Demonstration of MeV-Scale Physics in Liquid Argon Time Projection Chambers Using ArgoNeuT

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MeV-scale energy depositions by low-energy photons produced in neutrino-argon interactions have been identified and reconstructed in ArgoNeuT liquid argon time projection chamber (LArTPC) data. ArgoNeuT data collected on the NuMI beam at Fermilab were analyzed to select isolated low-energy depositions in the TPC volume. The total number, reconstructed energies and positions of these depositions have been compared to those from simulations of neutrino-argon interactions using the FLUKA Monte Carlo generator. Measured features are consistent with energy depositions from photons produced by de-excitation of the neutrino's target nucleus and by inelastic scattering of primary neutrons produced by neutrino-argon interactions. This study represents a successful reconstruction of physics at the MeV-scale in a LArTPC, a capability of crucial importance for detection and reconstruction of supernova and solar neutrino interactions in future large LArTPCs.

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I. INTRODUCTION

The Liquid Argon Time Projection Chamber 25 26 (LArTPC) is a powerful detection technology for neutrino experiments, as it allows for millimeter spatial 27 resolution, provides excellent calorimetric information 28 for particle identification, and can be scaled to large, 29 fully active, detector volumes. LArTPCs have been used 31 to measure neutrino-argon interaction cross sections 32 and final-state particle production rates in the case of ArgoNeuT [1-7] and MicroBooNE [8], neutrino 34 oscillations in the case of ICARUS [9], and charged 35 particle interaction mechanisms on argon in the case of 36 LArIAT [10].

LArTPCs are being employed to make important measu surements, e.g. understanding the neutrino-induced lowsu energy excess of electromagnetic events with Micro-

⁴⁰ BooNE [11] and will be used to search for sterile neu-⁴¹ trinos in the Fermilab SBN program [12] and for CP-⁴² violation in the leptonic sector with DUNE [13]. Precise ⁴³ measurements of neutrino-argon cross sections will be ⁴⁴ performed with SBN [12] and of charged hadron interac-⁴⁵ tions with ProtoDUNE [14]. In most of the existing mea-⁴⁶ surements, LArTPCs were placed in high energy neu-⁴⁷ trino beams to study GeV-scale muon and electron neu-⁴⁸ trinos as well as final-state products, generally with en-⁴⁹ ergies greater than 100 MeV. A smaller number of mea-⁵⁰ surements have investigated particles or energy deposi-⁵¹ tions in the < 100 MeV range [6, 15, 16], some using ⁵² scintillation light [17].

Few existing measurements have demonstrated LArTPC capabilities at the MeV scale for neutrino experiments, despite the wealth of physics studies that have been proposed for future large LArTPCs in this energy range. A number of studies have investigated expected supernova and solar neutrino interaction rates in the DUNE experiment: see Refs. [13] and [18] for reviews and relevant citations. Other studies have proposed using decay-at-rest neutrino interactions

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62 for short-baseline oscillation tests, coherent neutrino 115 Ionized charge drifted in the x-direction by means of an 63 scattering measurements and supernova-related stud- 116 electric field produced by a cathode biased at a negative 64 ies [19–23]. LArTPC experiments utilizing GeV-scale 117 high voltage of magnitude 23.5 kV. A field shaping cage 65 neutrino beamlines would also benefit from the ability 118 caused the electric field along the drift length to be uni- $_{66}$ to perform a reconstruction of MeV-scale features. This $_{119}$ form at 481 V/cm. The resulting drift velocity was 1.57 $_{67}$ ability would allow for a fuller reconstruction of beam $_{120}$ mm/ μ s, with a maximum drift time of 300.5 μ s. At the 68 69 released during de-excitation of the nucleus and of part 122 which two were instrumented (the innermost plane was 70 of the energy transferred to final-state neutrons. Fur- 123 a shield plane). The middle wire plane was the induction 71 thermore, MicroBooNE has shown that identifying and 124 plane; the outer one was the collection plane. Each of the 72 including full reconstructed energies at ends of showers 125 instrumented planes was comprised of 240 wires, with a 73 74 75 shower core [15].

76 77 ticles at MeV energies in a LArTPC is a challenging 100 hits identified from peaks above baseline. Triggering for ⁷⁸ task. At higher energies (> 100 MeV), charged particles ¹³¹ a readout was determined by the NuMI beam spill, at a 80 detectable signals on dozens to hundreds of TPC wires, 133 tional parameters of the ArgoNeuT detector are given in ⁸¹ producing an ionization track that can be utilized for re- ¹³⁴ [27]. ⁸² constructing the identity and kinematics of detected par-83 ticles. On the other hand, charged particles with kinetic 136 ND located immediately downstream of it. The MI-85 86 87 88 89 90 low-energy-specific methods. 91

We have used data acquired by the ArgoNeuT 145 92 ⁹³ LArTPC detector at Fermilab to search for small energy ¹⁴⁶ beam was operated in the low energy antineutrino mode: 94 depositions associated with neutrino events and com-⁹⁵ pared them to predictions from the FLUKA neutrino in- ¹⁴⁸ described in [2]. The composition of the beam was 58% 96 97 98 scattering of neutrons in the detector. 99

100 tor in Section II. We then overview nuclear de-excitation 154 (POT) acquired. 101 photon production, photon emission from inelastic scat-102 tering of neutrons, and photon propagation in argon in 103 Section III. We then describe utilized datasets and recon-104 struction in Sections IV and V. Final reconstructed sig-105 nal distributions are presented and compared to a Monte 106 Carlo (MC) simulation in Section VI. 107

II. THE ARGONEUT DETECTOR 108

ArgoNeuT was a LArTPC experiment which was 162 109 110 111 beamline at Fermilab for five months in 2009-2010. Ar- 164 nucleons and nuclear fragments may be emitted. The re-112 113 the MINOS near detector (MINOS ND). The TPC was 166 The nucleus de-excites by means of the emission of a $114 47(w) \times 40(h) \times 90(l)$ cm³ with a volume of 169 L. 167 photon or cascade of photons with energies ranging from

neutrino events by enabling reconstruction of photons 121 anode end of the TPC there were three wire planes, of is challenging and would benefit from the ability to $_{126}$ wire spacing of 4 mm and oriented at $\pm 60^{\circ}$ to the beam reconstruct Compton scatters of photons exiting the 127 direction. In each detector readout, each wire channel 128 was sampled every 198 ns, for a total readout window of Performing identification and reconstruction of par- $^{129}405\,\mu s$. The waveform for each wire was recorded with travel several centimeters to meters in distance, leaving 132 rate of 0.5 Hz. A more detailed description and opera-

ArgoNeuT benefited from the presence of the MINOS energies near the MeV scale travel a distance of the or- 137 NOS ND is a segmented magnetized steel and scintilder of or less than the distance between adjacent wires in 138 lator detector [28]. As a result, the momenta and signs of many LArTPCs (3-5 mm), leaving just one hit or a short 139 muons produced by neutrino interactions in ArgoNeuT cluster of a few consecutive hits. Thus, current analy- 140 and entering the MINOS ND could be determined by sis methods used to reconstruct physics quantities from 141 using reconstruction information from the MINOS ND. tracks made of large numbers of wire signals are ineffec- 142 ArgoNeuT also benefited from its placement 100 m untive in this energy regime, and there is a need for new, 143 derground; at this depth, cosmic rays are expected to be 144 seen in fewer than 1 in 7000 triggers.

During the majority of ArgoNeuT's run, the NuMI 147 neutrino fluxes produced during this operation mode are teraction generator [24-26]. Using new topological re- 149 muon neutrino, 40% muon antineutrino, and 2% elecconstruction tools, we find clear evidence of activity due 150 tron neutrino and antineutrino. The average energy for to de-excitation of the final-state nucleus and inelastic 151 muon neutrinos was 9.6 GeV, and the average energy of ¹⁵² muon antineutrinos was 3.6 GeV. The antineutrino mode We begin with a description of the ArgoNeuT detec- $_{153}$ run lasted 4.5 months with 1.25×10^{20} protons on target

III. PRODUCTION AND INTERACTION OF LOW-ENERGY PHOTONS IN NEUTRINO-ARGON **INTERACTIONS**

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MeV-energy photons can be produced in neutrino-158 159 argon interactions by two possible mechanisms, de-160 excitation of the target nucleus and inelastic scattering ¹⁶¹ of final-state particles. When a neutrino interacts with an ⁴⁰Ar nucleus, the target nucleon and the neutrino interplaced in the Neutrinos at the Main Injector (NuMI) 163 action products initiate a nuclear reaction during which goNeuT was located 100 m underground, in front of 165 maining residual nucleus is often left in an excited state.

 ~ 0.1 MeV – 10 MeV. Reaction products heavier than 221 are evaluated point-wise (for the exact neutron energy), 169 170 which inelastically scatter off an ⁴⁰Ar nucleus or are cap-²²⁴ following experimental energies and branching ratios. 171 tured by it will also produce photons in the energy range 225 172 of interest as the ⁴⁰Ar nucleus de-excites [29]. 173

174 175 176 177 178 179 180 182 183 185 ¹⁸⁶ in the energy range considered here.

Neutrino interactions and neutron scattering in 187 **FLUKA** 188

The only neutrino MC interaction generator that in-189 ¹⁹⁰ cludes the simulation of both mechanisms of low-energy photon production in GeV-scale neutrino interactions in 245 191 argon is FLUKA [24-26]. FLUKA is a multi-particle 192 transport and interaction code. Its neutrino interaction 246 193 194 196 197 198 199 200 201 202 203 204 205 206 207 account whenever available. 208

209 210 211 212 213 214 215 216 217 218 219 $_{220}$ been implemented for reactions on 40 Ar. Cross sections $_{273}$ sic 39 Ar activity, photons produced by entering neutrons

deuterons and the recoiling residual nucleus are gener- 222 correlations among reaction products are included, and ally not observable in a LArTPC. Final-state neutrons 223 gamma de-excitation is simulated as a photon cascade

Figure 1 shows the energies and numbers of pho-226 tons from charged current interactions of muon neutrinos As photons are neutral particles, they cannot be de- 227 from the NuMI beam interacting and depositing energy tected directly. Instead we detect electrons resulting from 228 in a volume of liquid argon with the dimensions of Ara photon interaction. The scale of the distance between 229 goNeuT, according to FLUKA simulation (see Section subsequent energy depositions for one photon is given 230 IV for details). A significant overlap in both the enby the radiation length (X_0) , which in liquid argon is 14 ²³¹ ergies and numbers of photons from the two processes cm. Over the $\sim 0.1 - 10$ MeV range of interest in this 232 (de-excitation of the target nucleus and inelastic neutron study, the most probable interaction process for photons 233 scattering) is visible, making separation of the source of in LAr is Compton scattering. In Compton scattering at 234 energy depositions difficult based on these metrics alone. this energy, each photon has a high probability of cre- 235 Considering ArgoNeuT's size, a photon could leave the ating multiple topologically isolated energy depositions 236 TPC with a significant amount of its energy undetected. within a LArTPC. Higher energy photons can also inter- 237 It is also notable that 24% of product nuclei in this simuact via pair-production, however this is still subdominant 238 lation are found in the ground state and produce no pho-239 tons.

> Typically, low energy photon-produced electrons are 240 ²⁴¹ expected to appear in a LArTPC event display as blips ²⁴² from isolated energy depositions around the neutrino in-²⁴³ teraction vertex. An example can be seen in Fig. 2, ²⁴⁴ where a typical ArgoNeuT neutrino event is shown.

IV. DATASETS

This analysis uses two primary real datasets from the generator, called NUNDIS [26], is embedded in the same 247 antineutrino mode run. Events with simple, low track nuclear reaction module of FLUKA used for all hadron- 248 multiplicity final-state topology have been selected for induced reactions. Quasi elastic, resonant (Δ produc- 249 the present analysis, as complex events make the selection only), and deep inelastic scattering interactions are 250 tion of isolated low-energy signatures more difficult. The modeled on single nucleons according to standard for- 251 first dataset, termed the neutrino dataset, is a subsample malisms. Initial state effects are accounted for by con- 252 of muon neutrino and antineutrino events from the Arsidering bound nucleons distributed according to a Fermi 253 goNeuT charged current pion-less (CC 0π) events sammomentum distribution. Final-state effects include a 254 ple, i.e. muon (anti)neutrino charged current events that generalized intranuclear cascade (G-INC), followed by a 255 do not produce pions in the final state. The selection pre-equilibrium stage and an evaporation stage. As men- 256 and analysis of these events [5], requires that a three tioned above, nucleons, mesons and nuclear fragments 257 dimensional (3D) track reconstructed in the LArTPC is can be emitted during these stages. Residual excitation is 258 matched to a MINOS ND muon track, and that any numdissipated through photon emission. Experimental data 259 ber of tracks at the vertex, identified as protons using the on nuclear levels and photon transitions are taken into 260 algorithm defined in [27], are present in the final state $_{261}$ ($\mu + Np$ events). In addition, we require that none of Neutron-induced reactions are treated as standard 262 the events contains a reconstructed 3D track identified as hadronic interactions for neutron energies above 20 MeV, 263 a charged pion or a reconstructed shower corresponding while for energies below 20 MeV a data-driven treat- 264 to a high-energy electron or photon. The threshold for ment is used, as in most low-energy neutron transport 265 proton (pion) identification is 21 (10) MeV [3]. From codes. Reaction cross sections, branching ratios and 266 the CC 0-pion sample we have selected a subsample of emitted particle spectra are imported from publicly avail- 267 events with one muon and up to one proton in the final able databases. Transport is based on a multi-group ap- $_{268}$ state (CC 0π , 0 or 1 proton events) for the present analyproach (neutron energies grouped in intervals, cross sec- 269 sis. The second dataset, termed the background dataset, tions averaged within groups), except for selected reac- 270 was obtained by examining "empty event" triggers which tions [24]. In the FLUKA version used for this work 271 do not appear to contain a neutrino interaction. These (FLUKA2017, not yet released), a special treatment has 272 readouts do contain ambient gamma ray activity, intrin-



FIG. 1. Energy (top) and multiplicity (bottom) of low-energy photons from charged current interactions of muon neutrinos from the NuMI beam interacting and depositing energy in a volume of liquid argon with the dimensions of ArgoNeuT. Color at 1.46 MeV corresponds to the first excited state of 40 Ar.

274 from neutrino interactions occurring upstream of the detector, and electronics noise. The beta emitter ³⁹Ar is a 275 radioactive isotope found in natural argon; at a rate of 276 277 1.38 Bq/L, it is not expected to be a large background in ArgoNeuT events. Electronics noise can be identified as 278 a hit if the deviation from the baseline is above a thresh-279 old. These features are also present in the neutrino events 280 previously described, so the background dataset is used 281 for a data-driven modeling of the background in the se-282 lected neutrino events. 283

ArgoNeuT data are compared with a MC dataset. We 284 produced simulated neutrino interactions in ArgoNeuT 285 using FLUKA and the energy spectrum of the NuMI 286 beamline. A simplified ArgoNeuT detector geometry 312 287 288 289 290 291



FIG. 2. A neutrino event (raw data) with one (longer) track reconstructed as a muon exiting the detector and one (shorter) track reconstructed as a proton. Possible photon activity (isolated blips) is visible in the event (e.g. collection plane wire 135, sample 700). The top image is the collection plane, and the bottom image is the induction plane. Wire number is indicated on the horizontal axis. The vertical axis indicates time sample number. Color indicates amount of charge collected.

292 of the final-state nucleus, resulting in the production 293 of final-state de-excitation photons. FLUKA was also ²⁹⁴ used to propagate final-state neutrons inside the LAr vol-²⁹⁵ ume, resulting in the simulation of energies and locations indicates source of photon (blue are de-excitation photons, red 296 of secondary neutron-produced photons. The FLUKAare photons produced by neutrons). For a photon to be tracked 297 determined properties of non-neutron final-state particles in the simulation, it must have an energy ≥ 0.2 MeV. The peak $_{298}$ and secondary neutron-produced photons were then used ²⁹⁹ as input to a LArSoft [30] MC simulation of ArgoNeuT 300 and propagated through the detector simulation, signal ³⁰¹ processing, and reconstruction stages as for real data. CC $_{302}$ 0π 0, 1 proton events, i.e. events with one muon track 303 entering the MINOS ND and up to one additional proton with kinetic energy > 21 MeV and no pions with kinetic energy > 10 MeV in the final state, compose the 305 306 selected MC samples for the present analysis. Electron-307 ics noise, ambient and internal radioactivity, and photons 308 from entering neutrons were not simulated; the back-309 ground dataset described above was instead used to di-³¹⁰ rectly include these contributions to the MC dataset.

EVENT RECONSTRUCTION V.

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As discussed in Section III, the radiation length in liqwas inserted into FLUKA. In addition to producing all 313 uid argon is 14 cm, and MeV photon-produced electrons the final-state particles emerging from the neutrino inter- 314 have ranges of a millimeter to a centimeter, as shown in action, including hadron re-interaction inside the nucleus 315 Fig. 3. Consequently, for the present analysis a signal (nuclear effects), FLUKA also simulates the physics 316 on the wire planes consists of a single hit or a very short 317 cluster of hits on consecutive wires on both active planes of the TPC, topologically isolated from the rest of the 318 event's features, possibly concentrated around the inter-319 action vertex, as shown in Fig. 2. 320

The same reconstruction procedure has been applied 321 to all the selected data and MC samples described in the 322 previous Section. The reconstruction proceeded through 323 two steps, one "standard" reconstruction step, followed 324 by a low-energy specific second step, described in Sec-325 tion VA. 326

First, the "standard" ArgoNeuT automated reconstruc-327 tion procedure, including hit finding, hit reconstruction 328 and track reconstruction, as described in detail in [7], 329 was applied. Events were required to have a recon-330 structed neutrino interaction vertex contained in the fidu-331 cial detector volume, defined as [3, 44] cm along the drift 332 direction, [-16, 16] cm vertically from the center of the 333 detector, and [6, 86] cm along the beam. The neutrino 334 and background datasets contain 552 and 1970 events, 335 336 respectively.

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Signal Selection

In the second step, a low-energy specific procedure to 338 339 identify and reconstruct isolated hits and clusters was applied. Since low-energy electrons will leave short 340 isolated features in the TPC, hits that are identified as 341 belonging to a reconstructed track longer than 1.5 cm 342 and beginning at the neutrino interaction vertex were re-343 moved. To also remove nearby wire activity associated 344 with a track (such as delta rays), all hits inside a 120° 345 cone around the first 2.4 cm of each reconstructed track 346 and a 5 cm cylinder along the remaining track length 347 were rejected. For tracks reconstructed as being longer 348 than 4 cm, the cylindrical rejection region was extended 349 past the end of the track, in case the automated recon-350 struction cuts the track short. 351

Then, several cuts were made on the remaining hits 352 found in each event. A threshold cut removed hits whose 353 fitted peak height is below a certain ADC count threshold 354 on the induction and collection planes (6 and 10 ADC, 355 respectively), corresponding to roughly 0.2 MeV of en-356 ergy deposited. Hits whose fitted peak height is above a 357 maximum ADC count (60 ADC, corresponding to ~ 1.2 358 MeV) were also removed, as they were unlikely to be 359 produced by photon energy depositions. As shown in 360 361 Fig. 3, such hits are more likely due to protons. For example, for a proton to travel a distance of 0.4 cm, the wire 362 spacing, it must have a kinetic energy of at least 21 MeV, 363 well above the maximum ADC cut. On the other hand, 364 an electron must have a kinetic energy of 1 MeV to travel 365 the same distance. Low energy protons with very short 366 range can result from a neutron-proton reaction on argon, 391 367 368 369



FIG. 3. Energy vs range for electrons and protons for the ranges of interest for this study. Red denotes protons, blue denotes electrons. The clear separation between electron and proton means it is unlikely a proton hit will be mistakenly identified as an electron hit. Data from [31].

370 then applied to remove all hits within 6 cm of the cath-371 ode and anode and hits near corners of the TPC. Real and 372 MC events were individually visually scanned to remove noisy wires and reconstruction failures. Individual wires 373 were removed on an event by event basis if it was clear 375 they had several hits due to electronics noise, with equiv-376 alent cuts applied to background events. Some hits were 377 also manually removed if it was clear they belonged to 378 a track that was not reconstructed properly. To suppress 379 hits originating from above-threshold electronics noise, 380 matching of hit times between induction and collection 381 planes was required. This plane matching also allowed 382 for reconstruction of the 3D space position for all hits ³⁸³ in the final sample passing the above selection criteria. Applied cuts are visually demonstrated in Fig. 4.

A summary of the level of hit removal achieved in each 386 cut for neutrino, background and MC datasets is found 387 in Table I. Once all cuts were applied and visual scan-³⁸⁸ ning was complete, the resulting neutrino (background) 389 datasets contained 716 (422) collection plane selected ³⁹⁰ hits in 552 (1970) events.

Cut	Percent of Hits Remaining			
	Neutrino	Background	MC	
Minimum Peak Height	65%	38%	94%	
Maximum Peak Height	58%	37%	84%	
Handscanning	54%	29%	78%	
Plane Matching	24%	10%	54%	

TABLE I. Impact of different cuts for collection plane hits. Cuts are applied sequentially. MC was simulated with no noise.

Following this selection, we grouped signal hits into however the FLUKA simulation indicates fewer than 1% 392 clusters and attempted a reconstruction of clusters' posiof hits passing cuts are due to protons. A fiducial cut was 393 tions and energies. A cluster is defined as a collection of



FIG. 4. Left: A raw data neutrino event display with one track reconstructed as a muon and with photon activity (isolated blips). The top image is the collection plane, and the bottom image is the induction plane. Wire number is indicated on the horizontal axis. The vertical axis indicates time sample number. Color indicates amount of charge collected. Right: The same event after hit finding and reconstruction. Each square denotes a reconstructed hit. Color indicates whether or not a hit was removed and by which cut (see text). Hits that pass all cuts are in red.

394 395 396 397 wire, each section was considered as a separate cluster. A 425 liquid argon during drift, as described in [7]. total number of 553, 319 and 4537 plane-matched clus-399 ters were reconstructed, yielding an average of 1.00, 0.16 400 and 1.12 clusters per event in the selected neutrino, background and MC events, respectively. In neutrino events, 402 ⁴⁰³ most of the clusters (75%) are composed of just one hit, 23% are two hit clusters, and only 2% are clusters with 404 more than two hits. 405

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B. Position Reconstruction

We reconstructed the 3D position of a cluster by 407 matching the furthest upstream collection plane hit in a 408 cluster to the furthest upstream induction plane hit in the 409 410 matched cluster. This yielded a coordinate on the yzplane. We then included the x-coordinate of the collec-411 tion plane hit to obtain a 3D position and calculated the distance of each cluster with respect to the neutrino inter-413 action vertex. While a cluster may span more than one 414 wire in a plane, the distance traveled by the presumed Compton-scattered electron creating the cluster is negli-416 gible when compared to the distance from the vertex. 417

Charge to Energy Conversion 418 С.

419 structed cluster, first the measured pulse area (ADC \times 453 ual hit forming the cluster. To test the efficacy of this 420

one or more signals on adjacent wires that occur within 421 time) of each hit was converted to charge (number of ion-40 samples on these wires. This value was determined 422 ization electrons) by an electronic calibration factor, then by examining a simulation of electrons with energies in 423 a lifetime correction was applied to account for ionizathe range of interest. If a cluster spans an unresponsive 424 tion electron loss due to attachment on impurities in the

> Calorimetric reconstruction in a LArTPC requires 426 427 converting the collected charge to the original energy de-428 posited in the ionization process. This requires applying 429 a recombination correction which depends on charge de-430 position per unit length dQ/dx [27]. The low-energy 431 photon-induced electrons in the present analysis result in 432 just isolated hits or clusters of very few hits, not extended 433 tracks, so the effective length of the electron track seen 434 by a wire cannot be determined.

A different method to estimate the energy from the de-435 436 posited charge which relies on the assumption that all ⁴³⁷ hits passing cuts are due to electrons has been developed. 438 The method uses the NIST table that provides the ac-439 tual track length for electrons in LAr at given energies 440 (ESTAR) [31], from 10 keV to 1 GeV. Using this table, ⁴⁴¹ we can thus approximate the deposited energy density $_{442} dE/dx$ by dividing the energy by the track length for 443 each row in the table. Using the Modified Box Equa-444 tion [32] to model the recombination effect, we can cal-445 culate the expected dQ/dx and by multiplying by the 446 track length (i.e. dx), we obtain the expected amount 447 of charge freed from ionization processes by an electron 448 at a given energy, as shown in Fig. 5 (left). By using 449 the result of a fit, also shown in the Figure, we can now 450 convert collected charge from the individual hit to de-⁴⁵¹ posited energy. The total energy in a cluster is the sum To reconstruct the energy associated with each recon- ⁴⁵² of the deposited energy reconstructed for each individ-



FIG. 5. Left: Energy deposited vs collected charge. Red curve indicates fit used to perform energy calculations from collected charge. Right: Reconstructed energy vs true electron energy using the charge method for a sample of simulated electrons with energies between 0 and 5 MeV. Events where the electron was not detectable are excluded.

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⁴⁵⁴ method, we applied it to a sample of GEANT4 simulated ⁴⁸⁶ electrons propagating in LAr in the energy range of in-455 terest. Figure 5 (right) indicates that it works well. We 487 find a detection efficiency of 50% and energy resolution 457 of 24% at 0.5 MeV, and an efficiency of almost 100% 458 $_{459}$ and energy resolution of 14% at 0.8 MeV.

D. Systematic Uncertainties 460

There are three primary sources of systematic uncer-461 tainty associated with hit and energy reconstruction in 462 this analysis. As the electron lifetime varies between 463 runs, we expect a variation and uncertainty in the num-464 ber of near-threshold hits that are selected as signal. De-465 spite having precise measurements of electron lifetime 466 for all runs, we conservatively account for electron life-467 time uncertainties by re-running FLUKA with a $\pm 25\%$ 468 change in electron lifetimes; the resultant spread in re-469 constructed multiplicities and energies is treated as the 470 systematic uncertainty from this source. A second sys-471 tematic uncertainty arises from the choice of a true un-472 derlying functional form for the recombination correc- 505 473 474 tion. To account for this uncertainty, we consider reconstruction of simulated events using the unmodified Box 507 simulation are shown in Figs. 6 and 7, respectively. 475 Model as described in [32]; deviation from the default 508 476 477 478 479 480 481 482 483 484 in this analysis. 485

VI. RESULTS

Comparison of Neutrino and Background Datasets

Table II shows a comparison of neutrino and background datasets. Comparing the different metrics leads 489 to the conclusion that we have observed a statistically 490 significant sample of neutrino-induced MeV-scale photons. Hit and cluster multiplicities are found to be sig-492 nificantly higher in the neutrino dataset than in the back-493 ground dataset, with 1.30 ± 0.07 and 0.21 ± 0.02 hits per 494 event, respectively. This difference corresponds to a 15σ 496 statistical excess of signal in the neutrino dataset. The 497 higher neutrino dataset multiplicity is also accompanied ⁴⁹⁸ by a larger per-event signal occupancy ($54 \pm 4\%$ in neu- $_{\rm 499}$ trino events versus $12\pm2\%$ in background events) and ⁵⁰⁰ total signal energy per event (1.1 MeV in neutrino events ⁵⁰¹ versus 0.19 MeV in background events). This can be 502 interpreted as evidence of neutrino-induced MeV-scale ⁵⁰³ energy depositions.

В. **Comparison to MC Simulations**

A comparison of reconstructed per-event signal multi-⁵⁰⁶ plicity and total signal energy for data and FLUKA MC

In both data and MC, around half of the events have selection is treated as an uncertainty contribution from 509 no signal clusters, as expected based on the small Arthis source. Finally, there is a 3% error associated with 510 goNeuT detector size and the previously-mentioned sizthe utilized calorimetric calibration constants, which are 511 able number of predicted product nuclei in the groundfully correlated between all runs. Any multiplicity or en- 512 state. Overall, there is good agreement between data and ergy variation arising from a $\pm 3\%$ shift in thresholds and 513 FLUKA MC predictions. We find a χ^2 /ndf of 7.81/12 reconstructed energies is treated as an uncertainty from 514 (p-value 0.80) for the total reconstructed energy disthis source. Systematic uncertainties in reconstructed po- 515 tributions, and a $\chi^2/ndf = 12.6/6$ (p-value 0.05) for sitions are expected to be small and were not considered 516 the cluster multiplicity distribution. Thus, we observe 517 that FLUKA, which incorporates low-level nuclear pro-

Metric	Neutrino Data	Background	
Number of hits per event	1.30	0.21	
Number of clusters per event	1.00	0.16	
Average total signal energy	1 11	0.10	
in an event (MeV)	1.11	0.15	
Percent of events with	54%	190%	
at least one signal hit	0470	1270	
Average cluster distance	22.4		
from vertex (cm)	22.4		

TABLE II. Comparison of neutrino and background datasets when examining hits passing all cuts. The difference in the first four metrics indicates neutrino-induced MeV-scale activity is visible.

Metric	De-excitation	Neutron	Total
Number of hits per event	0.48	0.98	1.46
Number of clusters per event	0.35	0.77	1.12
Average event energy (MeV)	0.41	0.76	1.17
Average cluster energy (MeV)	1.18	0.98	1.04
Average hit energy (MeV)	0.86	0.77	0.80
Average cluster distance	15.7	23.4	21.0
from vertex (cm)	10.1		

TABLE III. Relative contributions of de-excitation and neutron-produced photon components in FLUKA MC.

518 cesses that result in the production of MeV-scale energy 519 depositions following interactions of GeV-scale neutrinos in liquid argon, agrees well with the data. We ob-520 serve that the largest contributor to the χ^2 between the 521 data and MC multiplicity distributions is the difference 522 in high-multiplicity events. The modest excess in MC, 523 which spreads over multiple reconstructed energy bins, 524 could be indicative of flaws in the hit selection process, 525 or of imperfections in models or libraries utilized by 526 FLUKA. This feature can be better examined in future 527 high-statistics studies in larger LArTPCs. Finally, we 528 notice a dip in the first bin in Fig. 7, due to detector 529 thresholding, which can vary in data from event to event 530 due to different electron lifetime values. 531

Both components, de-excitation photons and photons 532 produced by interactions of final-state neutrons on ar-533 gon, are needed to have data-MC agreement. If de-534 excitation photons are removed from FLUKA distribu-536 537 538 tain $\chi^2/ndf = 194/12$ and $\chi^2/ndf = 197/6$ for these 539 same distributions, respectively. To confirm this, we also 550 540 541 542 543 eration of GENIE final states within the LArSoft frame- 553 these sources to the total activity in an event as given by 544



FIG. 6. Cluster multiplicity for neutrino data and FLUKA MC events. Data points include statistical error. Dark green line indicates FLUKA prediction with data-driven background added (see text). Dark green shaded area is statistical error in FLUKA, overlaid on total error (statistical + systematic) for FLUKA in light green shading. MC is normalized to the number of neutrino data events.



FIG. 7. Total signal reconstructed energy in an event for neutrino data and FLUKA MC events. Events with no reconstructed energy are not included. Data points include statistical error. Dark green line indicates FLUKA prediction with datadriven background added (see text). Dark green shaded area is statistical error in FLUKA, overlaid on total error (statistical + systematic) for FLUKA in light green shading. MC is normalized to the number of neutrino data events.

545 cedure as in FLUKA was applied to GENIE events. As tions, we obtain a $\chi^2/ndf = 82.6/12$ for reconstructed ⁵⁴⁶ an example, a comparison of reconstructed multiplicity energy and $\chi^2/ndf = 93.8/6$ for the cluster multiplic- ⁵⁴⁷ is shown in Fig. 8. The χ^2/ndf is 57.9/6. This disagreeity. If neutron-produced photons are removed, we ob- 548 ment is attributed to the lack of de-excitation photons in 549 the GENIE simulation of neutrino-argon interactions.

These results indicate that the observed MeV-scale compared ArgoNeuT data with a GENIE MC simula- 551 signals in ArgoNeuT contain both de-excitation and tion [33]; existing user interfaces allowed for easy gen- 552 neutron-produced photons. The contribution of each of work. The same event selection and reconstruction pro- 554 the FLUKA simulation is shown in Table III. We find



FIG. 8. Distribution of cluster multiplicity for neutrino data and GENIE events. Data points include statistical error. Dark blue indicates GENIE prediction (no de-excitation photons). Light tion. MC is normalized to the number of neutrino data events.

555 that we cannot distinguish between the two sources of photons by examining the energy of a hit or cluster alone, 583 556 but we do see a difference in the distance of a cluster 557 with respect to the neutrino interaction vertex. The dis-558 559 tribution of these distances is seen in Fig. 9. Photons produced by de-excitation of the final-state nucleus tend 560 to be concentrated at lower distances, while photons pro-561 562 distances. 563



FIG. 9. Distributions of cluster position with respect to the neutrino interaction vertex in neutrino data and FLUKA MC events. Data includes statistical error. Green indicates the contribution of photons from de-excitation of the final-state nucleus. Red indicates the contribution of photons from inelastic neutron scattering. MC is area normalized to data.

VII. CONCLUSION

The ability to reconstruct activity at the MeV scale in 565 a LArTPC is crucial for future studies of supernova, so-566 lar, and beam neutrino interactions. In addition, stud-567 568 ies of low scale new physics scenarios, such as mil-569 licharged particles, light mediators, and inelastic scatterings with small splittings (see e.g. Refs. [34-36]), could 570 invaluably profit from such low energy reconstruction. 571 By studying low-energy depositions produced by pho-572 tons in ArgoNeuT neutrino interactions and comparing 573 574 to simulation, we have shown that such a reconstruction 575 is possible. Performing this study required the creation of new techniques for low-energy LArTPC reconstruc-577 tion. By reconstructing photons produced by nuclear deblue shaded area indicates statistical error for GENIE predic- 578 excitation and inelastic neutron scattering, we have ex-579 tended the LArTPC's range of physics sensitivity down 580 to the sub-MeV level, reaching a threshold of 0.3 MeV ⁵⁸¹ in this analysis. This range now spans more than three ⁵⁸² orders of magnitude, up to the GeV level.

In our study of low-energy depositions in 552 Ar-584 goNeuT neutrino events, we found 553 clusters with an $_{585}$ average of 1.30 ± 0.07 hits per event and an average en- $_{586}$ ergy of 1.11 ± 0.16 MeV per event. Signal cluster multi-587 plicities in neutrino events outnumbered those in nearby 588 background events, establishing a clear neutrino-based duced by inelastic neutron scattering dominate at higher 589 origin for these MeV-scale features. These and other ⁵⁹⁰ cluster properties matched those predicted for photons 591 due to inelastic neutron scattering and de-excitation of the final-state nucleus in FLUKA using its model of nu-592 clear physics processes at the MeV-scale. Removal of ei-593 ther of these event classes significantly worsens the level 594 of data-simulation agreement. 595

> This analysis represents the first-ever reported de-596 tection of de-excitation photons or final-state neu-597 trons produced by beam neutrino interactions in argon. 598 Both of these particle classes could provide valuable 599 new avenues of investigation for physics reconstruction 600 in LArTPCs. Reconstruction of MeV-scale neutron-601 602 produced features may enable some level of direct reconstruction of final-state neutron energies or multiplicities, 603 604 which would provide a valuable new handle on one of 605 the dominant expected differences between neutrino and 606 antineutrino interactions in liquid argon. Precise recon-607 struction of de-excitation photon multiplicities and ener-608 gies will improve overall reconstruction of neutrino en-609 ergies, particularly for those at lower energies, such as 610 supernova and solar neutrinos. Future MC studies and 611 higher-statistics datasets from future large LArTPCs will 612 provide additional understanding of the value of these 613 MeV-scale features.

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- 636 161802 (2012). 637 682
- [2] R. Acciarri et al. (ArgoNeuT), Phys. Rev. D89, 112003 638 683 639 (2014).
- [3] R. Acciarri et al. (ArgoNeuT), Phys. Rev. D90, 012008 640 685 (2014).641
- 642 261801 (2014), [erratum: 113. Phys. Rev. 643 688 Lett.114,no.3,039901(2015)]. 644
- [5] O. Palamara (ArgoNeuT), JPS Conf. Proc. 12, 010017 690 [25] G. Battistoni et al., Annals of Nuclear Energy 82, 10 645 (2016).646
- [6] R. Acciarri et al. (ArgoNeuT), Phys. Rev. D96, 012006 647 (2017).648
- [7] R. Acciarri et al. (ArgoNeuT), (2018), arXiv:1804.10294 649 [hep-ex]. 650
- et al. (MicroBooNE), (2018), 696 Adams 651 [8] С. arXiv:1805.06887 [hep-ex]. 652
- [9] M. Antonello et al. (ICARUS), Eur.Phys.J. C73, 2345 653 (2013).654
- F. Cavanna, M. Kordosky, J. Raaf, and B. Rebel (LAr- 700 [29] 655 [10] IAT), (2014), arXiv:1406.5560 [physics.ins-det]. 656
- [11] R. Acciarri et al. (MicroBooNE), JINST 12, P02017 702 657 (2017).658
- [12] M. Antonello et al. (LAr1-ND, ICARUS-WA104, Micro-659
- BooNE), (2015), arXiv:1503.01520 [physics.ins-det]. 660
- B. Abi et al. (DUNE), (2018), arXiv:1807.10334 706 661 [13] [physics.ins-det]. 662 707
- [14] B. Abi et al. (DUNE), (2017), arXiv:1706.07081 663 708 [physics.ins-det]. 664
- [15] R. Acciarri et al. (MicroBooNE), JINST 12, P09014 710 665 (2017).666
- [16] S. Amoruso et al. (ICARUS), Eur. Phys. J. C33, 233 712 667 668 (2004).
- [17] W. Foreman (LArIAT), JINST 11, C01037 (2016). 669
- [18] K. Scholberg, Neutrino physics and astrophysics. Pro- 715 670
- ceedings, 19th International Conference, Neutrino 2000, 716 671
- Sudbury, Canada, June 16-21, 2000, Nucl. Phys. Proc. 717 [34] 672 Suppl. 91, 331 (2001), [331(2000)]. 673
- [19] C. Grant and B. Littlejohn, Proceedings, 38th Inter- 719 [35] 674
- national Conference on High Energy Physics (ICHEP 720 675
- 2016): Chicago, IL, USA, August 3-10, 2016, PoS 721 [36] 676 ICHEP2016, 483 (2016). 677 722
- [20] H. Berns et al. (CAPTAIN) (2013) arXiv:1309.1740 723 678 [physics.ins-det]. 679
- 680 [21] J. Spitz, Phys. Rev. D 85, 093020 (2012).

- [1] R. Acciarri et al. (ArgoNeuT), Phys. Rev. Lett. 108, 681 [22] D. Akimov et al. (CSI), in Proceedings, 2013 Community Summer Study on the Future of U.S. Particle Physics: Snowmass on the Mississippi (CSS2013): Minneapolis, MN, USA, July 29-August 6, 2013 (2013) arXiv:1310.0125 [hep-ex].
 - S. J. Brice et al., Phys. Rev. D89, 072004 (2014). 686 [23]
- [4] R. Acciarri et al. (ArgoNeuT), Phys. Rev. Lett. 687 [24] A. Ferrari, P. R. Sala, A. Fasso, and J. Ranft, FLUKA: A multi-particle transport code (Program version 2005), Tech. Rep. (2005).
 - (2015).
 - G. Battistoni, A. Ferrari, M. Lantz, P. R. Sala, and 692 [26] G. I. Smirnov, in CERN-Proceedings-2010-001 (2010) 693 pp. 387-394, proceedings of 12th International Confer-694 ence on Nuclear Reaction Mechanisms, Varenna, Italy, 695 15-19 June 2009.
 - 697 [27] C. Anderson et al. (ArgoNeuT), JINST 7, P10019 (2012).
 - 698 [28] D. G. Michael et al. (MINOS), Nucl. Instrum. Meth. A596, 190 (2008). 699
 - National Nuclear Data Center, information extracted from the Chart of Nuclides database, 701 http://www.nndc.bnl.gov/chart/.
 - E. L. Snider and G. Petrillo, Proceedings, 22nd Interna-[30] 703 tional Conference on Computing in High Energy and Nuclear Physics (CHEP2016): San Francisco, CA, October 14-16, 2016, J. Phys. Conf. Ser. 898, 042057 (2017).

704

705

709

711

713

- Berger, M.J., Coursey, J.S., Zucker, M.A., and Chang, [31] J. (2005), ESTAR, PSTAR, and ASTAR: Computer Programs for Calculating Stopping-Power and Range Tables for Electrons, Protons, and Helium Ions (version 1.2.3). [Online] Available: http://physics.nist.gov/Star [2017, December 8]. National Institute of Standards and Technology, Gaithersburg, MD.
- R. Acciarri et al. (ArgoNeuT), JINST 8, P08005 (2013). 714 [32]
 - [33] C. Andreopoulos et al., Nucl. Instrum. Meth. A614, 87 (2010).
- S. N. Gninenko, Phys. Lett. **B710**, 86 (2012), arXiv:1201.5194 [hep-ph]. 718
 - G. Magill, R. Plestid, M. Pospelov, and Y.-D. Tsai, (2018), arXiv:1806.03310 [hep-ph].
 - E. Bertuzzo, S. Jana, P. A. N. Machado, and R. Zukanovich Funchal, (2018), arXiv:1808.02500 [hepph].

614