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Trace anomaly of weyl fermions via the path integral

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ABSTRACT: We compute the trace, diffeomorphism and Lorentz anomalies of a free Weyl fermion in a gravitational background field by path integral methods. This is achieved by regularising the variation of the determinant of the Weyl operator building on earlier work by Leutwyler. The trace anomaly is found to be one half of the one of a Dirac fermion. Most importantly we establish that the potential parity-odd curvature term $R\tilde{R}$, corresponding to the Pontryagin density, vanishes. This is to the contrary of some recent findings in the literature which gave rise to a controversy. We verify, that the regularisation does not lead to (spurious) anomalies in the Lorentz and diffeomorphism symmetries. We argue that in $d=2 \pmod{4}$ P- and CP-odd terms cannot appear and that for $d=4 \pmod{4}$ they are absent at least at leading order.

KEYWORDS: Anomalies in Field and String Theories, Scale and Conformal Symmetries, Space-Time Symmetries

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1 Introduction

A classical theory free of explicit scales is invariant under scale transformation. In curved spacetime, this transformation is naturally generalised as a Weyl transformation, invariance implies that the trace of the energy-momentum tensor (EMT) vanishes, $T^{\rho}_{\ \rho}=0$, on physical states. At the quantum level this symmetry is anomalous as discovered by Capper and Duff [1]. Under the requirement of diffeomorphism (diffeo) invariance and on dimensional grounds, the trace anomaly takes on the form

$$g^{\alpha\beta}\langle T_{\alpha\beta}\rangle = a E_4 + b R^2 + c W^2 + d \square R + e R\tilde{R}, \qquad (1.1)$$

¹The formulation in terms of operators in gauge theories, which is useful in other contexts, has been worked out in refs. [2–4].

where \Box is the Laplacian, $W^2 = R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} - 2R_{\mu\nu}R^{\mu\nu} + \frac{1}{3}R^2$ is the Weyl tensor squared and $E_4 = R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} - 4R_{\mu\nu}R^{\mu\nu} + R^2$ the 4-dimensional topological Euler density see e.g. [5]. The last term is the topological (Pontryagin) density $R\tilde{R} \equiv \frac{1}{2}\epsilon^{\mu\nu\rho\sigma}R_{\alpha\beta\mu\nu}R^{\alpha\beta}_{\rho\sigma}$ which has been found to be non-vanishing in [6] and is known to satisfy the Wess-Zumino consistency condition [7]. It is analogous to the topological gauge field term $F\tilde{F} \equiv \frac{1}{2}\epsilon^{\mu\nu\rho\sigma}F_{\mu\nu}F_{\rho\sigma}$ which appears in the ABJ anomaly [8–11] and shares the same C-, P- and T- transformation properties

$$C \circ R\tilde{R} = +R\tilde{R}$$
, $P \circ R\tilde{R} = -R\tilde{R}$, $T \circ R\tilde{R} = -R\tilde{R}$, (1.2)

such that $R\tilde{R}$ is P-, CP-odd and CPT-even. Over the last few years a controversy has spanned in the literature since some authors have obtained a non-vanishing coefficient e [6, 12–18] whereas others have found it to be vanishing [19–23]. Weyl fermions are subtle and probe spacetime in their own way as in $d = 2 \pmod{4}$ they give rise to gravitational (diffeomorphism and Lorentz) anomalies [24], or more definitely to Lorentz anomalies [25, 26]. In the specific determinations of $e \neq 0$ there is something similarly unsettling in that the authors found it to be purely imaginary in Lorentzian signature. This in fact, together with (1.2), implies that its contribution is CPT-violating since $T \circ i = -i$ which would indicate a CPT anomaly. Whereas it was noted that an imaginary e would violate unitarity [6], the CPT-violation itself seems to have been overlooked. This would either mean that such theories have to be discarded [6] or supplemented by new particles such as three right-handed neutrinos in the Standard Model. This together with the fact that an $R\tilde{R}$ -term with real coefficient in the context P- and or CP-violation is of importance for phenomenological reasons (see for example the discussion in [28]), e.g. baryogenesis or gravitational waves as reviewed in [30], and could be observed experimentally [31], we consider it important to clarify the nature of this anomaly.

In this work we show that e = 0 when computed from the path integral by defining the Weyl determinant building on earlier work by Leutwyler [25, 26, 32, 33]. The key idea is that it is the variation of the determinant, which is well-defined, that enters both of our approaches i) proper time regularisation and ii) the Fujikawa method adapted to Weyl fermions. In both cases this is combined with a covariant derivative expansion (CDE) in curved spacetime [34], which is an alternative to the heat kernel, and has already proven useful to derive covariant and consistent anomalies in the context of effective field theories [35]. The diffeomorphism, the Lorentz and the Weyl anomalies are entangled by regularisation, since they can be related by local counterterms (see end of section 2 for further details); we therefore consider it essential to explicitly evaluate all three quantities.

The paper is organised as follows. In section 2 technicalities such as the Weyl determinant, the definition of the path integral and the anomalies are discussed. In sections 3 and 4 we compute anomalies applying the proper time and Fujikawa's method in conjunction with the CDE. In view of the variety of results in the literature, we make an attempt

²The parity anomaly in $QED_{d=3}$ [27] constitutes an example of an anomalous discrete symmetry. It is therefore not a priori clear that the Poyntryagin density could not be present.

³ CPT is a fundamental symmetry of quantum field theories and has been proven in the axiomatic context even in its curved space formulation [29].

to understand them in section 5. A generalisation of our results to even dimension is made in section 6. The paper ends with conclusions in section 7. Technical details are deferred to appendices A and B.

2 Technical preliminaries

2.1 The determinant of the Weyl operator

There are several challenges in defining the determinant for a Weyl fermion. For example if one starts with a Dirac fermion with only left-handed components then its associated Dirac operator $i \not \!\!\!D P_L$, where P_L projects on left-handed fermions, cannot be inverted. This problem can be avoided if one starts directly with a two-component Weyl fermion ψ_L . The Weyl operator \mathcal{D} , which is the Dirac operator acting on a Weyl fermion, is given by

$$\mathcal{D}\psi_L = i\bar{\sigma}^{\mu}(\partial + \omega_L)_{\mu}\psi_L , \qquad (2.1)$$

where $\bar{\sigma}^{\mu} = e^{\mu}_{\ a}\bar{\sigma}^{a}$, $\bar{\sigma}^{a} = (1, -\vec{\sigma})$, $e^{\mu}_{\ a}$ is the vierbein and ω_{L} is the spin-connection (A.6); more precise definitions can be found in the appendix A. Hereafter we work in Euclidian space as it is technically more convenient. The determinant of the Weyl operator appears formally in the effective action

$$W = -\log \det \mathcal{D} \,, \tag{2.2}$$

after performing the Gaussian path integral. Unfortunately, det \mathcal{D} is ill-defined, as emphasised by Álvarez-Gaumé and Witten [24], since it maps left onto right handed fermions which have different Hilbert space, i.e. $\mathcal{D}: (\frac{1}{2},0) \to (0,\frac{1}{2})$ and vice versa. This makes the phase of the functional determinant ambiguous, whereas the modulus is unaffected (and is also gauge invariant). However, the determinant itself is not an observable. That is Leutwyler and Mallik [25, 26, 32, 33] pointed out that the variation of (2.2)

$$\delta \log \det \mathcal{D} = \operatorname{Tr} \delta \mathcal{D} \mathcal{D}^{-1} , \qquad (2.3)$$

which formally holds for any operator, is well-defined since it maps fermions to fermions of the same chirality: $\delta \mathcal{D} \mathcal{D}^{-1}: (0, \frac{1}{2}) \to (0, \frac{1}{2})$. This is in line with the observation that the relative phase between two operators is well-defined [24]. In both the proper time regularisation and the Fujikawa method adapted for Weyl fermion it will be the formula (2.3) and not the determinant itself which will form the starting point of the evaluation.

The zero modes of \mathcal{D} consist in another problem for the definition of the determinant. In what follows we will assume that there are no zero modes, as is often done [33, 36], which is believed to be true in the realm of perturbation theory. On top of this the operator \mathcal{D}^{-1} is singular, as apparent from perturbation theory at short distances, but can be regularised to which we will turn to in sections 3 and 4 respectively.

2.2 Path integral formulation

For the remaining part of the paper, Fujikawa's diffeo-invariant path integral measure [37, 38] is used. It is free from spurious gravitational anomalies (i.e that can be removed by

local counterterms) in any even dimension. It has been shown for Dirac fermions [37, 38] and in this paper we show that it equally holds for Weyl fermions.

Since we will first compute the anomaly for a Dirac fermion,⁴ validating the method, we have to define the Dirac operator, in analogy to Weyl fermions in (2.1). It reads

$$\mathcal{D}\psi = i\mathcal{D}\psi = i\gamma^{\mu}(\partial + \omega)_{\mu}\psi , \qquad (2.4)$$

where the spin-connection is $\omega_{\mu} = \frac{1}{2}\omega_{\mu,ab}\Sigma^{ab}$, with $\omega_{\mu,ab}$ the tangent frame spin-connection and $\Sigma^{ab} = \frac{1}{4}[\gamma^a, \gamma^b]$. Latin indices denote the tangent space indices, whereas greek indices are the base manifold indices. The Dirac matrices in curved spacetime follow from the inverse vierbein $\gamma^{\mu} = e^{\mu}_{a}\gamma^{a}$. The covariant derivative is compatible with the vierbein, the vierbein determinant and the Dirac matrices: $D_{\mu}e^{\nu}_{a} = D_{\mu}e = D_{\mu}\gamma^{\nu} = 0$. It acts on tensors (with no spinorial indices) as

$$D_{\mu}t^{\nu} = \partial_{\mu}t^{\nu} + \Gamma^{\nu}_{\mu\rho}t^{\rho} , \qquad D_{\mu}t^{a} = \partial_{\mu}t^{a} + \omega^{a}_{\mu b}t^{b} ,$$
 (2.5)

and we follow [39] for the conventions in gravity.

Fujikawa's rescaled fermionic variables are defined by $(e = \det e^a_{\ \mu})$

$$\tilde{\psi} = \sqrt{e\psi} \;, \quad \tilde{\bar{\psi}} = \sqrt{e\bar{\psi}} \;, \tag{2.6}$$

for which the measure $\mathfrak{D}\tilde{\psi}\mathfrak{D}\tilde{\psi}$ is diffeo-invariant in $d=4\pmod{4}$. The Dirac operator associated with the adjusted spinors is

$$\mathcal{D}\tilde{\psi} = i\gamma^{\mu} \left(\partial_{\mu} + \omega_{\mu} - \frac{\partial_{\mu} \sqrt{e}}{\sqrt{e}} \right) \tilde{\psi} . \tag{2.7}$$

Similarly, the Weyl fermion invariant measure is $\mathfrak{D}\tilde{\psi}_L\mathfrak{D}\tilde{\psi}_L$, and the Weyl operator reads

$$\mathcal{D}\tilde{\psi}_L = i\bar{\sigma}^{\mu} \left(\partial_{\mu} + \omega_{\mu}^L - \frac{\partial_{\mu} \sqrt{e}}{\sqrt{e}} \right) \tilde{\psi}_L . \tag{2.8}$$

The corresponding path integrals assume the form

$$Z_{\text{Dirac}} = \int \mathfrak{D}\tilde{\bar{\psi}}\mathfrak{D}\tilde{\psi}e^{-\int d^4x\tilde{\bar{\psi}}\mathcal{D}\tilde{\psi}} , \quad Z_{\text{Weyl}} = \int \mathfrak{D}\tilde{\bar{\psi}}_L\mathfrak{D}\tilde{\psi}_Le^{-\int d^4x\tilde{\bar{\psi}}_L\mathcal{D}\tilde{\psi}_L} , \qquad (2.9)$$

where the $\tilde{\psi}$ - and $\tilde{\psi}_L$ -variables are to be treated as vierbein-independent.

2.3 Definition of the anomalies

Finally, we turn to the definition of the anomalies which we will also express in terms of the EMT. It is well-known that the anomalies follow from a variation of the effective quantum action (2.2) with respect to a symmetry. Concretely, applying $\delta_{\alpha(x)}$ to (2.2), using (2.3), the associated anomaly \mathcal{A} is formally defined by

$$\delta_{\alpha}W = -\operatorname{Tr}\delta_{\alpha}\mathcal{D}\mathcal{D}^{-1} = -\int d^4x \, e\,\alpha(x)\mathcal{A} \,.$$
 (2.10)

⁴In the absence of charges a Dirac fermion is equivalent to what is known as a vector-like fermion.

We define the three symmetry transformations, Weyl, diffeo and Lorentz, on the field variables.⁵ The Weyl variation of infinitesimal parameter $\sigma(x)$ reads

$$\delta_{\sigma}^{W} e^{\mu}{}_{a} = -\sigma e^{\mu}{}_{a} , \quad \delta_{\sigma}^{W} e = d \sigma e , \quad \delta_{\sigma}^{W} \omega_{\mu} = \frac{d-1}{2} (\partial_{\mu} \sigma) ,$$

$$\delta_{\sigma}^{W} \tilde{\psi} = \frac{1}{2} \sigma \tilde{\psi} , \qquad \delta_{\sigma}^{W} \tilde{\psi} = \frac{1}{2} \sigma \tilde{\psi} . \tag{2.11}$$

The (active) diffeo-transformation of infinitesimal parameter $\xi_{\mu}(x)$ reads

$$\delta_{\xi}^{d}\tilde{\psi} = \xi^{\mu} \partial_{\mu} \tilde{\psi} + \frac{1}{2} (\partial_{\mu} \xi^{\mu}) \tilde{\psi} , \quad \delta_{\xi}^{d} \tilde{\bar{\psi}} = \xi^{\mu} \partial_{\mu} \tilde{\bar{\psi}} + \frac{1}{2} (\partial_{\mu} \xi^{\mu}) \tilde{\bar{\psi}} ,
\delta_{\xi}^{d} e^{\mu}{}_{a} = \xi^{\nu} \partial_{\nu} e^{\mu}{}_{a} - e^{\nu}{}_{a} \partial_{\nu} \xi^{\mu} , \quad \delta_{\xi}^{d} e = (\partial_{\mu} e \xi^{\mu}) = e(D_{\mu} \xi^{\mu}) ,
\delta_{\xi}^{d} \omega_{\mu} = \xi^{\nu} \partial_{\nu} \omega_{\mu} + \omega_{\nu} \partial_{\mu} \xi^{\nu} .$$
(2.12)

The Lorentz transformation of infinitesimal parameter $\alpha_{ab}(x)$ ($\alpha_{ab} = -\alpha_{ba}$) reads

$$\delta_{\alpha}^{L}\tilde{\psi} = -\frac{1}{2}\alpha_{ab}\Sigma^{ab}\tilde{\psi} , \quad \delta_{\alpha}^{L}\tilde{\tilde{\psi}} = \frac{1}{2}\alpha_{ab}\tilde{\tilde{\psi}}\Sigma^{ab} , \quad \delta_{\alpha}^{L}e^{\mu}{}_{a} = e^{\mu}{}_{b}\alpha^{b}{}_{a} , \quad \delta_{\alpha}^{L}e = 0 ,
\delta_{\alpha}^{L}\omega_{\mu} = \frac{1}{2}[D_{\mu}, \alpha_{ab}\Sigma^{ab}] = \frac{1}{2}\Sigma^{ab}[D_{\mu}, \alpha_{ab}] .$$
(2.13)

As stated above we further define the anomalies from the path integral

$$\delta_{\alpha}W = -\frac{\int \mathcal{D}\tilde{\psi}\mathcal{D}\tilde{\bar{\psi}}\left(\delta_{\alpha}S\right)e^{-S}}{\int \mathcal{D}\tilde{\psi}\mathcal{D}\tilde{\bar{\psi}}e^{-S}} = -\langle \delta_{\alpha}S\rangle , \qquad (2.14)$$

where S is the action. Since W is a functional of the background fields only, δ_{α} acts solely on the background fields, which in the case at hand are the vierbeins. For the Weyl, the diffeo and the Lorentz transformations of respective parameters σ , ξ^{μ} and α_{ab} one obtains

$$\int d^4x \, e \, \sigma \, \mathcal{A}_{\text{trace}} = -\delta_{\sigma}^W W = \int d^4x \, e \, \sigma \, e^a_{\ \mu} \langle T^{\mu}_{\ a} \rangle \,,$$

$$\int d^4x \, e \, \xi_{\mu} \, \mathcal{A}^{\mu}_{\text{diffeo}} = -\delta_{\xi}^d W = \int d^4x \, e \, \xi^{\nu} \, \left\langle \omega_{\nu}^{\ ab} T_{ab} - D^{\mu} T_{\mu\nu} \right\rangle \,,$$

$$\int d^4x \, e \, \alpha_{ab} \, \mathcal{A}^{ab}_{\text{Lorentz}} = -\delta_{\alpha}^L W = \int d^4x \, e \, \alpha_{ab} \langle T^{ab} \rangle \,, \qquad (2.15)$$

where the EMT is obtained from $e T^{\mu}_{a} = \delta S/\delta e^{a}_{\mu}$, treating $\tilde{\psi}$, $\tilde{\psi}$ as vierbein-independent. The tensors $T_{\mu\nu}$ and T_{ab} follow from T^{μ}_{a} using the vierbein and the metric accordingly. Note that the EMT is not automatically symmetric, as it is when varying the action by the metric, since with fermions it is the vierbein that becomes the fundamental quantity. This allows for the Lorentz anomaly to be present in the first place and as a result the diffeo anomaly is not given solely by the divergence of the EMT [38, 40].

It is known that the Lorentz and diffeo anomalies are related by a local counterterm [41, 42]. However, as pointed out in [25, 26], this local counterterm is non-polynomial in the background fields and not permitted by the rules of renormalisation, which means that a

⁵We describe these transformations on the fields that occur in the Dirac case, from which we will infer the transformations involving a Weyl fermion and its Weyl operator later on.

priori one can not always transfer the Lorentz anomaly into the diffeo one and conversely. In principle it is known that in $d=4 \pmod 4$, but not $d=2 \pmod 4$, these anomalies are vanishing [24–26]. In fact in d=2 Leutwyler [25] has shown that for a Weyl fermion, there is only a Lorentz anomaly and no diffeo anomaly, and it is therefore likely that this pattern repeats for $d=2 \pmod 4$. However, in view of the controversy around the $R\tilde{R}$ -term we verify the absence of spurious Lorentz and diffeo anomalies explicitly.

3 Proper time regularisation

In order to perform a concrete computation the singular operator \mathcal{D}^{-1} , in (2.3), has to be regularised. Following [26] we use the proper time regularisation⁶

$$\mathcal{D}^{-1}|_{\Lambda} = \int_{\Lambda^{-2}}^{\infty} dt \, \mathcal{D}^{\dagger} e^{-t\mathcal{D}\mathcal{D}^{\dagger}} \,, \tag{3.1}$$

which is convergent in the infrared, that is $t \to \infty$, since $\mathcal{D}\mathcal{D}^{\dagger}$ is a manifestly positive operator, but still requires an ultraviolet regulator Λ . Inserting this expression above, the integral easily evaluates to

$$\delta \log \det \mathcal{D}|_{\Lambda} = \operatorname{Tr} \delta \mathcal{D} \int_{\Lambda^{-2}}^{\infty} dt \, \mathcal{D}^{\dagger} e^{-t\mathcal{D}\mathcal{D}^{\dagger}} = \operatorname{Tr} \delta \mathcal{D} \mathcal{D}^{-1} e^{-\frac{\mathcal{D}\mathcal{D}^{\dagger}}{\Lambda^{2}}} \,. \tag{3.2}$$

It is well-known that the anomalies are exactly marginal, that is Λ^0 -terms. The divergences in Λ^4 and Λ^2 are of no special interest and will therefore not be discussed any further. In what follows we will apply this regularisation to a Dirac and a Weyl fermion. The Dirac case is beyond doubt in the literature but serves to test and illustrate the method.

3.1 Dirac fermion

For a Dirac fermion the operator \mathcal{D} assumes the form

$$\mathcal{D} \to i e^{\mu}_{\ a} \gamma^a D_{\mu} = i \not \! D \ , \tag{3.3}$$

where D_{μ} is given by (2.7), when acting on $\tilde{\psi}$. The fact that it is hermitian $\mathcal{D}^{\dagger} = \mathcal{D}$, makes this case particularly simple since the regulator $e^{-t\mathcal{D}\mathcal{D}^{\dagger}} \to e^{-t\mathcal{D}^2}$ then commutes with both \mathcal{D} and \mathcal{D}^{-1} . The Weyl variation reads

$$\delta_{\sigma}^{W} \mathcal{D} = -\sigma \mathcal{D} - \frac{1}{2} [\mathcal{D}, \sigma] . \tag{3.4}$$

Therefore the trace anomaly is given by

$$\delta_{\sigma}^{W}W = \lim_{\Lambda \to \infty} \operatorname{Tr} \left(\sigma \mathcal{D} + \frac{1}{2} [\mathcal{D}, \sigma] \right) \mathcal{D}^{-1} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}}$$

$$= \lim_{\Lambda \to \infty} \operatorname{Tr} \left(\sigma e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}} - \frac{1}{2} \sigma [\mathcal{D}, \mathcal{D}^{-1} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}}] \right) = \lim_{\Lambda \to \infty} \operatorname{Tr} \sigma e^{-\frac{(i \not{D})^{2}}{\Lambda^{2}}},$$
(3.5)

⁶Alternatively, $\mathcal{D}^{-1}|_{\Lambda} = \int_{\Lambda^{-2}}^{\infty} dt \, e^{-t\mathcal{D}^{\dagger}\mathcal{D}} \mathcal{D}^{\dagger}$ could have been chosen.

where the cyclicity of the trace has been used. To evaluate the last term we use the CDE and obtain

$$\mathcal{A}_{\text{trace}}^{\text{Dirac}} = \frac{1}{16\pi^2} \left(\frac{1}{72} R^2 - \frac{1}{45} R_{\mu\nu} R^{\mu\nu} - \frac{7}{360} R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - \frac{1}{30} \Box R \right), \tag{3.6}$$

which is the well-known result [38, 40] first obtained in [1]. It remains to verify that there are no spurious Lorentz and diffeo anomalies. Using $\delta_{\alpha}^{L}\mathcal{D} = \frac{1}{2}[\mathcal{D}, \alpha_{ab}\Sigma^{ab}]$, the Lorentz-variation reads

$$\delta_{\alpha}^{L}W|_{\Lambda} = -\operatorname{Tr}\delta_{\alpha}^{L}\mathcal{D}\mathcal{D}^{-1}e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}} = \operatorname{Tr}\frac{1}{2}\alpha_{ab}\Sigma^{ab}[\mathcal{D},\mathcal{D}^{-1}e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}}] = 0, \qquad (3.7)$$

and thus $\mathcal{A}_{\mathrm{Lorentz}}^{ab} = 0$ follows. The diffeo-variation of \mathcal{D} reads

$$\delta_{\xi}^{d} \mathcal{D} = -[\mathcal{D}, \xi^{\mu}] \nabla_{\mu} - \xi^{\mu} [\mathcal{D}, \nabla_{\mu}] - \frac{1}{2} [\mathcal{D}, (D_{\mu} \xi^{\mu})], \qquad (3.8)$$

where ∇ is the covariant derivative deprived of spin-connection⁷

$$\nabla_{\mu} = D_{\mu} - \omega_{\mu} \ . \tag{3.9}$$

The diffeo-variation reads

$$\delta_{\xi}^{d}W|_{\Lambda} = -\operatorname{Tr} \delta_{\xi}^{d} \mathcal{D} \mathcal{D}^{-1} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}}$$

$$= -\operatorname{Tr} \left\{ \xi^{\mu} [\mathcal{D}, \nabla_{\mu} \mathcal{D}^{-1} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}}] - \xi^{\mu} [\mathcal{D}, \nabla_{\mu}] \mathcal{D}^{-1} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}} + \frac{1}{2} (\nabla_{\mu} \xi^{\mu}) [\mathcal{D}, \mathcal{D}^{-1} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}}] \right\}$$

$$= -\operatorname{Tr} \left\{ \xi^{\mu} \mathcal{D} \nabla_{\mu} \mathcal{D}^{-1} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}} - \xi^{\mu} \nabla_{\mu} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}} - \xi^{\mu} \mathcal{D} \nabla_{\mu} \mathcal{D}^{-1} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}} + \xi^{\mu} \nabla_{\mu} e^{-\frac{\mathcal{D}^{2}}{\Lambda^{2}}} \right\} = 0 .$$
(3.10)

It is noted that both the Lorentz and diffeo anomalies are vanishing prior to taking the limit $\Lambda \to \infty$. That the Lorentz and diffeo symmetry are not anomalous for Dirac fermions in any even dimension is a known result and further validates the method. For example, had we used the standard path integral measure $\mathfrak{D}\psi\mathfrak{D}\bar{\psi}$, instead of Fujikawa's (2.9), we would have obtained a non-vanishing but spurious diffeo anomaly and the wrong trace anomaly [38]. The correct one would then be obtained by adding a local counterterm that removes the spurious diffeo anomaly and leads to the correct trace anomaly. Note that no spurious Lorentz anomaly would arise since the determinant of the vierbein is Lorentz-invariant.

3.2 Weyl fermion

In this section and the remaining part of the paper, \mathcal{D} is the Weyl operator and is given by (2.8). In order to compute the variation of the Weyl operator $\delta \mathcal{D}$ it is easier to evaluate its Dirac counterpart $\delta i \not \!\!\!D$ as in the previous section and then project it in the left-right

 $^{^7\}nabla$ only contracts indices in the tangent and base spaces as in (2.5). In particular we have $D_\mu \xi^\nu = \nabla_\mu \xi^\nu$.

basis using eq. (A.8). We obtain

$$\delta_{\sigma}^{W} \mathcal{D} = -\sigma \mathcal{D} - \frac{1}{2} [\mathcal{D}, \sigma] ,$$

$$\delta_{\xi}^{d} \mathcal{D} = - [\mathcal{D}, \xi^{\mu}] \nabla_{\mu} - \xi^{\mu} [\mathcal{D}, \nabla_{\mu}] - \frac{1}{2} [\mathcal{D}, (D_{\mu} \xi^{\mu})] ,$$

$$\delta_{\alpha}^{L} \mathcal{D} = \left[\mathcal{D}, \frac{1}{2} \alpha_{ab} \mu^{ab} \right] + \frac{1}{2} \alpha_{ab} (\mu^{ab} - \lambda^{ab}) \mathcal{D} ,$$

$$(3.11)$$

where μ^{ab} and λ^{ab} are defined in appendix A, and we note that their form is equivalent to the Dirac case except for the Lorentz transformation.

Let us first verify that the diffeo and Lorentz anomalies vanish, such that the regularisation induces no spurious gravitational anomaly. The diffeo anomaly is given by

$$\delta_{\xi}^{d}W = -\lim_{\Lambda \to \infty} \operatorname{Tr} \delta_{\xi}^{d} \mathcal{D} \mathcal{D}^{-1} e^{-\frac{\mathcal{D}\mathcal{D}^{\dagger}}{\Lambda^{2}}} = -\lim_{\Lambda \to \infty} \operatorname{Tr} \left(\xi^{\mu} \nabla_{\mu} + \frac{1}{2} (D_{\mu} \xi^{\mu}) \right) \left(e^{-\frac{\mathcal{D}\mathcal{D}^{\dagger}}{\Lambda^{2}}} - e^{-\frac{\mathcal{D}^{\dagger}\mathcal{D}}{\Lambda^{2}}} \right),$$
(3.12)

where the cyclicity of the trace, eq. (A.7) and

$$\mathcal{D}^{-1}e^{-\frac{\mathcal{D}\mathcal{D}^{\dagger}}{\Lambda^{2}}}\mathcal{D} = e^{-\frac{\mathcal{D}^{\dagger}\mathcal{D}}{\Lambda^{2}}},$$
(3.13)

have been used. Noting that

$$i\not\!D = \begin{pmatrix} 0 & \mathcal{D} \\ \mathcal{D}^{\dagger} & 0 \end{pmatrix}, \qquad e^{-\frac{(i\not\!D)^2}{\Lambda^2}} = \begin{pmatrix} e^{-\frac{\mathcal{D}\mathcal{D}^{\dagger}}{\Lambda^2}} & 0 \\ 0 & e^{-\frac{\mathcal{D}^{\dagger}\mathcal{D}}{\Lambda^2}} \end{pmatrix} ,$$
 (3.14)

we can recast eq. (3.12) in Dirac space

$$\delta_{\xi}^{d}W = -\lim_{\Lambda \to \infty} \operatorname{Tr} \left(\gamma_{5} \left(\xi^{\mu} \nabla_{\mu} + \frac{1}{2} (D_{\mu} \xi^{\mu}) \right) e^{-\frac{(i \not D)^{2}}{\Lambda^{2}}} \right) . \tag{3.15}$$

A direct computation using the CDE in curved spacetime shows that it vanishes. Importantly, the diffeo anomaly may a priori not be covariant since ∇ in (3.15) is not the covariant derivative. This is where the CDE as carried out in [34] is useful since it easily allows for an expansion that is not manifestly covariant; with more details in appendix B. It is noted that the heat kernel with Seeley-DeWitt coefficients [43–45] and former CDE in curved spacetime approaches [46, 47] are designed to compute traces involving a quadratic operator, such as $\operatorname{Tr} a(x)e^{-D^2}$ where a is not a differential operator. With some work a trace of the form (3.15) can be brought to this form, but it is not straightforward and involves lengthy manipulations [26].

The Lorentz anomaly is given by

$$\delta_{\alpha}^{L}W = -\lim_{\Lambda \to \infty} \operatorname{Tr} \frac{1}{2} \alpha_{ab} \left(\mu^{ab} e^{-\frac{\mathcal{D}^{\dagger} \mathcal{D}}{\Lambda^{2}}} - \lambda^{ab} e^{-\frac{\mathcal{D} \mathcal{D}^{\dagger}}{\Lambda^{2}}} \right) , \qquad (3.16)$$

and once again it can be rewritten as a trace in Dirac space

$$\delta_{\alpha}^{L}W = \lim_{\Lambda \to \infty} \operatorname{Tr} \frac{1}{2} \alpha_{ab} \Sigma^{ab} \gamma_{5} e^{-\frac{(i\not D)^{2}}{\Lambda^{2}}}. \tag{3.17}$$

The direct computation using the CDE shows that it equally vanishes.

We now turn to the trace anomaly. Using (3.13) one obtains

$$\delta_{\sigma}^{W}W = \frac{1}{2} \lim_{\Lambda \to \infty} \operatorname{Tr} \sigma \left(e^{-\frac{\mathcal{D}^{\dagger}\mathcal{D}}{\Lambda^{2}}} + e^{-\frac{\mathcal{D}\mathcal{D}^{\dagger}}{\Lambda^{2}}} \right) = \frac{1}{2} \lim_{\Lambda \to \infty} \operatorname{Tr} \sigma e^{-\frac{(i\mathcal{D})^{2}}{\Lambda^{2}}}, \tag{3.18}$$

which is half the trace anomaly of a Dirac fermion

$$\mathcal{A}_{\text{trace}}^{\text{Weyl}} = \frac{1}{2} \mathcal{A}_{\text{trace}}^{\text{Dirac}},$$
 (3.19)

without spurious gravitational anomalies,

$$\mathcal{A}^{\mu}_{\text{diffeo}} = \mathcal{A}^{ab}_{\text{Lorentz}} = 0 \ . \tag{3.20}$$

In particular there is no Pontryagin density $R\tilde{R}$. We wish to emphasise that each term in (3.18) has an $R\tilde{R}$ -component, but it cancels in the sum of the two. This is in agreement with the computation of the heat kernel coefficient b_4 of the representation (1/2,0) of the Lorentz group [48], and we showed that the trace anomaly of a Weyl fermion is determined by $b_4(1/2,0) + b_4(0,1/2)$. We further note that the second term in (3.18) originates from the second one in eq. (3.11) which in turn is due to the spin-connection.

We note that this result has previously been obtained in a similar setting by Leutwyler and Mallik [26]. The main difference is in the evaluation of the expression in (3.15) for which they use the heat kernel which is rather laborious. In addition they use the fact that the Lorentz anomaly is not present in d=4 and do not proceed to evaluate the corresponding term. Hence we improve on their work in verifying the vanishing of the Lorentz anomaly explicitly and can do so in an economic manner.

4 The Fujikawa method adapted to Weyl fermions

The same results can be obtained adapting the path integral derivation of anomalies by Fujikawa [37, 49] (cf. also [38, 40]) to two-components Weyl fermions. As far as we know, only Dirac fermions with a projector P_L are considered in the literature, which suffer from an ill-defined path integral due to the non-invertibility of $i\not\!D P_L$. In the path integral, the anomaly arises from a non-trivial Jacobian which can be written as a fraction of determinants [35]⁸

$$J[\alpha] = e^{-\int d^4 x \, e \, \alpha(x) \mathcal{A}(x)}$$

$$= \frac{\det(\mathcal{D})}{\det(\mathcal{D} - \delta_\alpha \mathcal{D})} = \frac{1}{\det(\mathbb{1} - \delta_\alpha \mathcal{D} \mathcal{D}^{-1})} = \exp(\text{Tr}[\delta_\alpha \mathcal{D} \mathcal{D}^{-1}] + \mathcal{O}(\alpha^2)), \qquad (4.1)$$

and when expanded assumes the same form as in (2.3). This guarantees that it is well-defined, that is to say the global phase of the determinant cancels in this expression and the operator $\delta_{\alpha} \mathcal{D} \mathcal{D}^{-1}$ maps into the same Hilbert space as mentioned previously.

⁸Let us comment on the sign in det $(\mathcal{D} - \delta_{\alpha}\mathcal{D})$. In the Fujikawa approach one transforms the path integral variables only and since $\delta_{\alpha}S[\tilde{\psi},\tilde{\tilde{\psi}},e] = \delta_{\alpha}\tilde{\psi}\frac{\delta S}{\delta\tilde{\psi}} + \delta_{\alpha}\tilde{\psi}\frac{\delta S}{\delta\tilde{\psi}} + \delta_{\alpha}e^{\mu}_{\ a}\frac{\delta S}{\delta e^{\mu}_{\ a}} = 0$, this then implies that it is $-\delta_{\alpha}\mathcal{D}$, which stems from $\delta_{\alpha}e^{\mu}_{\ a}\frac{\delta S}{\delta e^{\mu}_{\ a}}$, that appears under the α-variation.

We proceed to construct the path integral measure. The Weyl operator (2.8) is not hermitian and does not have a well-defined eigenvalue problem. However, since $i\not\!D$ is hermitian and $i\not\!D$: $(\frac{1}{2},0) \oplus (0,\frac{1}{2}) \to (\frac{1}{2},0) \oplus (0,\frac{1}{2})$, the eigenvalue problem $i\not\!D\phi_n = \lambda_n\phi_n$ is well-posed. In particular $\lambda_n \in \mathbb{R}$ and the eigenfunctions $\{\phi_n\}$ define a complete orthonormal basis. These functions can be decomposed as follows

$$\phi_n = \begin{pmatrix} \phi_n^R \\ \phi_n^L \end{pmatrix} , \tag{4.2}$$

where $\{\phi_n^{R,L}\}$ define orthonormal eigenbases of right- and left-handed Weyl fermions respectively. The eigenvalue equations become

$$\mathcal{D}\phi_n^L = i\bar{\sigma}^\mu(\partial + \omega_L)_\mu \phi_n^L = \lambda_n \phi_n^R , \qquad \mathcal{D}^\dagger \phi_n^R = i\sigma^\mu(\partial + \omega_R)_\mu \phi_n^R = \lambda_n \phi_n^L . \tag{4.3}$$

We thus decompose the path integral measure into these eigenbases

$$\tilde{\psi}_L = \sqrt{e} \sum_n a_n \phi_n^L , \qquad \tilde{\bar{\psi}}_L = \sqrt{e} \sum_n \bar{b}_n \left(\phi_n^R\right)^{\dagger} , \qquad (4.4)$$

for which the measure assumes the form⁹

$$\mathfrak{D}\tilde{\psi}_L \mathfrak{D}\tilde{\bar{\psi}}_L = \prod_n \mathrm{d}a_n \mathrm{d}\bar{b}_n \ . \tag{4.5}$$

In order to evaluate the Jacobian, we need to determine how a variation δ_{α} acts on the Weyl fermion $\tilde{\psi}'_L \equiv \tilde{\psi}_L + \delta_{\alpha} \tilde{\psi}_L$ and its barred counterpart. We expand $\tilde{\psi}'_L$ and $\tilde{\psi}'_L$ into the eigenbases as in (4.4) to obtain

$$a'_{m} = \sum_{n} (\delta_{mn} + A_{mn}) a_{n} , \qquad \bar{b}'_{m} = \sum_{n} (\delta_{mn} + B_{mn}) \bar{b}_{n} .$$
 (4.6)

where A and B are of $\mathcal{O}(\alpha)$, and the resulting Jacobian of the Grassmann variables reads

$$\log J = -\operatorname{Tr} A - \operatorname{Tr} B + \mathcal{O}(\alpha^2) . \tag{4.7}$$

Using the orthonormality of the eigenbasis, we obtain the Jacobians of the transformations δ_{σ}^{W} , δ_{ξ}^{d} and δ_{α}^{L} ,

$$\log J_{\text{trace}}[\sigma] = -\int d^4x \, e^{\frac{\sigma}{2}} \sum_n \left((\phi_n^R)^{\dagger} \phi_n^R + \{R \leftrightarrow L\} \right) ,$$

$$\log J_{\text{diffeo}}[\xi^{\mu}] = -\int d^4x \, e^{-\frac{\sigma}{2}} \sum_n \left((\phi_n^R)^{\dagger} (\xi^{\mu} \nabla_{\mu} + \frac{1}{2} (D_{\mu} \xi^{\mu})) \phi_n^R - \{R \leftrightarrow L\} \right) ,$$

$$\log J_{\text{Lorentz}}[\alpha_{ab}] = -\int d^4x \, e^{-\frac{1}{2}} \alpha_{ab} \sum_n \left((\phi_n^R)^{\dagger} \lambda^{ab} \phi_n^R - \{R, \lambda \leftrightarrow L, \mu\} \right) . \tag{4.8}$$

These expression are ultraviolet divergent. They can be regularised by introducing $e^{-\frac{\lambda_n^2}{\Lambda^2}}$ into the sums, following Fujikawa. Note that J_{diffeo} and J_{Lorentz} are fully determined by

⁹This change of variable is defined up to a phase. As emphasised in section 2.1, this phase is irrelevant when dealing with the covariant form of anomalies which is the case here [38].

the zero modes whereas J_{trace} is determined by the zero modes only at lowest order of the loop expansion [38]. This means that the diffeo and Lorentz anomalies are of topological nature and determined at the one-loop level, as also argued by [24].

Using the eigenvalue equation eq. (4.3) the following replacement can be made

$$e^{-\frac{\lambda_n^2}{\Lambda^2}}\phi_n^L = e^{-\frac{\mathcal{D}^{\dagger}\mathcal{D}}{\Lambda^2}}\phi_n^L , \qquad e^{-\frac{\lambda_n^2}{\Lambda^2}}\phi_n^R = e^{-\frac{\mathcal{D}\mathcal{D}^{\dagger}}{\Lambda^2}}\phi_n^R , \tag{4.9}$$

where \mathcal{D} is the Weyl operator (2.8). As in the previous section the Jacobians can be recast in Dirac space using eq. (3.14)

$$\log J_{\text{trace}}[\sigma] = -\delta_{\sigma}^{W}W = -\frac{1}{2}\lim_{\Lambda \to \infty} \operatorname{Tr} \sigma e^{-\frac{(i\not D)^{2}}{\Lambda^{2}}},$$

$$\log J_{\text{diffeo}}[\xi] = -\delta_{\xi}^{d}W = +\frac{1}{2}\lim_{\Lambda \to \infty} \operatorname{Tr} \gamma_{5} \left(\xi^{\mu}\nabla_{\mu} + \frac{1}{2}(D_{\mu}\xi^{\mu})\right) e^{-\frac{(i\not D)^{2}}{\Lambda^{2}}} = 0,$$

$$\log J_{\text{Lorentz}}[\alpha] = -\delta_{\alpha}^{L}W = -\frac{1}{2}\lim_{\Lambda \to \infty} \operatorname{Tr} \alpha_{ab} \Sigma^{ab} \gamma_{5} e^{-\frac{(i\not D)^{2}}{\Lambda^{2}}} = 0.$$
(4.10)

These results are identical to eqs. (3.15), (3.17) and (3.18) found in the proper time regularisation and can thus be seen as a confirmation.

5 Comments on the literature

In view of the variety of methods applied and results obtained in the literature we make an attempt at understanding the results.

5.1 Approaches finding a non-vanishing RR-term

Let us first turn to the approaches finding $e \neq 0$ in eq. (1.1) which are in disagreement with our result.

• The perturbative approaches require an expansion in the metric $g_{\alpha\beta} = \eta_{\alpha\beta} + h_{\alpha\beta}$ followed by a "covariantisation" procedure, towards the end of the computation, to go beyond the linearisation. In practice this involves computing Feynman diagrams with a spectator Weyl fermion, that is $\mathcal{L} = e \, \bar{\psi} \, (i \not D P_L + i \not \partial P_R) \, \psi$, such that a propagator can be obtained by inversion. In a gauge theory this procedure has been successfully applied [41, 50] but in the context of gravity this is more subtle since the right-handed spectator seems to violate the universality of gravity (Lorentz invariance). Between the perturbative approaches the main open question is the choice of regularisation procedure. In [6, 12–17] dimensional, differential or Pauli-Villars regularisation have been employed, while verifying the vanishing of the diffeo anomaly, and a non-zero $R\tilde{R}$ -term has been found.

¹⁰In a gauge theory with a right-handed spectator, the gauge symmetry is preserved by adapting the gauge transformation $e^{i\alpha}\psi \to e^{i\alpha P_L}\psi$. However, the Lorentz transformation cannot be adapted in this manner since the transformation of the vierbein in $\partial P_R\psi = \gamma^a e^{\mu}_a \partial_{\mu} P_R\psi$ would spoil Lorentz invariance. In other words the spectator ceases to be a spectator in the context of gravity.

- The Fujikawa method was applied in [17] relying on a mass for a left-handed field $\bar{\psi}mP_L\psi$ which is known to vanish. In [18], a regularisation with a mass term $\bar{\psi}m^AP_L\psi$ such that $\{m^A, \gamma_5\} = [m^A, \gamma^\mu] = 0$ and $(m^A)^\dagger = m^A$ was introduced. We find that the only matrix satisfying these equations is $m^A = 0$ and hence the method seems contradictory.
- The heat kernel approaches were applied in [6, 15].
 - In [6] the heat kernel regulator $\exp(-\mathcal{D}^{\dagger}\mathcal{D}/\Lambda^2)$ was used where \mathcal{D} is the Weyl operator of ψ_L . This differs from (3.18), see also (4.8) (as well as eq. 4.4 in [26]), which corresponds to $\exp(-\mathcal{D}^{\dagger}\mathcal{D}/\Lambda^2) + \exp(-\mathcal{D}\mathcal{D}^{\dagger}/\Lambda^2)$ thereby explaining the difference. In our formalism the sum is a direct consequence of the variation (3.11) and the algebraic identity (3.13).¹¹
 - In [15] the metric-axial gravity (MAT) formalism was applied, where the metric is supplemented by an axial part $g_{\mu\nu} \to g_{\mu\nu} + \gamma_5 f_{\mu\nu}$, motivated by earlier work by Bardeen [11]. However, concerning the regularisation the same remarks apply as above.

5.2 Approaches finding a vanishing $R\tilde{R}$ -term

Let us discuss the approaches agreeing with our result e = 0 in eq. (1.1).

- The work of Leutwyler and Mallik [26] stands out in that it seems to be unknown to the current community. They invented the method (3.2) which we use throughout. They use point splitting and the proper time regularisation and evaluate the expression in terms of the heat kernel. We agree with their findings. As mentioned previously the difference in our paper is that we use the CDE which is more straightforward, and we also verify the vanishing of the Lorentz anomaly explicitly. We further adapt Fujikawa's method for Weyl fermions.
- In [19] Fujikawa's method was applied writing a Majorana type mass term for the Weyl fermion which then regulates the determinant. This is possible since the Weyl representation in Euclidian space is real in d=4 [24]. The problem of the determinant of the Weyl operator, which we have addressed in detail, has not been considered explicitly in that paper.¹²
- The MAT approach was applied with the Fujikawa method [20] again using a Majorana-type mass. The definition of the Weyl operator has not been addressed either.
- A non-perturbative approach employing the Hadamard regularisation was carried out in [21] working directly in the EMT. For the Weyl fermion the EMT is taken to be symmetric hence the vanishing of the Lorentz anomaly is assumed and not verified.

¹¹One may be tempted to think that regularising with $\exp\left(-\mathcal{D}^{\dagger}\mathcal{D}/\Lambda^{2}\right)$ only would be correct since it maps left- onto left-handed fermions. As emphasised earlier the \mathcal{D} -operator maps from one chirality into the other and this indicates that both $\exp\left(-\mathcal{D}^{\dagger}\mathcal{D}/\Lambda^{2}\right)$ and $\exp\left(-\mathcal{D}\mathcal{D}^{\dagger}/\Lambda^{2}\right)$ ought to appear in the regularisation.

¹²In [13] doubts were expressed that introducing a Majorana mass is legitimate. It remains however unclear to us why the Majorana mass term as first proposed in [24] should not be legitimate.

• In another series of perturbative approaches dimensional regularisation [22] (along with the Breitenlohner-Maison γ_5 -scheme) have been applied. These computations are further reviewed and clarified in [23]. Diffeomorphism invariance is verified assuming that there is no Lorentz anomaly.

The main mystery remains why some perturbative approaches obtain $e \neq 0$ and others e = 0, even though the vanishing of the diffeo anomaly at lowest order of the metric expansion is verified in each case. It either means that there is an error in one of the computations, which tend to be lengthy, or that there is a specific problem with some of the regularisations. As for the non-perturbative approaches, only Leutwyler and Mallik verify the vanishing of the diffeo anomaly in [25, 26] (but not the Lorentz anomaly), whereas the other papers do not address them at all.

6 Dimensions other than d=4

In d=4 we concluded that for free Weyl fermions the $R\tilde{R}$ -term is absent in the trace anomaly. It is therefore natural to ask whether P- and or CP-odd terms could be present in any other even dimension.¹³ In fact up to eqs. (3.15), (3.17) and (3.18) the expressions are independent of the dimension in our approach. In particular the factor $\frac{1}{2}$ in (3.18) makes it clear that the trace anomaly of a Weyl fermion remains half the trace anomaly of a Dirac fermion in any even dimensional spacetime. Even in the absence of concrete computations one can make interesting observations and come to the same conclusion:

- It is important to distinguish $d=2\pmod 4$ and $d=4\pmod 4$ as their (Euclidean) Weyl representation are complex and real respectively [24]. For example, Lorentz and diffeo anomalies cannot arise in $d=4\pmod 4$ dimensions since the reality of the representation allows for a Pauli-Villars regulator mass term which is symmetry-preserving and this means that no diffeo and Lorentz anomaly can appear [24]. Another way to look at it is that in $d=4\pmod 4$, unlike in $d=2\pmod 4$, the CPT operation flips the chirality such that the Weyl fermions effectively look vector-like [24] (and [52] for more detail).
- In $d=2 \pmod{4}$ a P- and CP-odd term should not violate CPT since the appearance of the factor of i is dimension-dependent. For example in Minkowski space one has $\text{Tr}[\gamma_{\alpha}\gamma_{\beta}\gamma_{5}] = 2\epsilon_{\alpha\beta}$ in d=2 whereas $\text{Tr}[\gamma_{\alpha}\gamma_{\beta}\gamma_{\gamma}\gamma_{\delta}\gamma_{5}] = 4i\epsilon_{\alpha\beta\gamma\delta}$ in d=4, and this is related to the complex and real representations mentioned above. Hence in $d=2 \pmod{4}$ we would not expect an imaginary prefactor and thus no CPT-violation.
- However, one can argue that in $d = 2 \pmod{4}$ one cannot write down a P- or CP-odd diffeo-invariant scalar of mass dimension d (which is relevant for the trace anomaly). Firstly, let us note that due to the Bianchi identity, $\epsilon^{\dots\mu\nu\rho}R_{\alpha\mu\nu\rho} = 0$ holds in any

¹³Odd dimensional spaces are beyond the scope of this paper and we refer the reader to [51] for intricate relations with the even dimensional case.

dimension, that is to say the Levi-Civita tensor cannot contract more than two indices of a Riemann tensor. For concreteness let us first focus on d = 6. Using the Bianchi identities one can show that a parity-odd operator of mass dimension 6 can only be of the form,

$$\epsilon^{\alpha_1...\alpha_6} R_{\alpha_1\alpha_2..} R_{\alpha_3\alpha_4..} R_{\alpha_5\alpha_6..} . \tag{6.1}$$

The only way to contract these Riemann tensors together is

$$\epsilon^{\alpha_1...\alpha_6} R_{\alpha_1\alpha_2\mu}{}^{\nu} R_{\alpha_3\alpha_4\rho\nu} R_{\alpha_5\alpha_6}{}^{\rho\mu} = 0 , \qquad (6.2)$$

which vanishes since $\epsilon^{\alpha_1...\alpha_6}R_{\alpha_1\alpha_2\mu}{}^{\nu}R_{\alpha_3\alpha_4\rho\nu}$ is symmetric under $\mu\leftrightarrow\rho$ whereas $R_{\alpha_5\alpha_6}{}^{\rho\mu}$ is antisymmetric. On the other hand, in d=8 for example, there is an even number of Riemann tensors, which can then be contracted in pairs

$$\epsilon^{\alpha_1 \dots \alpha_8} R_{\alpha_1 \alpha_2 \mu \nu} R_{\alpha_3 \alpha_4}^{\ \mu \nu} R_{\alpha_5 \alpha_6 \rho \sigma} R_{\alpha_7 \alpha_8}^{\ \rho \sigma} \neq 0. \tag{6.3}$$

This generalises straightforwardly to any even dimension

$$\epsilon^{\alpha_1...\alpha_{2n}} R_{\alpha_1\alpha_2..} \dots R_{\alpha_{2n-1}\alpha_{2n}..}$$

$$\begin{cases}
= 0 \ d = 2 \ (\text{mod } 4) \\
\neq 0 \ d = 4 \ (\text{mod } 4)
\end{cases}$$
(6.4)

since it involves an odd number of Riemann tensors in $d = 2 \pmod{4}$ but an even number in $d = 4 \pmod{4}$. In fact, the absence of such a term in d = 6 has been inferred from a cohomology-type argument given a long time ago [7].

In summary the absence of parity-odd terms can be established in $d = 2 \pmod{4}$ without explicit computation whereas in $d = 4 \pmod{4}$ a computation is required.

7 Conclusion

We investigated the trace anomaly of a free Weyl fermion in a curved space which gave rise to some controversy recently as some authors found a term proportional to the Pontryagin density $R\tilde{R}$, as reviewed in [16]. The results found in the literature for d=4, as we point out, do not only violate P and CP but also CPT because of the purely imaginary prefactor. Handling Weyl fermions is technically subtle and requires care as the Weyl determinant is ill-defined. We use the method of proper time regularisation building on earlier work by Leutwyler (section 3) and develop the Fujikawa method for a two-component Weyl fermion (section 4), where it is not the Weyl determinant but its well-defined variation which is used. This allows us to perform the computation in a rather compact manner from which our main result follows: the trace anomaly of the Weyl fermion is half the one of a Dirac fermion and therefore no Pontryagin density $R\tilde{R}$ is present. We have further checked that the diffeomorphism and the Lorentz anomalies are absent which is a consistency check on our regularisation since these anomalies are known to be absent in $d = 4 \pmod{4}$ [24]. In section 6 we have concluded that no parity-odd term can appear in any even dimension for a free Weyl fermion. In section 5 we have attempted to compare to the literature and thereby hope that our findings are helpful in settling the controversy. An investigation

of the anomalies from a completely different viewpoint is postponed to a forthcoming paper [53].

Our findings do not mean that a $R\tilde{R}$ -term could not play a role in fundamental physics. It can arise from sources other than Weyl fermions as studied in extensions of gravity [30], it may appear in conjunction with axions [54, 55] or more generally in connection with P-and CP-violation e.g. [56]. Further exploration is left to future work.

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A Weyl fermions in curved spacetime

The Dirac matrices can be expressed as

$$\gamma^a = \begin{pmatrix} 0 & \bar{\sigma}^a \\ \sigma^a & 0 \end{pmatrix}, \qquad \gamma_5 = \begin{pmatrix} \mathbb{1}_2 & 0 \\ 0 & -\mathbb{1}_2 \end{pmatrix} , \tag{A.1}$$

where $\sigma^a=(\mathbb{1}_2,\sigma^i)$, $\bar{\sigma}^a=(\mathbb{1}_2,-\sigma^i)$, and $\sigma^{i=1,2,3}$ are the Pauli matrices. As for the Dirac matrices, we have $\sigma^\mu=e^\mu{}_a\sigma^a$. In an Euclidian space, it is possible to choose a representation of the σ -matrices such that $\bar{\sigma}^\mu=(\sigma^\mu)^\dagger$, i.e $(\gamma^\mu)^\dagger=\gamma^\mu$ which is the convention used throughout. From $\{\gamma^\mu,\gamma^\nu\}=2g^{\mu\nu}\mathbb{1}_4$ one deduces

$$\sigma^{\mu}\bar{\sigma}^{\nu} + \sigma^{\nu}\bar{\sigma}^{\mu} = \bar{\sigma}^{\mu}\sigma^{\nu} + \bar{\sigma}^{\nu}\sigma^{\mu} = 2g^{\mu\nu}\mathbb{1}_2. \tag{A.2}$$

The generator of rotations can be written as

$$\Sigma^{ab} = \frac{1}{4} [\gamma^a, \gamma^b] = \begin{pmatrix} \lambda^{ab} & 0 \\ 0 & \mu^{ab} \end{pmatrix} , \tag{A.3}$$

where

$$\lambda^{ab} = \frac{1}{4}(\bar{\sigma}^a\sigma^b - \bar{\sigma}^b\sigma^a), \qquad \mu^{ab} = \frac{1}{4}(\sigma^a\bar{\sigma}^b - \sigma^b\bar{\sigma}^a) \ . \tag{A.4}$$

Using $D_{\mu}\psi = (\partial_{\mu} + \frac{1}{2}\omega_{\mu,ab}\Sigma^{ab})\psi$ and decomposing ψ and D_{μ} as¹⁴

$$\psi = \begin{pmatrix} \psi_R \\ \psi_L \end{pmatrix}, \qquad D_\mu = \begin{pmatrix} D_\mu^R & 0 \\ 0 & D_\mu^L \end{pmatrix} , \tag{A.5}$$

 $^{^{14}}D_{\mu}$ is diagonal since it does not change the chirality of fermions.

one obtains

$$D_{\mu}^{R}\psi_{R} = (\partial_{\mu} + \omega_{\mu}^{R})\psi_{R} = \left(\partial_{\mu} + \frac{1}{2}\omega_{\mu,ab}\lambda^{ab}\right)\psi_{R},$$

$$D_{\mu}^{L}\psi_{L} = (\partial_{\mu} + \omega_{\mu}^{L})\psi_{L} = \left(\partial_{\mu} + \frac{1}{2}\omega_{\mu,ab}\mu^{ab}\right)\psi_{L}.$$
(A.6)

The Weyl operator, used in the main text, can be identified from the second line. Note that when acting on a scalar with Lorentz index such as ξ^{μ} one has $(D_{\mu}^{R}\xi^{\nu}) = (D_{\mu}^{L}\xi^{\nu}) = (D_{\mu}\xi^{\nu}) = (\partial_{\mu}\xi^{\nu}) + \Gamma_{\mu\rho}^{\nu}\xi^{\rho}$.

From the compatibility with Dirac matrices $[D_{\mu}, \gamma^{\nu}] = 0$ one obtains

$$D_{\mu}^{R}\bar{\sigma}^{\nu} = \bar{\sigma}^{\nu}D_{\mu}^{L}, \qquad D_{\mu}^{L}\sigma^{\nu} = \sigma^{\nu}D_{\mu}^{R}.$$
 (A.7)

Finally, since

$$i\not\!D = \begin{pmatrix} 0 & i\bar{\sigma}^{\mu}D^{L}_{\mu} \\ i\sigma^{\mu}D^{R}_{\mu} & 0 \end{pmatrix} , \qquad (A.8)$$

is hermitian in Euclidian space, it follows that

$$\left(i\bar{\sigma}^{\mu}D_{\mu}^{L}\right)^{\dagger} = i\sigma^{\mu}D_{\mu}^{R} \,. \tag{A.9}$$

B Covariant derivative expansion — Computations

In this appendix, we outline the computation of the anomalies of a Weyl fermion using the CDE in curved spacetime [34].

Let us first note that we regularised the functional traces using the function e^{-x} in both sections. The result is however independent of this choice. In fact, any smooth function f such that f(0) = 1 and $x^n f^{(n)}(x) \to 0$ for $x \to \infty$ and for all $n \ge 0$ can be used instead. This is well-known for Fujikawa's approach [38] and we establish it here for Leutwyler's approach. For a function f satisfying the criteria above the following equation holds

$$\mathcal{D}^{-1} = \mathcal{D}^{\dagger} \int_0^{\infty} dt \, f'(t \mathcal{D} \mathcal{D}^{\dagger}) , \qquad (B.1)$$

for which (3.1) is a special case.

and Λ^4 -divergences are subtracted.

Having established this universality in the function f we turn to the CDE for which it is convenient to use f(x) = 1/(1+x). We thus consider the following functional traces¹⁵

$$T_{1}[a] = \lim_{\Lambda \to \infty} \operatorname{Tr} \left[a(x) f\left(\frac{(i\cancel{D})^{2}}{\Lambda^{2}}\right) \right] = \lim_{\Lambda \to \infty} \operatorname{Tr} \left[a(x) \frac{\Lambda^{2}}{-\cancel{D}^{2} + \Lambda^{2}} \right],$$

$$T_{2}[b] = \lim_{\Lambda \to \infty} \operatorname{Tr} \left[b^{\mu}(x) \nabla_{\mu} f\left(\frac{(i\cancel{D})^{2}}{\Lambda^{2}}\right) \right] = \lim_{\Lambda \to \infty} \operatorname{Tr} \left[b^{\mu}(x) \nabla_{\mu} \frac{\Lambda^{2}}{-\cancel{D}^{2} + \Lambda^{2}} \right],$$
(B.2)

where a(x) and $b^{\mu}(x)$ are local functions. As in the main text it is understood that the Λ^2 -

The interval of the Dirac matrices such that $(\gamma^{\mu})^{\dagger} = -\gamma^{\mu}$ in Euclidean space, we would obtain $f\left(\frac{\cancel{p}^2}{\Lambda^2}\right)$ instead of $f\left(\frac{(i\cancel{p})^2}{\Lambda^2}\right) = f\left(\frac{\cancel{p}^2}{-\Lambda^2}\right)$. The anomaly being of order Λ^0 does not depend on that choice.

B.1 Computation of the trace and Lorentz anomalies

From T_1 the trace and the Lorentz anomalies of a Weyl fermion follow

$$\delta_{\sigma}^{W}W = \frac{1}{2}T_{1}\left[\sigma\right], \quad \delta_{\alpha}^{L}W = \frac{1}{2}T_{1}\left[\alpha_{ab}\Sigma^{ab}\gamma_{5}\right]. \tag{B.3}$$

The functional trace can be recast in momentum space as

$$T_{1}[a] = \lim_{\Lambda \to \infty} \int d^{4}x \frac{d^{4}q}{(2\pi)^{4}} \operatorname{tr} a(x) \Lambda^{2} \frac{1}{-(\cancel{D} + i\cancel{q})^{2} + \Lambda^{2}}$$

$$= -\lim_{\Lambda \to \infty} \int d^{4}x \frac{d^{4}q}{(2\pi)^{4}} \operatorname{tr} a(x) (-\Lambda^{2}) \sum_{n \geq 0} \left[\Delta (D^{2} + g^{\mu\nu} \{D_{\mu}, iq_{\nu}\} + \Sigma \cdot F) \right]^{n} \Delta ,$$
(B.4)

where $\Delta = 1/(q^2 + \Lambda^2)$, $\Sigma^{\mu\nu} = [\gamma^{\mu}, \gamma^{\nu}]/4$, and tr denotes the trace in internal space. Note that we can maintain $\partial_{\mu} q_{\nu} = 0$, but contrary to the CDE in flat spacetime we have $D_{\mu}q_{\nu} = -\Gamma^{\rho}_{\mu\nu}q_{\rho} \neq 0$ and $[D_{\mu}, \Delta] \neq 0$ [34]. F is the fermion field strength due to the spin-connection such that for any fermion $\tilde{\psi}$ we have $[D_{\mu}, D_{\nu}]\tilde{\psi} = F_{\mu\nu}\tilde{\psi}$ with 16

$$F_{\mu\nu} = \frac{1}{4} \gamma^{\rho} \gamma^{\sigma} R_{\mu\nu\rho\sigma} , \qquad \Sigma \cdot F = -\frac{R}{4} \mathbb{1}_{Dirac} . \tag{B.5}$$

The expansion is then carried out with the help of Mathematica and the package xAct [57]. The result reads

$$T_1[a] = \frac{1}{16\pi^2} \int d^4x \operatorname{tr} a(x) \left\{ -\frac{1}{6} \Box (\Sigma \cdot F) - \frac{1}{12} F^2 - \frac{1}{72} R^2 + \frac{1}{180} R_{\mu\nu} R^{\mu\nu} - \frac{1}{180} R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - \frac{1}{30} \Box R - \frac{1}{6} R \Sigma \cdot F - \frac{1}{2} (\Sigma \cdot F)^2 \right\},$$
(B.6)

from which, by using (B.3) and (B.5), the trace anomaly and the vanishing of the Lorentz anomaly follows.

In fact once we know, from (B.6), that the Lorentz anomaly is covariant we can infer its vanishing in yet another way in d=4. It can be shown, using intrinsically 4-dimensional identities [58–60], that the only parity-odd (covariant) 2-tensor of mass dimension 4 is $g^{\mu\nu}R\tilde{R}$, which is symmetric in its indices and thus vanishes when contracted with the Lorentz parameter $\alpha_{\mu\nu}$. In particular, one has in d=4

$$\frac{1}{2}g^{\alpha\beta}R\tilde{R} = \epsilon^{\alpha\nu\rho\sigma}R^{\beta}_{\ \nu}{}^{\lambda\chi}R_{\rho\sigma\lambda\chi} = \epsilon^{\mu\nu\rho\sigma}R^{\alpha\lambda}_{\ \mu\nu}R^{\beta}_{\ \lambda\rho\sigma} , \qquad (B.7)$$

and

$$\epsilon^{\alpha\nu\rho\sigma}R^{\beta\lambda}_{\rho\sigma}R_{\nu\lambda} = 0 \; , \tag{B.8}$$

for the Ricci-tensor contraction. Every other parity-odd 2-tensor of mass dimension 4 is related to these by Bianchi identities and algebra.

 $^{^{16}}$ Note that ψ and $\tilde{\psi}$ have the same field strength.

B.2 Computation of the diffeomorphism anomaly

The diffeo anomaly is given by

$$\delta_{\xi}^{d}W = -T_{2}[\xi\gamma_{5}] - \frac{1}{2}T_{1}[(D_{\mu}\xi^{\mu})\gamma_{5}], \qquad (B.9)$$

a sum of a T_1 - and a T_2 -term (B.2). Let us focus on the T_1 -term first. Using (B.5) the only term that is non-vanishing under the Dirac trace in $T_1[(D_\mu \xi^\mu)\gamma_5]$ is

$$T_1[(D_{\mu}\xi^{\mu})\gamma_5] = \frac{-1}{16\pi^2} \int d^4x \, e(D_{\mu}\xi^{\mu}) \operatorname{tr} \gamma_5 \left(-\frac{1}{12}F^2\right)$$

$$= \frac{-1}{16\pi^2} \int d^4x \, e\xi^{\mu} \operatorname{tr} \gamma_5 \left(\frac{1}{12}[D_{\mu}, F^2]\right) ,$$
(B.10)

where integration by parts was applied using the fact that $\xi^{\mu}(x)$ vanishes at infinity.

Finally, let us turn to the T_2 -term which is less straightforward. In eq. (B.2), it is convenient to rewrite $\nabla = D - \omega$, such that

$$T_2[b] = \lim_{\Lambda \to \infty} \operatorname{Tr} \left[b^{\mu}(x) D_{\mu} \frac{\Lambda^2}{-\cancel{D}^2 + \Lambda^2} \right] - \lim_{\Lambda \to \infty} \operatorname{Tr} \left[b^{\mu}(x) \omega_{\mu} \frac{\Lambda^2}{-\cancel{D}^2 + \Lambda^2} \right] . \tag{B.11}$$

We can notice that the second term is the Lorentz anomaly with $\alpha_{ab} = b^{\mu}\omega_{\mu ab}$ (with $b^{\mu} \propto \gamma_5$), hence vanishes as we just verified. We are left with

$$T_{2}[b] = \lim_{\Lambda \to \infty} \text{Tr} \left[b^{\mu}(x) D_{\mu} \frac{\Lambda^{2}}{-\not D^{2} + \Lambda^{2}} \right]$$

$$= \lim_{\Lambda \to \infty} \int d^{4}x \frac{d^{4}q}{(2\pi)^{4}} \text{tr } b^{\mu}(x) (D_{\mu} + iq_{\mu}) (-\Lambda^{2}) \sum_{n > 0} \left[\Delta (D^{2} + g^{\mu\nu} \{D_{\mu}, iq_{\nu}\} + \Sigma \cdot F) \right]^{n} \Delta.$$
(B.12)

The computation must not be carried out in a manifestly covariant manner, and one cannot use a specific choice of coordinate (for example Riemann normal coordinates). Since the diffeo anomaly involves $b^{\mu}(x) = \xi^{\mu}(x)\gamma_5$, the computation can be simplified using $\operatorname{Tr} \gamma_5 = \operatorname{Tr} \gamma^{\mu} \gamma^{\nu} \gamma_5 = 0$, and noticing that the only source of Dirac matrices are in the covariant derivatives via the spin-connection. We finally obtain a very compact result

$$T_2[\xi \gamma_5] = \frac{-1}{16\pi^2} \int d^4x \, e \, \xi^\mu \, \text{tr} \, \gamma_5 \left(-\frac{1}{24} [D_\mu, F^2] \right) , \qquad (B.13)$$

as every other term vanishes by lack of Dirac matrices.

Using (B.9), (B.10) and (B.13) we see the canceling of terms and finally conclude that the diffeo anomaly of a Weyl fermion vanishes.

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References

- [1] D.M. Capper and M.J. Duff, Trace anomalies in dimensional regularization, Nuovo Cim. A 23 (1974) 173 [INSPIRE].
- [2] S.L. Adler, J.C. Collins and A. Duncan, Energy-Momentum-Tensor Trace Anomaly in Spin 1/2 Quantum Electrodynamics, Phys. Rev. D 15 (1977) 1712 [INSPIRE].
- [3] J.C. Collins, A. Duncan and S.D. Joglekar, *Trace and Dilatation Anomalies in Gauge Theories*, *Phys. Rev. D* **16** (1977) 438 [INSPIRE].
- [4] N.K. Nielsen, The Energy Momentum Tensor in a Nonabelian Quark Gluon Theory, Nucl. Phys. B 120 (1977) 212 [INSPIRE].
- [5] M.J. Duff, Twenty years of the Weyl anomaly, Class. Quant. Grav. 11 (1994) 1387[hep-th/9308075] [INSPIRE].
- [6] L. Bonora, S. Giaccari and B. Lima de Souza, *Trace anomalies in chiral theories revisited*, JHEP 07 (2014) 117 [arXiv:1403.2606] [INSPIRE].
- [7] L. Bonora, P. Pasti and M. Bregola, Weyl cocycles, Class. Quant. Grav. 3 (1986) 635[INSPIRE].
- [8] J.S. Bell and R. Jackiw, A PCAC puzzle: $\pi^0 \to \gamma \gamma$ in the σ model, Nuovo Cim. A **60** (1969) 47 [INSPIRE].
- [9] S.L. Adler and W.A. Bardeen, Absence of higher order corrections in the anomalous axial vector divergence equation, Phys. Rev. 182 (1969) 1517 [INSPIRE].
- [10] S.L. Adler, Axial vector vertex in spinor electrodynamics, Phys. Rev. 177 (1969) 2426 [INSPIRE].
- [11] W.A. Bardeen, Anomalous Ward identities in spinor field theories, Phys. Rev. 184 (1969) 1848 [INSPIRE].
- [12] L. Bonora, A.D. Pereira and B. Lima de Souza, Regularization of energy-momentum tensor correlators and parity-odd terms, JHEP 06 (2015) 024 [arXiv:1503.03326] [INSPIRE].
- [13] L. Bonora et al., Axial gravity, massless fermions and trace anomalies, Eur. Phys. J. C 77 (2017) 511 [arXiv:1703.10473] [INSPIRE].
- [14] L. Bonora et al., Pontryagin trace anomaly, EPJ Web Conf. 182 (2018) 02100 [INSPIRE].
- [15] L. Bonora et al., Axial gravity: a non-perturbative approach to split anomalies, Eur. Phys. J. C 78 (2018) 652 [arXiv:1807.01249] [INSPIRE].
- [16] L. Bonora, Elusive anomalies, EPL 139 (2022) 44001 [arXiv:2207.03279] [INSPIRE].
- [17] C.-Y. Liu, Investigation of Pontryagin trace anomaly using Pauli-Villars regularization, Nucl. Phys. B 980 (2022) 115840 [arXiv:2202.13893] [INSPIRE].
- [18] C.-Y. Liu, The trace anomaly for a chiral fermion, arXiv:2304.06507 [INSPIRE].
- [19] F. Bastianelli and R. Martelli, On the trace anomaly of a Weyl fermion, JHEP 11 (2016) 178 [arXiv:1610.02304] [INSPIRE].
- [20] F. Bastianelli and M. Broccoli, Axial gravity and anomalies of fermions, Eur. Phys. J. C 80 (2020) 276 [arXiv:1911.02271] [INSPIRE].
- [21] M.B. Fröb and J. Zahn, Trace anomaly for chiral fermions via Hadamard subtraction, JHEP 10 (2019) 223 [arXiv:1904.10982] [INSPIRE].

- [22] S. Abdallah, S.A. Franchino-Viñas and M.B. Fröb, Trace anomaly for Weyl fermions using the Breitenlohner-Maison scheme for γ_* , JHEP **03** (2021) 271 [arXiv:2101.11382] [INSPIRE].
- [23] S. Abdallah, S.A. Franchino-Viñas and M.B. Fröb, Trace anomalies for Weyl fermions: too odd to be true?, J. Phys. Conf. Ser. 2531 (2023) 012004 [arXiv:2304.08939] [INSPIRE].
- [24] L. Alvarez-Gaume and E. Witten, *Gravitational Anomalies*, *Nucl. Phys. B* **234** (1984) 269 [INSPIRE].
- [25] H. Leutwyler, Gravitational anomalies: a soluble two-dimensional model, Phys. Lett. B 153 (1985) 65 [Erratum ibid. 155 (1985) 469] [INSPIRE].
- [26] H. Leutwyler and S. Mallik, Gravitational anomalies, Z. Phys. C 33 (1986) 205 [INSPIRE].
- [27] A.N. Redlich, Gauge Noninvariance and Parity Violation of Three-Dimensional Fermions, Phys. Rev. Lett. **52** (1984) 18 [INSPIRE].
- [28] Y. Nakayama, CP-violating CFT and trace anomaly, Nucl. Phys. B 859 (2012) 288 [arXiv:1201.3428] [INSPIRE].
- [29] S. Hollands and R.M. Wald, Axiomatic quantum field theory in curved spacetime, Commun. Math. Phys. 293 (2010) 85 [arXiv:0803.2003] [INSPIRE].
- [30] S. Alexander and N. Yunes, Chern-Simons Modified General Relativity, Phys. Rept. 480 (2009) 1 [arXiv:0907.2562] [INSPIRE].
- [31] C.-S. Chu and R.-X. Miao, Chiral current induced by torsional Weyl anomaly, Phys. Rev. B 107 (2023) 205410 [arXiv:2210.01382] [INSPIRE].
- [32] H. Leutwyler, On the determinant of the Weyl operator, BUTP-84/33-BERN (1984) [INSPIRE].
- [33] H. Leutwyler, Chiral fermion determinants and their anomalies, Phys. Lett. B 152 (1985) 78 [INSPIRE].
- [34] R. Larue and J. Quevillon, The universal one-loop effective action with gravity, JHEP 11 (2023) 045 [arXiv:2303.10203] [INSPIRE].
- [35] B. Filoche, R. Larue, J. Quevillon and P.N.H. Vuong, Anomalies from an effective field theory perspective, Phys. Rev. D 107 (2023) 025017 [arXiv:2205.02248] [INSPIRE].
- [36] E. Witten and K. Yonekura, Anomaly Inflow and the η-Invariant, in the proceedings of the The Shoucheng Zhang Memorial Workshop, (2019) [arXiv:1909.08775] [INSPIRE].
- [37] K. Fujikawa, Energy Momentum Tensor in Quantum Field Theory, Phys. Rev. D 23 (1981) 2262 [INSPIRE].
- [38] K. Fujikawa and H. Suzuki, *Path integrals and quantum anomalies*, (2004) [DOI:10.1093/acprof:oso/9780198529132.001.0001] [INSPIRE].
- [39] C.W. Misner, K.S. Thorne and J.A. Wheeler, *Gravitation*, W. H. Freeman, San Francisco (1973) [INSPIRE].
- [40] R.A. Bertlmann, Anomalies in quantum field theory, (1996) [INSPIRE].
- [41] L. Alvarez-Gaume and P.H. Ginsparg, *The Topological Meaning of Nonabelian Anomalies*, *Nucl. Phys. B* **243** (1984) 449 [INSPIRE].
- [42] W.A. Bardeen and B. Zumino, Consistent and Covariant Anomalies in Gauge and Gravitational Theories, Nucl. Phys. B 244 (1984) 421 [INSPIRE].

- [43] J.S. Schwinger, On gauge invariance and vacuum polarization, Phys. Rev. 82 (1951) 664 [INSPIRE].
- [44] B.S. DeWitt, Quantum Theory of Gravity. 1. The Canonical Theory, Phys. Rev. 160 (1967) 1113 [INSPIRE].
- [45] B.S. DeWitt, Quantum Theory of Gravity. 2. The Manifestly Covariant Theory, Phys. Rev. 162 (1967) 1195 [INSPIRE].
- [46] P. Binetruy and M.K. Gaillard, The Leading Divergent Part of the Effective Action for the Nonlinear σ Model in n-dimensions, Nucl. Phys. B 312 (1989) 341 [INSPIRE].
- [47] R. Alonso, A covariant momentum representation for loop corrections in gravity, JHEP 05 (2020) 131 [arXiv:1912.09671] [INSPIRE].
- [48] S.M. Christensen and M.J. Duff, New Gravitational Index Theorems and Supertheorems, Nucl. Phys. B 154 (1979) 301 [INSPIRE].
- [49] K. Fujikawa, Path Integral Measure for Gauge Invariant Fermion Theories, Phys. Rev. Lett. 42 (1979) 1195 [INSPIRE].
- [50] L. Alvarez-Gaume and P.H. Ginsparg, The Structure of Gauge and Gravitational Anomalies, Annals Phys. 161 (1985) 423 [Erratum ibid. 171 (1986) 233] [INSPIRE].
- [51] L. Alvarez-Gaume, S. Della Pietra and G.W. Moore, Anomalies and Odd Dimensions, Annals Phys. 163 (1985) 288 [INSPIRE].
- [52] L. Alvarez-Gaume, An introduction to anomalies, HUTP-85/A092 (1985) [INSPIRE].
- [53] R. Larue, J. Quevillon and R. Zwicky, *Gravity-gauge Anomaly Constraints on the Energy-momentum Tensor*, forthcoming (2023).
- [54] G. Dvali, S. Folkerts and A. Franca, How neutrino protects the axion, Phys. Rev. D 89 (2014) 105025 [arXiv:1312.7273] [INSPIRE].
- [55] R. Alonso and A. Urbano, Wormholes and masses for Goldstone bosons, JHEP **02** (2019) 136 [arXiv:1706.07415] [INSPIRE].
- [56] S. Deser, M.J. Duff and C.J. Isham, *Gravitationally induced CP effects*, *Phys. Lett. B* **93** (1980) 419 [INSPIRE].
- [57] J.M. Martín-García, xAct: Efficient tensor computer algebra for the Wolfram Language, 2002–2022, http://xact.es/.
- [58] E. Remiddi and L. Tancredi, Schouten identities for Feynman graph amplitudes; The Master Integrals for the two-loop massive sunrise graph, Nucl. Phys. B 880 (2014) 343 [arXiv:1311.3342] [INSPIRE].
- [59] M. Chala, Á. Díaz-Carmona and G. Guedes, A Green's basis for the bosonic SMEFT to dimension 8, JHEP 05 (2022) 138 [arXiv:2112.12724] [INSPIRE].
- [60] Y. Chung, C.-O. Hwang and H.S. Yang, Algebraic properties of Riemannian manifolds, Gen. Rel. Grav. 55 (2023) 92 [arXiv:2206.08108] [INSPIRE].