

OBSERVING BEAM MOTION USING INFRARED INTERFEROMETRY *

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Abstract

Phase noise in the RF master oscillator driving synchrotron oscillations of the beam has been identified as one of the dominant sources of noise in the infrared beamline at the Advanced Light Source. We present measurements of the effect of the electron beam motion in a Fourier transform interferometer (FTIR) detector. This form of detector may be sensitive to very small beam motions.

1 INTRODUCTION

Interest in infrared (IR) synchrotron radiation has grown dramatically in the past few years. However, measurements at the synchrotron IR source have proven to be very sensitive to both beam motion and mechanical motion of the endstation[1, 2, 3]. The beam motion can result from a combination of self-excited and driven motion, ranging in frequency from a few hertz to hundreds of kilohertz. IR beamline 1.4.3 at the Advanced Light Source (ALS) located at Lawrence Berkeley National Laboratory is no exception to these problems.

Over the past year, we have made an effort to identify and resolve noise issues in the beamline as well as understand the mechanisms for observing the noise. In particular, we have made a detailed study of beam motion excited by phase noise in the master oscillator (MO) of the RF system since this was determined to be one of the dominant noise sources. This has led us to the conclusion that the interferometer used in the endstation may serve as an excellent general diagnostic for observing very small beam motion. This paper summarizes our measurements to date and provides several hypotheses for the sensitivity to beam motion. Section 2 provides a description of the IR source, beamline, and interferometer. Section 3 presents measurements of the beam motion excited by MO phase noise.

2 MEASUREMENT SETUP

The ALS is a 1.5–1.9 GeV electron storage ring optimized for producing high brightness synchrotron radiation. The synchrotron light is collected from the 1.4 bending magnet with source parameters listed in Table 1.

The light passes through a 10mrad vertical and 40mrad horizontal opening, as schematically drawn in Figure 1.

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Parameter	Description	Value
E	Beam energy	1.5–1.9 GeV
I	nominal current range	100–400 mA
$\epsilon_{x,y}$	x,y emittance@1.9 GeV	6,0.06 nm-rad
$\beta_{x,y}$	x,y beta functions	0.45,19.0 m
$\alpha_{x,y}$	x,y alpha functions	0.43,-7.7
η_x	dispersion function	4.18 cm
η'_x	dispersion slope	0.13

Table 1: ALS IR source parameters.

This light is deflected vertically by 0.5 meter, then is re-focused outside the shield wall by an ellipsoidal mirror, m2. A “switchyard” then contains a series of optics to collimate the IR beam and distribute the beam to one of three end stations.

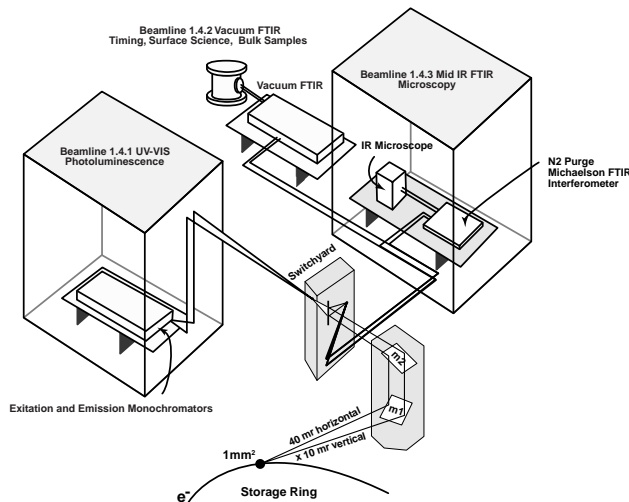


Figure 1: Schematic layout of the ALS infrared beamline.

BL 1.4.3 directs the collimated synchrotron light into a Nicolet Magna 760 FTIR bench. The modulated light is then passed through a Nic-Plan IR microscope that can perform both transmission and reflection measurements. BL 1.4.2 uses the synchrotron light for the input of a Bruker IFS 66v/S vacuum FTIR spectrometer. This IFS 66v/S instrument has a wide spectral range, 50 cm⁻¹ to 25,000 cm⁻¹, and it has step-scan capabilities in addition to rapid-scan to enable fast timing measurements. BL 1.4.1 uses the UV part of the spectrum from the 1.4 front end to do photoluminescence and related studies up to 6eV.

noisy MO at very low beam current, where we could independently determine the amplitude of synchrotron oscillations and thus the amplitude of energy oscillations at the horizontally dispersive IR source point. Shown in Fig. 6 is a plot of the FTIR noise at low current for the old and new MO. We independently determined that the amplitude of beam energy oscillations was 2.8×10^{-5} (relative to the nominal energy), corresponding the horizontal position and angle oscillations of $1.6 \mu\text{m}$ and $3.8 \mu\text{rad}$, respectively.

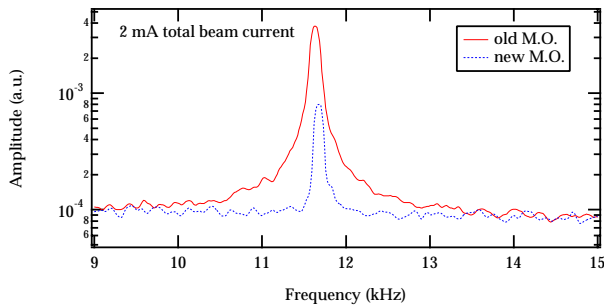


Figure 6: Motion driven by MO phase noise at low beam current for two MO sources. The reduction is equal to the reduction in the phase noise.

Given such small beam motion with a signal/noise of about 40 indicated that the interferometer was sensitive to beam motion of about 40 nm! This suggests that the interferometer may be used as a beam motion detector with extremely high sensitivity. Following are two possible noise mechanisms which we are currently studying. Consider the schematic diagram of a Michelson interferometer as shown in Fig. 7. If the incoming light is not normal to the mirror (M2), there is a path length difference in between the two arms $L1$ and $L2$ of $\Delta L \theta^2$, where ΔL is the difference $L2 - L1$ and θ is the incident light angle. Angular variations of the incident light can occur from position variations of the light incident of the ellipsoidal focussing mirror or deviations from optical flatness of any of the subsequent mirrors. A second mechanism is nonflatness of the beam splitter. In this case, transverse motion of the incident light beam will result in a path length difference in the two arms.

4 CONCLUSIONS

Phase noise in the RF master oscillator has been identified as one of the dominant sources of noise in the ALS infrared beamline. The problem has been substantially reduced by replacing the MO with a lower noise source. Interferometric measurements of synchrotron radiation are very sensitive to beam motion and may be used as a sensitive beam diagnostic. The means by which the noise is observed at the ALS is still under investigation. We would like to thank H. Zyngier for many useful discussions.

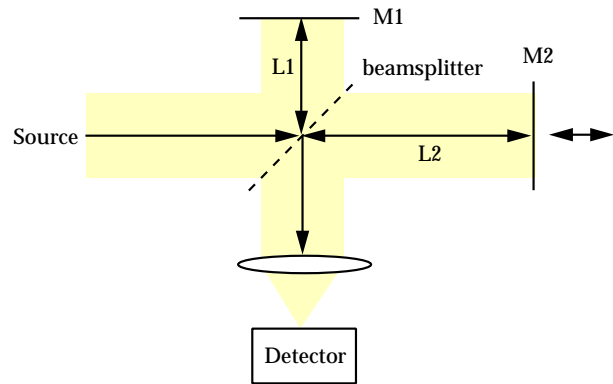


Figure 7: Schematic diagram of a Michelson interferometer. Path length differences can arise from variations of the incident light angle or position variations combined with a nonflat beamsplitter.

5 REFERENCES

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