$(N_{\Lambda \leftarrow \Xi^-})_{\text{data}} = \frac{(N_{\Xi^-})_{\text{data}}}{r_{\text{feed-down}}}$$

Besides the Ξ contribution, other sources may feed the Λ population resulting in additional systematic uncertainties. In Monte Carlo simulations, Λ particles possibly generated in the detector material induce a 1.7% uncertainty. The same uncertainty in the case of $\bar{\Lambda}$ is below 1%. The contribution from Ω decays is found to be negligible. It should be noted that since Λ ($\bar{\Lambda}$) from electromagnetic Σ^0 ($\bar{\Sigma}^0$) decays cannot be distinguished from the direct ones, the identified Λ ($\bar{\Lambda}$) include these contributions.

3.6 p_T spectra and global yield extraction

The K_S^0 spectrum is first shown on a linear scale in Fig. 13 and compared with charged kaon spectra [16]. Within uncertainties, good agreement is found between K_S^0 and K^+ in the measured p_T range.

Figure 14 presents the corrected p_T spectra for all species, including both statistical errors and systematic uncertainties. The spectra are fitted with two different functional forms in order to extract the global integrated yields:

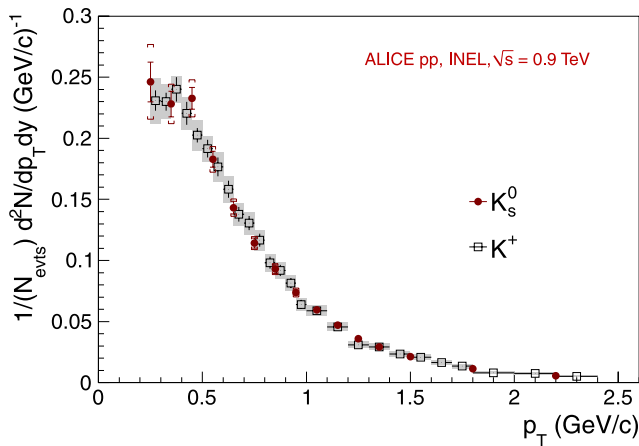
$$\frac{d^2N}{dy dp_T} = A \times p_T \times e^{-\frac{p_T}{T}} \tag{1}$$

$$\frac{d^2N}{dy dp_T} = \frac{(n-1)(n-2)}{nT[nT+m(n-2)]} \times \frac{dN}{dy} \times p_T \times \left(1 + \frac{m_T - m}{nT}\right)^{-n} \tag{2}$$

where $m_T = \sqrt{m^2 + p_T^2}$. The p_T exponential has two parameters: the normalization A and the inverse slope parameter T . The Lévy function [Eq. (2)], already used at lower

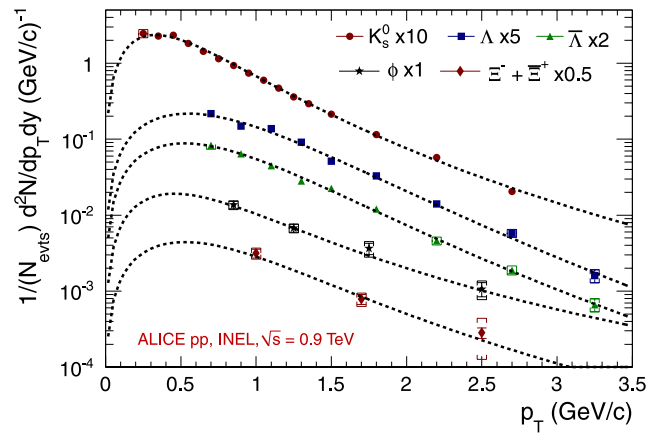
Table 5 Summary of the parameters extracted from the fits to the measured transverse momenta spectra using p_T exponential (1) and Lévy (2) functional forms and including point-to-point systematic uncertainties

Particles		p_T exponential (1)		Lévy (2)		
		T (MeV)	χ^2/NDF	T (MeV)	n	χ^2/NDF
Mesons	K_S^0	325 ± 4	117.6/14	168 ± 5	6.6 ± 0.3	10.8/13
	ϕ	438 ± 31	1.3/2	164 ± 91	4.2 ± 2.5	0.6/1
Baryons	Λ	392 ± 6	10.2/7	229 ± 15	10.8 ± 2.0	9.6/6
	$\bar{\Lambda}$	385 ± 6	5.1/7	210 ± 15	9.2 ± 1.4	3.7/6
	$\Xi^- + \bar{\Xi}^+$	421 ± 42	2.0/1	175 ± 50	5.2 ± 2.3	–

**Fig. 13** Comparison of the corrected yields as a function of p_T for K_S^0 (circle) and charged kaons (K^+) (open squares), identified via energy loss in the TPC and ITS, and via time of flight in the TOF. The points are plotted at the centre of the bins. The full vertical lines associated to the K_S^0 points, as well as the gray shaded areas associated to the K^+ points, correspond to the statistical and systematic uncertainties summed in quadrature whereas the inner vertical lines contain only the statistical uncertainties (i.e. the number of reconstructed particles) and the systematics from the signal extraction

energies [28] and alternatively referred to as the Tsallis function [29], is shown to be useful when the p_T range is wide: it includes both an exponential shape for low p_T (which can be characterized by an inverse slope parameter T) and a power law component (governed by the power parameter n) for the higher p_T region. The results of these fits to the spectra, where statistical and systematic uncertainties are added in quadrature, are shown in Fig. 14 and in Table 5. In the case of K_S^0 for which the statistics and the p_T range are larger than for other species, the χ^2/NDF indicates clearly that the p_T exponential parameterization cannot properly reproduce the spectrum shape.

For the spectra of the ϕ , Λ and $\bar{\Lambda}$ both functions give similar and acceptable χ^2/NDF . Within uncertainties, Λ and $\bar{\Lambda}$ have the same fit parameters. In the case of the $\Xi^- + \bar{\Xi}^+$ spectrum, the low number (i.e. 3) of p_T bins cannot constrain the Lévy function and therefore its χ^2/NDF in Table 5 is not defined. Nevertheless, for consistency and

**Fig. 14** Particle spectra (corrected yields) as a function of p_T for K_S^0 (circles), Λ (squares), $\bar{\Lambda}$ (triangles), ϕ (stars) and $\Xi^- + \bar{\Xi}^+$ (diamonds). The data points are scaled for visibility and plotted at the centre of the bins. Uncertainties corresponding to both statistics (i.e. the number of reconstructed particles) and systematics from the signal extraction are shown as vertical error bars. Statistical uncertainties and systematics (summarized in Table 4) added in quadrature are shown as brackets. The fits (dotted curves) using Lévy functional form [see Eq. (2)] are superimposed

in order to extract particle ratios, a Lévy fit is performed to obtain the integrated yields and particle ratios for all species. It must be noted that the rapidity range is slightly different for each species (cf. Table 6). However, the rapidity dependence of particle production at mid-rapidity is weak enough to allow direct comparisons of the spectra [23].

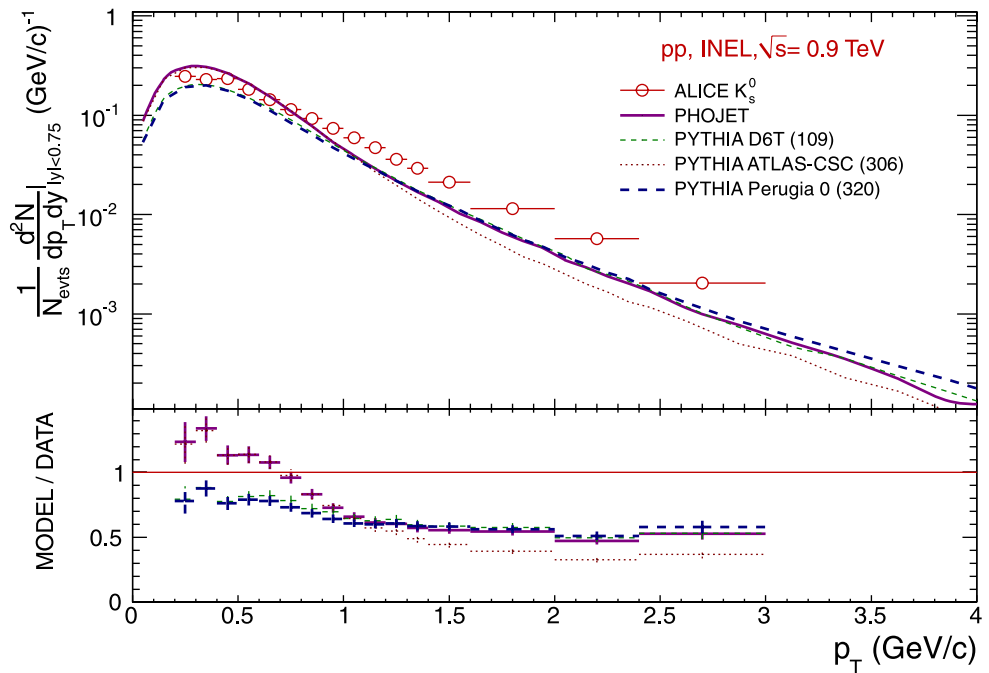
4 Results and discussion

The p_T spectra for K_S^0 , Λ , $\bar{\Lambda}$ and ϕ are shown in Fig. 14 along with the Lévy fits. When comparing the different spectra, it is found that the inverse slope parameter T increases with the mass of the particle. For example, it changes from 168 ± 5 MeV for K_S^0 to 229 ± 15 MeV for Λ when the Lévy fit is used. The $\Xi^- + \bar{\Xi}^+$ apparently do not follow this trend.

Table 6 Rapidity and p_T ranges, $\langle p_T \rangle$, corrected yields and extrapolated fraction at low p_T using the Lévy function (2)

Particles	$ y $	p_T range (GeV/c)	$\langle p_T \rangle$ (GeV/c)	dN/dy (INEL)	Extrapolation (%)
Mesons	K_S^0	<0.75	[0.2–3.0]	$0.184 \pm 0.002 \pm 0.006$	$12 \pm 0.4 \pm 0.5$
	ϕ	<0.60	[0.7–3.0]	$0.021 \pm 0.004 \pm 0.003$	$48 \pm 18 \pm 7$
Baryons	Λ	<0.75	[0.6–3.5]	$0.048 \pm 0.001 \pm 0.004$	$36 \pm 2 \pm 4$
	$\bar{\Lambda}$	<0.75	[0.6–3.5]	$0.047 \pm 0.002 \pm 0.005$	$39 \pm 3 \pm 4$
	$\Xi^- + \bar{\Xi}^+$	<0.8	[0.6–3.0]	$0.0101 \pm 0.0020 \pm 0.0009$	$35 \pm 8 \pm 4$

Fig. 15 Comparison of the transverse momentum differential yield for the K_S^0 particles for INEL pp collisions with PHOJET and PYTHIA tunes 109, 306 and 320



However, this is most likely because the very limited statistics do not allow for a well-constrained fit. The shapes of the p_T spectra are also compared to PHOJET and PYTHIA models. For PYTHIA, several tunes (109 [20], 306 [30] and 320 [31]) are presented. For all species, the p_T spectra are found to be slightly harder (i.e. they have a slower decrease with p_T) than the models as presented in Figs. 15, 16, 17 and 18. For transverse momenta larger than ~ 1 GeV/c, the strange particle spectra are strongly underestimated by all models, by a factor of ~ 2 for K_S^0 and even ~ 3 for hyperons. The discrepancy is smaller in the case of the ϕ .

The integrated yields (dN/dy) are obtained using the spectra in the measured range and integrating the Lévy function for the extrapolated regions at low and high p_T . The uncertainties for the dN/dy and $\langle p_T \rangle$ values are computed from the errors on the fit parameters, where both the point-to-point statistical and systematic uncertainties of the p_T spectra are taken into account. Due to the rapid decrease of the spectra, most of the extrapolation is done in the low p_T region and amounts to 12% for K_S^0 and 48% for the ϕ (smallest and highest values respectively).

Therefore, an additional uncertainty is added for the dN/dy to account for the uncertainty in the shape of the spectra outside the measured range, which differs for each particle: it corresponds to 25% of the yield fraction that is extrapolated at low p_T . The measured p_T ranges are specified in Table 6 for each particle species.

Using the particle integrated yields presented in this paper along with the yields of charged π , K, p and \bar{p} [16] and the measured \bar{p}/p ratio [23], a comparison with STAR feed-down corrected particle ratios at $\sqrt{s} = 0.2$ TeV [28] is shown in Fig. 19. With the centre of mass energy increasing from $\sqrt{s} = 0.2$ TeV to 0.9 TeV the measured ratios are similar except the \bar{p}/π^- ratio which decreases slightly from 0.068 ± 0.011 to 0.051 ± 0.005 . The strange to non-strange particle ratios seem to increase but stay compatible within uncertainties: the K^-/π^- from 0.101 ± 0.012 to 0.121 ± 0.013 and the $\bar{\Lambda}/\pi^+$ from 0.027 ± 0.004 to 0.032 ± 0.003 .

The yields and $\langle p_T \rangle$ obtained with the ALICE experiment are compared for each particle with existing data at the same energy and also with results at lower and higher en-

Fig. 16 Comparison of the transverse momentum differential yield for the Λ particles for INEL pp collisions with PHOJET and PYTHIA tunes 109, 306 and 320

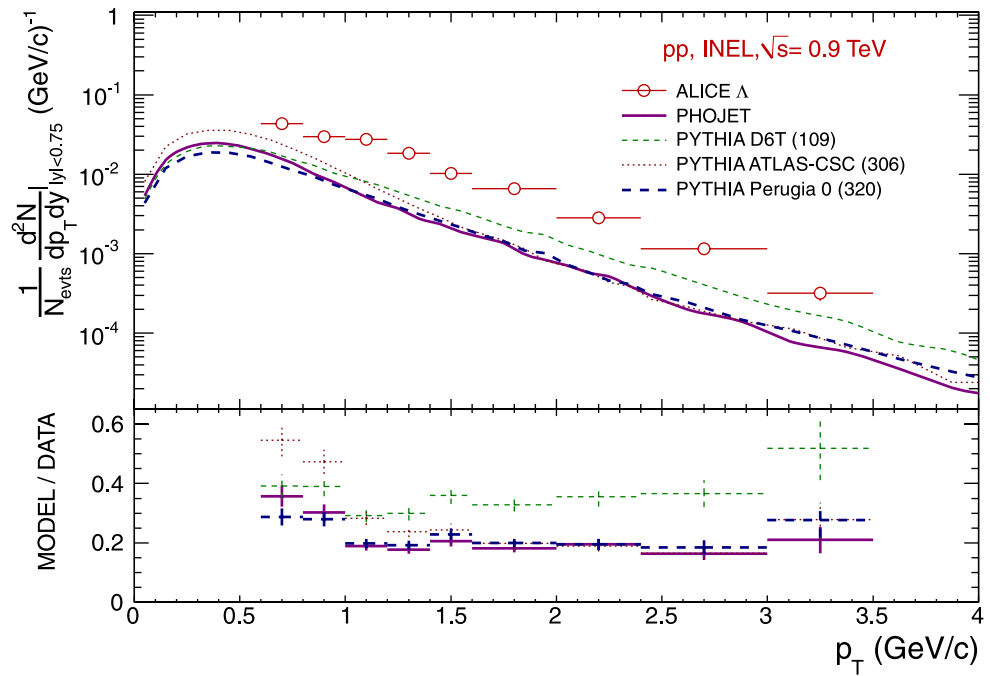
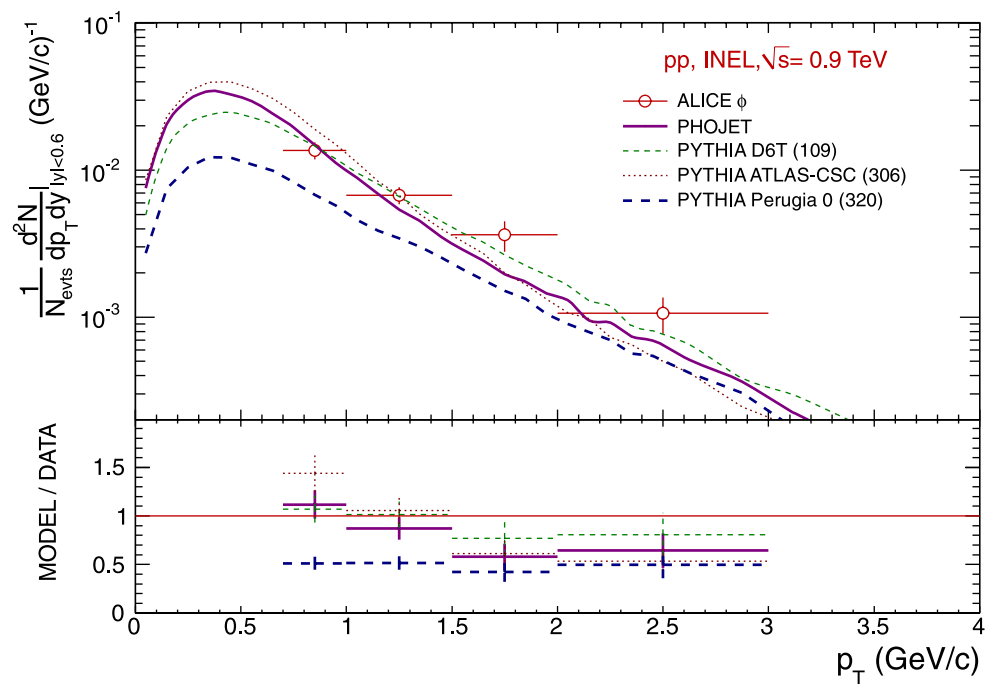


Fig. 17 Comparison of the transverse momentum differential yield for the ϕ particle for INEL pp collisions with PHOJET and PYTHIA tunes 109, 306 and 320



ergies. The various experiments differ in acceptance and in event selection, i.e. INEL, non-single-diffractive (NSD) or non-diffractive (ND). The dependence of $\langle p_T \rangle$ with respect to these variables is found to be negligible though. Consequently the $\langle p_T \rangle$ values are directly comparable, whereas the comparison of the yields can require further scaling because of different (pseudo)rapidity coverages. Figure 20 reports ALICE $\langle p_T \rangle$ measurements along with those of the STAR experiment [28, 32]. It is remarkable that the $\langle p_T \rangle$ re-

mains close to the ISR parameterization [33] although the collision energy increased by a factor 36. Table 7 summarizes the K_S^0 measurements performed by the UA5 [34], CDF [35] and ALICE Collaborations for INEL events, and by the STAR [28] Collaboration for NSD events. The ALICE K_S^0 yield at mid-rapidity, as well as the $\langle p_T \rangle$, are in good agreement with UA5 results at 900 GeV albeit with improved precision. The comparison of $(\Lambda + \bar{\Lambda})$ measurements are presented in Table 8 for NSD events. ALICE

Fig. 18 Comparison of the transverse momentum differential yield for the $\Xi^- + \Xi^+$ particle for INEL pp collisions with PHOJET and PYTHIA tunes 109, 306 and 320

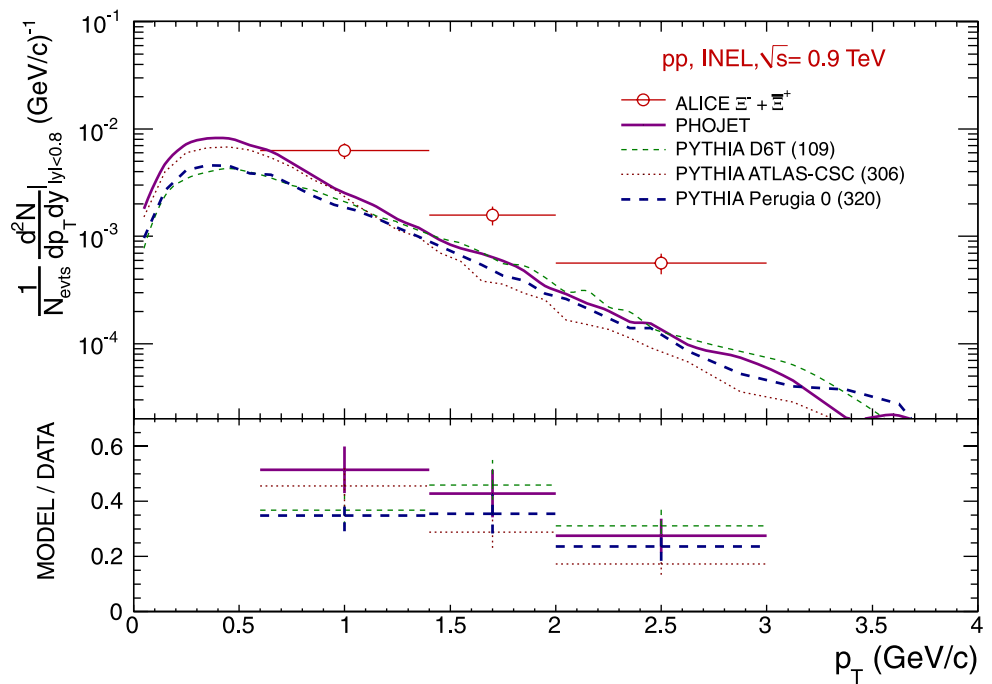


Table 7 The K_S^0 mean transverse momentum and yields in INEL events from UA5, CDF, and ALICE and in NSD events in STAR for various \sqrt{s} . STAR results are taken from [28], CDF ones and yield values with "*" are from [35]. Other UA5 values concerning $\langle p_T \rangle$ are from [34]

Experiment	\sqrt{s} (GeV)	Acceptance	$\langle p_T \rangle$ (GeV/c)	dN/dy
STAR	200	$ y < 0.5$	0.61 ± 0.02	0.134 ± 0.011
UA5	200	$ y < 2.5$	0.53 ± 0.07	$0.14 \pm 0.02^*$
UA5	546	$ y < 2.5$	0.57 ± 0.03	$0.15 \pm 0.02^*$
CDF	630	$ y < 1.0$	0.5 ± 0.1	$0.2 \pm 0.1^*$
UA5	900	$ y < 2.5$	0.62 ± 0.08	$0.18 \pm 0.02^*$
ALICE	900	$ y < 0.75$	$0.65 \pm 0.01 \pm 0.01$	$0.184 \pm 0.002 \pm 0.006$
CDF	1800	$ y < 1.0$	0.60 ± 0.03	$0.26 \pm 0.03^*$

Table 8 The $(\Lambda + \bar{\Lambda})$ mean transverse momentum and yields for NSD events and different \sqrt{s} . STAR results are from [28] and UA5 results are from [36, 37]. ALICE and STAR results are feed-down corrected. The yields measured by ALICE for INEL events have been scaled to NSD with Eq. 3 then to match UA5 acceptance ($|y| < 2.0$) using the method explained in Sect. 4

Experiment	\sqrt{s} (GeV)	Acceptance	$\langle p_T \rangle$ (GeV/c)	dN/dy $_{ y=0}$ (NSD)	$\langle n_{\Lambda+\bar{\Lambda}} \rangle$ per event	
					Measured	Scaled to UA5 $ y $
STAR	200	$ y < 0.5$	0.77 ± 0.04	0.074 ± 0.005	–	0.24 ± 0.02
UA5	200	$ y < 2.0$	$0.80^{+0.20}_{-0.14}$	–	0.27 ± 0.07	–
UA5	546	$ y < 2.0$	0.62 ± 0.08	–	0.25 ± 0.05	–
UA5	900	$ y < 2.0$	0.74 ± 0.09	–	0.38 ± 0.08	–
ALICE	900	$ y < 0.75$	$0.85 \pm 0.01 \pm 0.01$	$0.113 \pm 0.003 \pm 0.008$	—	$0.45 \pm 0.01 \pm 0.03$

yields, measured in $|y| < 0.75$ for INEL events, are scaled to the UA5 [36, 37] acceptance ($|y| < 2.0$) using PYTHIA simulations. The $\Lambda + \bar{\Lambda}$ yield in NSD events is estimated by scaling the measured yield in inelastic events with the

known ratio R of charged particle multiplicities in NSD and INEL events [16]:

$$R = \frac{(dN_{ch}/d\eta)_{NSD}}{(dN_{ch}/d\eta)_{INEL}} = 1.185 \pm 0.008 \quad (3)$$

Table 9 The $(\Xi^- + \Xi^+)$ mean transverse momentum and yields for NSD events and different \sqrt{s} . STAR results are from [28] and UA5 results are from [37]. UA5 measures $(\Xi^- + \Xi^+)$ for $p_T > 1$ GeV/c. The yields measured by ALICE for INEL events have been scaled to NSD with Eq. 3 then to match the UA5 acceptance ($|y| < 3.0$) using the method explained in Sect. 4

Experiment	\sqrt{s} (GeV)	Acceptance	$\langle p_T \rangle$ (GeV/c)	$dN/dy _{y=0}$ (NSD)	$\langle n_{\Xi^- + \Xi^+} \rangle$ per event	
					Measured	Scaled to UA5 $ y $
STAR	200	$ y < 0.5$	0.90 ± 0.01	0.006 ± 0.001	–	0.022 ± 0.006
UA5	200	$ y < 3.0$	$0.80^{+0.20}_{-0.14}$	–	$0.03^{+0.04}_{-0.02}$	–
UA5	546	$ y < 3.0$	1.10 ± 0.02	–	$0.08^{+0.03}_{-0.02}$	–
UA5	900	$ y < 3.0$	$0.7^{+0.2}_{-0.1}$	–	$0.05^{+0.04}_{-0.02}$	–
ALICE	900	$ y < 0.8$	$0.95 \pm 0.14 \pm 0.03$	$0.0120 \pm 0.0024 \pm 0.0011$	–	$0.071 \pm 0.014 \pm 0.006$

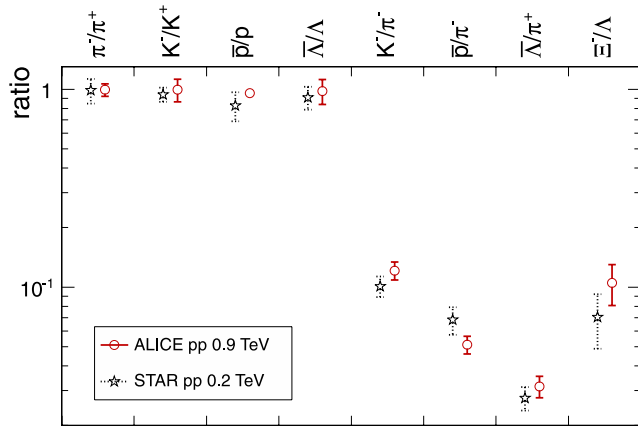


Fig. 19 Ratios of integrated yields including $\pi^{(\pm)}$, $K^{(\pm)}$, p and \bar{p} performed with the ALICE experiment [16, 23] and compared with STAR values for pp collisions at $\sqrt{s} = 0.2$ TeV [28]. All ratios are feed-down corrected. For the ratio Ξ^-/Λ of ALICE, the $dN/dy|_{y=0}$ for $\Xi^- + \Xi^+$ is divided by 2. Statistical and systematic uncertainties are added in quadrature

This scaling factor is also used for the ALICE $(\Xi^- + \Xi^+)$ yield presented in Table 9. The ALICE yields and $\langle p_T \rangle$ for both $(\Lambda + \bar{\Lambda})$ and $(\Xi^- + \Xi^+)$ are in good agreement with the UA5 measurements [37]. Table 10 shows the evolution of dN/dy and $\langle p_T \rangle$ with the collision energy for the ϕ particle scaled to NSD events using also Eq. 3. It includes the ALICE measurements, which are the first ϕ measurements at 900 GeV, and compares them to the results from the STAR experiment [38, 39] at 200 GeV and the E735 experiment [40] at 1800 GeV.

The baryon to meson ratio as a function of p_T obtained with the $(\Lambda + \bar{\Lambda})$ and K_S^0 spectra measured by ALICE is presented in Fig. 21. It includes the $(\Lambda + \bar{\Lambda})/2K_S^0$ ratio in pp collisions at 200 GeV measured by STAR [28], and the ratios in $p\bar{p}$ collisions at 630 GeV and 1800 GeV computed with the $(\Lambda + \bar{\Lambda})$ and K_S^0 spectra published by CDF [41] and UA1 [42]. UA1 and CDF Collaborations provide inclusive spectra whereas the ALICE and STAR ones are feed-

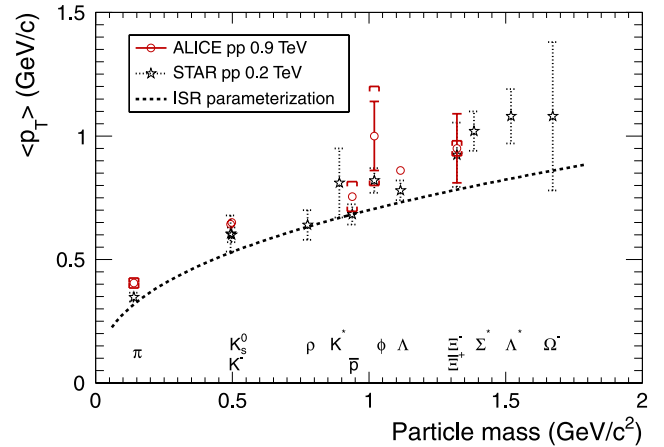


Fig. 20 $\langle p_T \rangle$ vs. particle mass for the measurements performed with the ALICE experiment and compared with STAR values for pp collisions at $\sqrt{s} = 0.2$ TeV [28, 32] and the ISR parameterization [33]. Both statistical (vertical error bars) and systematic (brackets) uncertainties are shown for ALICE data

down corrected.¹ The acceptance windows of these experiments differ significantly: ALICE measures Λ , $\bar{\Lambda}$ and K_S^0 in $|y| < 0.75$, STAR in $|y| < 0.5$, CDF in $|\eta| < 1.0$, whereas UA1 reconstructs $(\Lambda + \bar{\Lambda})$ in $|\eta| < 2.0$ and K_S^0 in $|\eta| < 2.5$. The ALICE ratio agrees very well with the STAR results in the measured p_T range, which would suggest little or no energy dependence of $(\Lambda + \bar{\Lambda})/2K_S^0$. A similar conclusion can be drawn when comparing only the ratios measured by CDF at 630 GeV and 1800 GeV. However, the absolute ratio found by CDF for $p_T > 1.5$ GeV/c is significantly higher than the one observed with ALICE and STAR. Also the ratio computed from UA1 spectra shows a clear disagreement with the ALICE and STAR data above $p_T \approx 1.5$ GeV/c and in addition differs from CDF in the intermediate p_T range (1.5–3 GeV/c). PYTHIA simulations show that this discrep-

¹The ALICE ratio would increase by about 20% if all produced Λ and $\bar{\Lambda}$ were included.

Table 10 The ϕ mean transverse momentum and yields for different \sqrt{s} . STAR results (NSD) are from [38, 39] and E735 non-diffractive results (ND) are from [40]. The E735 Collaboration provided two values of $\langle p_T \rangle$ depending on the functional form used to fit the data points and the uncertainties associated with each value are only statistical. The yields measured by ALICE for INEL events have been scaled to NSD with Eq. 3 as explained in Sect. 4

Experiment	\sqrt{s} (GeV)	Acceptance	$\langle p_T \rangle$ (GeV/c)	$dN/dy _{y=0}$
STAR (NSD)	200	$ y < 0.5$	$0.82 \pm 0.03 \pm 0.04$	$0.018 \pm 0.001 \pm 0.003$
ALICE (NSD)	900	$ y < 0.6$	$1.00 \pm 0.14 \pm 0.20$	$0.025 \pm 0.005 \pm 0.004$
E735 (ND)	1800	$-0.4 < y < 1.0$	1.06 ± 0.18 0.94 ± 0.26	0.0186 ± 0.0041

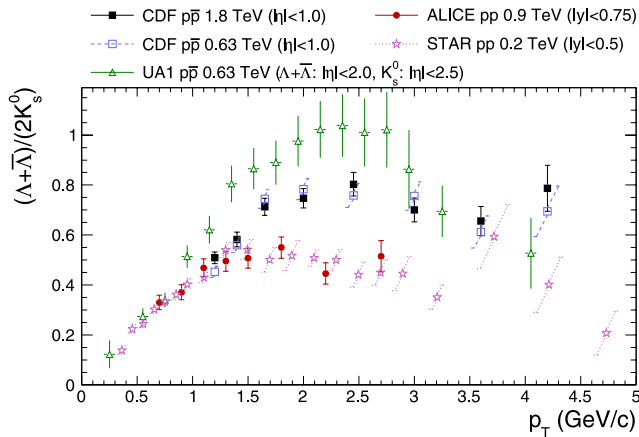


Fig. 21 $(\Lambda + \bar{\Lambda})/2K_S^0$ as a function of p_T for different collision energies in pp and $p\bar{p}$ minimum bias events. The STAR ratio is taken from [28] whereas the CDF and UA1 ratios are computed with the $(\Lambda + \bar{\Lambda})$ and K_S^0 spectra published in [41] and [42] respectively. The ALICE and STAR ratios are feed-down corrected. Because the K_S^0 and $(\Lambda + \bar{\Lambda})$ spectra from UA1 have incompatible binning, the K_S^0 differential yield has been calculated for each $(\Lambda + \bar{\Lambda})$ p_T data point using the fit function published by UA1. Such a choice is motivated by the fact that the χ^2 value for the K_S^0 spectrum fit is better than that for the $(\Lambda + \bar{\Lambda})$ spectrum

any can not be attributed to the differences in the acceptance or in the colliding system (i.e. $p\bar{p}$ instead of pp).

5 Conclusion

Measurements of mesons containing strange quarks (K_S^0 and ϕ) and hyperons (Λ , $\bar{\Lambda}$ and $\Xi^- + \bar{\Xi}^+$) have been performed for inelastic pp collisions at $\sqrt{s} = 0.9$ TeV with the ALICE experiment at the LHC. The Lévy function gives a good description of the transverse momentum spectra which have been compared with pQCD-based models. The K_S^0 transverse momentum spectrum is overestimated by PYTHIA tune ATLAS-CSC and PHOJET below 0.75 GeV/c but is higher by a factor of ~ 2 in the p_T range [1–3] GeV/c. Within uncertainties, the ϕ meson spectrum is reasonably described by these models and the best agreement is obtained by

PYTHIA tune D6T. We find that strange baryons are significantly under-predicted in both PYTHIA and PHOJET by a factor of ~ 3 . The feed-down corrected ratio of baryon to meson as a function of p_T , illustrated by the Λ/K_S^0 , is consistent with the STAR measurements at $\sqrt{s} = 0.2$ TeV but lower than UA1 and CDF results at $\sqrt{s} = 0.63$ TeV and $\sqrt{s} = 1.8$ TeV. The integrated yields and average transverse momenta have been compared with earlier data collected in pp and $p\bar{p}$ interactions at various energies. These results provide a useful baseline for comparisons with recent tunes of the PYTHIA model and a reference for future measurements in heavy-ion collisions at the LHC. These studies demonstrate the precision with which ALICE can measure resonances and topologically reconstructed weakly decaying particles. Measurements of these particles will be a substantial part of the ALICE programme in both pp and Pb–Pb collisions. The measurement of the ϕ resonance provides an unprecedented reference at this energy.

Acknowledgements The ALICE Collaboration would like to thank all its engineers and technicians for their invaluable contributions to the construction of the experiment and the CERN accelerator teams for the outstanding performance of the LHC complex. The ALICE Collaboration acknowledges the following funding agencies for their support in building and running the ALICE detector: Calouste Gulbenkian Foundation from Lisbon and Swiss Fonds Kidagan, Armenia; Conselho Nacional de Desenvolvimento Científico e Tecnológico (CNPq), Financiadora de Estudos e Projetos (FINEP), Fundação de Amparo à Pesquisa do Estado de São Paulo (FAPESP); National Natural Science Foundation of China (NSFC), the Chinese Ministry of Education (CMOE) and the Ministry of Science and Technology of China (MSTC); Ministry of Education and Youth of the Czech Republic; Danish Natural Science Research Council, the Carlsberg Foundation and the Danish National Research Foundation; The European Research Council under the European Community’s Seventh Framework Programme; Helsinki Institute of Physics and the Academy of Finland; French CNRS-IN2P3, the ‘Region Pays de Loire’, ‘Region Alsace’, ‘Region Auvergne’ and CEA, France; German BMBF and the Helmholtz Association; Greek Ministry of Research and Technology; Hungarian OTKA and National Office for Research and Technology (NKTH); Department of Atomic Energy and Department of Science and Technology of the Government of India; Istituto Nazionale di Fisica Nucleare (INFN) of Italy; MEXT Grant-in-Aid for Specially Promoted Research, Japan; Joint Institute for Nuclear Research, Dubna; National Research Foundation of Korea (NRF); CONACYT, DGAPA, México, ALFA-EC and the HELEN Program (High-Energy

physics Latin-American–European Network); Stichting voor Fundamenteel Onderzoek der Materie (FOM) and the Nederlandse Organisatie voor Wetenschappelijk Onderzoek (NWO), Netherlands; Research Council of Norway (NFR); Polish Ministry of Science and Higher Education; National Authority for Scientific Research—NASR (Autoritatea Națională pentru Cercetare Științifică—ANCS); Federal Agency of Science of the Ministry of Education and Science of Russian Federation, International Science and Technology Center, Russian Academy of Sciences, Russian Federal Agency of Atomic Energy, Russian Federal Agency for Science and Innovations and CERN-INTAS; Ministry of Education of Slovakia; CIEMAT, EELA, Ministerio de Educación y Ciencia of Spain, Xunta de Galicia (Consellería de Educación), CEADEN, Cubaenergía, Cuba, and IAEA (International Atomic Energy Agency); The Ministry of Science and Technology and the National Research Foundation (NRF), South Africa; Swedish Research Council (VR) and Knut & Alice Wallenberg Foundation (KAW); Ukraine Ministry of Education and Science; United Kingdom Science and Technology Facilities Council (STFC); The United States Department of Energy, the United States National Science Foundation, the State of Texas, and the State of Ohio.

Open Access This article is distributed under the terms of the Creative Commons Attribution Noncommercial License which permits any noncommercial use, distribution, and reproduction in any medium, provided the original author(s) and source are credited.

References

1. T. Sjöstrand, S. Mrenna, P.Z. Skands, J. High Energy Phys. **0605**, 026 (2006)
2. R. Engel, J. Ranft, S. Roesler, Phys. Rev. D **52**, 1459 (1995)
3. F. Becattini, U.W. Heinz, Z. Phys. C **76**, 269 (1997)
4. I. Kraus, J. Cleymans, H. Oeschler, K. Redlich, Phys. Rev. C **79**, 014901 (2009)
5. F. Becattini, P. Castorina, A. Milov, H. Satz, [arXiv:0912.2855](https://arxiv.org/abs/0912.2855) [hep-ph]
6. T. Sjöstrand, P.Z. Skands, Eur. Phys. J. C **39**, 129 (2005)
7. L. Evans, P. Bryant (eds.), J. Instrum. **3**, S08001 (2008)
8. K. Aamodt et al. (ALICE Collaboration), Eur. Phys. J. C **65**, 111 (2010)
9. K. Aamodt et al. (ALICE Collaboration), Eur. Phys. J. C **68**, 89 (2010)
10. K. Aamodt et al. (ALICE Collaboration), J. Instrum., **64**(3), S08002 (2008)
11. K. Aamodt et al. (ALICE Collaboration), J. Instrum. **5**, P03003 (2010)
12. J. Alme et al., Nucl. Instrum. Methods A **622**, 316 (2010)
13. A. Akhondinov et al., Eur. Phys. J. C **68**, 601 (2010) and references therein
14. B. Alessandro et al. (eds.) (ALICE Collaboration), J. Phys. G **32**, 1295 (2006)
15. K. Aamodt et al. (ALICE Collaboration), Phys. Lett. B **693**, 53 (2010)
16. K. Aamodt et al. (ALICE Collaboration), [arXiv:1101.4110](https://arxiv.org/abs/1101.4110) [hep-ex]. Submitted to Eur. Phys. J. C
17. J. Podolanski, R. Armenteros, Phil. Mag. **45**, 13 (1954)
18. K. Nakamura et al. (Particle Data Group), J. Phys. G **37**, 075021 (2010)
19. W. Blum, L. Rolandi, *Particle Detection with Drift Chambers* (Springer, Berlin, 1998)
20. M.G. Albrow et al., *Tevatron-for-LHC conference report of the QCD Working Group*, Fermilab-Conf-06-359, [hep-ph/0610012](https://arxiv.org/abs/hep-ph/0610012)
21. R. Brun et al., *GEANT3 User Guide* (CERN Data Handling Division DD/EE/841)
22. R. Brun et al., *CERN. Program Library Long Write-up*, W5013, GEANT Detector Description and Simulation Tool
23. K. Aamodt et al. (ALICE Collaboration), Phys. Rev. Lett. **105**, 072002 (2010)
24. P. Christakoglou et al., ALICE Internal Note ALICE-INT-2008-010 (2008)
25. P. Christakoglou et al., ALICE Internal Note ALICE-INT-2010-006 (2010)
26. P. Christakoglou et al. (ALICE Collaboration), PoS (EPS-HEP 2009) 405
27. G. Bendiscioli, D. Kharzeev, Riv. Nuovo Cim. **17**(6), 1 (1994)
28. B.I. Abelev et al. (STAR Collaboration), Phys. Rev. C **75**, 064901 (2007)
29. C. Tsallis, J. Statist. Phys. **52**, 479 (1988)
30. A. Moraes (ATLAS Collaboration), ATLAS Note ATL-COM-PHYS-2009-119 (2009)
31. P.Z. Skands, Contribution to the 1st international workshop on multiple partonic interactions at the LHC, Perugia, Italy, Oct. 2008, Fermilab-Conf-09-113-T. [arXiv:0905.3418](https://arxiv.org/abs/0905.3418) [hep-ph] and [arXiv:1005.3457](https://arxiv.org/abs/1005.3457) [hep-ph]
32. J. Adams et al. (STAR Collaboration), Phys. Rev. Lett. **97**, 132301 (2006)
33. M. Bourquin, J.M. Gaillard, Nucl. Phys. B **114**, 334 (1976)
34. R.E. Ansorge et al. (UA5 Collaboration), Z. Phys. C **41**, 179 (1988)
35. F. Abe et al. (CDF Collaboration), Phys. Rev. D **40**, 3791 (1989), RC
36. G.J. Alner et al. (UA5 Collaboration), Phys. Rep. **154**, 247 (1987)
37. R.E. Ansorge et al. (UA5 Collaboration), Nucl. Phys. B **328**, 36 (1989)
38. J. Adams et al. (STAR Collaboration), Phys. Lett. B **612**, 181 (2005)
39. B.I. Abelev et al. (STAR Collaboration), Phys. Rev. C **79**, 064903 (2009)
40. T. Alexopoulos et al. (E735 Collaboration), Z. Phys. C **67**, 411 (1995)
41. D. Acosta et al. (CDF Collaboration), Phys. Rev. D **72**, 052001 (2005)
42. G. Bocquet et al. (UA1 Collaboration), Phys. Lett. B **366**, 441 (1996)